









Extreme Light Infrastructure-Nuclear Physics (ELI-NP) - Phase II

The ELI-NP nuclear photonics program: A new source for nuclear structure data

Dimiter L. Balabanski

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IFIN-HP







Extreme Light Infrastructure – Nuclear Physics (ELI-NP)

<u>Mission:</u> Nuclear Physics studies with high-intensity lasers and brilliant γ beams

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 Nuclear Physics experiments to characterize laser – target interaction

- Photonuclear Physics
- Exotic Nuclear Physics and astrophysics
 complementary to other ESFRI Large Scale Physics Facilities (FAIR- Germany, SPIRAL2- France)
- Applications based on high intensity laser and very brilliant γ beams
 - ELI-NP in 'Nuclear Physics Long Range Plan in Europe' as a major facility

NP laboratory building





Platform supported on dampers

Anti–vibration platform ±1 μm @ < 10 Hz

Thermalized building $22^0 \pm 0.2^0$

Clean rooms

civil construction was commissioned in September 2016



Nuclear



ELI-NP Gamma Beam System (GBS)





Gamma Beam System

low-energy accelerator section: 0.2-3.5 MeV factory acceptance in Dec. 2015

high-energy accelerator section: 3.0-19.5 MeV



et Experiments with high-brilliance gamma beams at ELI-NP

Nuclear Physics

S. Gales et al., Phys. Scr. 91, 093004 (2016)



Nuclear Resonance Fluorescence (NRF) – Rom. Rep. Phys. 68, S483 (2016) Giant/Pigmy Resonances (GANT) – Rom. Rep. Phys. 68, S539 (2016) Photodisintegration (γ ,n), (γ ,p), (γ , α) – Rom. Rep. Phys. 68, S699 (2016) Photofission (γ ,ff) – Rom. Rep. Phys. 68, S621 (2016) Applications – Rom. Rep. Phys. 68, S735 (2016), *ibid* 68, S799 (2016), *ibid* 68, S847 (2016)



Rom. Rep. Phys. 68, S483 (2016)



ELI-NP NRF physics cases

- Self-absorption measurements (Γ_0/Γ_i)
- Low-energy dipole response (e.g. Actinides)
- Dipole response and parity measurements for weakly-bound nuclei
- Investigation of the Pigmy Dipole Resonance
- Rotational 2⁺ states of the scissor mode
- Constraints on the $0\nu\beta\beta$ -decay matrix elements of the scissors mode decay channel: ^{150}Sm



p-nuclei and actinides



Sensitivity frontier

week channels



Rom. Rep. Phys. 68, S539 (2016)

Nuclear Physics

GANT experiment at ELI-NP



ELIGANT-GN array 30 LaBr₃ or CeBr₃ 20 ⁷Li glasses 30 Lq. Scint.



Day ONE: studies of GDR and PDR decay (⁹⁰Zr, ²⁰⁸Pb)

- combine with information from (γ,n) experiments
- combine with information from (γ,γ') experiments (*e.g.* polarization)
- γ-decay to gs and excited states as a function of excitation energy



ELIADE





Neutron stars, equation of state and dipole polarizability @ELI-NP

-Neutron stars (NS) properties depend sensitively on the equation of state (EOS) of nuclear matter -EOS can affect many NS properties: mass-radius relationship, moment of inertia, cooling rates, Urca process, ... -It has been suggested that the slope (L) of the symmetry energy term of the EOS is closely related to the dipole polarizability α_{r} through the neutron skin thickness [1,2,3]



ELI-NP: experimental photo-nuclear reaction facility - The dipole polarizability is obtained from the photo-absorption cross section

$$\alpha_{D} = \frac{\hbar c}{2\pi^{2}} \int_{0}^{\infty} \frac{\sigma_{abs}}{\omega} d\omega = \frac{8\pi}{9} \int_{0}^{\infty} \frac{dB(E1)}{\omega}$$

150

-Strongly dependent on the low-energy strength, e.g. Pygmy resonance (see also FIG. 2) -ELI-NP will provide (accurate and unambiguous) measures of E1 strength below and above the neutron-threshold -Model independent results: pure electromagnetic excitation process

[1]P.-G. Reinhard and W. Nazarewicz, Phys. Rev. C81, 051303® (2010) [2] J. Piekarewicz, Phys. Rev. C83, 034319 (2011) [3] X. Roca-Maza et al., Phys. Rev. Lett. 106, 252501 (2011)



RCNP Osaka vs. ELI-NP experiments

RCNP

High-resolution (p,p') measurement at 0^o and forward angles A. Tamii, NIM A605, 326 (2009)

ELI-NP

High-resolution $(\gamma, \gamma') + (\gamma, n)$ measurement

<u>experiment</u>: polarized (>99%) γ beam simultaneous (γ , γ') + (γ ,n) measurement



Rom. Rep. Phys. 68, S539 (2016)

(γ,n) cross-section experiment at ELI-NP

Nuclear Physics





P-PROCESS NUCLEOSYNTHESIS FOR ¹⁸⁰Ta AND MEASUREMENTS OF THE PHOTO-NEUTRON CROSS SECTION

¹⁸⁰Ta characteristics

➤ Lowest natural abundancy (0.012%)
 ➤ Short-lived (T_{1/2} = 8.15h) J^π = 1⁺ ground state (¹⁸⁰Ta^g)
 ➤ Very long-lived (T^{1/2} > 10¹⁵ yr) J^π = 9⁻ isomeric state (¹⁸⁰Ta^m)
 ➤ ¹⁸¹Ta(γ,n)¹⁸⁰Ta and ¹⁸⁰Ta(γ,n)¹⁷⁹Ta photo-disintegration reactions





NuPECC LOng Range Plan 2016-2020 – Astrophysics

- Correct prediction of the ¹⁸⁰Ta^m yield highly requires both ¹⁸¹Ta(y,n)¹⁸⁰Ta and ¹⁸⁰Ta(y,n)¹⁷⁹Ta cross section measurements.
- The measurements for the (y,n) cross sections related to the p-nuclides destruction requires gamma ray beam three orders of magnitude higher than the existing ones.
- Measurements of the ¹⁸⁰Ta(y,n)¹⁷⁹Ta reaction are foreseen in the Day 1 experiment at ELI-NP facility by using the maximum available gamma ray energy of 19 MeV.



ELITPC

flagship experiment: ${}^{16}O(\gamma, \alpha){}^{12}C$



Detector upside-down view



The mini-eTPC detector with 256-channel readout was built and successfully tested in-beam at the IFIN Tandem in 2016





nuclear astrophysics with ELISSA

ELISSA:

- 3 rings of 12 position sensitive X3 silicon-strip detectors by Micron
- 2 end cap detectors from 4 QQQ3 segmented detectors by Micron
- 320 channels readout with GET electronics

⁷Li(γ,t)α

- reaction could still be a game changer in resolving the "Li problem"
- experimental measurements below 1.5 MeV are 30 yrs. old and disagree with theoretical predications
- higher energy measurements can restrict the extrapolation to astrophysically important energies
 - C. Matei et al., exp. at $HI\gamma S$ in March 2017



in collaboration with LNS Catania





DSSD testing at ELI-NP





ALTO, ARIEL, etc



ELI-NP





IGISOL beamline: Location





Location 1:

- CSC at 7m from IP \rightarrow A \approx 0.7cm
- maximum CSC length 1.5m
- crowded exp hall!

Location 2:

- CSC at 40m from IP \rightarrow A \approx 4cm
- plenty of space!



Rom. Rep. Phys. 68, S621 (2016)



IGISOL facility at ELI-NP







P. Constantin et al, NIM B 378, 78 (2016), ibid (2016) submitted



CSC Simulations: Fragment Slowing Down in the Gas Cell

Geant4: He, T=70K, p=300mbar (ρ =0.206mg/cm³) \rightarrow >95% of fragments stop in





CSC Simulations: Space Charge (I)

Divide CSC in 1x1x1 cm³ cells: 24x24x100 for ρ =0.21 mg/cm³, 40x40x100 for ρ =0.12 mg/cm³, 90x90x100 for ρ =0.05 mg/cm³;

Cummulate dE/dx deposited in 1s of beam and divide by W_i=41 eV.





Expected Rates

Rom. Rep. Phys. 68, S699 (2016)

Conservative "day-one": beam $5 \cdot 10^{10} \gamma/s$, target release eff. 25%, CSC extraction eff. 50% $\rightarrow \sim 10^7$ photofissions/s and $\sim (0.8-2) \cdot 10^6$ extracted ions/s

Optimal estimate: beam $10^{12} \gamma/s$, twice CSC extraction eff. \rightarrow expect ~2 orders of magnitude more!



Next phases of ELI-NP







The European initiative for Extreme Light Infrastructure laboratories in Romania (ELI-NP), will shortly provide tunable energy y-rays from inverse Compton scattering of laser light on a high-energy electron beam. This will allow Nuclear Resonance Object Witness foil Fluorescence studies of isotope-specific ELI-NP element distributions to be trace γ beam with performed unprecedented sensitivity. It is planned to use this Transmission powerful tool for cultural heritage object detector studies.

Rom. Rep. Phys. 68, S849 (2016)



Medical radioisotopes at ELI-NP

test case

•¹⁹⁵mPt: In chemotherapy of tumors it can be used to exclude "non responding" patients from unnecessary chemotherapy and optimizing the dose of all chemotherapy





feasibility study: Wen Luo et al., Appl. Phys. B 122, 8 (2016)



Human Resources



http://www.eli-np.ro/jobs.php







Sectoral Operational Programme "Increase of Economic Competitiveness" "Investments for Your Future!"



Extreme Light Infrastructure - Nuclear Physics (ELI-NP) - Phase II www.eli-np.ro



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Thank you!