neUtron nuclear reaction theory and application Four papers by Chinese authors

## NEUTRON NUCLEAR REACTION THEORY AND APPLICATION

Four papers by Chinese authors

## Contents

## Page

1. Systematics of average total radiative widths for s-and p-wave resonances

Zhuang Youxiang, Wang Shunuan, Zhou Delin Inst. of Atomic Energy, Academia Sinica, Beijing

Jia Zhize
Res. Inst. of Science and Technology Information, the Ministry of Nuclear Industry, Beijing
2. A new set of level density parameters for fermi gas model . . 11

Su Zongdi, Wang Cuilan, Zhuang Youxiang, Zhou Chunmei Institute of Atomic Energy, Academia Sinica, Beijing
3. Calculations of the double differential cross section for partical emissions in the continuum region19

Li Cha, Su Zongdi
Institute of Atomic Energy, Academia Sinica, Beijing
4. The statistical theory of compound nucleus reaction and nuclear data calculations25

Su Zongdi, Shi Xiangjun, Li Cha, Zhuang Youxiang Institute of Atomic Energy, Academia Sinica, Beijing

# Systematics of Average Total Radiative Widths for s- and p-Wave Resonances <br> Zhuang Youxiang, Wang Shunuan, Zhou Delin Institute of Atomic Energy, Academia Sinica P.O.B.275(41), Beijing, China 

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## Abstract

The experimental values of total radiative widths, $\left\langle\Gamma_{\gamma}\right\rangle$, for 208 nuclides ( $Z=11-98, A=23-249$ ) are compiled and analysed. Two empirical formulas for the radiative widths which are dependent on atomic number $Z$, mass number $A$ and neutron binding energy $S n$ of compound nucleus are obtained: $\left\langle\Gamma_{\gamma}^{\bullet}\right\rangle=31 \mathrm{Z}^{-2.5} \mathrm{~A}^{0.56} \mathrm{Sn}^{0.83}$ and $\left\langle\Gamma_{\gamma}^{\prime}\right\rangle=24$ $Z^{1.3} \mathrm{~A}^{-2.6} \mathrm{Sn}^{0.95}$ (in eV). Using the strong-coupling dipole model, the coupling coefficients $C r$ which are varied smooth with the mass numbers A are also obtained. The values Cr from the smoothed curves can be used to calculate the radiative widths $\left\langle\Gamma_{\gamma}^{\circ}\right\rangle$ for those nuclides which are deficient in the experimental dat.A.

## Introduction

Average total radiative widths of neutron resonance are one of the resonance parameters which are necessary for application in astrophysics, reactor physics and fission physics. But there are quite a number of nuclides which lack experimental values $\left\langle\Gamma_{\gamma}\right\rangle$. Thus it is necessary to systematically studyaverage radiative widths to derive < $\left.\Gamma_{\gamma}\right\rangle$ for those nuclides. Moreover, the determination of the dependence of radiative widths on nutron spacing, spin and parity of the initial state, excitation energy, nuclear size and nuclear structure effects is of particular interest from a theoretical and evaluation point of view.

In 1956 J.S.Levin and D.J.Hughes ${ }^{(1)}$ carried out the first attempt to study systematically average radiative widths and to interpret them in terms of Blatt-Weisskopf's estimates of partial radiative widths. In 1957 by a similar study A.Stolovy and J.A.Harrey ${ }^{(2)}$ arrived at the following empirical expression for $\left\langle\Gamma_{\gamma}\right\rangle$ :

$$
\begin{equation*}
\left\langle\Gamma_{\gamma}\right\rangle=5.3 \cdot 10^{-4} \mathrm{~A}^{2 / 3}[\mathrm{D}(\mathrm{U})]^{0.25} \mathrm{U}^{4.3} \quad \text { (in } \mathrm{meV} \text { ), } \tag{1}
\end{equation*}
$$

Where $U$ and $D(U)$ are the effective excitation energy in $M e V$ and the spacing of levels with the same spin and parity as the radiating levels around the excitation energy $U$, respectively.

In 1971 H.Malecky et al. (3) noted that the previously derived relations of total radiative widths can be expressed in terms of a combination of the parameters $A, a, U$ and $T$. A least squares fit of the experimental data for 108 nuclides ( $Z=27-96, A=59-246$ ) was carried out. It has the following form

$$
\begin{equation*}
\left\langle\Gamma_{\gamma}\right\rangle=9 \mathrm{U}^{0.9} \mathrm{~A}^{-0.9} \mathrm{a}^{-0.57}\left(1-0.011 \mathrm{I}^{2}\right) \quad(\text { in } \mathrm{eV}), \tag{2}
\end{equation*}
$$

where I stands for the spin of the target nuclides, a is the level density parameter, and $U$ and $a^{-1}$ are expressed in MeV. As pointed out by H.Malecky et al., (3) there is a gemeral agreement between the experimental values and those calculated on the basis of Eq. (2) with a few exceptions. Because those $\left\langle\Gamma_{\gamma}\right\rangle$ compiled by H.Malecky et al. are published in 1966-1969, and in addition, there are no average radiative widths for p-wave resonance, we have carried out a further systematical study.

The purpose of this work is to extend the range of nuclides, to compile more recent experimental values $\left\langle\Gamma_{\gamma}\right\rangle$ as much as possible, to fit them using an empirical expression for $\left\langle\Gamma_{\gamma}\right\rangle$, to carry out theoretical analysis and calculation of average radiative widths and from which some regularity may be found out for setting up a more complete systematics of average radiative widths.

Experimental Systematics of Average Total Radiative Wiaths
The experimental values of average radiative widths for neutron resonance are collected according to CINDA 1982, and average radiative widths of s- and p-wave neutron resonances for 208 nuclides ( $Z=11-98$, $A=23-249$ ) are selected, in which 207 nuclides for s-wave and 56 nuc-
lides for p-wave. These experimental or evaluated data are published in 1973-1981, most of them are taken from BNL-325 (the fourth edition, 1981) (4) and our evaluation (1981) (5).

The variation of the average s-wave radiative widths with mass number $A$ is illustrated in Fig.1. Two interesting features emerge from this plot: (a) the general monotonic decrease of $\left\langle\Gamma_{\gamma}^{0}\right\rangle$ with $A$ in the mass region 60-190, this is favourable to the study of systematics, and (b) the maxima at about the mass numbers $A=50$ and 208, which are related to the magic numbers $Z=28$ and 82 , these should be the results of shell effect.

The variation of p-wave radiative widths with mass number is shown in Fig.2, which demonstrates the same qualitative features as for s-wave radiative widths.


Figure 1. The average s-wave radiative widths plotted versus mass number $A$ of target nucleus


Figure 2. Variation of the average p-wave radiative widths with mass number $A$ of target nucleus

By means of an empirical experession $\left\langle\Gamma_{\gamma}\right\rangle=K Z^{\alpha} A^{\beta} \operatorname{Sn}^{\gamma}$, which is similar to reference (3), and the least squares analysis, average radiative widths are fitted with the following forms:

$$
\begin{array}{lll}
\left\langle\Gamma_{\gamma}^{0}\right\rangle=31 & \mathrm{z}^{-2.5} \mathrm{~A}^{0.56} \mathrm{Sn}^{0.83} & \text { (in eV). } \\
\left\langle\Gamma_{\gamma}^{\prime}\right\rangle=24 & \mathrm{Z}^{1.3} \mathrm{~A}^{-2.6} \mathrm{Sn}^{0.95} & \text { (in eV). } \tag{4}
\end{array}
$$

where the neutron binaing energy Sn is experessed in MeV. Generally, the experimental values are in resonable agreement with those calculated on the basis of Eq. (3) and (4) except 16 nuclides which are around $A=50$ and 208 .as well as ${ }^{114, ~}{ }^{116} \mathrm{Cd}$ for s-wave radiative widths and ${ }^{23} \mathrm{Na},{ }^{32} \mathrm{~S},{ }^{60} \mathrm{Ni},{ }^{88} \mathrm{Sr},{ }^{140} \mathrm{Ce},{ }^{146} \mathrm{Nd}$ and ${ }^{207} \mathrm{~Pb}$ for p-wave.

## Analysis of the Strong-Coupling Dipole Model

## for Average Total Radiative Widths $\left\langle\Gamma_{\dot{\gamma}}^{\circ}\right\rangle$

In the fast neutron capture calculations there are three reac-
tion mechanisms which are usually used, i.e. compound-nucleus, direct and semidirect interation theories. From a qualitative point of view, the relative importance of these contributions to ( $n, \gamma$ ) cross sections is given in Fig.3. For orientation, the rapid falloff of the compound-nucleus contribution typically occurs near En=1 MeV , and the peak in the semidirect contribution is in the neighborhood of 14 MeV , where the ( $n, \gamma$ ) cross section is $\sim 1 \mathrm{mb}$. For most applications the contribution of compound-nucleus is the most important among the three ones.


Figure 3. Schematic view of the relative importance of different reaction mechnisms to neutron capture in a medium-weight nucleus.

There are two models which are commonly used to determine radiative widths in compound-nucleus theory calculations, they are the Weisskopf strong coupling model ${ }^{(8)}$ and the giant dipole resonance model ${ }^{(9)}$. The former was adopted by us because it has a more complete set of parameters. For example, using Cameron type (10) level density formula and parameters, 202 nuclides among 207 ones can be analysed and calculated.

In terms of compund-nucleus theory ${ }^{(11)}$ an estimate of average total radiative width can be expressed as:

$$
\begin{equation*}
\left\langle\Gamma_{\gamma}(E, A)\right\rangle=\frac{1}{P(E, A)} \int_{0}^{E} f\left(E, \varepsilon_{\gamma}\right) P\left(E-\varepsilon_{\gamma}, A\right) d \varepsilon_{\gamma}, \tag{5}
\end{equation*}
$$

where $P(E, A)$ and $P\left(E-\varepsilon_{\gamma}, A\right)$ are the level densities of the compound nucleus and the residnal nucleus respectively, the gamma-ray strength function $f\left(E, \varepsilon_{\gamma}\right)$ contains the energy dependent transition matrix elements and phase space factor $\varepsilon_{\gamma}^{2 l+1}$ which is dependent on the transition multipolarity. For electric dipole emission, which usually gives the largest contribution to the radiative decay process, the radiative width for s-wave neutron at the binding energy $S n$ is

$$
\begin{equation*}
\left\langle\Gamma_{\gamma}^{0}\left(S_{n}, A\right)\right\rangle=\frac{C_{\gamma}}{P\left(S_{n}, A\right)} \int_{0}^{S_{n}} f_{E l}\left(\varepsilon_{\gamma}\right) \varepsilon_{\gamma}^{3} P\left(S_{n}-\varepsilon_{\gamma}, A\right) d \varepsilon_{\gamma}, \tag{6}
\end{equation*}
$$

where $f_{E l}\left(\mathcal{E}_{G}\right)$ stands for the strength function of electric dipole radiation and $C_{r}$ is the coupling coefficient introduced.

Puting, $f_{E 1}\left(\varepsilon_{\gamma}\right)=1$, the formula (6) becomes the Weisskopf formula, as described by J.E.Iynn,

$$
\begin{equation*}
\left\langle\Gamma_{\gamma}^{0}\left(S_{n}, A\right)\right\rangle=\frac{C_{\gamma}}{\rho\left(S_{n}, A\right)} \int_{0}^{S_{n}} \varepsilon_{\gamma}^{3} \rho\left(S_{n}-\varepsilon_{\gamma}, A\right) d \varepsilon_{\gamma}, \tag{7}
\end{equation*}
$$

The strength function for electric dipole radiation is usually taken as Lorentzian in shape. More recently D.G.Gardner et al. (12,13) have investigated the use of Breit-Wigner shapes and have developed expressions for $f_{E 1}$ based on systematics, covering the mass range A>40, that is an energy dependent Breit-Wigner (EDBV) model. The Cameron type level density formula and parameters ${ }^{(10)}$ are used by us.

Firstly, the known experimental values $\left\langle\Gamma_{\gamma}^{0}\right\rangle$ are utilized to determine a set of coupling coefficients $\mathrm{C}_{\mathrm{r}}$. For 202 nuclides the values $C_{r}$ from formulas (6) and (7) are between $10^{0} \sim 10^{1} \mathrm{MeV}^{-3}$ and $10^{-8} \sim 10^{-7}$ $\mathrm{MeV}^{-3}$ respectively (See Fig. 4 and 5). The values $C_{r}$ from the smoothed curves are used to calculate the radiative widths $\left\langle\Gamma_{\gamma}^{\circ}\right\rangle$ in order to test the status of agreement with the experimental data. The radiative wiaths of the nuclides which are deficient in the experimertal data could be calculated from the smoothed values of $C_{r}$.


Figure 4. The coupling coefficients $C_{r}$ from formula (6) plotted versus mass number $A$ of compcnid nucleus.


Figure 5. Variation of the coupling coefficients $C_{r}$ from formula (7) with mass number $A$ of compound nucleus.

## Result and Discuss

In the above passages the expirical expression and the strongcoupling dipole model are utilized to fit and analyse the experimental values $\left\langle\Gamma_{\dot{\gamma}}\right\rangle$ and/or $\left\langle\Gamma_{\gamma}^{\prime}\right\rangle$. The goodness of fit can be expressed by the value $X_{r}^{2}$ :

$$
\begin{equation*}
X_{r}^{2}=\frac{\sum_{i=1}^{N}\left(\Gamma_{\gamma i}-\Gamma_{\gamma f i}\right)^{2} /\left(\Delta \Gamma_{\gamma i}\right)^{2}}{V} \tag{8}
\end{equation*}
$$

where $\Gamma_{\gamma i}$ and $\Delta \Gamma_{\gamma_{i}}$ are experimental values and their corresponding errors, $\Gamma_{r f i}$ is the fitting or calculated value, and $V$ is the number of degrees of freedom. The $X_{r}^{2}$ values are listed in Table 1.

Table 1. the values $\chi_{r}^{2}$ of the results

| fitting method | experimerimental expression |  | strong-coupling dipole <br> model |  |
| :---: | :---: | :---: | :---: | :---: |
|  | formula (3) | formula (4) | formula (6) formula(7) |  |
|  | 191 | 50 | 202 | 202 |
| $X_{r}^{2}$ | 7.9 | 4.2 | 7.3 | 4.0 |

All of the formulas (3),(4).(6) and (7) can be used to calculate average total radiative widths of nuclides which lack experimental values. Formula (3) and (4) are suitable to calculate the values $\left\langle\Gamma_{\gamma}^{0}\right\rangle$ and $\left\langle\Gamma_{\gamma}^{\prime}\right\rangle$ for all nuclides except those around $A=50$ and 208 respectively; formulas (6) and (7) are appropriate to all nuclides from Na to Cm . And it can be noticed that for $\left\langle\Gamma_{\gamma}^{0}\right\rangle$ the value $\mathcal{X}_{r}^{2}$ of formula (7) is the smallest.

We have respectively utilized formulas (6) and (7) as well as their correspounding values $C_{r}$ interpolated to calculate the values $\left\langle\Gamma_{\gamma}^{0}\right\rangle$ of 83 nuclides which have the neutron binding energy but have not experimental data. They are shown in Table 2.

The level density is an important physical magnitude in the calculations of neutron reaction cross sections, spectrums and angular distributions. Unforturately, there is very deficient in direct experimental information about it. For this reason it is interesting to compare the calculated $\left\langle\Gamma_{\gamma}(S n)\right\rangle$ values using statistical model with the experimental ones. In this comparison one may notice follwing things:

In terms of the smoothed values $\left\langle\Gamma_{\gamma}\right\rangle$ from formulas (6) and (7) respectively, it may be seen that the most result of ever-even and odd-odd or odd $A$ nuclides don't show obvious deviation. This inaicates that the parameters of level density for these nuclides are rather suitable. However, there are a few even-even or odd nuclides whose $C_{r}$ deviate obviousiy from those of odd-odd nuclides. For these nuclides one may conclude that either experimental values $\left\langle\Gamma_{\gamma}^{\circ}\right\rangle$ and/ or their binding energies are unreliable, or the pairing energies are unsuitable. Illustrate with examples, the pairing energies for even-even nuciides $A=78$, 114 etc. seem to be larger; perhaps those of even-even and odd $A$ nuclides $A=136,138$, 181 are smaller.

If level density formula is reasonable, the parameters in this formula are suitable for pairing energies and shefl corrections and

Table 2. the values $\left\langle\Gamma_{\gamma}^{0}\right\rangle$ from the systematics for those nuclides which lack experimental data.

| target nuclide | $\left\langle\Gamma_{\gamma}^{\circ}\right\rangle(\mathrm{meV})$ |  | target <br> nuclide | $\left\langle\Gamma_{j}^{\circ}\right\rangle$ (meV) |  | target nuclide | $\left\langle\Gamma_{\gamma}^{\circ}\right\rangle \quad(\mathrm{meV})$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $(7)^{*}$ | (6)* |  | (7)* | $(6)^{*}$ |  | (7)* | (6)* |
| 12-Mg-24 | 2700 | 3000 | 46-1'd-107 | 110 | 110 | 68-Er-164 | 84 | 80 |
| 12-Mp-25 | 4000 | 6800 | 50-Sn-115 | 110 | 110 | 68-Er-170 | 70 | 67 |
| 12-Mg-26 | 1900 | 1700 | $50-\mathrm{Sn}-116$ | 80 | 70 | 70-Yb-168 | 75 | 70 |
| 13-A1-27 | 2700 | 3100 | 50-Sn-118 | 76 | 67 | 70-Yb-170 | 72 | 69 |
| 14-Si-28 | 3600 | 4800 | 50-Sn-120 | 78 | 68 | 70-Yb-176 | 57 | 59 |
| 14-5i-29 | 3800 | 6200 | 50-Sn-122 | 98 | 92 | 71-Lu-176 | 85 | 97 |
| 14-Si-30 | 1800 | 1700 | 50-Sn-124 | 110 | 110 | 72-Hf-174 | 60 | 61 |
| 15-P -31 | 2500 | 3000 | 52-Te-124 | 88 | 84 | 72-Hf-176 | 64 | 64 |
| 16-S -33 | 2800 | 4800 | 52-Te-128 | 110 | 120 | 72-HP-178 | 56 | 58 |
| 16-S -34 | 1500 | 1500 | 52-Te-130 | 190 | 230 | 72-Hf-179 | 69 | 75 |
| 17-Cl-35 | 2000 | 2400 | 53-I -129 | 120 | 130 | 72-Hf-180 | 57 | 59 |
| 17-Cl-37 | 920 | 730 | 54-Xe-124 | 62 | 50 | 73-T'a-182 | 77 | 80 |
| 18-Ar-36 | 2000 | 2500 | 54-Xe-126 | 67 | 56 | 74-W -180 | 62 | 65 |
| 18-Ar-40 | 690 | 510 | 54-Xe-128 | 70 | 62 | 76-0s-187 | 69 | 63 |
| 19-K - 39 | 1300 | 1300 | 54-Xe-130 | 83 | 81 | 78-Pt-196 | 110 | 95 |
| 19-K-41 | 600 | 350 | 54-Xe-132 | 120 | 150 | 78-It-198 | 140 | 130 |
| 20-Ca-46 | 8880 | 790 | 54-Xe-134 | 220 | 310 | $80-\mathrm{Hg}-196$ | 130 | 120 |
| 20-Ca-48 | 850 | 730 | 54-Xe-136 | 210 | 260 | 80- $\mathrm{Hg} \mathrm{l}-200$ | 210 | 220 |
| 23-V-50 | 1900 | 2300 | 56-Ba-132 | 61 | 56 | $80-\mathrm{Hg}-202$ | 370 | 430 |
| 23-V-51 | 1100 | 920 | 57-Ia-138 | 240 | 330 | 81-T1-204 | 1100 | 1800 |
| 27-60-60 | 890 | 820 | 58-Ce-136 | 76 | 70 | $82-\mathrm{Pb}-206$ | 2100 | 3800 |
| 30-2n-70 | 230 | 200 | 58-Ce-138 | 110 | 120 | $82-\mathrm{Pb}-207$ | 3800 | 9000 |
| 34-Se-8,2 | 190 | 160 | 58-Ce-142 | 61 | 59 | $82-\mathrm{Pb}-208$ | 540 | 430 |
| 40-Zr-93 | 140 | 140 | 61-Pm-148 | 72 | 75 | 90-Th-228 | 26 | 26 |
| 40-zr-96 | 836 | 73 | 62-Sm-150 | 45 | 42 | 90-Th-229 | 35 | 38 |
| 46-19-102 | 160 | 180 | 66-Dy-156 | 72 | 75 | 92-U -232 | 34 | 37 |
| 46-Pd-104 | 120 | 120 | 66-Dy-158 | 80 | 83 | 92-U -237 | 31 | 32 |
| 46-1’d-106 | 91 | 82 | $68-E r-162$ | 85 | 83 |  |  |  |

[^0]strength fuction of electric dipole emission are reliable, the variation of the values $C_{r}$ from theoretical formula (6) with mass number A would be a constant nears one. From Fig. 4 and 5 the values $C_{r}$ of formula (6) are between 1.0~10; however those of formula (7) are between $3.0 \times 10^{-8} 3.0 \times 10^{-7}$. Therefore the strength function fE ( $\left(\varepsilon_{\gamma}\right)$ of elactric dipole radiation within $40<A<208$ given by M.A.Gardner et al. is appropriate and can be extend to more wide range which we have dealed with.

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A Now Set of Level Density Parameters for fermi has fouel

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#### Abstract

In order to recommend more accurate and reliable parameters of nuclear level density, three kinds of experimental data have been collected and evaluated. Based on the newest and more kinds of experimental data a new set of level density parameters has been obtained. It may be seen from the results that our parameters are in good agreement with these more accurate and newer measurements, especially to level spacings.


Nuclear level densities are a crucial ingredient in the statistical models, for instance, in the calculations of the widths, cross sections, emitted particle spectra etc. for various reaction channels. Up to now, the Gilbert-Cameron type level density formula ${ }^{l}$ as well as parameters is still applied to the calculations on the statistical models widely. Based on their work, Cook et al. 2 have also got a set of parameters. Because these parameters only depend on $Z$ and $N$, the numbers of proton and neutron in the nucleus respectively, Gilber-Cameron's formula is universally applicable. However, their parameters are based on the available experimental data at that time. The present work attempt to obtain a new set of parameters based on the newest and more kinds of experimental data.

In order to recommend more accurate and reliable parameters of level density, three kinds of experimental data have been collected and evaluated: average level spacing $D_{0}$ and radiative capture width $\Gamma_{\gamma}^{0}$ at the neutron binding energy and cumulative number $N_{0}$ of level (all spin) at the low excitation energy in the amount of 667 which reported during the period of 1973 to 19833 . While that in Gilbert-Cameron's are 129 level spacings which are report-
ed durine the peropd of 1959 to 1964. So the experimental daid adopted in this work are of high precision, wide distribution of nuclei and more kinds of data.

Based on these experimental data especilly $D_{o}$ and $N_{o}$, the parameters $S(Z), S(N), P(Z)$ and $P(N)$ have been adjusted. And a new set of level density parameters has been obtained (denoted as work 1). Next the empirical parameters have also been reobtained to fit the experimental data which we have collected. In the Gilbert-Cameron formula, the corresponding quantities including the empirical parameters are as follows:

$$
\begin{align*}
& a=[0.00917(S(Z)+S(N))+Q B] A \\
& U_{X}=2.5+150 / A \tag{1}
\end{align*}
$$

The new ones of $0.00880,1.4$ and 263 are adopted respectively instead of $0.00917,2.5$ and 150 used in the formula. Thus, the formula are rewriten as follows (denoted as work 2):

$$
\begin{align*}
& a=[0.00880(S(Z)+S(N))+Q B] A \\
& U_{x}=1.4+263 / A \tag{2}
\end{align*}
$$

Where the values of $Q B$ are taken as:

$$
\hat{Q B}=\left\{\begin{array}{l}
0.142,29 \leqslant Z \leqslant 62 \text { (and } N \leqslant 89) ; 79 \leq Z \leqslant 85 ; \\
0.120, Z<28 ; 62 \text { (and } N>89) \leqslant Z \leqslant 78 ; Z \geqslant 86
\end{array}\right.
$$

A detailed comparison with Gilbert-Cameron's and Cook's is shown in Table 1 , where the percentages of the nuclei whose relative deviations of calculated values to the experimental data are within a certain scope are given.

Table 1 A comparison between three results

| Relative Diviation | GilbertCameron | Cook et al | Work 1 | Work 2 |
| :---: | :---: | :---: | :---: | :---: |
| $\leqslant \pm 20 \%$ | 21.5 | 22.4 | 47.8 | 48.7 |
| $\delta D_{0} \leqslant \pm 50 \%$ | 29.4 | 28.1 | 36.4 | 37.3 |
| \% $\leqslant 100 \%$ | 24.1 | 27.6 | 15.4 | 13.2 |
| $> \pm 100 \%$ | 25.0 | 21.9 | 0.4 | 0.8 |
| $\chi^{2}{ }_{\text {d }}$ | 175.9 | 120.9 | 11.3 | 11.4 |
| $\leqslant \pm 20 \%$ | 13.0 | 13.9 | 18.5 | 18.9 |
| $\delta N^{*} \leqslant \pm 50 \%$ | 30.7 | 23.9 | 36.1 | 31.5 |
| ONo $\leqslant \pm 100 \%$ | 49.2 | 50.0 | 41.2 | 44.5 |
| $> \pm 100 \%$ | 7.1 | 12.2 | 4.2 | 5.1 |
| $\chi^{2} \times$ | 13.9 | 47.9 | 11.5 | 10.1 |
| $\leqslant \pm 20 \%$ | 26.9 | 26.4 | 27.9 | 30.8 |
| $\delta \Gamma^{\circ} \leqslant \pm 50 \%$ | 39.8 | 40.8 | 37.3 | 35.8 |
| \%17 $\leqslant \pm 100 \%$ | 30.3 | 29.8 | 31.8 | 30.8 |
| $1> \pm 100 \%$ | 3.0 | 3.0 | 3.0 | 2.6 |
| $\chi_{r i}^{2}$ | 9.6 | 9.7 | 11.0 | 9.3 |
| $x_{D_{0}}^{2}+x_{N_{e}}^{2}+x_{r_{i}}^{2}$ | 199.4 | 178.5 | 33.8 | 30.8 |

From T'able l, the good aqreement with these more accurate and newer measurements, especially to level spacings $D_{0}$ should be obtained. The number of nuclei whose relative deviations are less than $\pm 20 \%$ is nearly a half of the total one and the deviations which are more than $\pm 100 \%$ is reduced to $1 \%$ of the total one. Intuitively this new set of level density parameters seems to be more accurate. Of course, it should be tested by the calculations on statistical models. The parameters adopted by our work are listed in Table 2.

The giant dipole resonance parameters are necessary for calculating radiative capture widths, which are taken from measured values ${ }^{6}$. But lack of experimental data for some nuclei, usually their values are substituted by those of adjacent nuclides. In this work, the adopted giant dipole resonance parameters 7 are matched with recommended level density parameters and also they are worth recommending to be used in theoretical calculations.

It is worth noting that there is not a distinct improvement in the calculations of $N_{0}$ and $\Gamma_{\gamma}^{0}$. And the negative relative deviations of $N_{o}$ are much more than the positive ones (even improved in this work). It turns out there are some problems remained at the low excitation energies in Gilbert-Cameron's formula to be improved.

The outhors are grateful to Profs. Zhou Delin and Ding Dazhao for their help with this work.

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Table 2 the level density parameters of this work

| N OR | Z | PN | SN | PZ | S2 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 11 |  | 0.00 | 6.80 | 0.00 | -5.40 |
| 12 |  | 2.67 | 7.53 | 2.00 | -3.80 |
| 13 |  | 0.00 | 7.55 | 0.00 | -5.72 |
| 14 |  | 1.80 | 7.21 | 2.09 | -7.80 |
| 15 |  | 0.00 | 7.44 | 0.00 | -8.97 |
| 16 |  | 1.67 | 8.07 | 1.45 | -9.50 |
| 17 |  | 0.00 | 9.10 | 0.00 | -10.10 |
| 18 |  | 1.86 | 9.81 | 1.62 | -10.70 |
| 19 |  | 0.00 | 10.60 | 0.00 | -11.38 |
| 20 |  | 2.01 | 11.39 | 1.75 | -12.00 |
| 21 |  | 0.00 | 13.85 | 0.00 | -12.10 |
| 22 |  | 1.64 | 13.68 | 1.53 | -12.80 |
| 23 |  | -0.10 | 14.50 | -0.26 | -12.20 |
| 24 |  | 1.40 | 14.10 | 1.20 | -14.30 |
| 25 |  | -0.10 | 14.93 | 0.00 | -14.00 |
| 26 |  | 1.36 | 13.70 | 1.30 | -14.40 |
| 27 |  | -0.20 | 14.55 | -0.10 | -14.75 |
| 28 |  | 1.20 | 12.50 | 1.20 | -16.57 |
| 29 |  | 0.00 | 13.70 | 0.00 | -17.10 |
| 30 |  | 1.15 | 13.45 | 1.06 | -17.95 |
| 31 |  | -0.20 | 15.45 | 0.00 | -17.68 |
| 32 |  | 1.32 | 14.98 | 1.06 | -16.78 |
| 33 |  | -0.15 | 15.80 | 0.00 | -16.65 |
| 34 |  | 1.30 | 16.38 | 1.35 | -16.35 |
| 35 |  | -0.18 | 17.25 | 0.00 | -16.70 |
| 36 |  | 1.50 | 17.55 | 1.10 | -15.20 |
| 37 |  | -0.10 | 18.05 | 0.20 | -16.20 |
| 38 |  | 1.90 | 17.65 | 1.24 | -16.41 |
| 39 |  | 0.00 | 18.59 | 0.29 | -17.05 |
| 40 |  | 1.43 | 18.71 | 1.35 | -16.75 |
| 41 |  | -0.28 | 18.95 | 0.00 | -15.40 |
| 42 |  | 1.50 | 18.99 | 1.17 | -15.85 |
| 43 |  | -0.10 | 18.60 | -0.15 | -15.50 |
| 44 |  | 1.70 | 18.25 | 1.15 | -17.29 |
| 45 |  | -0.10 | 18.20 | -0.10 | -18.44 |
| 46 |  | 1.57 | 17.38 | 1.20 | -17.82 |


| N OR Z | PN | SN | PZ | SZ |
| :---: | :---: | :---: | :---: | :---: |
| 47 | 0.00 | 16.72 | 0.34 | -18.62 |
| 48 | 1.60 | 14.25 | 1.36 | -18.58 |
| 49 | -0.28 | 14.88 | 0.28 | -19.65 |
| 50 | 1.20 | 12.88 | 1.47 | -19.95 |
| 51 | 0.10 | 12.95 | 0.00 | -19.44 |
| 52 | 0.75 | 13.65 | 1.00 | -17.75 |
| 53 | -0.30 | 14.80 | 0.00 | -17.70 |
| 54 | 1.12 | 15.20 | 1.12 | -16.72 |
| 55 | 0.00 | 15.89 | 0.26 | -16.95 |
| 56 | 1.15 | 17.15 | 1.20 | -14.37 |
| 57 | 0.00 | 17.00 | 0.10 | -14.44 |
| 58 | 0.80 | 17.73 | 1.40 | -14.00 |
| 59 | -0.25 | 19.75 | 0.93 | -12.30 |
| 60 | 1.35 | 18.80 | 1.36 | -11.91 |
| 61 | -0.32 | 19.45 | 0.36 | -11.59 |
| 62 | 1.25 | 19.51 | 1.22 | -10.78 |
| 63 | -0.44 | 20.17 | -0.20 | -11.40 |
| 64 | 1.14 | 19.85 | 0.97 | -10.21 |
| 65 | 0.08 | 19.98 | -0.20 | -10.00 |
| 66 | 1.32 | 20.05 | 0.78 | -9.35 |
| 67 | -0.11 | 20.28 | 0.00 | -9.30 |
| 68 | 1.15 | 19.83 | 0.52 | -8.85 |
| 69 | -0.16 | 20.20 | 0.00 | -8.91 |
| 70 | 1.06 | 19.44 | 0.68 | -8.13 |
| 71 | 0.22 | 18.64 | 0.00 | -7.90 |
| 72 | 1.65 | 17.31 | 0.68 | -7.45 |
| 73 | -0.20 | 17.08 | 0.00 | -7.24 |
| 74 | 1.37 | 16.16 | 0.69 | -7.30 |
| 75 | -0.16 | 16.65 | 0.00 | -7.05 |
| 76 | 1.20 | 15.33 | 0.58 | -6.68 |
| 77 | -0.54 | 15.35 | 0.28 | -8.28 |
| 78 | 0.92 | 13.54 | 0.78 | -7.57 |
| 79 | -0.57 | 13.17 | 0.20 | -7.80 |
| 80 | 1.17 | 10.75 | 0.73 | -7.89 |
| 81 | -0.30 | 10.40 | 0.20 | -8.70 |
| 82 | 0.85 | 9.15 | 0.80 | -8.50 |
| 83 | -0.37 | 10.00 | 0.00 | -7.71 |
| 84 | 1.40 | 10.07 | 0.89 | -6.38 |


| N OR Z | PN | SN | PZ | S2 |
| :---: | :---: | :---: | :---: | :---: |
| 85 | -0.11 | 12.50 | 0.00 | -5.47 |
| 86 | 1.12 | 11.80 | 0.79 | -4.78 |
| 87 | -0.16 | 12.40 | 0.00 | -4.37 |
| 88 | 1.09 | 12.35 | 0.80 | -4.05 |
| 89 | -0.35 | 13.60 | 0.00 | -4.12 |
| 90 | 0.72 | 13.47 | 0.58 | -4.00 |
| 91 | -0.20 | 13.00 | 0.00 | -5.00 |
| 92 | 0.92 | 12.11 | 0.61 | -5.24 |
| 93 | -0.30 | 12.15 | 0.00 | -5.60 |
| 94 | 0.73 | 11.21 | 0.51 | -6.15 |
| 95 | 0.30 | 11.25 | 0.00 | -6.45 |
| 96 | 0.79 | 9.80 | 0.72 | -6.93 |
| 97 | 0.00 | 10.43 | 0.00 | $-7.20$ |
| 98 | 0.80 | 9.70 | 0.77 | -7.74 |
| 99 | 0.00 | 9.75 |  |  |
| 100 | 0.80 | 8.65 |  |  |
| 101 | 0.00 | 8.72 |  |  |
| 102 | 0.80 | 8.23 |  |  |
| 103 | 0.00 | 7.80 |  |  |
| 104 | 0.85 | 7.50 |  |  |
| 105 | 0.00 | 7.40 |  |  |
| 106 | 0.50 | 7.35 |  |  |
| 107 | 0.00 | 7.15 |  |  |
| 108 | 0.75 | 6.96 |  |  |
| 109 | 0.00 | 6.94 |  |  |
| 110 | 0.75 | 6.55 |  |  |
| 111 | 0.00 | 6.72 |  |  |
| 112 | 0.86 | 6.49 |  |  |
| 113 | -0.10 | 6.69 |  |  |
| 114 | 1.10 | 6.00 |  |  |
| 115 | -0.26 | 5.60 |  |  |
| 116 | 0.84 | 4.53 |  |  |
| 117 | $-0.33$ | 4.25 |  |  |
| 118 | 0.72 | 2.60 |  |  |
| 119 | 0.03 | 2.39 |  |  |
| 120 | 0.96 | 1.58 |  |  |
| 121 | 0.10 | 2.15 |  |  |
| 122 | 0.52 | 0.46 |  |  |


| N OR Z | PN | SN | FZ | S2 |
| :---: | :---: | :---: | :---: | :---: |
| $12 \%$ | -0.05 | -0.85 |  |  |
| 124 | 0.41 | -1.69 |  |  |
| 125 | -0. 20 | -2.63 |  |  |
| 126 | 0.38 | -3.16 |  |  |
| 127 | 0.00 | -1.37 |  |  |
| 128 | 0.45 | -0.41 |  |  |
| 129 | 0.00 | 0.71 |  |  |
| 130 | 0.61 | 1.66 |  |  |
| 131 | 0.00 | 2.62 |  |  |
| 132 | 0.78 | 3.22 |  |  |
| 133 | 0.00 | 3.76 |  |  |
| 134 | 0.67 | 4.10 |  |  |
| 135 | 0.00 | 4.46 |  |  |
| 136 | 0.67 | 4.83 |  |  |
| 137 | 0.00 | 5.09 |  |  |
| 138 | 0.79 | 5.18 |  |  |
| 139 | 0.00 | 5.25 |  |  |
| 140 | 0.80 | 4.75 |  |  |
| 141 | 0.00 | 5.09 |  |  |
| 142 | 0.60 | 4.95 |  |  |
| 143 | 0.00 | 5.08 |  |  |
| 144 | 0.49 | 4.87 |  |  |
| 145 | 0.00 | 5.10 |  |  |
| 146 | 0.52 | 4.88 |  |  |
| 147 | 0.00 | 5.58 |  |  |
| 148 | 0.58 | 5.18 |  |  |
| 149 | 0.00 | 5.80 |  |  |
| 150 | 0.39 | 5.50 |  |  |

# CALCULATIONS OF THE DOUBLE DIFFERENTIAL CROCS SECTION FOR PAFTICAL EMISSIONS IM THE CONTINUUA REGION 

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## Abstract

Based on the framework of the width fluctuation-corrected lauser-Feshbach theory, this work will sive the formulae for calculating the double differential cross section in the continuum region. Taking the neutron-induced inelastic scattering as an example, we calculated the spectra for given outgoing angles, the angular distributions and the spectra. All the calculated cross sections are found reasonable and self-consistent.

To study the double differential cross sections of the secondary particles in the continuum region based on the framework of the width-fluctuation-corrected Hauser-Feshbach theory, we should pay attention to the extensions from the original theory. The discretization method has been employed to deal with the width fluctuation correction factor and the statistical description has been used to treat the continuum levels of the residual nucleus as the two essential points.

For the neutron induced reactions the inelastic scattering cross section and angular distribution of kth level ( $u_{k}, I_{k}, h_{k}$ ) can be written in, the form.

$$
\begin{align*}
& \times W_{i, l_{k} i_{k}}^{\prime{ }_{k}}\left(U_{k}\right) . \quad k=0,1, \cdots, K_{m}  \tag{1}\\
& \sigma\left(U_{k}, I_{k}, \pi_{k} ; \theta_{c}\right)=\sum_{t=0,2, \ldots} B_{c}\left(L, U_{k}\right) P_{L}\left(\cos \theta_{c}\right) . \tag{2}
\end{align*}
$$

$$
\begin{aligned}
& B_{c}\left(L, U_{k}\right)=\frac{\hbar_{n}^{2}}{S\left(2 I_{0}+1\right)} \sum_{,: H_{k} i_{k}}(2 J+1) Q_{L}^{\prime}\left(I_{0} I_{k} j j_{k}\right)
\end{aligned}
$$

$$
\begin{align*}
& k=0,1, \cdots, K_{m} \tag{3}
\end{align*}
$$

In particular if the residual nucleus heeps its excitation enerry in the continum region, the inetastic cross section can be written in the form

$$
\begin{equation*}
\sigma_{l}\left(E_{\varepsilon}\right)=\frac{\pi \lambda_{n}^{2}!}{2\left(2 I_{0}+1\right)} \sum_{j \pi i j}(2 J+1) \frac{T_{i}^{\prime \pi}\left(E_{\varepsilon}\right) T_{n}^{\prime n}}{T^{\prime \pi}} \times W_{i j}^{\prime \pi} \tag{4}
\end{equation*}
$$

Where $T_{n}^{J \pi}$ is the transmition coefficient of neutron emission in the continuum region. which reads

Where $\rho_{A}^{J \pi}$ is the level density of residual nucleus $A$ with the excitation energy $u$.

Thus the total transmition coefficient can be obtained by

$$
\begin{equation*}
T^{\prime n}=\sum_{n=0}^{K_{m}} \sum_{i_{n} J_{k}} T_{j_{n} / N_{k}^{\prime}}^{\prime}\left(E_{v}-U_{h}\right)+T_{n}^{\prime n}+T_{j^{\prime}}^{\prime}+T_{v}^{\prime n}+\cdots \tag{6}
\end{equation*}
$$

Where $W_{I_{k} j}\left(I_{k}\right)$ and $W_{I_{j}}$ in (1), (3) and (4) are the width fluctuation correction factor?

$$
\begin{align*}
& W_{l i, i_{h} i_{k}}^{\prime,}\left(U_{h}\right)=\frac{1+2 \delta_{U, U_{k}} \delta_{l l_{h}} \delta_{i_{k}}}{2} \int_{0}^{1} d y \\
& \times \exp \left[-\frac{T_{n}^{\prime \prime}+T_{I^{\prime \prime}}^{\prime x}+T_{y}^{\prime x}}{2 T^{\prime x}}\left(\frac{1}{y}-1\right)\right] y^{-2}\left[1+\frac{T_{i i^{\prime \prime}}^{\prime \prime}\left(E_{c}\right)}{T^{\prime x}}\left(\frac{1}{y}-1\right)\right]^{-1} \\
& \times\left[1+\frac{T_{t_{k} h_{k}}^{\prime}\left(E_{c}-U_{k}\right)}{T^{\prime *}}\left(\frac{1}{y}-1\right)\right]^{-1} \\
& \times \prod_{k^{\prime}=0}^{\kappa_{m}} \prod_{i_{A^{\prime}}} \prod_{i_{A^{\prime}}}\left[1+\frac{r_{h^{\prime}}^{\prime} i_{A^{\prime}}\left(E_{c}-U_{k^{\prime}}\right)}{T^{\prime *}}\left(\frac{1}{y}-1\right)\right]^{-\frac{1}{2}} .  \tag{7}\\
& W_{i i}^{\prime T}=\frac{1}{2} \int_{0}^{1} \mathrm{~d} y \exp \left[-\frac{T^{\prime \prime}+T_{j_{i}^{\prime}}^{\prime}+T_{i x}^{\prime *}}{2 T^{\prime \pi}}\left(\frac{1}{y}-1\right)\right] y^{-2} \\
& \times\left[1+\frac{T_{t}^{\prime \prime}\left(E_{c}\right)}{T^{\prime *}}\left(\frac{1}{y}-1\right)\right]^{-1} \prod_{i=0}^{n} \prod_{i_{n}} \prod_{i_{n}}\left[1+\frac{T_{A_{A_{i}^{\prime}}^{\prime *}}^{\prime *}\left(E_{c}-U_{k^{\prime}}\right)}{T^{\prime \prime}} \times\left(\frac{1}{y}-1\right)\right]^{-\frac{1}{2}} . \tag{8}
\end{align*}
$$

 sion in the continum rerion at a riven outroins ansle, the whole continum refion should be divided into $N$ intervals $\Delta u_{n}(n=1,2, \ldots$ iv). Tinich implies that all of the lovels in $\Delta u_{n}$ are concentrated at the point $u_{n}$. This is the discretization method.

Hence one can get the double differential cross section as

$$
\begin{equation*}
\sigma\left(E_{c}-U_{n}, \theta_{c}\right)=\sum_{I_{n}=1} \rho_{1}^{\prime} n^{n} \cdot\left(U_{0}\right) \sigma\left(U_{n}, I_{n}, \pi_{n} ; \theta_{c}\right) \tag{9}
\end{equation*}
$$

Where $E_{c}-u_{n}$ is the enersy of outcoing neutrons.
The Gilbert-Cameron type level density formula is employed to make the discretization level density as mentioned above.

The width fluctuation correction factor must be modified in our model. The whole particle emissions in the continuum region were treated as one channel only as did before. Then the width fluctuation correction factor in (7), (8) are available. But in our model the discretization method is used to treat the whole continuum region. Thus the width fluctuation correction factor becomes into

$$
\begin{aligned}
& \times y^{-2}\left[1+\frac{T_{i=}^{\prime *}\left(E_{0}\right)}{T_{1}^{\prime *}}\left(\frac{1}{y}-1\right)\right]_{i}^{-1}\left[1+\frac{T_{i_{q}^{\prime}}^{\prime \prime},\left(E_{0}-U_{q}\right)}{T^{\prime 2}}\left(\frac{1}{y}-1\right)\right]^{-1} \\
& \times \prod_{n=0}^{K} \prod_{i_{n},} \prod_{i_{n}}\left[1+\frac{\dot{T}_{i_{n}}^{I_{n}^{n} i_{n}}\left(E E_{j}-U_{h^{\prime}}\right)}{T^{\prime N}}\left(\frac{1}{y}-1\right)\right]^{-\frac{\pi}{2}} \\
& \times \prod_{n=1}^{N} \prod_{I_{n}^{\prime}} \prod_{n_{n}^{\prime}} \exp \left[-\frac{1}{2} \rho_{A_{n}^{\prime} n^{\prime} n_{n}^{\prime}\left(U_{n^{\prime}}\right) \Delta U_{n^{\prime}} \ln \prod_{n^{\prime} I_{n}} .}\right. \\
& \left.\left(1+\frac{T i_{n}^{\prime \prime} j_{n}\left(E_{c}-U_{n}^{\prime}\right)}{T^{\prime \pi}}\left(\frac{1}{y}-1\right)\right)\right] . \\
& q= \begin{cases}k & \text { the discrete energy level } \\
n & \text { the continuum region }\end{cases}
\end{aligned}
$$

The width fluctuation factor for radiation capture channel and fission channel is

$$
\begin{aligned}
& W_{1}^{\prime}:=\frac{1}{2} \int_{0}^{1} \mathrm{~d} y \operatorname{cxp}\left[-\frac{T^{\prime x}+T^{\prime *}}{2 T^{\prime \prime}}\left(\frac{1}{y}-1\right)\right] y^{-:}
\end{aligned}
$$

$$
\begin{align*}
& \left.\left(1+\frac{T_{n}^{\prime} i_{n}\left(E_{c}-U_{n} \cdot\right.}{T^{\prime *}}\left(\frac{1}{y}-1\right)\right)\right] . \tag{11}
\end{align*}
$$

It is easy to extend tive formulae mentioned above into the situation in which the other types of particale imissions are Eakera into account. As before if we need to calculated the spectra of these channels for the fixed outgoing angles only, the particale emissions in whole continuum region was considered as one reaction channel, the formulae (7) and (8) are available. Otherwise, if the double differential cross section is needed to be calculated then formulae (10) and (11) should be used in stead.

The angular distribution for particale emission in continuum will be given by the integration over the energy $E^{\prime}$.

$$
\begin{equation*}
\sigma_{s}\left(\theta_{0}\right)=\int_{0}^{E_{c}-u_{k_{m}}} \mathrm{~d} E^{\prime} \sigma\left(E^{\prime}, \theta_{0}\right) \tag{12}
\end{equation*}
$$

Also the inelastic scattering cross section can be obtained by the integration over the whole solid angle.

$$
\begin{align*}
\sigma_{t}\left(E_{s}\right) & =\int_{I}\left(\theta_{c}\right) \mathrm{d} \Omega_{c} \\
= & 2 \pi \int_{0}^{x} \sin \theta_{c} \mathrm{~d} \theta_{c} \int_{0}^{E_{c}-U_{h_{h}}} \mathrm{~d} E^{\prime} \sigma\left(E^{\prime}, \theta_{c}\right) \tag{13}
\end{align*}
$$

Then the normalized spectra of neutron emission is given by

$$
\begin{equation*}
\frac{\mathrm{d} \sigma}{\mathrm{~d} E^{\prime}}=2 \pi \int_{0}^{\pi} \sin \theta_{c} \mathrm{~d} \theta_{c} \sigma\left(E_{c}-U_{n}, \theta_{c}\right) / \sigma_{l}\left(E_{c}\right) \tag{14}
\end{equation*}
$$

Finally, we took ${ }^{179} \mathrm{Hf}\left(\mathrm{n}, \mathrm{n}^{\prime}\right)$ and ${ }^{180_{\mathrm{Hf}} \mathrm{H}\left(\mathrm{n}, \mathrm{n}^{\prime}\right)}$ as the examples. The inelastic scattering spectra for given outgoing, anfles, the angular distributions and the spectra are shown in Fig. 1.

Fig. 1.



the ongular distributions for ${ }^{180}\left(n, n^{\prime}\right)$
in be calculations the excitation encrese $u_{k m}$ os the firnest discrete enercy level are 0.4761 NeV for ${ }^{179} \mathrm{Hf}, 1.1416$ 离eV for 180 iff respectively.

As shown in fig. i. the spectras of emitted particlae have the characteristices of Marxwell distribution, and the angular distribution are nearly istropic. Obersely, this is caused by the theoretical model limited in the statistical theory for compound nucleus and the mechanism of preequilibriun stare is exclusive.
(1). Su Zongai et al., China Phys., 1, 952(1981).

#  <br> AND NUCLEAR DATA CAi,CUATMLOR <br> Su Zongdi, Shi Xiangjun, Li Cha, Zhuang Youxiang (Institute of Atomic Energy, Academia Sinica, P. O. Box 275, Beijing, China) 

Abstract Our work about the statistical theory of compound nucleus reaction and its application in nuclear data calculations is summarized. The work involves deriving some formulas, adjusting the level density parameters and calculating neutron reaction data.

## 1. INTRODUCTION

The optical model, the statistical theory of the nuclear reaction etc. are the most important and widely used theories in the calculation of the neutron reaction data. This paper will only concern the work about the statistical theory of the compound nucleus reaction and its applications. The different effects in compound nucleus reaction have been compared. The unified treatment of the width fluctuation correction (WFC) have been made for including all kinds of reaction channels in a compound nucleus reaction. We adjusted a new set of level density parameters, which are worth recomnending for use, Finally, we shall present briefly our calculated results.

## 2. THE COMPARISON OF THE DIFFERENT EFFECTS

The statistical theory of nuclear reaction has gone through a series of steps, which are the Hauser-Feshbach (HF) theory ${ }^{\prime}$, the width fluctuation corrected Hauser-Feshbach (WHF) theory 2, the Moldauer theory 3 , the Weidenmuller theory 4 etc. respectively. Today it has become even perfect, accurate and self-consistent. In this work we have studied and compared the effects
of the different corrections. For example, the wFC can increase compound elastic scattering cross section and decrease the other cross sections. In contrast the level-level correlation which is taken into account in the Moldauer theory will decrease the compound elastic cross section and increase others. But these effects can not be just canceled with each other. The compound elastic cross section based on the WHF theory is twice or three times larger than one on the HF theory. The calculated results of various cross sections of $n+235 \mathrm{U}$ reaction are given in Table 1 as a special example

Table 1 The comparison of the results of the WHF theory and the HF theory

| $\begin{gathered} \mathrm{En} \\ (\mathrm{MeV}) \end{gathered}$ | Calculation Method | $\sigma_{a b s}$ <br> (b) | $\sigma_{c e}$ <br> (b) | $\sigma_{\text {non }}$ <br> (b) | $\sigma_{i n}$ <br> (b) | $\sigma_{n, F}$ <br> (b) | $\sigma_{n, r}$ <br> (b) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 0.01 | HF | 4.18 | 0.45 | 3.73 | 0 | 2.36 | 1.37 |
|  | WHF | 4.18 | 0.90 | 3.28 | 0 | 2.13 | 1.15 |
| 0.1 | HF | 3.27 | 0.55 | 2.72 | 0.57 | 1.58 | 0.56 |
|  | WHF | 3.27 | 0.98 | 2.29 | 0.44 | 1.38 | 0.47 |
| 0.3 | HF | 3.31 | 0.48 | 2.83 | 1.59 | 1.00 | 0.24 |
|  | WHF | 3.31 | 0.84 | 2.47 | 1.33 | 0.92 | 0.22 |
| 0.5 | HF | 3.14 | 0.36 | 2.78 | 1.69 | 0.91 | 0.19 |
|  | WHF | 3.14 | 0.60 | 2.54 | 1.54 | 0.82 | 0.17 |

However, the level-level correlätion will affect the compound elastic cross section only by twenty percent as compared with that of the WHF theory. In order to compare the level-level correlation with the WFC, the calculated results and the experimental values of ${ }^{238} \mathrm{U}\left(\mathrm{n}, \mathrm{n}^{\prime}\right)^{238} \mathrm{U}_{\mathrm{U}}{ }^{\left(2^{+}, 0.0449 \mathrm{MeV}\right)}$ inelastic scattering are given in Fig 1.


Fig1. The inelastic scattering excitation curves calculated by two kinds of theories (the WHF theory and the Moldauer theory) and the experimental values for ${ }^{238} \mathrm{U}\left(\mathrm{n}, \mathrm{n}^{\prime}\right)^{238} \mathrm{U}^{*}$ $\left(2^{+}, 0.0449 \mathrm{MeV}\right)$ reaction.

It is seen from the above results that these effects only have a significance in the energy region where the compound elastic cross section is big enough. And the WFC is dominant in the sta $=$ tistical theory of the nuclear reaction. In fact, these secondary are concealed by the parameterization in the nuclear reaction calculation. Therefore, the WHF theory is commonly used to calculate the nuclear data today.

## 3. THE WIDTH FLACTUATION CORRECTION

We have developed the WHF theory and applied to a nuclear reaction process including radiative capture, fission and particle emission with the residual nucleus in the continuum. The formula of the double differential cross section for particle emission in the continuum region has been derived within the framework of the WHF theory. The cumbersome problems in the derivation of the double differential cross section are how to deal with the continuum states. First, we use a discretization method. The whole continuum region is divided to $n$ intervals. The levels in each interval $\Delta U_{n}$ are accumulated on one point $U_{n}$. Therefore, a discretized and degenerate spectrum is formed. Secondly, the continuum levels of residual nucleus are described by a statistical method. Thus, the spectra for given outgoing angles can be written as

$$
\sigma\left(E,-U_{n}, 0_{c}\right)=\sum_{I_{n} x_{n}} \rho_{A^{n}}^{\prime} \cdot\left(U_{n}\right) \sigma\left(U_{n}, I_{n}, \pi_{n} ; \theta_{c}\right) .
$$

Where $E_{C}-U_{n}$ is the energy of the outgoing particles, $\rho_{A}^{I_{n} \delta_{n}^{\prime}}\left(U_{n}\right)$ is the level density of the residual nucleus. The $\sigma\left(U_{n}, I_{n}, \pi_{n}, \theta_{c}\right)$ can be obtained from the angular distribution calculation of the discrete energy level, which have been simplified 5 . In the angular distribution and the cross section formulas for particle emissions in discrete states and in continuum the WFC factor should

$$
\begin{aligned}
& W_{c c_{q}^{\prime}}= \frac{1+2 \delta_{c c_{q}^{\prime}}}{2} \int_{0}^{1} d y \exp \left[-\frac{T_{f}^{\mu}+T_{\gamma}^{\mu}}{2 T^{\mu}}\left(\frac{1}{y}-1\right)\right] \\
& \times y^{-2}\left[1+\frac{T_{c}\left(E_{2}\right)}{T^{\mu}}\left(\frac{1}{y}-1\right)\right]^{-1}\left[1+\frac{T_{c_{q}^{\prime}}\left(E_{c}-U_{q}\right)}{T^{\mu}}\left(\frac{1}{y}-1\right)\right]^{-1} \\
& x \prod_{k}^{K_{M}} \prod_{c_{k}^{\prime}}\left[1+\frac{T_{c_{k}^{\prime}}\left(E_{c}-U_{k^{\prime}}\right)}{T^{\mu}}\left(\frac{1}{y}-1\right)\right]^{-\frac{1}{2}} \\
& \times \prod_{n^{\prime} I_{n^{\prime}} \pi_{n^{\prime}}}^{n}\left[-\frac{1}{2} \rho_{A}^{I_{n^{\prime}} \cdot \pi_{n^{\prime}}}\left(U_{n^{\prime}}\right) \Delta U_{n^{\prime}} \ln \prod_{c_{n^{\prime}}}\left[1+\frac{T_{c_{n}}\left(E_{c}-U_{n^{\prime}}\right)}{T^{\mu}}\left(\frac{1}{y}-1\right)\right],\right. \\
&-q= \begin{cases}k & \text { the discrete energy level, }, \\
n & \text { the continuum region. }\end{cases}
\end{aligned}
$$

For the radiative capture and fission channels, the WFC factor can be similarly obtained

$$
\begin{align*}
W_{c} & =\frac{1}{2} \int_{0}^{1} d y \exp \left[-\frac{T_{t}^{\mu}+T_{r}^{\mu}}{2 T^{\mu}}\left(\frac{1}{y}-1\right)\right] y^{-2} \\
& \times\left[1+\frac{T_{c}\left(E_{c}\right)}{T^{\mu}}\left(\frac{1}{y}-1\right)\right]^{-1} \prod_{k}^{K_{M}} \prod_{c_{k}}\left[1-\frac{T_{c^{\prime}}\left(E_{c}-U_{k^{\prime}}\right)}{T^{\mu}}\left(\frac{1}{y}-1\right)\right]^{-\frac{1}{2}}  \tag{3}\\
& \times \prod_{n^{\prime}}^{n} \prod_{I_{n} \pi_{n}}\left[-\frac{1}{2} \int_{A}^{I_{n^{\prime}} \pi_{n^{\prime}}}\left(U_{n^{\prime}}\right) \Delta U_{n} \ln \prod_{c_{n^{\prime}}}\left[1+\frac{T_{c_{n}}\left(E_{c}-U_{n^{\prime}}\right)}{T^{\mu}}\left(\frac{1}{y}-1\right)\right] .\right.
\end{align*}
$$

Making use of the above formulas, the spectra for given outgoing angles and various reaction cross sections can be calculated.

Therefore, the angular distributions, the spectra and inelastic cross sections for particle emissions in continuum recion are readily calculated by the following formulas, respectively

$$
\begin{align*}
\sigma_{I}\left(\theta_{c}\right) & =\int_{0}^{E_{c}-U_{k M}} d E^{\prime} \sigma\left(E^{\prime}, \theta_{c}\right)  \tag{4}\\
\frac{d \sigma_{I}}{d E^{\prime}} & =2 \pi \int_{0}^{\pi} \sin \theta_{c} d \theta_{c} \sigma\left(E_{c}-U_{n}, \theta_{c}\right) / \sigma_{I}\left(E_{c}\right)  \tag{5}\\
\sigma_{I}\left(E_{c}\right) & =\int_{\Omega} \sigma_{I}\left(\theta_{c}\right) d \Omega \\
& \equiv 2 \pi \int_{0}^{\pi} \sin \theta_{c} d \theta_{c} \int_{0}^{E_{c}-U_{k_{M}}} d E^{\prime} \sigma\left(E^{\prime}, \theta_{c}\right) \tag{6}
\end{align*}
$$

We take the neutron-induced inelastic scattering as an example, and calculate the spectra for given outgoing angles, the angular distributions, and the spectra. All the calculated cross sections are found reasonable and consistent with each other. The calculated results have the typical feature of the compound nucleus reaction. The spectra are Maxwell type spectra. And the angular distributions are nearly isotropic.

The particle emissions in the continuum region can be described by one transmition coefficient $\mathrm{T}_{\mathrm{p}}^{\mu}$ when the double differential cross section need not be calculated. In this way a unified formula of the WFC was obtained, which is available for fission, radiative capture and particle emission in continuum simultaneously ${ }^{6}$. The WFC factor of the particle emissions with the residual nucleus in a discrete energy level region can be written as

$$
\begin{align*}
& W_{c c_{k}^{\prime}}=\frac{1+2 \delta_{c c_{k}^{\prime}} \int_{0}^{1} d y \exp \left[-\frac{T_{p}^{\mu}+T_{f}^{\mu}+T_{x}^{\mu}}{2}\left(\frac{1}{y}-1\right)\right]}{} \begin{array}{l}
T^{\mu} y^{-2}\left[1+\frac{T_{c}\left(E_{c}\right)}{T^{\mu}}\left(\frac{1}{y}-1\right)\right]^{-1}\left[1+\frac{T_{c_{k}^{\prime}}\left(E_{c}-U_{k}\right)}{T^{\mu}}\left(\frac{1}{y}-1\right)\right]^{-1} \\
\quad \times \prod_{k}^{k_{M}} \prod_{C_{k^{\prime}}}\left[1-\frac{T_{c k}\left(E_{c}-U_{k^{\prime}}\right)}{T^{\mu}}\left(\frac{1}{y}-1\right)\right]^{-1 / 2}
\end{array} .
\end{align*}
$$

And for radiative capture, fission and the particle emission in continuum, the WFC factor can be similarly written as

$$
\begin{align*}
& W_{c}^{\mu}=\frac{1}{2} \int_{0}^{1} d y \exp \left[-\frac{T_{p}^{\mu}+T_{f}^{\mu}+T_{\gamma}^{\mu}}{2 T^{\mu}}\left(\frac{1}{y}-1\right)\right] \\
& \quad \times y^{-2}\left[1+\frac{T_{c}\left(E_{c}\right)}{T^{\mu}}\left(\frac{1}{y}-1\right)\right]^{-1} \prod_{k^{\prime}}^{K_{\mu}} \prod_{C_{k}^{\prime}}\left[1+\frac{T_{C_{k}^{\prime}}\left(E_{c}-U_{k}\right)}{T^{\mu}}\left(\frac{1}{y}-1\right)\right]^{-1 / 2} . \tag{8}
\end{align*}
$$

These formulas are successful in practical application and can also be extended to other theory (for example, the Moldauer theory).
4. A NEW SET OF LEVEL DENSITY PARAMETERS

The nuclear level density plays an important role in the statistical theory of nuclear reaction. It affects directly the calculation results of the widths, cross sections, emitted particle spectra etc. for various reaction channels. The Gilbert-Cameron ? type level density formula as well as parameters has been widely employed in the statistical theory. Based on their work, Cook et al ${ }^{8}$. have also got a set of parameters. Because these parameters only depend on $Z$ and $N$, i.e. the numbers of proton and neutron in the nucleus respectively, Gilbert-Cameron's formula is universally applicable. However, their parameters are based on the experimental data at that time. The present work attempts to obtain a new set of parameters based on the newest and more kinds of experimental data.

In order to recommend more accurate and reliable parameters of level density, we have collected and evaluated three kinds of experimental data: average level spacing $D_{0}$ and radiative capture width $\Gamma_{\gamma}^{0}$ at the neutron binding energy and the cumulative number $N_{0}$ of the levels at the low exitation energy, which amount to 667 and were reported from 1973 to 1983.

Based on these experimental data, especially $D_{0}$ and $N_{O}$, the parameters $S(Z), S(N), P(Z)$ and $P(N)$ have been readjusted (denoted as work 1). Next the empirical parameters have also been reobtained to fit the experimental data which we have collected. In the Gilbert-Cameron's formula, the corresponding quantities
including the empirical parameters are as follows:

$$
\begin{align*}
& a=[0.00917(S(Z)+S(N))+Q B] A \\
& U_{x}=2.5+150 / A \tag{9}
\end{align*}
$$

The new ones of $0.00880,1.4$ and 263 (denoted as work 2) are adopted respectively instead of $0.00917,2.5$ and 150 used in the formulas. Thus, the formulas are rewritten as

$$
\begin{align*}
& \mathrm{a}=[0.00880(\mathrm{~S}(\mathrm{Z})+\mathrm{S}(\mathrm{~N}))+\mathrm{QB}] \mathrm{A} \\
& \mathrm{U}_{\mathrm{x}}=1.4+263 / \mathrm{A} \tag{10}
\end{align*}
$$

where the values of $Q B$ are taken as:

$$
Q B= \begin{cases}0.142, & 29 \leqslant Z \leqslant 62(\text { and } N \leqslant 89) ; 79 \leqslant Z \leqslant 85 ;  \tag{11}\\ 0.120 & Z \leqslant 29 ; 62(\text { and } N>89) \leqslant Z \leqslant 78 ; Z \geqslant 86 .\end{cases}
$$

The relative deviations of the calculated values in our work to the experimental data for each nucleus whose data have been collected are obtained. The percentages of those nuclei with the relative deviation in a certain scope are compared with that of Gilbert-Cameron's and Cook's which are shown in Table 2.

Table 2 A comparison between three results

| relative deviations | Gilbertcameron | Cook et al | work 1 | work 2 |
| :---: | :---: | :---: | :---: | :---: |
| \$ $\pm 20 \%$ | 21.5 | 22.4 | 47.8 | 48.7 |
| 8D $\leqslant \pm 50 \%$ | 29.4 | 28.1 | 36.4 | 37.3 |
| $0 \leqslant \pm 100 \%$ | 24.1 | 27.6 | 15.4 | 13.2 |
| $> \pm 100 \%$ | 25.0 | 21.9 | 0.4 | 0.8 |
| $\boldsymbol{x}^{2}{ }_{0}$ | 175.9 | 120.9 | 11.3 | 11.4 |
| $\leqslant \pm 20 \%$ | 13.0 | 13.9 | 18.5 | 18.9 |
| $\leqslant \pm 50 \%$ | 30.7 | 23.9 | 36.1 | 31.5 |
| $0 \leqslant \pm 100 \%$ | 49.2 | 50.0 | 41.2 | 44.5 |
| $> \pm 100 \%$ | 7.1 | 12.2 | 4.2 | 5.1 |
| $x^{2}{ }_{0}$ | 13.9 | 47.9 | 11.5 | 10.1 |
| $1 \leqslant \pm 20 \%$ | 26.9 | 26.4 | 27.9 | 30.8 |
| $\int \Gamma^{0} \leqslant \pm 50 \%$ | 39.8 | 40.8 | 37.3 | 35.8 |
| Or $\gamma \leqslant \pm 100 \%$ | 30.3 | 29.8 | 31.8 | 30.8 |
| $> \pm 100 \%$ | 3.0 | 3.0 | 3.0 | 2.6 |
| $x_{r}$ | 9.6 | 9.7 | 11.0 | 9.3 |
| $x_{D_{0}}^{2}+\chi_{N_{0}}^{2}+x_{r_{2}^{*}}^{2}$ | 199.4 | 178.5 | 33.8 | 30.8 |

It is shown in Table 2 that our parameters are in a good agreement with these more accurate and newer experimental data, especially the level spacings $D_{0}$. Intuitively this new set of level density parameters seems to be more accurate. Of course, it should be tested by the calculations on statistical models. The new set of parameters $S(Z), S(N), P(Z)$ and $P(N)$ are listed in Table 3. But it is worth noting that there are some problems remained at the low excitation energies in the Gilbort-Cameron formula to be improved.

## 5. SOME CALCULATED RESULTS

The neutron reaction data (including all kinds of cross sections, the outgoing neutron spectra and the neutron angular distributions) for a series of uranium, plutonium isotopes and some natural elements (e.g. Hf, $\mathrm{Zr}, \mathrm{Cr}$ etc.) have been calculated with the WHF theory and the Moldauer theory. The calculated neutron energy region is from 1 keV to several MeV . The comparison of some calculated results with the experimental data is given in Fig 2-3. It can be seen from these figures, that the results are in a good agreement with the evaluated experimental data. We also extended the calculation procedure to the regions where lack of the experimental data after we had made a systematic review of the relavant parameters. Some results for the fissionable nuclei are given Fig. 4.

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Fig 2. The comparison, between the calculated cross sections and the experimental data for $n+{ }^{235} U, n+{ }^{238} U$ and $n+{ }^{239} \mathrm{Pu}$ reactions.


Fif 3. The comparison between the calculated cross sections and the experimental data for $n+{ }^{N} H f, n+{ }^{i N}$ Zr reactions





 sections rot some $\dot{\sim}$ and fu isotopes whicit are lacitime in the experimertal data

Table 3 the level dersity paranaber of this work

| $\therefore$ OR Z | FH | $9 \%$ | F | 9\% |
| :---: | :---: | :---: | :---: | :---: |
| 11 | 0.00 | 6.80 | 0.00 | -5.40 |
| 12 | 2.67 | 7.93 | 2.00 | -3.80 |
| 13 | 0.00 | 7.55 | 0.00 | -5.72 |
| 14 | 1.80 | 7.21 | 2.09 | -7.80 |
| 15 | 0.00 | 7.44 | 0.00 | -8.97 |
| 16 | 1.67 | 8.07 | 1.45 | -0.50 |
| 17 | 0.00 | 9.10 | 0.00 | -10.10 |
| 18 | 1.86 | 9.81 | 1.62 | -10.70 |
| 19 | 0.00 | 10.50 | 0.00 | -11.38 |
| 20 | 2.01 | 11.39 | 1.75 | -12.00 |
| 21 | 0.00 | 13.85 | 0.00 | -12.10 |
| 22 | 1.64 | 13.68 | 1.53 | -12.80 |
| 23 | -0.30 | 14.50 | -0.26 | -12.20 |
| 24 | 1.40 | 14.10 | 1.20 | -14.30 |
| 25 | -0.10 | 14.93 | 0.00 | -14.00 |
| 26 | 1.36 | 13.70 | 1.30 | -14.40 |
| 27 | -0.20 | 14.55 | -0.10 | -14.75 |
| 28 | 1.20 | 12.50 | 1.20 | -16.57 |
| 29 | 0.00 | 13.70 | 0.00 | -17.10 |
| 30 | 1.15 | 13.45 | 1.06 | -17.95 |
| 31 | -0.20 | 15.45 | 0.00 | -17.68 |
| 32 | 1.32 | 14.98 | 1.06 | -16.78 |
| 33 | -0.15 | 15.80 | 0.00 | -16.65 |
| 34 | 1.30 | 16.38 | 1.35 | -16.35 |
| 35 | -0.18 | 17.25 | 0.00 | -16.70 |
| 36 | 1.50 | 17.55 | 1.10 | -15.20 |
| 37 | -0.10 | 18.05 | 0.20 | -16.20 |
| 38 | 1.90 | 17.65 | 1.24 | -16.41 |
| 39 | 0.00 | 18.59 | 0.29 | -17.05 |
| 40 | 1.43 | 18.71 | 1.35 | -16.75 |
| 41 | -0.28 | 18.95 | 0.00 | -15.40 |
| 42 | 1.50 | 18.99 | 1.17 | -15.85 |
| 43 | -0.10 | 18.60 | -0.15 | -15.50 |
| 44 | 1.70 | 18.25 | 1.15 | -17.29 |
| 45 | -0.10 | 18.20 | -0.10 | -18.44 |
| 46 | 1.57 | 17.38 | 1.20 | -17.82 |


| $\because \cap \mathrm{R} \mathrm{Z}$ | PN | S | PZ | 32 |
| :---: | :---: | :---: | :---: | :---: |
| 47 | 0.00 | 16.72 | 0.34 | -18.62 |
| 48 | 1.60 | 14.25 | 1.36 | -18.58 |
| 49 | -0.28 | 14.88 | 0.28 | -19.65 |
| 50 | 1.20 | 12.88 | 1.47 | -19.95 |
| 51 | 0.10 | 12.95 | 0.00 | -19.44 |
| 52 | 0.75 | 13.65 | 1.00 | -17.75 |
| 53 | -0.30 | 14.80 | 0.00 | -17.70 |
| 54 | 1.12 | 15.20 | 1.12 | -16.72 |
| 55 | 0.00 | 15.89 | 0.26 | -16.95 |
| 56 | 1.15 | 17.15 | 1.20 | -14.37 |
| 57 | 0.00 | 17.00 | 0.10 | -14.44 |
| 58 | 0.80 | 17.73 | 1.40 | -14.00 |
| 59 | -0.25 | 19.75 | 0.93 | -12.30 |
| 60 | 1.35 | 18.80 | 1.36 | -11.91 |
| 61 | -0.32 | 19.45 | 0.36 | -11.59 |
| 62 | 1.25 | 19.51 | 1.22 | -10.78 |
| 63 | -0.44 | 20.17 | -0.20 | -11.40 |
| 64 | 1.14 | 19.85 | 0.97 | -10.21 |
| 65 | 0.08 | 19.98 | -0.20 | -10.00 |
| 66 | 1.32 | 20.05 | 0.78 | -9.35 |
| 67 | -0.11 | 20.28 | 0.00 | -9.30 |
| 68 | 1.15 | 19.83 | 0.52 | -8.85 |
| 69 | -0.16 | 20.20 | 0.00 | -8.91 |
| 70 | 1.06 | 19.44 | 0.68 | -8.13 |
| 71 | 0.22 | 18.64 | 0.00 | -7.90 |
| 72 | 1.65 | 17.31 | 0.68 | -7.45 |
| 73 | -0.20 | 17.08 | 0.00 | -7.24 |
| 74 | 1.37 | 16.16 | 0.69 | -7.30 |
| 75 | -0.16 | 16.65 | 0.00 | -7.05 |
| 76 | 1.20 | 15.33 | 0.58 | -6.68 |
| 77 | -0.54 | 15.35 | 0.28 | -8.28 |
| 78 | 0.92 | 13.54 | 0.78 | -7.57 |
| 79 | -0.57 | 13.17 | 0.20 | -7.80 |
| 80 | 1.17 | 10.75 | 0.73 | -7.89 |
| 81 | -0.30 | 10.40 | 0.20 | -8.70 |
| 82 | 0.85 | 9.15 | 0.80 | -8. 50 |
| 83 | -0.37 | 10.00 | 0.00 | -7.71 |
| 84 | 1.40 | 10.07 | 0.89 | -6.38 |


| YOR 2 | Pi | 3N | PZ | S2 |
| :---: | :---: | :---: | :---: | :---: |
| 85 | -0.11 | 12.50 | 0.00 | -5.47 |
| 6 | 1.12 | 1. 80 | 0.79 | -4.78 |
| 87 | -0.16 | 12.40 | 0.00 | -4.37 |
| 88 | 1.09 | 12.35 | 0.80 | -4.05 |
| 89 | -0.35 | 13.60 | 0.00 | -4.12 |
| 90 | 0.72 | 13.47 | 0.58 | -4.00 |
| 91 | -0.20 | 13.00 | 0.00 | -5.00 |
| 92 | 0.92 | 12.11 | 0.61 | -5.24 |
| 93 | -0.30 | 12.15 | 0.00 | -5.60 |
| 94 | 0.73 | 11.21 | 0.51 | -6.15 |
| 85 | 0.30 | 11.25 | 0.00 | -6.45 |
| 95 | 0.79 | 9.80 | 0.72 | -6.93 |
| 97 | 0.00 | 10.43 | 0.00 | -7.20 |
| 98 | 0.80 | 9.70 | 0.77 | -7.74 |
| 99 | 0.00 | 9.75 |  |  |
| 100 | 0.80 | 8.65 |  |  |
| 101 | 0.00 | 8.72 |  |  |
| 102 | 0.80 | 8.23 |  |  |
| 103 | 0.00 | 7.80 |  |  |
| 104 | 0.85 | 7.50 |  |  |
| 105 | 0.00 | 7.40 |  |  |
| 106 | 0.50 | 7.35 |  |  |
| 107 | 0.00 | 7.15 |  |  |
| 108 | 0.75 | 6.96 |  |  |
| 109 | 0.00 | 6.94 |  |  |
| 110 | 0.75 | 6.55 |  |  |
| 111 | 0.00 | 6.72 |  |  |
| 112 | 0.86 | 6.49 |  |  |
| 113 | -0.10 | 6.69 |  |  |
| 114 | 1.10 | 6.00 |  |  |
| 115 | -0.26 | 5.60 |  |  |
| 116 | 0.84 | 4.53 |  |  |
| 117 | -0.33 | 4.25 |  |  |
| 118 | 0.72 | 2.60 |  |  |
| 119 | 0.03 | 2.39 |  |  |
| 120 | 0.96 | 1.58 |  |  |
| 121 | 0.10 | 2.15 |  |  |
| 122 | 0.52 | 0.46 |  |  |




[^0]:    * The values calculated by use of formula (7) or (6).

