# Evaluation and Use of the Prompt Fission Neutron Spectrum and Spectra Covariance Matrices in Criticality and Shielding

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## **Outline**

Constrained sensitivity method for the calculation of fission spectra sensitivity.

 Yasunobu Nagaya, Ivan Kodeli, Go Chiba, Makoto Ishikawa, "Evaluation of sensitivity coefficients of effective multiplication factor with respect to prompt fission neutron spectrum", Nuclear Instruments and Methods in Physics Research A 603 (2009) 485–490

Evaluation of PFNS and spectra covariance matrices for <sup>235</sup>U, <sup>238</sup>U and <sup>239</sup>Pu.

- 2) Ivan Kodeli, Andrej Trkov, Roberto Capote, Yasunobu Nagaya, Vladimir Maslov, "*Evaluation and use of the prompt fission neutron spectrum and spectra covariance matrices in criticality and shielding*," Nuclear Instruments and Methods in Physics Research A **610** (2009) 540–552
- 3) **OECD/NEA WPEC Subgroup 26** Final Report, Uncertainty and Target Accuracy Assessment for Innovative Systems Using Recent Covariance Data Evaluations (Feb. 2008) APPENDIX M: I. Kodeli, M.Ishikawa, G. Aliberti, Evaluation of Fission Spectra Uncertainty and their Propagation

# **Objectives**

### Investigate:

- 1. Validation of fission spectrum covariance preparation and processing (analytical and MC)
- 2. Study practical implications of sensitivity profile normalisation
- 3. Apply uncertainty analysis to criticality  $(k_{eff})$  and shielding of different reactor systems;
- 4. Differences between ENDF/B-VII, Watt, Kornilov model spectra, and their consistency with the estimated uncertainties.

Recommendations of WPEC-WG26

## Data Formats for Cross Section Covariances in Evaluated Data Files

- *MF*=31: covariance of average number of neutrons per fission (v MT=452, 455, 456)
- *MF*=32: Shape and area of individual resonances
- *MF*=33: covariance of neutron cross section
- MF=34: covariance of angular distribution of secondary neutron (currently MT=2/P<sub>1</sub> only)
- **MF=35** covariance of energy distribution of secondary particles (currently MT=18 only)
- No processing available:
- *MF*=30: Covariances obtained from parameter covariances and sensitivities
- *MF*=40: Covariances for production of radioactive nuclei *Processing available (NJOY-ERRORR and ERRORJ)*

# **Prompt fission neutron spectra**



## **Prompt fission neutron spectrum properties**

- Fission spectrum is normalised to 1 sum of bin probabilities equals 1:
- To assure that the perturbed spectrum remains normalised the covariance matrix should comply with the **«zero sum» rule**: sum of absolute matrix elements in each line and column must be zero:

 Covariances are given as absolute covariances of the bin probabilities (not average probability distributions)

$$\sum_{g} \chi_{g} = 1$$



# Normalisation applied to the sensitivity coefficients

 If the matrix does not sattisfy the "zero sum" rule, the ENDF-6 manual suggests the correction formula:

$$\widetilde{V}_{i,j} = V_{i,j} - \chi_i \sum_k V_{k,j} - \chi_j \sum_k V_{k,i} + \chi_i \chi_j \sum_{k,k'} V_{k,k'}$$

or in matrix notation:  $V = S_{\chi}^{T} \cdot V \cdot S_{\chi} = S_{g,g'}^{\chi} - \chi_{g'}$ 

 Instead of "correcting" the matrices, we can apply the «correction» to the sensitivities :

$$(\Delta R)^{2} = S_{R}^{t} \cdot \widetilde{V} \cdot S_{R} = S_{R}^{t} \cdot \left(S_{\chi}^{t} \cdot V \cdot S_{\chi}\right) \cdot S_{R}$$
  

$$= \left(S_{\chi} \cdot S_{R}\right)^{t} \cdot V \cdot \left(S_{\chi} \cdot S_{R}\right) = S_{RN}^{t} \cdot V \cdot S_{RN}$$
  
SAGEP, SUSD3D

## Constructing fission spectra covariance matrices- analytic method

$$M_{\chi} = S_{ab}^{T} \cdot M_{ab} \cdot S_{ab}$$

$$\Box \text{ Covariance matrix of the parameters a and b of the watt spectrum:}$$

$$\delta a/a = 1.2\%$$

$$\delta b/b = 5.9\% \qquad R = \begin{vmatrix} 1 & 0 \\ 0 & 1 \end{vmatrix}$$

$$(\Delta R)^{2} = S_{R}^{T} \cdot \tilde{V} \cdot S_{R} = S_{R}^{T} \cdot \left(S_{\chi}^{T} \cdot S_{ab}^{T} \cdot V_{ab} \cdot S_{ab} \cdot S_{\chi}\right) \cdot S_{R}$$

$$= \left(S_{ab} \cdot S_{\chi} \cdot S_{R}\right)^{T} \cdot V_{ab} \cdot \left(S_{ab} \cdot S_{\chi} \cdot S_{R}\right) = S_{RN}^{T} \cdot V_{ab} \cdot S_{RN}$$

$$ENDF File-30$$

## Constructing fission spectra covariance matrices – Monte Carlo method

 Gaussian probability density distribution of the parameters a and b (non-correlated):

$$P(a,b) = P(a) \cdot P(b) = \frac{1}{2\pi \cdot \delta a \cdot \delta b} \cdot e^{-\frac{1}{2} \left[ \left( \frac{a - \overline{a}}{\delta a} \right)^2 + \left( \frac{b - \overline{b}}{\delta b} \right)^2 \right]}$$

- Mean:  $\overline{\chi}_{g} = \sum_{i} P(a_{i}, b_{i}) \chi_{g}(a_{i}, b_{i}) = \frac{1}{n} \sum_{i=1}^{n} \chi_{g,i}$ 

- Covariance:  $\overline{\delta \chi_g \cdot \delta \chi_{g'}} = \sum_{i=1}^n P(a_i, b_i) \left( \chi_g(a_i, b_i) - \overline{\chi_g} \right) \left( \chi_g'(a_i, b_i) - \overline{\chi_{g'}} \right)$ 

 $=\frac{1}{n}\sum_{i=1}^{n} \left(\chi_{g,i} - \overline{\chi_g}\right) \left(\chi_{g',i} - \overline{\chi_{g'}}\right)$ 

# Parameters (r,α) of Kornilov PFNS and parameters (a,b) of Watt model

Parameter	U-235	U-238	Pu-239
r=T <sub>l</sub> /T <sub>h</sub>	1.1215±3.0%	1.248±3.1%	1.1215±3.0%
α	0.905±3.0%	0.880±3.0%	0.823±5.1%
а	0.988±1.2%	0.881±1.2%	0.966±1.2%
b	2.249±5.9%	3.401±5.9%	2.842±5.9%

# Covariance matrices of U-235 fission spectrum



## **VENUS-3 benchmark**

#### **Fission averaged detector cross-sections**

Detector	M	leasured	Nev Kor	v nilov	Watt		JENDL-	3.3	ENDF/B- VII.0
<sup>27</sup> Al(n,a) [mb]	0.	706	0.699 ± 11.1%		0.742 ± 11.1%		0.732 ± 14.1%		0.746
<sup>58</sup> Ni(n,p) [b]	0.	1085	0.10 3.0%	56 ± %	0.1085 4.0%	±	0.1072 ± 5.2%		0.1075
<sup>115</sup> ln(n,n') [b]	0.	1903	0.18 1.2%	851 ± %	0.1871 1.7%	±	0.1883 ± 2.5%		0.1881
Reaction		Uncertainty (%)					RR ratio		
rates		New Kornilov		Watt		JEN	DL-3.3	(Nev Kor	w nilov/B7)
<sup>27</sup> Al(n,a)		11.5 11.5		11.5		14.5		0.95	
<sup>58</sup> Ni(n,p)		4.9 5.		5.9		7.6		0.98	3
<sup>115</sup> ln(n,n')		2.8		3.3 5.2		5.2 0.98		3	

# **Pressure Vessel Dosimetry**

Reaction rate	Uncertainty (	RR ratio		
	New Kornilov	Watt	JENDL-3.3	(New Kornilov/B7)
<sup>56</sup> Fe(n,p)	6.0	6.8	8.7	0.989
<sup>58</sup> Ni(n,p)	5.7	6.5	8.1	0.986
<sup>63</sup> Cu(n,)	11.0	10.9	13.9	0.981
<sup>237</sup> Np(n,f)	2.8	3.7	5.0	0.987
<sup>238</sup> U(n,f)	3.7	4.6	6.0	0.987

#### **IRPhE-International Reactor Physics Experiments Database (OCDE/AEN)**

#### KRITZ-2 –thermal benchmarks 6 critical configurations with UO2 and MOX fuel.

#### SNEAK-7 – fast benchmarks 2 configurations MOX

![](_page_13_Figure_3.jpeg)

SNEAK-LMFR-EXP-001 CRIT-SPEC-COEF-KIN-RRATE-MISC

#### SNEAK 7A AND 7B PU-FUELLED FAST CRITICAL ASSEMBLIES IN THE KARLSRUHE FAST CRITICAL FACILITY

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![](_page_13_Figure_8.jpeg)

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Analysis based on déterministic transport (TWODANT, THREEDANT) and cross section sensitivity and uncertainty code (SUSD3D)

## Normalised and un-normalised sensitivities

Sensitivity to <sup>235</sup>U fission spectrum (KRITZ-201C benchmark)

![](_page_14_Figure_2.jpeg)

## Normalised and un-normalised sensitivities

![](_page_15_Figure_1.jpeg)

# **Normalised sensitivity TO PFNS**

![](_page_16_Figure_1.jpeg)

17

![](_page_17_Figure_0.jpeg)

## Uncertainty in k-eff (pcm) due to Pu-239 fission spectra uncertainty (SNEAK-80g and KRITZ-18g)

![](_page_18_Figure_1.jpeg)

Incertitudes (pcm)

# Conclusions

- **MC method** was used to produce covariances for <sup>235</sup>U, <sup>238</sup>U & <sup>239</sup>Pu **fission spectra** based on the Watt and Kornilov models. Spectra were validated against an analytical approach restricted to linear approximation.
- MC method can be applied with confidence to non-linear models, such as optical model.
- "Normalised" sensitivity method (discussed during WPEC-SG26) was implemented in SUSD3D. Incorrectly normalised matrices can be detected comparing the uncertainties based on "normalised" and "un-normalised" sensitivities.
- "Normalised" sensitivity method and new covariance matrices were tested on sets of thermal and fast critical experiments.
- Uncertainty in k-eff due to the fission spectra uncertainties are ~10 30 pcm for thermal and ~200 300 pcm for fast systems.
- Differences due to different spectra are consistent with the uncertainties.
- Difefrences between thermal and fast PFNS uncertainties (fast<th?).</li>