

Deformation in $^{28}\text{Si}^*$ produced via the $^{16}\text{O} + ^{12}\text{C}$ reactionS. Kundu,^{*} C. Bhattacharya, S. Bhattacharya, T. K. Rana, K. Banerjee, S. Mukhopadhyay, D. Gupta,[†] A. Dey, and R. Saha
Variable Energy Cyclotron Centre, 1/AF, Bidhan Nagar, Kolkata 700 064, India

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The energy spectra of the α particles emitted in the reactions ^{16}O (7–10 MeV/nucleon) + ^{12}C have been measured in the center-of-mass angular range of $25^\circ \lesssim \theta_{\text{c.m.}} \lesssim 70^\circ$. The experimental energy spectra have been compared with those obtained from the statistical model calculation with “deformability” parameters predicted by rotating liquid drop model (RLDM) and also fitted the same with optimized deformability parameters, which are quite different from the respective RLDM values. The data have also been found to be explained quite well using “frozen” deformation approximation, where the deformability parameters have been kept fixed at RLDM values of the parent nucleus throughout the decay process. The effective radius in the latter case is smaller compared to that obtained using the optimized parameters; however, in both cases, the deformations (effective radii) are larger than the corresponding RLDM values. So, considering the uncertainties in the estimation of actual compound nucleus deformations, it can, only qualitatively, be said that equilibrium orbiting, which is similar to particle evaporation in time scale, could also be one of the contributing factors for the observed deformation.

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I. INTRODUCTION

Exploring the role of clustering in fragment emission mechanism is a subject of current interest in low-energy nuclear reaction. Several attempts have recently been made to understand the effect of cluster structure on the reaction mechanism of light α -cluster nuclei, e.g., $^{20}\text{Ne} + ^{12}\text{C}$ [1–4], $^{24}\text{Mg} + ^{12}\text{C}$ [5], $^{28}\text{Si} + ^{12}\text{C}$ [6,7], etc. In each case, an enhancement in the yield and/or resonancelike excitation function in a few outgoing channels not very different from the entrance channel has been observed, which was indicative of significant contribution from the deep-inelastic orbiting (DIO) mechanism. Through this process, a long-lived dinuclear composite is formed [8], which is not fully equilibrated in all (in particular, shape) degrees of freedom and therefore decays preferentially around the entrance channel; at higher excitation energy, however, the decay is expected to be of fully equilibrated in nature. It is thus interesting to probe into the limit of survival of such dinuclear composites at higher excitations.

The study of deformation of the excited composite is important to differentiate the DIO composite from the fully equilibrated compound nucleus (CN) because the deformation should be larger in the former case. Detailed studies of the deformation of $^{40}\text{Ca}^*$ [9,10] and $^{32}\text{S}^*$ [11] were done recently with the help of light charged particle (LCP) spectroscopy to extract “anomalous” deformation in these cases confirming the presence of DIO mechanism. For these systems, orbiting ($^{40}\text{Ca}^*$ [6,7], $^{32}\text{S}^*$ [1–3]) had already been conjectured from the fragment emission studies. Recently, we have reported that, for α -cluster system $^{16}\text{O} + ^{12}\text{C}$ [12], there was enhancement in the boron product yield at excitation energies ~ 67 – 85 MeV (beam energy in the range of 7–10 MeV/nucleon), which indicated

the survival of long-lived dinuclear orbiting composites at such high excitation energies. An orbiting dinuclear complex is assumed to be more deformed than a fully shape equilibrated CN; this motivated us to look for such enhanced deformation in a $^{16}\text{O} + ^{12}\text{C}$ dinuclear system at same excitation energies using the LCP spectroscopy technique. In this technique, LCPs, evaporated from the excited composites, are used to study the properties of hot rotating nuclear systems as a function of excitation energy, angular momentum, and deformation [11,13–25] using statistical model codes, e.g., CASCADE [26], GANES [27], CACARIZO [28].

The article has been arranged as follows. The experimental setup has been described in Sec. II. The experimental results and analysis have been presented in Sec. III. The details of the statistical model calculation has been given in Sec. IV. The extracted values of the deformation of $^{28}\text{Si}^*$ and their possible implications have been discussed in Sec. V. Finally, the summary and conclusion have been given in Sec. VI.

II. EXPERIMENTAL DETAILS

The experiment was performed at the Variable Energy Cyclotron Centre, Kolkata, using ^{16}O ion beams of energies 117, 125, 145, and 160 MeV. Self-supporting ^{12}C of thickness $\sim 514 \mu\text{g}/\text{cm}^2$ was used as a target. The α particles were detected using a Si(SB) telescope ($\sim 10 \mu\text{m} \Delta E$, $\sim 5 \text{mm} E$). The typical solid angle covered by the telescope was ~ 0.3 msr. The calibration of the telescope was done using elastically scattered ^{16}O ion from a Au target, as well as a ^{229}Th - α source. Inclusive energy distributions for the LCPs were measured in the laboratory angular range of 11° – 29° , which covered the angular range $\sim 25^\circ$ – 70° in the center-of-mass frame.

III. EXPERIMENTAL RESULTS**A. Energy spectra**

The center-of-mass (c.m.) energy spectra of the α particles at different beam energies are shown in Fig. 1. Because of the

^{*} skundu@vecc.gov.in[†]Present address: Department of Physics and Centre for Astroparticle Physics and Space Science, Bose Institute, Block EN, Sector V, Bidhan Nagar, Kolkata 700 091, India.

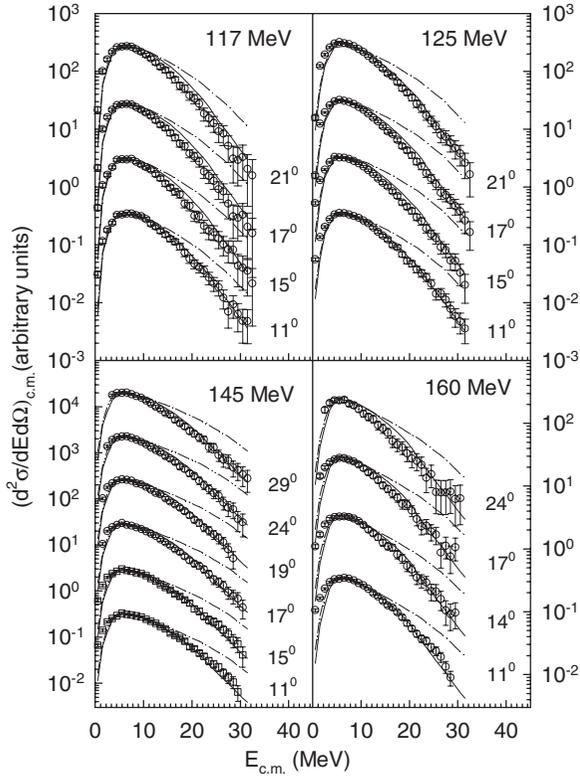


FIG. 1. Energy spectra (c.m.) of α particles obtained at different angles for different beam energies. The symbols represent experimental data. The dash-dot-dashed and solid lines represent CASCADE calculations with RLDM and optimized values of spin-dependent “deformability” parameters, respectively (see Table II). All experimental and calculated spectra, starting from the lowest angle, were multiplied by 10^{-2} , 10^{-1} , 10^0 , 10^1 , 10^2 , 10^3 , respectively.

inclusive nature of the spectra, there may be some admixture of contributions of other direct reaction mechanisms (such as pre-equilibrium emission, etc.) to the equilibrium emission spectra at forward angles in particular. In Fig. 2, the energy

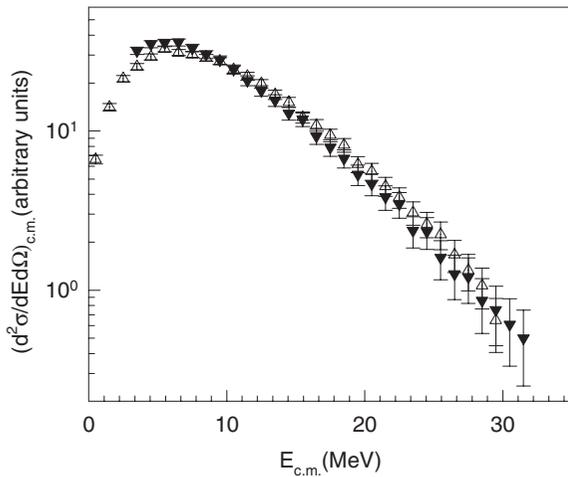


FIG. 2. Comparison of the energy spectra (c.m.) of α particles obtained at c.m. angles 26° (open triangles) and 66° (inverted triangles) for beam energy 145 MeV.

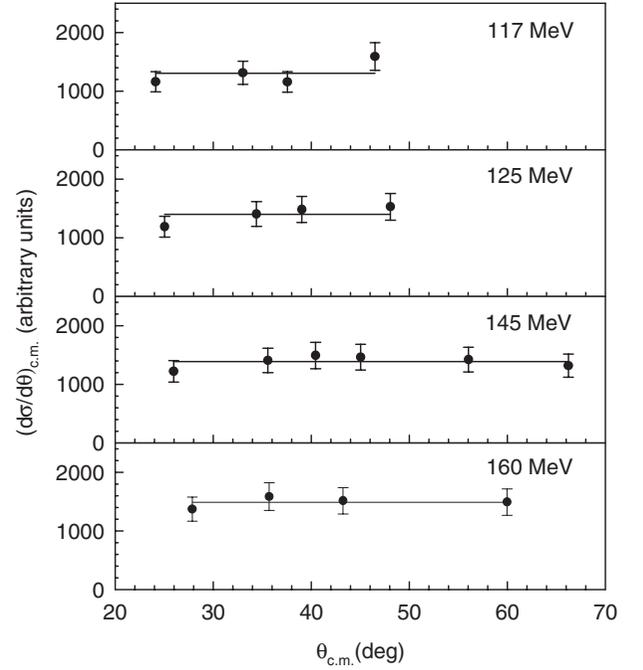


FIG. 3. Angular distribution of α particles as a function of c.m. angles, $\theta_{c.m.}$. Symbols represent the experimental data and solid lines show fit to the data obtained using $(d\sigma/d\theta)_{c.m.} = \text{constant}$.

spectra obtained at laboratory angles 11° ($\theta_{c.m.} = 26^\circ$, open triangles) and 29° ($\theta_{c.m.} = 66^\circ$, solid inverted triangles) at $E_{\text{lab}} = 145$ MeV are compared; it is seen that the slopes of the both spectra are completely matched with each other. It is thus clear that the effects of other direct reaction mechanisms are not significant at the measured angles.

B. Angular distribution

The c.m. angular distributions, $(d\sigma/d\theta)_{c.m.}$, obtained by integration of the c.m. energy distributions for the beam energies of 117, 125, 145, and 160 MeV, are shown in Fig. 3 as a function of $\theta_{c.m.}$. In all cases, the values of $(d\sigma/d\theta)_{c.m.}$ are found to be constant over the whole range of observed c.m. angles. So, $(d\sigma/d\theta)_{c.m.} \propto 1/\sin\theta_{c.m.}$, which is characteristic of the emission from an equilibrated composite nucleus.

C. Average velocity

The average velocities, v_{av} , of the α particles emitted at different angles have been extracted from the average energy, E_{av} , which has been calculated using the expression

$$E_{\text{av}} = \frac{\sum_i E_i (d^2\sigma/dE d\Omega)_i}{\sum_i (d^2\sigma/dE d\Omega)_i}, \quad (1)$$

where the index i covers the whole energy spectrum in the laboratory. The parallel (v_{\parallel}) and perpendicular (v_{\perp}) components (with respect to the beam direction) of v_{av} at each angle are plotted in Fig. 4. It is seen that they fall on a circle with the center at CN velocity, v_{CN} , and radius of average velocity in c.m., $v_{\text{av}}^{\text{c.m.}}$, which implies that the average velocities (as well as energies) of the α particles are independent of the c.m. emission angles. It again indicates that the α particles are

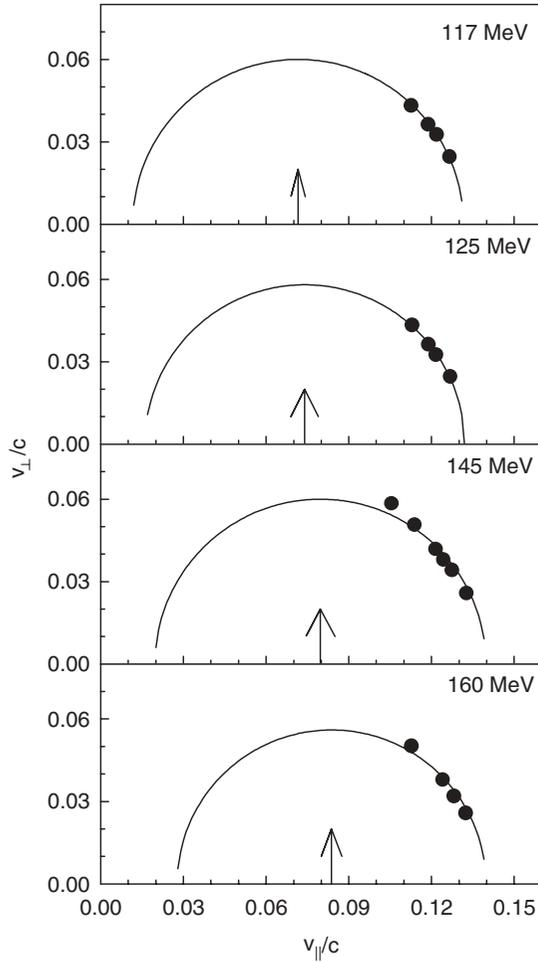


FIG. 4. Average velocity curve. Solid circles correspond to experimental data and solid lines show fit to the data obtained using the equation $v_{\perp}^2 = (v_{\text{av}}^{\text{c.m.}})^2 - (v_{\parallel} - v_{\text{CN}})^2$, where $v_{\text{av}}^{\text{c.m.}}$ is the average velocity of α particles in the c.m. frame. The arrow indicates the position of v_{CN} .

emitted from a fully energy equilibrated source moving with the velocity v_{CN} .

IV. STATISTICAL MODEL CALCULATIONS

It is confirmed from the experimental results that the α particles are emitted from a fully energy equilibrated composite which may be a CN. So the statistical model code CASCADE [26] has been used to explain the present experimental data because this code is based on the assumption that the evaporation of particles takes place from an excited CN which is in full equilibrium with respect to all degrees of freedom.

A. CASCADE calculations procedure

The spin distribution in CASCADE is usually derived from the known fusion cross section using the strong-absorption model wherever data are available; otherwise, it is calculated theoretically. The decay of the CN by particle emission is calculated in the framework of Hauser-Feshbach formalism

TABLE I. Input parameters used for CASCADE calculations for the reactions $^{16}\text{O} + ^{12}\text{C}$ at beam energies 117, 125, 145, and 160 MeV.

Angular momentum distribution

- Critical angular momentum J_{cr} as in Table II
- Diffuseness parameter $\Delta J = 1\hbar$

OM potentials of the emitted LCP and neutrons

- (1) Neutrons: Wilmore and Hodgson [31]
- (2) Protons: Perey and Perey [32]
- (3) α particles: Huizenga and Igo [33]
- (4) Multiplication factor of the OM radius: RFACT = 1

Level-density parameters at low excitation: $E \leq 7.5$ MeV

- (1) Fermi-gas level-density formula with empirical parameters from Dilg *et al.* [34]

Level-density parameters at high excitation: $E \geq 15$ MeV

- (1) Fermi-gas level-density formula with parameters from LDM (Myers and Swiatecki [35])
- (2) Level-density parameter $a = A/8 \text{ MeV}^{-1}$

Yrast line

$$I_{\text{eff}} = I_0(1 + \delta_1 J^2 + \delta_2 J^4), \delta_1, \delta_2 \text{ are given in Table II}$$

γ -ray width (Weisskopf units)

- (1) $E1 = 0.001$
- (2) $M1 = 0.01$
- (3) $E2 = 5.0$

[26,29]. The probability that the parent nucleus (excitation energy E_1 , spin J_1 , parity π_1) emits a particle x with an orbital angular momentum l , kinetic energy ϵ_x ($\epsilon_x = E_1 - E_2$), and spin s is given by

$$P_x d\epsilon_x = \frac{1}{\hbar} \Gamma(\epsilon_x) = \frac{\rho_2(E_2, J_2, \pi_2)}{2\pi\hbar\rho_1(E_1, J_1, \pi_1)} \times \sum_{S=|J_2-s|}^{J_2+s} \sum_{l=|J_1-S|}^{J_1+S} T_l(\epsilon_x) d\epsilon_x, \quad (2)$$

where E_2 , J_2 , and π_2 are the excitation energy, spin, and parity of the daughter nucleus, respectively, ρ_1 and ρ_2 are the level densities of the excited parent nucleus and daughter nucleus, respectively, and, $S = J_2 + s$ is the channel spin. The transmission coefficients $T_l(\epsilon_x)$ for the scattering of particle x on the daughter nucleus (inverse process) are obtained by using standard optical model (OM) potential for elastic scattering. In the OM transmission coefficient calculation, the parameters have been taken from the references given in the Table I.

B. Input parameters

The standard form of CASCADE is quite successful in explaining the LCP evaporation in the light-ion-induced reaction, in general, where the CN is assumed to be nearly spherical. However, in case of heavy-ion-induced reaction, there is appreciable deviation between the experimental and the predicted LCP evaporation spectra. This deviation is attributed to the deformation of the excited compound system which is angular momentum dependent. Therefore, to explain the LCP energy spectra, the effects of the deformation of the CN

should be included in the statistical model calculations. The deformation affects the particle spectra in two ways. First, it lowers the effective emission barrier, and second, it increases the moment of inertia. The first effect modifies the transmission coefficients for the evaporated particles which may be taken care of by increasing the radius parameter of OM potential. However, the change in moment of inertia affects the level density and the slope of the particle spectrum. This can be taken care of by incorporating the spin-dependent deformability parameters [14,22,30]. For level density calculations, excited energies have been divided into three regions.

Region I (low excitation energy, $E \leq 3$ to 4 MeV): Here, the experimentally known discrete levels are used for all the nuclei produced in the cascade. In some cases, known high-spin states at higher excitation energy are included as yrast levels in region II.

Region II (medium excitation energy, $4 \text{ MeV} \leq E \leq 7.5 \text{ MeV}$): Analytical level density formula is used in this region. The parameters a and Δ are deduced empirically for each nucleus from the work of Vonach *et al.* [36] and Dilg *et al.* [34]. The excitation energy is corrected for the parity effects.

Region III (high excitation energy, $E \geq E_{\text{LDM}}$): Shell effects and parity corrections are neglected in this region. The same formula is then used but with LDM parameters taken from Ref. [35].

Between regions II and III, the level density parameters are interpolated linearly. The parameters are given in Table I. The level density used in regions II and III for a given angular momentum J and excitation energy E is given by well-known Fermi-gas expression [11] with equidistant single-particle levels and a constant level density parameter a :

$$\rho(E, J) = \frac{(2J+1)}{12} a^{1/2} \left(\frac{\hbar^2}{2I_{\text{eff}}} \right)^{3/2} \frac{1}{(E+T-\Delta-E_J)^2} \times \exp[2\{a(E-\Delta-E_J)\}^{1/2}], \quad (3)$$

where T is the thermodynamic temperature and Δ is the pairing correction. The rotational energy, E_J , is expressed as

$$E_J = \left(\frac{\hbar^2}{2I_{\text{eff}}} \right) J(J+1). \quad (4)$$

The effective moment of inertia, I_{eff} , is written as

$$I_{\text{eff}} = I_0(1 + \delta_1 J^2 + \delta_2 J^4), \quad (5)$$

where $I_0 (= \frac{2}{5} A^{5/3} r_0^2)$ is the rigid-body moment of inertia, δ_1 and δ_2 are deformability parameters, r_0 is radius parameter, and a is the level density parameter. So, from the above equations, it is clear that by changing r_0 , a , δ_1 , and δ_2 , it may be possible to reproduce the experimental spectra. By increasing r_0 , both transmission coefficient and level density will be affected. It reduces the potential barrier, leading to the increase of transmission coefficient. Simultaneously, the increase in r_0 leads to increase of the I_{eff} and vis-à-vis the available phase space. We have chosen $r_0 = 1.29$ which reproduced the lower energy part of the spectra. A similar value was used in Ref. [37]. The level density parameter a was taken to be $A/8$ as in previous works [13,30]. To reproduce the present experimental spectra, we have only changed the deformability parameters δ_1 and δ_2 , as in the previous works [14,28,30].

TABLE II. The values different sets of deformability parameters: A, obtained from RLDM; B, obtained by fitting the experimental data (see text). E_{lab} , E , J_{cr} , and J_{av} are the beam energy, excitation energy, critical angular momentum, and average angular momentum, respectively.

E_{lab}	E	J_{cr}	J_{av}	δ_1^A	δ_2^A	δ_1^B	δ_2^B
117	67	20	13	3.7×10^{-4}	1.1×10^{-6}	1.9×10^{-3}	2.0×10^{-8}
125	70	21	14	3.7×10^{-4}	1.1×10^{-6}	2.1×10^{-3}	2.0×10^{-8}
145	79	22	15	3.7×10^{-4}	1.1×10^{-6}	2.3×10^{-3}	2.0×10^{-8}
160	85	23	15	3.7×10^{-4}	1.1×10^{-6}	2.5×10^{-3}	2.0×10^{-8}

C. Experimental spectra and CASCADE predictions

The measured energy spectra have been compared with the respective CASCADE [26] calculations which have been shown in Fig. 1. The critical angular momenta, J_{cr} , used in the calculation, are 20, 21, 22, and 23(\hbar) for the bombarding energies of 117, 125, 145, and 160 MeV, respectively [12]. The other input parameters are given in Table I. The dash-dot-dashed lines represent the results of the CASCADE calculation with the radius parameter $r_0 = 1.29$ and the default values of the deformability parameters, $\delta_1^A = 3.7 \times 10^{-4}$ and $\delta_2^A = 1.1 \times 10^{-6}$, at all beam energies which have been predicted by the rotating liquid drop model [38]. The low-energy parts of the experimental α particle spectra are found to match with the theoretical spectra, but higher energy parts do not. To reproduce the whole spectra, we have followed the procedure proposed by Huizenga *et al.* [30]. The deformability parameters δ_1 and δ_2 have been suitably optimized to reproduce the experimental spectra, which in effect modified the phase space for statistical decay by relocation of the yrast line. The results with these modified deformability parameters, are shown by solid lines in Fig. 1 and the optimized values of deformability parameters, δ_1^B and δ_2^B , are given in Table II.

V. DISCUSSION

It is apparent from Table II that the deformability parameters had to be modified from the corresponding RLDM values considerably to fit the experimental spectra. Following the empirical procedure given in Refs. [10,13], it can be shown that the above change in the deformability parameters is indicative of substantial enhancement of the deformation of the excited composite over the corresponding ground state (RLDM) deformation. Before we look further into the origin of such deformation, it may be worthwhile to see if there is any other alternative explanation of the observed deviation of the experimental spectra from the respective standard statistical model prediction.

Though the present formalism of angular momentum dependent level density is largely successful in explaining experimental LCP spectra, the magnitude of enhancement required is quite large and lacking proper explanation [39,40]. So an alternative approach, based on frozen degrees of freedom, has been proposed, which has been shown to reproduce the data quite well [41,42]. In this approach, it is assumed that the deformation of the CN is frozen during the decay, i.e., there is no

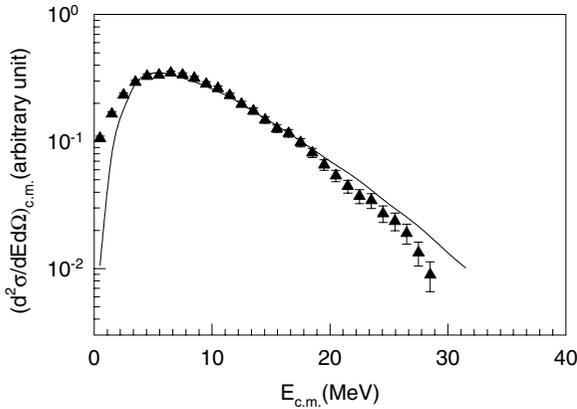


FIG. 5. Comparison of the experimental energy spectra (c.m.) of α particles (triangles) with the same obtained by theoretical calculations for beam energy 160 MeV. Solid curve represents the CASCADE calculations done using ‘frozen’ deformation (I_{eff}/I_0) obtained using RLDM deformability parameters for $J = 23\hbar$.

change of shape of the nascent final nucleus; so the phase space is calculated using RLDM deformation of the parent nucleus, rather than that of the usual daughter nucleus. This indicates that the dynamical effects such as shape relaxation should be taken into account to properly understand the phenomenon of particle evaporation from a deformed CN.

The formalism of “frozen deformation” has been applied in the present CASCADE calculation. In the case of α particle emission, the effect of frozen deformation on the energy spectrum may be taken into account in the following way. As the α particles are emitted predominantly at the initial stage of the decay cascade, the deformation may be “frozen” at its value corresponding to the highest angular momentum that the CN may have, which is $\approx J_{\text{cr}}$. So in the present CASCADE calculation, the deformation has been kept fixed throughout by freezing the value of I_{eff}/I_0 , which has been calculated using a fixed value of $J \approx J_{\text{cr}}$ using Eq. (5) with δ_1 and δ_2 obtained from RLDM (see Table II). The results of the CASCADE calculation with frozen deformation have been shown in Fig. 5 along with the experimental data. It is observed that the CASCADE prediction with frozen deformation [41,42] for $J = J_{\text{cr}} = 23\hbar$ (solid curve) is also in fair agreement with the experimental data.

To investigate further into the nature of deformation, we have computed the values of average effective radius parameter, $\langle r_{\text{eff}} \rangle$, for all sets of deformability parameters (Table II). Following [11], the effective radius parameter, r_{eff} , is defined as $r_{\text{eff}} = r_0 \sqrt{1 + \delta_1 J^2 + \delta_2 J^4}$, and the average effective radius parameter is calculated using the expression, $\langle r_{\text{eff}} \rangle = \sum_0^{j_{\text{cr}}} r_{\text{eff}}(2j + 1) / \sum_0^{j_{\text{cr}}} (2j + 1)$. Then the deviation of $\langle r_{\text{eff}} \rangle$ from r_0 is the measure of deformation, and the deviation of the same from the corresponding RLDM value may be considered as an enhancement. The values of r_{eff} estimated

in the case of frozen deformation approach varied between 1.48 (at 117 MeV) and 1.58 (at 160 MeV), which are, though higher than the corresponding RLDM values ($\langle r_{\text{eff}} \rangle^{\text{RLDM}} \sim 1.38\text{--}1.42$), quite lower than the values obtained with the best fitted parameter set ($\langle r_{\text{eff}} \rangle^{\text{fitted}} \sim 1.52\text{--}1.66$). From the above discussion, it may be said that, though the present study indicates deformation of the excited composite which is higher than the corresponding RLDM value, uncertainty still remains about the actual magnitude of deformation. This limitation of the present study notwithstanding, there is, at least qualitative, indication about some enhanced deformation, which may be linked with orbiting as one of the contributing factors.

VI. SUMMARY AND CONCLUSION

The energy and angular distribution of α particles, emitted in the reactions $^{16}\text{O}(117, 125, 145, \text{ and } 160 \text{ MeV}) + ^{12}\text{C}$, have been measured. It has been observed from the angular distribution and average velocity plots that the α particles are emitted from fully energy equilibrated sources, at all beam energies. The measured energy spectra have been compared with the same predicted by the standard statistical model calculations. It has been found that the experimental α particle energy spectra are properly explained by CASCADE by optimizing the deformability parameters, which are quite different from the respective RLDM values. This appears to be indicative of appreciable change (enhancement) of deformation. However, the understanding does not seem to be so simple and straightforward when one compares the above with the results of CASCADE calculation with frozen deformation, which has also been shown to be almost equally effective in explaining the data. In this case, the effective frozen deformation turned out to be smaller than that obtained using the optimized parameters. So, the uncertainties about the magnitudes of the actual CN deformations notwithstanding, it can, only qualitatively, be said that equilibrium orbiting, which is similar to particle evaporation in time scale, could also be one of the contributing factors for the observed deformation. However, the present models are too simplistic to predict the actual deformation of the CN; more realistic event-by-event Monte Carlo calculations, taking into account the initial deformation, spin distribution of the CN and their subsequent evolutions, should be performed to have a proper understanding of the CN deformation. In addition, new experimental inputs (such as measurement of deformation from GDR studies [43]) are needed for a more comprehensive understanding of the process.

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[1] D. Shapira, J. L. C. Ford, Jr., J. Gomez del Campo, R. G. Stokstad, and R. M. DeVries, *Phys. Rev. Lett.* **43**, 1781 (1979).

[2] D. Shapira, J. L. C. Ford, Jr., and J. Gomez del Campo, *Phys. Rev. C* **26**, 2470 (1982).

- [3] C. Bhattacharya, A. Dey, S. Kundu, K. Banerjee, S. Bhattacharya, S. Mukhopadhyay, D. Gupta, T. Bhattacharjee, S. R. Banerjee, S. Bhattacharyya, T. Rana, S. K. Basu, R. Saha, S. Bhattacharjee, K. Krishan, A. Mukherjee, D. Bandopadhyay, and C. Beck, *Phys. Rev. C* **72**, 021601(R) (2005).
- [4] Aparajita Dey, C. Bhattacharya, S. Bhattacharya, S. Kundu, K. Banerjee, S. Mukhopadhyay, D. Gupta, T. Bhattacharjee, S. R. Banerjee, S. Bhattacharyya, T. K. Rana, S. K. Basu, R. Saha, K. Krishan, A. Mukherjee, D. Bandopadhyay, and C. Beck, *Phys. Rev. C* **76**, 034608 (2007).
- [5] W. Dünneweber, A. Glaesner, W. Hering, D. Konnerth, R. Ritzka, W. Trombik, J. Czakanski, and W. Zipper, *Phys. Rev. Lett.* **61**, 927 (1988).
- [6] D. Shapira, R. Novotny, Y. D. Chan, K. A. Erb, and J. L. C. Ford, Jr., *Phys. Lett. B* **114**, 111 (1982).
- [7] D. Shapira, D. Schull, J. L. C. Ford, Jr., B. Shivakumar, R. L. Parks, R. A. Cecil, and S. T. Thornton, *Phys. Rev. Lett.* **53**, 1634 (1984).
- [8] S. J. Sanders, A. Szanto de Toledo, and C. Beck, *Phys. Rep.* **311**, 487 (1999), and reference therein.
- [9] B. Fornal *et al.*, *Phys. Rev. C* **44**, 2588 (1991).
- [10] M. Rousseau *et al.*, *Phys. Rev. C* **66**, 034612 (2002).
- [11] A. Dey, S. Bhattacharya, C. Bhattacharya, K. Banerjee, T. K. Rana, S. Kundu, S. Mukhopadhyay, D. Gupta, and R. Saha, *Phys. Rev. C* **74**, 044605 (2006).
- [12] S. Kundu, A. Dey, K. Banerjee, T. K. Rana, S. Mukhopadhyay, D. Gupta, R. Saha, S. Bhattacharya, and C. Bhattacharya, *Phys. Rev. C* **78**, 044601 (2008).
- [13] C. Bhattacharya *et al.*, *Phys. Rev. C* **65**, 014611 (2001).
- [14] I. M. Govil, J. R. Huizenga, W. U. Schröder, and J. Töke, *Phys. Lett. B* **197**, 515 (1987).
- [15] R. K. Choudhury, P. L. Gonthier, K. Hagel, M. N. Namboodiri, J. B. Natowitz, L. Adler, S. Simon, S. Kniffen, and G. Berkowitz, *Phys. Lett. B* **143**, 74 (1984).
- [16] G. LaRana, D. J. Moses, W. E. Parker, M. Kaplan, D. Logan, R. Lacey, J. M. Alexander, and R. J. Welberry, *Phys. Rev. C* **35**, 373 (1987).
- [17] G. LaRana, R. Moro, A. Brondi, P. Cuzzocrea, A. D'Onofrio, E. Perillo, M. Romano, F. Terrasi, E. Vardaci, and H. Dumont, *Phys. Rev. C* **37**, 1920 (1988).
- [18] B. Fornal, G. Prete, G. Nebbia, F. Trotti, G. Viesti, D. Fabris, K. Hagel, and J. B. Natowitz, *Phys. Rev. C* **37**, 2624 (1988).
- [19] B. Fornal, F. Gramegna, G. Prete, G. Nebbia, R. Smith, G. D'Erasmus, L. Fiore, A. Pantaleo, G. Viesti, P. Blasi, F. Lucarelli, I. Iori, and A. Moroni, *Phys. Rev. C* **41**, 127 (1990).
- [20] C. Bhattacharya, M. Rousseau, C. Beck, V. Rauch, R. Nouicer, R. M. Freeman, O. Stezowski, D. Mahboub, S. Belhabib, A. Hachem, E. Martin, A. Dummer, S. J. Sanders, and A. Szanto de Toledo, *Nucl. Phys. A* **654**, 841c (1999).
- [21] D. Bandyopadhyay, S. K. Basu, C. Bhattacharya, S. Bhattacharya, K. Krishan, A. Chatterjee, S. Kailas, A. Navin, and A. Srivastava, *Phys. Rev. C* **59**, 1179 (1999).
- [22] D. Bandyopadhyay, C. Bhattacharya, K. Krishan, S. Bhattacharya, S. K. Basu, A. Chatterjee, S. Kailas, A. Shrivastava, and K. Mahata, *Eur. Phys. J. A* **14**, 53 (2009).
- [23] W. E. Parker, M. Kaplan, D. J. Moses, G. LaRana, D. Logan, R. Lacey, J. M. Alexander, D. M. deCastroRizzo, P. DeYoung, R. J. Welberry, and J. T. Boger, *Phys. Rev. C* **44**, 774 (1991).
- [24] N. N. Ajitanand, G. La Rana, R. Lacey, David J. Moses, Louis C. Vaz, G. F. Peaslee, D. M. de Castro Rizzo, M. Kaplan, and J. M. Alexander, *Phys. Rev. C* **34**, 877 (1986).
- [25] M. Kildir, G. La Rana, R. Moro, A. Brondi, E. Vardaci, A. D'Onofrio, D. Fessas, E. Perillo, V. Roca, M. Romano, and F. Terrasi, G. Nebbia, G. Viesti, and G. Prete, *Phys. Rev. C* **51**, 1873 (1995).
- [26] F. Pühlhofer, *Nucl. Phys. A* **280**, 267 (1977).
- [27] N. N. Ajitanand, R. Lacey, G. F. Peaslee, E. Duek, and J. M. Alexander, *Nucl. Instr. Meth. A* **243**, 111 (1986).
- [28] G. Viesti, B. Fornal, D. Fabris, K. Hagel, J. B. Natowitz, G. Nebbia, G. Prete, and F. Trotti, *Phys. Rev. C* **38**, 2640 (1988).
- [29] W. Hauser and H. Feshbach, *Phys. Rev.* **87**, 366 (1952).
- [30] J. R. Huizenga, A. N. Behkami, I. M. Govil, W. U. Schröder, and J. Toke, *Phys. Rev. C* **40**, 668 (1989).
- [31] D. Wilmore and P. E. Hogson, *Nucl. Phys.* **55**, 673 (1964).
- [32] C. M. Perey and F. G. Perey, *At. Data Nucl. Data Tables* **17**, 1 (1976).
- [33] J. R. Huizenga and G. Igo, *Nucl. Phys.* **29**, 462 (1962).
- [34] W. Dilg, W. Schantl, H. Vonach, and M. Uhl, *Nucl. Phys. A* **217**, 269 (1973).
- [35] W. D. Myers and W. J. Swiatecki, *Nucl. Phys.* **81**, 1 (1966).
- [36] H. Vonach and M. Hille, *Nucl. Phys. A* **127**, 289 (1969).
- [37] D. Mahboub *et al.*, *Phys. Rev. C* **69**, 034616 (2004).
- [38] S. Cohen, F. Plasil, and W. J. Swiatecki, *Ann. Phys.* **82**, 557 (1974).
- [39] W. E. Ormand, P. F. Bortignon, A. Bracco, and R. A. Broglia, *Phys. Rev. C* **40**, 1510 (1989).
- [40] B. Lauritzen and G. Bertsch, *Phys. Rev. C* **39**, 2412 (1989), and references therein.
- [41] B. Fornal, G. Viesti, G. Nebbia, G. Prete, and J. B. Natowitz, *Phys. Rev. C* **40**, 664 (1989).
- [42] B. Fornal, F. Gramegna, G. Prete, R. Burch, G. D'Erasmus, E. M. Fiore, L. Fiore, A. Pantaleo, V. Paticchio *et al.*, *Phys. Lett. B* **255**, 325 (1991).
- [43] D. Pandit, S. Mukhopadhyay, S. Bhattacharya, S. Pal, A. De, S. Bhattacharya, C. Bhattacharya, K. Banerjee, S. Kundu, T. K. Rana, A. Dey, G. Mukherjee, T. Ghosh, D. Gupta, and S. R. Banerjee, *Phys. Rev. C* **81**, 061302 (2010).