# UPDATE OF THE EVALUATION OF THE CROSS SECTION OF THE NEUTRON DOSIMETRY REACTION ${ }^{103} \mathbf{R h}\left(n, n^{\prime}\right)^{103 m} \mathbf{R h}$ 

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# Update of the evaluation of the cross section of the neutron-dosimetry reaction ${ }^{103} \mathbf{R h}\left(n, n^{\prime}\right){ }^{103 \mathrm{~m}} \mathbf{R h}$ 

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#### Abstract

On the occasion of a new measurement of the excitation function of the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right)^{103 \mathrm{~m}} \mathrm{Rh}$ in the energy range between 5.69 and 12.0 MeV performed at the present institute in collaboration with the PTB Braunschweig, the cross section of this reaction, which is part of the International Reactor Dosimetry File (IRDF-90), was re-evaluated. Whereas the energy range of the evaluation, namely from threshold to 20 MeV , was kept unchanged with respect to IRDF-90, the underlying data base was extended by the experiment mentioned as well as by another measurement, and revised with regard to judgement and normalization of older data in the light of recent information. Based on the experimental data upgraded in this way, new model calculations were carried out, which in the energy region $14-20 \mathrm{MeV}$ served to supplement the experimental cross sections for this evaluation. The cross sections and their uncertainties were evaluated in energy groups with widths of 0.2 to 1.0 MeV , and the relative correlation matrix of the evaluated cross sections at the different energies was calculated. The results presented here supersede the corresponding values published in Physics Data 13-5 and included to the IRDF-90.


## 1. Introduction

The excitation function for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{1}\right)^{103 \mathrm{~m}} \mathrm{Rh}$ was first evaluated at the Institut für Radiumforschung und Kernphysik der Universität Wien (IRK) in the energy region from threshold ( 40 keV ) to 20 MeV in 1980 [1]. In the energy ranges 6.5 to 12.5 MeV and beyond 17 MeV this evaluation was based entirely on the results of nuclear model calculations. For an updated version of the evaluation of neutron-dosimetry cross sections [2], also meant for inclusion in the International Reactor Dosimetry File (IRDF90) [3], the recommended cross sections for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 m} \mathrm{Rh}$ were taken over unchanged, as no new measurements could be retrieved from EXFOR [4] in early 1990. In 1993, two of the present authors started a precision measurement of the ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ cross section in collaboration with the PTB Braunschweig. A short description of this experiment with preliminary results was presented at the Gatlinburg Conference in 1994 [5], a more complete report is available now [6] and will be published shortly [7]. - Apart from this measurement, another experiment between 3.74 and 5.18 MeV incident neutron energy, presented by Wu et al. [8] at the Mito

[^0]Conference in 1988 and included to EXFOR [4] in 1990 was added to the experimental data base.

The new experiment [ $5,6,7$ ] covers the energy range 5.69 to 12.0 MeV and triggered off a critical review of the ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 m} \mathrm{Rh}$ cross sections which had been measured previously and formed the data base for the evaluation in 1980 [1]. Consequently, also the reaction model calculations which had been part of the previous evaluation [1,2] were repeated as both the experimental ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ excitation function to be reproduced and the features adopted for the reaction models required some revision. The reaction model calculations were extended to comprise the major neutron-induced reactions on ${ }^{103} \mathrm{Rh}$ for which experimental data existed in a consistent theoretical description [9].

The present evaluation was carried out applying the procedures adopted previously for evaluations at the $\operatorname{IRK}[1,2,10]$. The experimental data underlying the evaluation, together with short summaries of newly included data sets as well as considerations regarding the older data sets, are presented in Sec. 2. The nuclear reaction model calculations used for extension of the evaluation beyond 14 MeV are discussed in Sec. 3. In Sec. 4, the evaluation and its results are described. In Sec. 5 we present the relative covariance matrix of the evaluated group cross sections and in Sec. 6 the integral cross section in the ${ }^{252} \mathrm{Cf}$ fission neutron spectrum.

## 2. Experimental data base

A literature search was performed using the bibliographical index CINDA [11] and the EXFOR data library [4]. In addition, the proceedings of recent international nuclear data conferences were also checked for relevant papers. The deadline for the literature search was the end of the year 1994. Beside the already mentioned experiments performed in collaboration between IRK and PTB Braunschweig (Miah 95) and by Wu et al. (Wu 88) no additional data sets for inclusion in the data base were identified.

The most important information on all experimental papers is summarized in Table 1. Two entries appear for the papers Santry 74, Barnard 78, and Paulsen 80. These authors had used two different methods to determine the neutron fluence. According to the evaluation procedures applied (see Ref. 10, Sec. II.), these papers were split into two parts for further processing. For experiments which are based on the measurement of the characteristic K x rays of Rh following the highly internally converted isomeric decay of ${ }^{103 \mathrm{~m}} \mathrm{Rh}$ the value of $0.0766 \pm 0.0014$ [12] for the number of $\mathrm{K} x$ rays emitted per decay and a half-life of $56.114 \pm 0.020 \mathrm{~min}$ [13] were adopted as standards.
A number of original values given for the cross sections or their uncertainties had to be renormalized according to the general procedures outlined in Ref. 10, Sec. II. Some obviously erroneous data points had to be rejected from the final data base. As the recommended values for decay data and reference cross sections have been changed
since 1980 renormalizations done by Strohmaier et al. [1] had to be revised in several cases. A critical review of all available data sets showed that corrections significantly different from those carried out in 1980 should be applied to the experiments Santry 74 and Paulsen 80. In the following paragraphs we describe all renormalization procedures in detail.

Nagel 66 and Kimura 69: Both data sets had already been rejected in the course of the work of Strohmaier et al. [1] due to strong deviations from all other measurements.
Pazsit 72: In this experiment both $\mathrm{K} x$ rays and $\gamma$ rays from the ${ }^{103 \mathrm{~m}} \mathrm{Rh}$ decay were measured. A renormalization due to changes in the decay scheme was not performed as no value for the fluorescence yield used by Pazsit et al. in the analysis of their data was given in the paper. Their result was, however, normalized in order to refer to the evaluated value of the reference cross section $\left({ }^{27} \mathrm{Al}(\mathrm{n}, \alpha){ }^{24} \mathrm{Na}\right)$ as given by Wagner et al. [2].
Santry 74: Santry and Butler had calibrated their x ray detector by a conversion electron - $x$ ray coincidence experiment using a chemically purified ${ }^{103 \mathrm{~m}} \mathrm{Rh}$ source. In this way, the "detector efficiency" (given by the number of detected $x$ rays per ${ }^{103 m} \mathrm{Rh}$ decay) was measured. With this technique the measurement became independent of conversion coefficients and fluorescence yields. On the other hand, it was possible to measure the $x$ ray emission probability comparing the measured "detector efficiency" with an efficiency value determined independently from x ray absorption coefficients. An x ray emission probability of $0.0697 \pm 0.0028$ was determined this way which is about $9 \%$ lower than the standard value used for this evaluation. According to our judgement this discrepancy stems most probably from a too low detector efficiency measured in the coincidence experiment. Therefore, we decided to apply a renormalization factor of 0.9099 to all data points of Santry 74 and keep the relative uncertainties as given in the original papers.
The cross section data measured relative to the ${ }^{32} \mathrm{~S}(\mathrm{n}, \mathrm{p})^{32} \mathrm{P}$ cross section (Santry 74A) were not corrected for changes in the reference cross section as we think the used reference cross sections are a better choice than using the cross section values recommended in ENDF/B-Vl. (See the detailed discussion in Ref. 2, p. 10.) As the uncertainty of the reference cross section is not included in the errors given, an uncertainty of $\pm 5 \%$ was added quadratically in order to obtain the total error.
Santry and Butler had used different neutron production reactions ( ${ }^{7} \mathrm{Li}(\mathrm{p}, \mathrm{n})^{7} \mathrm{Be}$, $T(p, n)^{3} \mathrm{He}, \mathrm{D}(\mathrm{d}, \mathrm{n})^{3} \mathrm{He}$ and $\mathrm{T}(\mathrm{d}, \mathrm{n})^{4} \mathrm{He}$ ) for different (overlapping) energy regions in their experiment. The cross sections for neutron energies above 8 MeV measured with the $\mathrm{D}(\mathrm{d}, \mathrm{n})^{3} \mathrm{He}$ source reaction had been corrected for activation of the Rh samples by deuterium-breakup neutrons. The correction factors given in Santry 74 were replaced by correction factors estimated in the recent IRK/PTB experiment [5,6]. With these corrections we found the decision made in the course of the 1980 evaluation to disregard the data points measured with the $\mathrm{D}(\mathrm{d}, \mathrm{n})^{3} \mathrm{He}$ source reaction no longer justified. These data now show reasonable agreement with the (also corrected, see below) data of Paulsen et al. (Paulsen 80) as well as with the recent results (Miah 95). On the other hand, the excitation function measured with the $\mathrm{T}(\mathrm{p}, \mathrm{n})^{3} \mathrm{He}$ source reaction exhibits a nearly
vanishing slope for neutron energies between 2.5 and 6 MeV . In this way, it deviates from the results given in Paulsen 80 in shape as well as in magnitude. Around 6 MeV , where both neutron producing reactions were used, disagreement exists between the cross sections measured with the $\mathrm{D}(\mathrm{d}, \mathrm{n})^{3} \mathrm{He}$ and the $\mathrm{T}(\mathrm{p}, \mathrm{n})^{3} \mathrm{He}$ source reaction. This had already been noted in the 1980 evaluation; however, the fact that owing to the renormalization to the new value of the x ray emission probability the part from the $\mathrm{D}(\mathrm{d}, \mathrm{n})^{3} \mathrm{He}$ source reaction is confirmed by the data of Paulsen 80 and Miah 95 for neutron energies greater than about 6 MeV now suggests to disregard all data points measured with the $T(p, n)^{3} \mathrm{He}$ source reaction at neutron energies above 2.5 MeV from the data sets Santry 74A and Santry 74B.
Pazsit 75: The cross section results were renormalized in order to refer to the x ray emission probability given in Ref. 12 and the reference cross section ( ${ }^{155} \mathrm{In}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{115 \mathrm{~m}} \mathrm{In}$ ) recommended by ENDF/B-VI [14]. The uncertainties were adjusted accordingly.
Barnard 78: The cross section values given in the original paper remained unchanged. The uncertainties of the data set Barnard 78 A had already been increased in the course of the work of Strohmaier et al. [1] to account for the uncertainties of the reference cross section.

Paulsen 80: The calibration of their x ray detector is finally based on an absolute measurement of the activity of a ${ }^{103 \mathrm{~m}} \mathrm{Rh}$ source by Vaninbroukx using liquid scintillation counting techniques [15]. This absolute measurement was also used to determine the K x ray emission probability in the isomeric decay of ${ }^{103 \mathrm{~m}} \mathrm{Rh}$ by calibrating the x ray detector conventionally using standard reference sources. The $\mathrm{K} x$ ray emission probability of $0.0843 \pm 0.0013$ reported by Vaninbroukx and Zehner [16] is about $10 \%$ higher than the standard value used in the present evaluation. Like in the case of Santry 74 we think that the discrepancy in the x ray emission probability stems from the absolute activity measurement on which the detector calibration for the cross section measurement is based. Therefore, we decided to apply a renormalization factor of 1.1005 to all cross sections from the data sets Paulsen 80 A and Paulsen $80 B$ and keep the relative uncertainties. With these corrections we now achieve good agreement within the uncertainty limits of the data given by Santry 74, Paulsen 80 and Miah 95, which are now based on the same value of the $\mathrm{K} x$ ray emission probability.
The cross section curve measured in the experiment Paulsen $80 B$ shows unrealistically high values at neutron energies above 14 MeV , probably caused by activation due to background neutrons of lower energy. Therefore, we rejected the data points at 15.0 $\mathrm{MeV}, 16.0 \mathrm{MeV}$, and 16.7 MeV and decided to use the results of the model calculations only in the energy range 15 to 20 MeV .
$W u 88$ : Wu et al. stated their results to be tentative and preliminary in their paper published in 1988, as higher accuracy might be achieved by further calibration measurements of their x ray detector. In 1990 these tentative data have become available in the EXFOR data library and to our knowledge no final results were ever published. As the uncertainty of the detector calibration is included in the final cross section uncertainty
given by Wu et al. we accepted this data set for our evaluation without further renormalizations. As no information on the decay parameters used is given in the paper of Wu et al. we could not decide if a normalization with respect to the x ray emission probability might be necessary.
Miah 95: No normalization procedures were applied to the data given in this paper.
In Fig. l, all accepted and renormalized data points are displayed. Here, the tag Strohmaier 95 indicates the results of the model calculations described in the following section. In Fig. 2, the same data are shown on an expanded scale in the energy range from 0 to 6 MeV .

## 3. Nuclear model calculations

Due to the features of the experimental data base, the 1980 evaluation [1,2] of the ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ excitation function included statistical model calculations in the energy regions $6.0-13.5 \mathrm{MeV}$ and $16.0-20.0 \mathrm{MeV}$. The parameters of the model calculations were adjusted to reproduce the experimental values in the $14-15 \mathrm{MeV}$ range.

For the updated evaluation [2] performed for the IRDF-90 [3], we checked not only the status of the experimental data, but also that of the calculations supplementing the evaluations. At that time, part of the parameters or information entering the ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{1}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ calculations appeared to be not quite to the state of the art. E.g., refined measuring techniques had yielded more comprehensive and more reliable data on low-lying levels and their $\gamma$-decay than had been available in 1980, optical-model analyses of neutron differential cross sections on ${ }^{103} \mathrm{Rh}$ had resulted in optical potentials appropriate especially for this nuclide, and the use of a pairing correction in preequilibrium-model particle-hole state densities is now common. - Also, with regard to $\gamma$-ray strength functions, prescribing the ratios of strengths of radiation of all multipole types to that of El radiation at the neütron binding energy may attribute too much strength to radiation of multipole types other than El at small $\gamma$-ray energies, if the strength function for El is derived from a giant-dipole resonance, but for the other multipole types is assumed energy independent according to the Weisskopf model. However, as in the 1980 calculations no reactions competing with ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ were considered, these possible deficiencies were evidently alleviated by an appropriate choice of the rest of the parameters, as the calculated excitation function described the accepted experimental data very well. Therefore, in connection with the evaluation update for Refs. 2 and 3, it was decided not to redo the calculations as long as no esssential changes in the data base occurred.

When this was the case in 1994, model calculations were carried out again, superior to the older ones as the improvements mentioned above were considered, and aiming at a consistent description of the experimental cross sections for the ( $\mathrm{n}, 2 \mathrm{n}$ ), ( $\mathrm{n}, 3 \mathrm{n}$ ), ( $\mathrm{n}, \mathrm{p}$ ) and $\left(\mathrm{n}, \alpha\right.$ ) reactions up to 30 MeV together with the $\left(\mathrm{n}, \mathrm{n}^{\prime}\right)^{\mathrm{m}}$ reaction.

The details of the choice of the model parameters are discussed in Ref. 9. The main changes with respect to the 1980 calculations are the use of the coupled-channel optical model together with a neutron optical potential adjusted individually for ${ }^{103} \mathrm{Rh}$ [17] for generating the neutron transmission coefficients in the incoming channels. In the outgoing neutron channels, the same optical potential with an increased imaginary part was used in single-channel mode. For charged particles, the spherical optical model was applied with the global optical potentials of Mani et al. [18] for protons and of Huizenga and Igo [19] for $\alpha$-particles.

In the exciton model, we used a pairing shift in the particle-hole state densities and an exciton-number dependent matrix element in the internal transition rates.
The level schemes and decay properties were updated on the basis of Nuclear Data Sheets [13,20]. For determining the $\gamma$-ray strength functions, a giant dipole model was applied for E1 as well as M1 radiation, with normalization factors chosen such as to achieve agreement with the corresponding strength function values compiled in Ref. 21.

The resulting excitation functions, also for the other reactions considered, are displayed in Ref. 9. For the ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{1}\right)^{103 \mathrm{~m}} \mathrm{Rh}$ reaction, the excitation function as evaluated from all accepted experimental data is well described by the theoretical result in the energy regions from threshold to 2 MeV and between 6 and 14 MeV . In Ref. 9, the calculation overestimates the experimental evaluation in the energy range $2-6 \mathrm{MeV}$, where the latter has a shape problem. This has meanwhile been solved by omission of part of the data points measured with the $T(p, n)^{3} H e$ source reaction of the set Santry $74 B$ as expounded in Sec. 2. The overshoot of the experimental values over the theoretical ones above 14 MeV in Fig. 1 of Ref. 9 is due to the inclusion of the data points of Paulsen $80 B$ in this energy range which were rejected for the present evaluation.
The influence of the model calculations on the evaluation is reduced with respect to the 1980 version [ $1,2,3$ ] as the theoretical curve is used between 14 and 20 MeV only.
The parameter variations which had been performed in 1980 to derive uncertainties of the calculated cross sections were not repeated for the new choice of parameters. Instead, we retained the relative uncertainties of the 1980 calculations [1,2,3] for the 1995 results.

## 4. Evaluation and results thereof

All accepted experimental data points as well as the results of the model calculation included in the data base are listed in Table 2. For each data point the following quantities are given:

- the average neutron energy, the energy spread (half-width at half maximum) and the uncertainty of the average neutron energy;
- the cross section value and uncertainty as given by the authors;
- an indication which renormalization procedures had been applied, and finally,
- the renormalized cross section and its uncertainty.

The neutron energy spread and the uncertainty of the average neutron energy, which were not reported in several papers, were estimated according to the experimental conditions.
For the evaluation procedure, the energy range between 0.1 and 20.0 MeV was divided into 31 groups. The widths of these groups are based on the density of the data points, the shape of the excitation function and the average neutron energy resolution in the respective group. The results of the evaluation, i.e., the cross section averages over the individual groups and their uncertainties are listed in Table 3. The last column of the table gives the ratios of the external to the internal errors, indicating good agreement amongst the data. In Fig. 3, the results of the present evaluation are compared with those of the previous one [ $1,2,3$ ]. There is an essential increase in magnitude of the newly evaluated excitation function compared to the preceding edition in the plateau and the descending part, viz. between 2 and 12 MeV , which is plausible from the revision of the data base used previously and the effect of the newly added data set from the IRK/PTB collaboration. In Fig. 4, we compare the evaluated ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right)^{103 \mathrm{~m}} \mathrm{Rh}$ cross sections with those of the nuclear reaction-model calculations. The overall agreement is quite satisfactory, with a slight deviation remaining between 2 and 4 MeV incident neutron energy.

## 5. Relative covariance matrix

Relative correlation coefficients and covariances between the cross sections in the energy groups used in the evaluation were computed for the excitation function of the ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ reaction. The correlation coefficients were derived from the quantities $\mathrm{B}_{\text {nnk }}$, the average correlation coefficients within each data set k (see Ref. 10, Sec. II.2.). Estimates of these coefficients $\mathrm{B}_{\text {nnk }}$ are compiled in Table 4 for all included data sets which contain more than one data point; in case of measurements they were based on the information on experimental uncertainties relevant in the respective works.
Whereas in the 1980 evaluation the elements of

Table 4: Average relative correlation coefficients $\mathrm{B}_{\text {mnk }}$ assumed for the various data sets.

| Reference | $\mathrm{B}_{\text {nnk }}$ |
| :--- | :--- |
| Santry 74A | 0.30 |
| Santry 74B | 0.65 |
| Pazsit 75 | 0.40 |
| Barnard 78A | 0.70 |
| Barnard 78B | 0.70 |
| Paulsen 80A | 0.71 |
| Paulsen 80B | 0.97 |
| Wu 88 | 0.80 |
| Miah 95 | 0.50 |
| Strohmaier 95 | 0.95 | the covariance matrix in the higher-energy part of the excitation function, where model calculations entered the evaluation, were derived from the results of the cross-section

variations according to a procedure described in Sec. II.1.2. of Ref. 1, in the present work, all covariance-matrix elements were computed from correlation coefficients throughout the energy region. The $\mathrm{B}_{\text {nnk }}$ value for the theoretical calculation was assumed to be 0.95 , expressing nearly full correlation between the calculated cross sections in different energy groups.
The relative covariance matrix for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ is given in Table 5 .

## 6. Comparison of spectrum averaged cross sections in the ${ }^{252} \mathbf{C f}$ fission neutron spectrum

The evaluated excitation function for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ was used for calculating a fission spectrum average, $\langle\sigma\rangle$, in the neutron spectrum of ${ }^{252} \mathrm{Cf}$ for comparison with measured values of this quantity. For the description of the Cf neutron spectrum, we used the evaluation of Mannhart [22]. The experimental cross sections < $\sigma>$ were again converted to a common value of $0.0766 \pm 0.0014$ [12] for the number of $\mathrm{K} x$ rays per decay of ${ }^{103 m} \mathrm{Rh}$. In the case of the result of Pazsit et al. [23], renormalization for the cross sections of the neutron flux monitor reaction was performed, too, using the evaluated value for the spectrum-averaged cross section of the monitor reaction ${ }^{115} \operatorname{In}\left(n, n^{\prime}\right){ }^{1115 m}$ In given by Mannhart [24]. In Table 6, the integral cross section calculated from the present evaluation is given together with various experimental values. The uncertainty of $\langle\sigma\rangle$ computed from the evaluated excitation function is made up of the uncertainties of the ${ }^{252} \mathrm{Cf}$ neutron spectrum and those of the evaluated cross sections taking into account the correlations between the data for the different energy groups. As can be seen from Table 6, the discrepancy between the experimental values of $\langle\sigma\rangle$ is increased by the renormalization. Our calculated result is a fair reproduction of the mean value of the measured < $\sigma>$ data.

Table 6: Calculated average ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right)^{103 \mathrm{~m}} \mathrm{Rh}$ cross section in the ${ }^{252} \mathrm{Cf}$ neutron spectrum compared with experimental data for $\langle\sigma>$.

| $<\sigma>_{\text {orig }}$ <br> $[\mathrm{mb}]$ | $<\sigma\rangle_{\text {renorm }}$ <br> $[\mathrm{mb}]$ | Origin | $1^{\text {st }}$ author, yr | Ref. |
| :---: | :---: | :--- | :--- | :---: |
| $647 \pm 70$ | $621 \pm 67$ | exp. | Kirouac 74 | $[25]$ |
| $757 \pm 53$ | $741 \pm 39$ | exp. | Pazsit 75 | $[23]$ |
| $739 \pm 22$ | $813 \pm 26$ | exp. | Lamaze 88 | $[26]$ |
| $750 \pm 27$ |  | calc. | Pavlik 95 | this work |

Table 1: Summary of experiments for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right)^{103 \mathrm{~m}} \mathrm{Rh}$.

| Energy range [MeV] |  | Nr. of data points | Average accur. [\%] | Method | Flux, reference cross section | Ref. | Nr . |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 14.20 | 14.20 | 2 |  | Act., NaI, x rays | ${ }^{65} \mathrm{Cu}(\mathrm{n}, 2 \mathrm{n}) 945 \mathrm{mb},{ }^{56} \mathrm{Fe}(\mathrm{n}, \mathrm{p}) 109 \mathrm{mb}$ | Nagel 66 | 210 |
| 0.18 | 4.60 | 18 |  | Act., Nal, x rays | Calculated from prot. or deut. beam current | Kimura 69 | 209 |
| 14.70 | 14.70 | 1 |  | Act., $\mathrm{Si}(\mathrm{Li})$, x rays | ${ }^{27} \mathrm{Al}(\mathrm{n}, \alpha) 114 \pm 4 \mathrm{mb}$ | Pazsit 72 | 203 |
| 4.80 | 14.74 | 33 |  | Act., NaI, x rays | ${ }^{32} \mathrm{~S}(\mathrm{n}, \mathrm{p})$ excitation function | Santry 74A | 201 |
| 0.12 | 6.00 | 46 |  | Act., Nal, x rays | Calibrated long counter | Santry 74B | 202 |
| 2.70 | 14.80 | 2 |  | Act., Si(Li), x rays | ${ }^{115} \mathrm{In}\left(\mathrm{n}, \mathrm{n}^{\prime}\right) 340^{\circ} \pm 30 \mathrm{mb} ; 63 \pm 6 \mathrm{mb}$ | Pazsit 75 | 207 |
| 0.55 | 1.50 | 10 |  | Inel. scatt. neutr. in liq. sc. | Carbon elastic scattering | Barnard 78A | 205 |
| 1.10 | 1.93 | 8 |  | Inelast. $\gamma$ with $\mathrm{Ge}(\mathrm{Li})$ | ${ }^{92} \mathrm{Zr}\left(\mathrm{n}, \mathrm{n}^{\prime} \gamma\right)$ and ${ }^{94} \mathrm{Zr}\left(\mathrm{n}, \mathrm{n}^{\prime} \gamma\right)$ | Barnard 78B | 206 |
| 0.20 | 6.10 | 62 |  | Act., Nal, x rays | Rel. exc. fct., norm. to n-p scatt. at 1.8 MeV | Paulsen 80A | 204 |
| 3.00 | 16.70 | 5 |  | Act., NaI, x rays | Rel. exc. fct., norm. to Paulsen 80A | Paulsen 80B | 208 |
| 3.74 | 5.18 | 9 | 9.7 | Act., Si(Li), x rays | Proton recoil telescope | Wu 88 | 212 |
| 5.69 | 12.00 | 14 | 4.6 | Act., Si(Li), x rays | ${ }^{238} \mathrm{U}(\mathrm{n}, \mathrm{f}), \mathrm{ENDF} / \mathrm{B}-\mathrm{VI}$ | Miah 95 | 213 |
| 14.14 | 19.79 | 15 |  | Model calculation | Code "STAPRE" | Strohmaier 95 | 214 |

Table 2: Cross section data for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$.

| NR. | E-NEUTR [MEV] | WIDTH [MEV] | ERR.CENTR [MEV] | X-SECT(ORIG) [MB] | $\begin{gathered} \text { ERR (ORIG) } \\ {[\mathrm{MB}]} \end{gathered}$ | CORR.APPL. | $\underset{[M B]}{X-\operatorname{SECT}(F I N)}$ | ERR(FIN) [MB] | REF. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 0.12 | 0.020 | 0.002 | 12.900 | 1.000 | 1) | 11.738 | 0.910 | SANTRY 74B |
| 2 | 0.15 | 0.040 | 0.004 | 26.400 | 2.700 | 1) | 24.021 | 2.457 | SANTRY 74B |
| 3 | 0.15 | 0.020 | 0.002 | 21.600 | 1.300 | 1) | 19.654 | 1.183 | SANTRY 74B |
| 4 | 0.20 | 0.024 | 0.003 | 74.000 | 5.900 | 1) | 81.437 | 6.515 | PAULSEN 80a |
| 5 | 0.20 | 0.020 | 0.002 | 40.500 | 3.000 | 1) | 36.851 | 2.730 | SANTRY 74B |
| 6 | 0.22 | 0.040 | 0.004 | 43.100 | 3.000 | 1) | 39.217 | 2.730 | SANTRY 74B |
| 7 | 0.25 | 0.020 | 0.002 | 63.100 | 3.100 | 1) | 57.415 | 2.821 | SANTRY 74B |
| 8 | 0.25 | 0.026 | 0.003 | 103.000 | 7.400 | 1) | 113.352 | 8.161 | PAULSEN 80A |
| 9 | 0.28 | 0.040 | 0.004 | 83.800 | 3.300 | 1) | 76.250 | 3.003 | SANTRY 748 |
| 10 | 0.30 | 0.020 | 0.002 | 81.700 | 4.500 | 1) | 74.339 | 4.095 | SANTRY 74B |
| 11 | 0.30 | 0.025 | 0.004 | 119.000 | 8.300 | 1) | 130.960 | 9.167 | PAULSEN 80A |
| 12 | 0.30 | 0.030 | 0.003 | 94.500 | 5.400 | 1) | 85.986 | 4.913 | SANTRY 74B |
| 13 | 0.35 | 0.040 | 0.004 | 100.000 | 6.000 | 1) | 90.990 | 5.459 | SANTRY 74B |
| 14 | 0.35 | 0.021 | 0.003 | 124.000 | 9.300 | 1) | 136.462 | 10.235 | PAULSEN 80A |
| 15 | 0.35 | 0.030 | 0.003 | 113.000 | 5.000 | 1) | 102.819 | 4.550 | SANTRY 74B |
| 16 | 0.40 | 0.027 | 0.004 | 129.000 | 9.200 | 1) | 141.964 | 10.124 | PAULSEN 80A |
| 17 | 0.40 | 0.040 | 0.004 | 123.000 | 6.000 | 1) | 111.918 | 5.459 | SANTRY 74B |
| 18 | 0.40 | 0.030 | 0.003 | 122.000 | 6.000 | 1) | 111.008 | 5.459 | SANTRY 74B |
| 19 | 0.41 | 0.030 | 0.003 | 135.00 | 10.000 | 1) | 122.837 | 9.099 | SANTRY 74B |
| 20 | 0.45 | 0.040 | 0.004 | 131.000 | 6.000 | 1) | 119.197 | 5.459 | SANTRY 74B |
| 21 | 0.45 | 0.020 | 0.003 | 146.000 | 7.000 | 1) | 132.845 | 6.369 | SANTRY 74B |
| 22 | 0.46 | 0.020 | 0.003 | 144.000 | 7.000 | 1) | 131.026 | 6.369 | SANTRY 748 |
| 23 | 0.50 | 0.030 | 0.003 | 146.000 | 7.000 | 1) | 132.845 | 6.369 | SANTRY 74B |
| 24 | 0.50 | 0.058 | 0.005 | 151.000 | 9.100 | 1) | 166.175 | 10.014 | PAULSEN 80A |
| 25 | 0.55 | 0.015 | 0.020 | 205.000 | 20.000 | 2) | 205.000 | 20.000 | BARNARD 78A |
| 26 | 0.55 | 0.030 | 0.003 | 180.000 | 8.000 | 1) | 163.782 | 7.279 | SANTRY 74B |
| 27 | 0.57 | 0.100 | 0.010 | 247.000 | 12.000 | 1) | 224.745 | 10.919 | SANTRY 74B |
| 28 | 0.60 | 0.015 | 0.020 | 260.000 | 20.000 | 2) | 260.000 | 20.000 | BARNARD 78A |
| 29 | 0.60 | 0.040 | 0.005 | 238.000 | 12.100 | 1) | 261.919 | 13.316 | PAULSEN 80a |
| 30 | 0.65 | 0.030 | 0.003 | 300.000 | 14.000 | 1) | 272.970 | 12.739 | SANTRY 74B |
| 31 | 0.70 | 0.030 | 0.003 | 379.000 | 18.000 | 1) | 344.852 | 16.378 | SANTRY 74B |
| 32 | 0.70 | 0.069 | 0.005 | 368.000 | 16.900 | 1) | 404.984 | 18.598 | PAULSEN 80A |
| 33 | 0.77 | 0.100 | 0.010 | 450.000 | 22.000 | 1) | 409.455 | 20.018 | SANTRY 74B |
| 34 | 0.79 | 0.030 | 0.003 | 546.000 | 27.000 | 1) | 496.805 | 24.567 | SANTRY 74B |
| 35 | 0.80 | 0.054 | 0.007 | 525.000 | 24.200 | 1) | 577.763 | 26.632 | PAULSEN 80A |
| 36 | 0.90 | 0.030 | 0.003 | 622.000 | 27.000 | 1) | 565.958 | 24.567 | SANTRY 74B |
| 37 | 0.90 | 0.098 | 0.008 | 593.000 | 27.300 | 1) | 652.596 | 30.044 | PAULSEN 80A |
| 38 | 0.92 | 0.015 | 0.020 | 555.000 | 29.000 | 2) | 555.000 | 44.000 | BARNARD 78A |
| 39 | 0.99 | 0.030 | 0.003 | 623.000 | 28.000 | 1) | 566.868 | 25.477 | SANTRY 74B |
| 40 | 1.00 | 0.090 | 0.009 | 633.000 | 28.000 | 1) | 575.967 | 25.477 | SANTRY 74B |
| 41 | 1.01 | 0.106 | 0.008 | 613.000 | 28.200 | 1) | 674.607 | 31.034 | PAULSEN 80A |
| 42 | 1.02 | 0.015 | 0.020 | 593.000 | 29.000 | 2) | 593.000 | 48.000 | BARNARD 78A |
| 43 | 1.10 | 0.015 | 0.020 | 606.000 | 35.000 | 2) | 606.000 | 50.000 | BARNARD 78A |
| 44 | 1.10 | 0.030 | 0.003 | 505.000 | 76.000 | NONE | 505.000 | 76.000 | BARNARD 78B |
| 45 | 1.10 | 0.110 | 0.010 | 576.000 | 28.200 | 1) | 633.888 | 31.034 | PAULSEN 80A |
| 46 | 1.15 | 0.030 | 0.003 | 673.000 | 30.000 | 1) | 612.363 | 27.297 | SANTRY 74B |
| 47 | 1.15 | 0.015 | 0.020 | 619.000 | 29.000 | 2) | 619.000 | 50.000 | BARNARD 78A |
| 48 | 1.18 | 0.030 | 0.003 | 533.000 | 80.000 | NONE | 533.000 | 80.000 | BARNARD 78B |
| 49 | 1.19 | 0.080 | 0.008 | 675.000 | 30.000 | 1) | 614.182 | 27.297 | SANTRY 74B |
| 50 | 1.20 | 0.015 | 0.020 | 590.000 | 29.000 | 2) | 590.000 | 48.000 | BARNARD 78A |
| 51 | 1.20 | 0.110 | 0.010 | 631.000 | 30.900 | 1) | 694.415 | 34.005 | PAULSEN 80A |
| 52 | 1.30 | 0.015 | 0.020 | 643.000 | 29.000 | 2) | 643.000 | 52.000 | BARNARD 78A |
| 53 | 1.30 | 0.030 | 0.003 | 585.000 | 88.000 | NONE | 585.000 | 88.000 | BARNARD 78B |
| 54 | 1.30 | 0.110 | 0.010 | 619.000 | 29.100 | 1) | 681.209 | 32.024 | PAULSEN 80A |
| 55 | 1.35 | 0.030 | 0.003 | 706.000 | 106.000 | NONE | 706.000 | 106.000 | BARNARD 78B |
| 56 | 1.36 | 0.030 | 0.003 | 672.000 | 31.000 | 1) | 611.453 | 28.207 | SANTRY 74B |
| 57 | 1.40 | 0.015 | 0.020 | 698.000 | 30.000 | 2) | 698.000 | 56.000 | BARNARD 78A |
| 58 | 1.40 | 0.140 | 0.010 | 726.000 | 33.400 | 1) | 798.963 | 36.757 | PAULSEN 80A |
| 59 | 1.50 | 0.080 | 0.008 | 747.000 | 33.000 | 1) | 679.695 | 30.027 | SANTRY 74B |
| 60 | 1.50 | 0.015 | 0.020 | 605.000 | 35.000 | 2) | 605.000 | 50.000 | BARNARD 78A |
| 61 | 1.50 | 0.030 | 0.003 | 607.000 | 91.000 | NONE | 607.000 | 91.000 | BARNARD 78B |
| 62 | 1.50 | 0.090 | 0.010 | 729.000 | 34.300 | 1) | 802.264 | 37.747 | PAULSEN 80A |
| 63 | 1.55 | 0.030 | 0.003 | 668.000 | 100.000 | NONE | 668.000 | 100.000 | BARNARD 788 |
| 64 | 1.60 | 0.080 | 0.010 | 718.000 | 33.000 | 1) | 790.159 | 36.316 | PAULSEN 80A |
| 65 | 1.70 | 0.030 | 0.003 | 714.000 | 107.000 | NONE | 714.000 | 107.000 | BARNARD 78B |

Table 2: Cross section data for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ (contd.).

| NR. | E-NEUTR [MEV] | HIDTH <br> [MEV] | ERR.CENTR [MEV] | $\begin{gathered} X-S E C T(O R I G) \\ {[M B]} \end{gathered}$ | ERR(ORIG) [MB] | CORR.APPL. | $\begin{gathered} \mathrm{X}-\mathrm{SECT}(F I N) \\ \text { [MB] } \end{gathered}$ | ERR(FIN) <br> [MB] | REF. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 66 | 1.70 | 0.170 | 0.010 | 805.000 | 37.000 | 1) | 885.902 | 40.718 | PAULSEN 80A |
| 67 | 1.80 | 0.180 | 0.010 | 817.000 | 29.400 | 1) | 899.108 | 32.355 | PAULSEN 80A |
| 68 | 1.85 | 0.070 | 0.007 | 849.000 | 38.000 | 1) | 772.505 | 34.576 | SANTRY 74B |
| 69 | 1.90 | 0.120 | 0.010 | 800.000 | 37.600 | 1) | 880.400 | 41.379 | PAULSEN 80A |
| 70 | 1.93 | 0.030 | 0.003 | 622.000 | 93.000 | NONE | 622.000 | 93.000 | BARNARD 78B |
| 71 | 2.00 | 0.100 | 0.010 | 844.000 | 39.700 | 1) | 928.822 | 43.689 | PAULSEN 80A |
| 72 | 2.10 | 0.180 | 0.010 | 867.000 | 39.900 | 1) | 954.133 | 43.910 | PAULSEN 80A |
| 73 | 2.20 | 0.100 | 0.010 | 885.000 | 40.700 | 1) | 973.943 | 44.790 | PAULSEN 80A |
| 74 | 2.30 | 0.140 | 0.010 | 916.000 | 44.000 | 1) | 1008.060 | 48.422 | PAULSEN 80A |
| 75 | 2.40 | 0.170 | 0.010 | 927.000 | 42.600 | 1) | 1020.160 | 46.881 | PAULSEN 80A |
| 76 | 2.50 | 0.160 | 0.010 | 898.000 | 41.300 | 1) | 988.249 | 45.451 | PAULSEN 80A |
| 77 | 2.60 | 0.150 | 0.010 | 994.000 | 45.700 | 1) | 1093.900 | 50.293 | PAULSEN 80A |
| 78 | 2.70 | 0.200 | 0.020 | 999.000 | 111.000 | 1)3) | 941.957 | 66.530 | PAZSIT 75 |
| 79 | 2.70 | 0.130 | 0.010 | 974.000 | 44.800 | 1) | 1071.890 | 49.302 | PAULSEN 80A |
| 80 | 2.80 | 0.170 | 0.010 | 950.000 | 43.700 | 1) | 1045.480 | 48.092 | PAULSEN 80a |
| 81 | 2.90 | 0.180 | 0.010 | 1059.000 | 47.700 | 1) | 1165.430 | 52.494 | PAULSEN 80A |
| 82 | 3.00 | 0.190 | 0.020 | 1034.000 | 52.700 | 1) | 1137.920 | 57.996 | PAULSEN 80A |
| 83 | 3.10 | 0.200 | 0.020 | 1039.000 | 50.900 | 1) | 1143.420 | 56.015 | PAULSEN 80A |
| 84 | 3.20 | 0.210 | 0.020 | 1052.000 | 50.500 | 1) | 1157.730 | 55.571 | PAULSEN 80A |
| 85 | 3.30 | 0.210 | 0.020 | 1000.000 | 51.000 | 1) | 1100.500 | 56.125 | PAULSEN 80A |
| 86 | 3.40 | 0.220 | 0.020 | 1038.000 | 50.900 | 1) | 1142.320 | 55.974 | PAULSEN 80A |
| 87 | 3.50 | 0.230 | 0.020 | 1049.000 | 51.400 | 1) | 1154.420 | 56.566 | PAULSEN 80A |
| 88 | 3.60 | 0.240 | 0.020 | 1022.000 | 49.100 | 1) | 1124.710 | 53.986 | PAULSEN 80A |
| 89 | 3.70 | 0.250 | 0.030 | 1031.000 | 50.500 | 1) | 1134.620 | 55.596 | PAULSEN 80A |
| 90 | 3.74 | 0.130 | 0.013 | 1170.000 | 116.000 | NONE | 1170.000 | 116.000 | WU B8 |
| 91 | 3.80 | 0.260 | 0.030 | 1075.000 | 51.600 | 1) | 1183.040 | 56.785 | PAULSEN 80A |
| 92 | 3.90 | 0.270 | 0.030 | 1045.000 | 50.200 | 1) | 1150.020 | 55.201 | PAULSEN 80A |
| 93 | 4.00 | 0.280 | 0.030 | 1077.000 | 50.600 | 1) | 1185.240 | 55.706 | PAULSEN 80A |
| 94 | 4.07 | 0.140 | 0.014 | 1013.000 | 103.000 | NONE | 1013.000 | 103.000 | WU 88 |
| 95 | 4.10 | 0.290 | 0.030 | 1094.000 | 51.400 | 1) | 1203.950 | 56.585 | PAULSEN 80A |
| 96 | 4.13 | 0.120 | 0.012 | 1202.000 | 121.000 | NONE | 1201.000 | 121.000 | WU 88 |
| 97 | 4.20 | 0.340 | 0.030 | 1140.000 | 69.500 | 1) | 1254.570 | 76.529 | PAULSEN 80A |
| 98 | 4.28 | 0.110 | 0.011 | 1164.000 | 101.000 | NONE | 1164.000 | 101.000 | WU 88 |
| 99 | 4.30 | 0.340 | 0.030 | 1250.000 | 73.800 | 1) | 1375.630 | 81.162 | PAULSEN 80A |
| 100 | 4.37 | 0.160 | 0.016 | 1124.000 | 115.000 | NONE | 1124.000 | 115.000 | WU 88 |
| 101 | 4.40 | 0.340 | 0.030 | 1122.000 | 70.700 | 1) | 1234.760 | 77.789 | PAULSEN 80A |
| 102 | 4.50 | 0.340 | 0.030 | 1066.000 | 62.900 | 1) | 1173.130 | 69.215 | PAULSEN 80A |
| 103 | 4.57 | 0.100 | 0.010 | 1243.000 | 122.000 | NONE | 1243.000 | 122.000 | WU 88 |
| 104 | 4.60 | 0.340 | 0.030 | 1064.000 | 61.700 | 1) | 1170.930 | 67.914 | PAULSEN 80A |
| 105 | 4.70 | 0.330 | 0.030 | 1139.000 | 66.100 | 1) | 1253.470 | 72.701 | PAULSEN 80A |
| 106 | 4.76 | 0.060 | 0.006 | 1170.000 | 109.000 | NONE | 1170.000 | 109.000 | W 88 |
| 107 | 4.80 | 0.330 | 0.030 | 1080.000 | 63.700 | 1) | 1188.540 | 70.123 | PAULSEN 80A |
| 108 | 4.90 | 0.330 | 0.030 | 1194.000 | 68.100 | 1) | 1314.000 | 74.898 | PAULSEN 80A |
| 109 | 4.96 | 0.110 | 0.011 | 1135.000 | 117.000 | NONE | 1135.000 | 117.000 | WU 88 |
| 110 | 5.00 | 0.320 | 0.030 | 1163.000 | 67.500 | 1) | 1279.880 | 74.233 | PAULSEN 80A |
| 111 | 5.00 | 0.150 | 0.015 | 1231.000 | 55.000 | 1)2) | 1120.090 | 75.106 | SANTRY 74A |
| 112 | 5.10 | 0.310 | 0.030 | 1164.000 | 72.200 | 1) | 1280.980 | 79.421 | PAULSEN 80A |
| 113 | 5.18 | 0.050 | 0.005 | 1164.000 | 102.000 | NONE | 1164.000 | 102.000 | W 88 |
| 114 | 5.20 | 0.300 | 0.030 | 1017.000 | 59.000 | 1) | 1119.210 | 64.914 | PAULSEN 80A |
| 115 | 5.30 | 0.290 | 0.030 | 1003.000 | 57.200 | 1) | 1103.800 | 62.917 | PAULSEN 80A |
| 116 | 5.30 | 0.160 | 0.016 | 1204.000 | 48.000 | 1)2) | 1095.520 | 70.057 | SANTRY 74A |
| 117 | 5.40 | 0.280 | 0.030 | 1136.000 | 62.500 | 1) | 1250.170 | 68.759 | PAULSEN 80A |
| 118 | 5.50 | 0.270 | 0.030 | 1117.000 | 61.400 | 1) | 1229.260 | 67.609 | PAULSEN 80A |
| 119 | 5.60 | 0.250 | 0.030 | 1259.000 | 71.800 | 1) | 1385.530 | 78.975 | PAULSEN 80A |
| 120 | 5.69 | 0.104 | 0.020 | 1314.270 | 57.565 | NONE | 1314.270 | 57.565 | MIAH 95 |
| 121 | 5.70 | 0.230 | 0.020 | 1150.000 | 62.100 | 1) | 1265.580 | 68.341 | PAULSEN 80A |
| 122 | 5.80 | 0.210 | 0.020 | 1077.000 | 60.300 | 1) | 1185.240 | 66.373 | PAULSEN 80A |
| 123 | 5.90 | 0.190 | 0.020 | 1111.000 | 60.000 | 1) | 1222.660 | 66.024 | PAULSEN 80A |
| 124 | 5.94 | 0.130 | 0.013 | 1368.000 | 53.000 | 1)2) | 1244.740 | 78.734 | SANTRY 74A |
| 125 | 6.00 | 0.150 | 0.010 | 1239.000 | 66.900 | 1) | 1363.520 | 73.630 | PAULSEN 80A |
| 126 | 6.04 | 0.095 | 0.020 | 1295.560 | 47.158 | NONE | 1295.560 | 47.158 | MIAH 95 |
| 127 | 6.10 | 0.110 | 0.010 | 1207.000 | 65.200 | 1) | 1328.300 | 71.728 | PAULSEN 80A |
| 128 | 6.40 | 0.120 | 0.012 | 1478.000 | 56.000 | 1)2) | 1344.830 | 84.367 | SANTRY 74A |
| 129 | 6.51 | 0.120 | 0.012 | 1433.000 | 57.000 | 1)2) | 1303.890 | 83.308 | SANTRY 74A |
| 130 | 6.53 | 0.086 | 0.020 | 1311.540 | 48.789 | NONE | 1311.540 | 48.789 | MIAH 95 |

Table 2: Cross section data for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ (contd.).

| NR. | E-NEUTR [MEV] | WIDTH [MEV] | ERR.CENTR [MEV] | $X-\text { SECT (ORIG) }$ [MB] | $\begin{gathered} \text { ERR(ORIG) } \\ \text { [MB] } \end{gathered}$ | CORR.APPL. | $\begin{gathered} X-\operatorname{SECT}(F I N) \\ {[M B]} \end{gathered}$ | ERR(FIN) [MB] | REF. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 131 | 6.75 | 0.120 | 0.012 | 1406.000 | 63.000 | 1)2) | 1279.320 | 85.893 | SANTRY 74A |
| 132 | 6.93 | 0.080 | 0.020 | 1320.810 | 46.492 | NONE | 1320.810 | 46.492 | MIAH 95 |
| 133 | 7.25 | 0.100 | 0.010 | 1324.000 | 59.000 | 1)2) | 1204.710 | 80.686 | SANTRY 74A |
| 134 | 7.47 | 0.100 | 0.010 | 1385.000 | 62.000 | 1)2) | 1260.210 | 84.576 | SANTRY 74A |
| 135 | 7.58 | 0.072 | 0.020 | 1203.340 | 43.079 | NONE | 1203.340 | 43.079 | MIAH 95 |
| 136 | 7.80 | 0.100 | 0.010 | 1393.000 | 63.000 | 1)2) | 1267.490 | 85.454 | SANTRY 74A |
| 137 | 7.98 | 0.068 | 0.020 | 1247.200 | 44.280 | NONE | 1247.200 | 44.280 | MIAH 95 |
| 138 | 8.51 | 0.080 | 0.008 | 1409.000 | 63.000 | 1)2)6) | 1278.590 | 85.763 | SANTRY 74A |
| 139 | 8.55 | 0.065 | 0.020 | 1293.100 | 48.490 | NONE | 1293.100 | 48.490 | MIAH 95 |
| 140 | 8.97 | 0.062 | 0.020 | 1220.600 | 44.180 | NONE | 1220.600 | 44.180 | MIAH 95 |
| 141 | 9.02 | 0.080 | 0.008 | 1287.000 | 58.000 | 1)2)6) | 1175.960 | 79.156 | SANTRY 74A |
| 142 | 9.55 | 0.060 | 0.020 | 1227.900 | 47.520 | NONE | 1227.900 | 47.520 | MIAH 95 |
| 143 | 9.80 | 0.080 | 0.008 | 1309.000 | 59.000 | 1)2)6) | 1212.140 | 81.597 | SANTRY 74A |
| 144 | 9.92 | 0.057 | 0.020 | 1195.400 | 49.850 | NONE | 1195.400 | 49.850 | MIAH 95 |
| 145 | 10.03 | 0.080 | 0.008 | 1272.000 | 57.000 | 1)2)6) | 1186.560 | 79.668 | SANTRY 74A |
| 146 | 10.50 | 0.080 | 0.008 | 1099.000 | 49.000 | 1)2,6) | 965.781 | 64.699 | SANTRY 74A |
| 147 | 10.58 | 0.056 | 0.020 | 937.680 | 48.010 | NONE | 937.680 | 48.010 | MIAH 95 |
| 148 | 10.75 | 0.080 | 0.008 | 1026.000 | 46.000 | 1)2)6) | 885.945 | 59.498 | SANTRY 74A |
| 149 | 11.00 | 0.080 | 0.008 | 981.000 | 46.000 | 1)2,6) | 838.787 | 57.497 | SANTRY 74A |
| 150 | 11.04 | 0.055 | 0.020 | 771.030 | 46.880 | NONE | 771.030 | 46.880 | MIAH 95 |
| 151 | 11.42 | 0.054 | 0.020 | 720.900 | 49.020 | NONE | 720.900 | 49.020 | MIAH 95 |
| 152 | 11.80 | 0.080 | 0.008 | 767.000 | 43.000 | 1)2,6) | 599.909 | 45.065 | SANTRY 74A |
| 153 | 12.00 | 0.080 | 0.008 | 665.000 | 42.000 | 1)2,6) | 500.887 | 40.348 | SANTRY 74A |
| 154 | 12.00 | 0.054 | 0.020 | 564.760 | 45.180 | NONE | 564.760 | 45.180 | MIAH 95 |
| 155 | 12.10 | 0.080 | 0.008 | 718.000 | 47.000 | 1)2)6) | 533.099 | 43.912 | SANTRY 74A |
| 156 | 12.60 | 0.080 | 0.008 | 546.000 | 49.000 | 1)2,6) | 395.556 | 40.636 | SANTRY 74A |
| 157 | 13.00 | 0.180 | 0.006 | 398.000 | 37.400 | 1) | 437.999 | 41.172 | PAULSEN 80B |
| 158 | 13.58 | 0.100 | 0.013 | 359.000 | 16.000 | 1)2) | 326.654 | 21.879 | SANTRY 74A |
| 159 | 13.58 | 0.090 | 0.009 | 373.000 | 92.000 | 1)2,6) | 254.544 | 64.060 | SANTRY 74A |
| 160 | 13.89 | 0.060 | 0.014 | 325.000 | 25.000 | 1)2) | 295.717 | 27.131 | SANTRY 74A |
| 161 | 14.00 | 0.250 | 0.004 | 286.000 | 27.200 | 1) | 314.743 | 29.900 | PAULSEN 808 |
| 162 | 14.14 | 0.001 | 0.001 | 260.700 | 14.562 | 8) | 260.700 | 14.562 | STROHMAIER 95 |
| 163 | 14.24 | 0.080 | 0.014 | 297.000 | 13.000 | 1)2) | 270.240 | 17.958 | SANTRY 74A |
| 164 | 14.50 | 0.150 | 0.015 | 281.000 | 13.000 | 1)2) | 255.682 | 17.417 | SANTRY 74A |
| 165 | 14.54 | 0.001 | 0.001 | 236.200 | 13.193 | 8) | 236.200 | 13.193 | STROHMAIER 95 |
| 166 | 14.70 | 0.150 | 0.030 | 280.000 | 25.000 | 9) | 276.276 | 22.710 | PAZSIT 72 |
| 167 | 14.74 | 0.200 | 0.020 | 272.000 | 12.000 | 1)2) | 247.493 | 16.503 | SANTRY 74A |
| 168 | 14.80 | 0.200 | 0.020 | 216.000 | 26.000 | 1)3) | 187.639 | 14.998 | PAZSIT 75 |
| 169 | 14.94 | 0.001 | 0.001 | 216.200 | 12.076 | 8) | 216.200 | 12.076 | STROHMAIER 95 |
| 170 | 15.35 | 0.001 | 0.001 | 199.600 | 32.202 | 8) | 199.600 | 32.202 | STROHMAIER 95 |
| 171 | 15.75 | 0.001 | 0.001 | 185.700 | 29.960 | 8) | 185.700 | 29.960 | STROHMAIER 95 |
| 172 | 16.16 | 0.001 | 0.001 | 174.000 | 32.770 | 8) | 174.000 | 32.770 | STROHMAIER 95 |
| 173 | 16.56 | 0.001 | 0.009 | 163.900 | 30.868 | 8) | 163.900 | 30.868 | STROHMAIER 95 |
| 174 | 16.96 | 0.001 | 0.001 | 155.200 | 29.230 | 8) | 155.200 | 29.230 | STROHMAIER 95 |
| 175 | 17.37 | 0.001 | 0.001 | 147.500 | 30.174 | 8) | 147.500 | 30.174 | STROHMAIER 95 |
| 176 | 17.77 | 0.001 | 0.001 | 140.700 | 28.783 | 8) | 140.700 | 28.783 | STROHMAIER 95 |
| 177 | 18.17 | 0.001 | 0.001 | 134.700 | 28.659 | 8) | 134.700 | 28.659 | STROHMAIER 95 |
| 178 | 18.58 | 0.001 | 0.001 | 129.200 | 27.489 | 8) | 129.200 | 27.489 | STROHMAIER 95 |
| 179 | 18.98 | 0.001 | 0.001 | 124.300 | 26.447 | 8) | 124.300 | 26.447 | STROHMAIER 95 |
| 180 | 19.39 | 0.001 | 0.001 | 119.800 | 25.822 | 8) | 119.800 | 25.822 | STROHMAIER 95 |
| 181 | 19.79 | 0.009 | 0.001 | 115.800 | 24.959 | 8) | 115.800 | 24.959 | STROMMAIER 95 |

CORRECTION CODES:

1) Cross section renormalized to present decay data (half-life, branching ratios etc.)
2) ERROR GIVEN IN PUBLICATION DID NOT INCLUDE ERROR OF REFERENCE CROSS SECTION.
3) CRoss section renormalized to endf/b-vi values of reference cross section used in measurement. errors taken from the associated file 33 included in final error.
4) Cross section renormalized to angular distribution of source neutrons of liskien and paulsen.
5) error has been reduced by a factor two or three in order to represent i standard deviation.
6) SPECIAL CORRECTION. SEE TEXT FOR DETAILS.
7) cross section from measured relative excit. function, normalized to present evaluation.
8) CROSS SECTION FROM THEORETICAL CALCULATION.
9) renormalization using reference cross section evaluated at irk, see physics data 13-5.

Table 3: Evaluated group cross sections for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{1}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$.

| Energy group <br> $[\mathrm{MeV}]$ to $[\mathrm{MeV}]$ | Cross section <br> $[\mathrm{mb}]$ | Error <br> $[\mathrm{mb}]$ | Error <br> $[\%]$ | Ratio <br> $\Delta_{\text {ex }} / \Delta_{\text {int }}$ |  |
| :---: | ---: | ---: | ---: | ---: | ---: |
| 0.10 | 0.30 | 37.71 | 13.64 | 36.2 | 3.103 |
| 0.30 | 0.50 | 116.05 | 13.79 | 11.9 | 1.500 |
| 0.50 | 0.70 | 227.94 | 15.41 | 6.8 | 0.879 |
| 0.70 | 0.90 | 545.00 | 40.91 | 7.5 | 1.694 |
| 0.90 | 1.10 | 619.67 | 34.54 | 5.6 | 1.854 |
| 1.10 | 1.30 | 629.78 | 20.44 | 3.2 | 1.080 |
| 1.30 | 1.50 | 645.36 | 24.41 | 3.8 | 1.055 |
| 1.50 | 1.75 | 731.96 | 41.27 | 5.6 | 1.913 |
| 1.75 | 2.00 | 816.23 | 60.20 | 7.4 | 2.373 |
| 2.00 | 2.50 | 1009.32 | 47.33 | 4.7 | 0.552 |
| 2.50 | 3.00 | 1030.96 | 57.58 | 5.6 | 1.448 |
| 3.00 | 3.50 | 1148.88 | 57.14 | 5.0 | 0.592 |
| 3.50 | 4.00 | 1154.02 | 50.25 | 4.4 | 0.162 |
| 4.00 | 4.50 | 1208.80 | 59.50 | 4.9 | 0.901 |
| 4.50 | 5.00 | 1205.64 | 60.49 | 5.0 | 0.277 |
| 5.00 | 5.75 | 1219.33 | 40.90 | 3.4 | 1.100 |
| 5.75 | 6.50 | 1294.44 | 36.20 | 2.8 | 0.152 |
| 6.50 | 7.25 | 1303.72 | 41.30 | 3.2 | 0.339 |
| 7.25 | 8.00 | 1222.56 | 38.53 | 3.2 | 0.107 |
| 8.00 | 9.00 | 1267.73 | 41.01 | 3.2 | 0.140 |
| 9.00 | 10.00 | 1261.44 | 46.89 | 3.7 | 0.524 |
| 10.00 | 11.00 | 967.11 | 39.56 | 4.1 | 0.324 |
| 11.00 | 12.00 | 684.74 | 33.02 | 4.8 | 0.198 |
| 12.00 | 13.00 | 452.82 | 26.95 | 6.0 | 0.792 |
| 13.00 | 14.00 | 357.47 | 28.46 | 8.0 | 0.685 |
| 14.00 | 15.00 | 237.90 | 13.23 | 5.6 | 1.707 |
| 15.00 | 16.00 | 193.28 | 31.18 | 16.1 | $--*^{*}$ |
| 16.00 | 17.00 | 164.90 | 31.06 | 18.8 | $--*$ |
| 17.00 | 18.00 | 144.99 | 29.66 | 20.5 | $-\ldots *)$ |
| 18.00 | 19.00 | 130.06 | 27.67 | 21.3 | $-\ldots *$ |
| 19.00 | 20.00 | 117.67 | 25.36 | 21.6 | $-\ldots *)$ |
|  |  |  |  |  |  |

[^1]
# Table 5: Relative covariance matrix for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$. 

## Correlations given in \%

group energy group



```
    100464756 54 57 54 58 20 16 20 18 16 17 8 8 9 0 0 0 0
```




```
                10066 63 66 6044 3341 37 35 35 17 19 00 0
                    10065686341 334137 34 35 17 19 0 0 0 0 0 0 0 0 0 0 0 0 0
                    10065 58 24 19 24 21 20 20 10 11 0 0 0 0 0 0 0 0 0 0 0 0
                    1006240324036 33 34 17 19 0 0 0 0 0 0 0 0 0
                    10045364540 38 381921 0 0 0 0 0 0 0 0 0 0 0 0 0 0
                            100 57 71 64 60 61 30 33 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0
                                    100575248492427 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0}0
```



```
                                    10073734030}000000,0\mp@code{0
                                    100744428}0
                                    1004128
                                    100 46 36 36 36 36 35 35 34 9, 7 0 0 0 0 0 0 0 5.00 5.75
```



```
                                    100454546444241 9 7 7 0 0 0
                                    10046464444241}8486\mp@code{0
                                    10046444241
                                    100444241
                                    100424410
                                    100 40 12 9
                                    100}121210 0 0 0 0 0 0 0 12.00 13.00
                                    100}33000000 0 0 0 13.00 14.00
                                    100 56 56 56 56 56 14.00 15.00
                                    10095959595 15.00 16.00
                                    100959595 16.00 17.00
                                    1009595 17.00 18.00
                                    100 95 18.00 19.00
                                    100 19.00 20.00
```



Fig. 1: Experimental cross section data for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right)^{103 m} \mathrm{Rh}$. (The displayed data points represent the renormalized cross sections and their effective l $\sigma$ uncertainties.)


Fig. 2: Experimental cross section data for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ : expanded display of the energy range $0-6 \mathrm{MeV}$.


Fig. 3: Results of the present evaluation of the cross section for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ compared to the previous evaluation reported in Ref. 2.


Fig. 4: The evaluated cross section for the reaction ${ }^{103} \mathrm{Rh}\left(\mathrm{n}, \mathrm{n}^{\prime}\right){ }^{103 \mathrm{~m}} \mathrm{Rh}$ compared with the results of model calculations [9].

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[^1]:    *) Not applicable (only data points from theoretical calculation in this energy group).

