

# Second Advanced Workshop on Model Codes for Spallation Reactions

Saclay, February 8-11, 2010

## Results of the de-excitation code ABLA07

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# **ABLA07**

**see proceedings of the  
“Joint ICTP-IAEA Advanced Workshop on Model Codes for  
Spallation Reactions,,  
held in Trieste, Italy, 4-8 January 2008**

# ABLA07

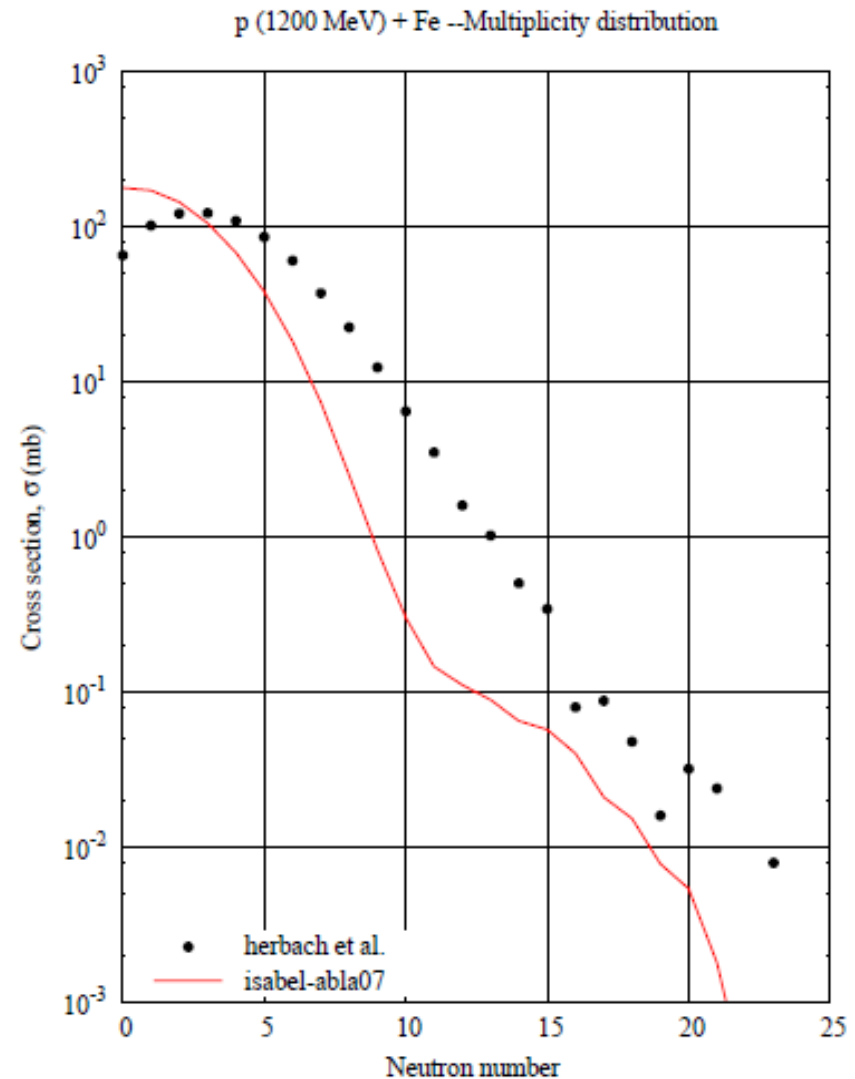
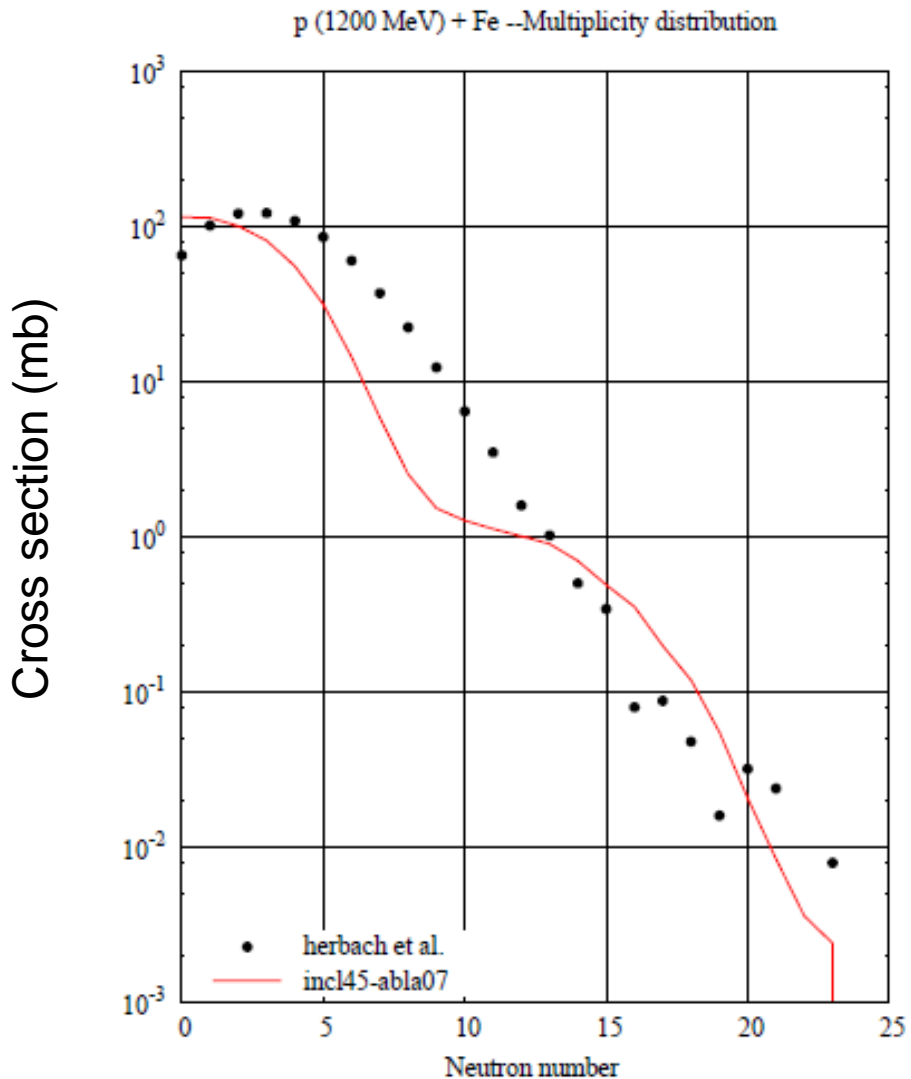
1. Emission of neutrons, LCP ( $Z=1, 2$ ), IMF ( $Z>2$ ) and  $\gamma$  is considered.
2. Particle decay widths based on Weisskopf-Ewing formalism, with:
  - Energy dependent inverse cross sections based on nuclear potential
  - Barriers for charged particles are calculated using the Bass potential
  - Thermal expansion of the source is taken into account.
  - Change of angular momentum due to particle emission is considered.
3. The fission decay width is described by including:
  - An analytical time-dependent approach to the solution of the Fokker-Planck equation,
  - The influence of the initial deformation on the fission decay width,
  - The double-humped structure in the fission barriers of actinides,
  - Symmetry classes in low-energy fission.
4. Particle emission on different stages, i.e. between ground state and saddle point, between the saddle and scission point, and from two separate fission fragments, of the fission process is calculated separately.
5. Kinetic-energy spectra from Maxwell-Boltzmann distribution
6. A stage of simultaneous break-up (MF) in the decay of hot excited systems is treated.

# Neutron multiplicities

# p(1200 MeV) + Fe – Neutron multiplicity distribution

## INCL45-ABLA07

## ISABEL-ABLA07

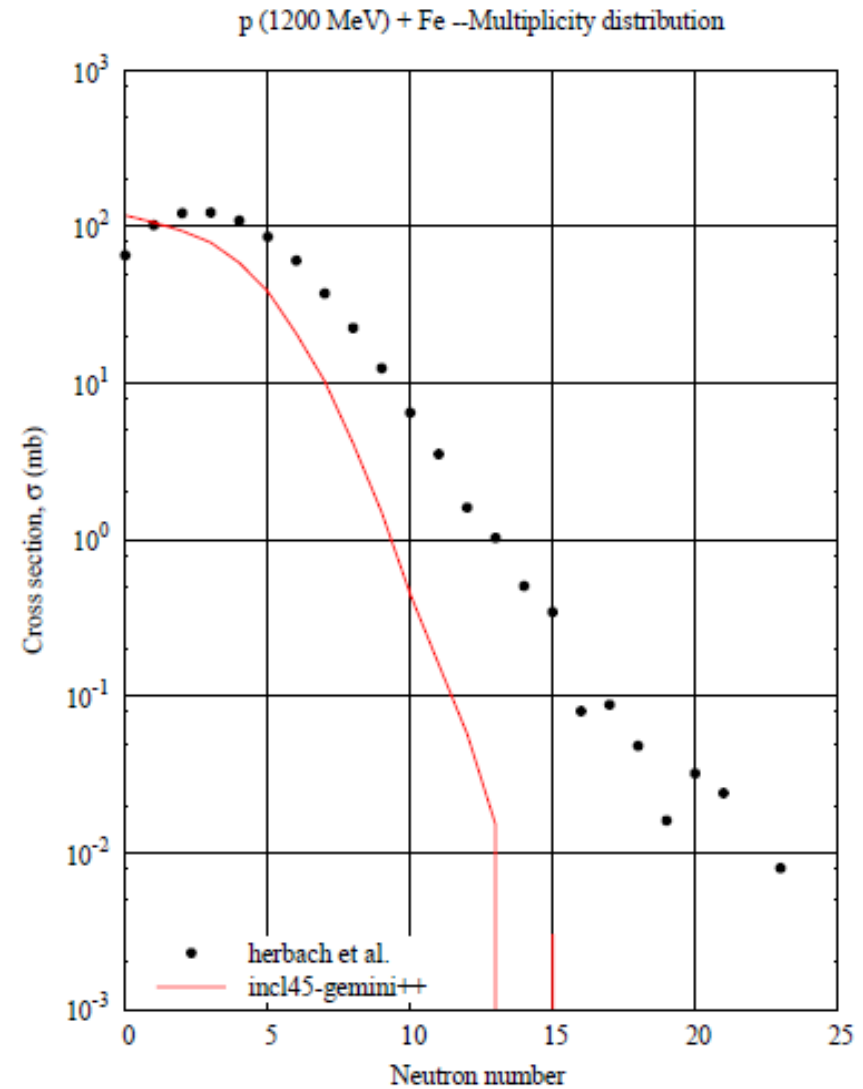
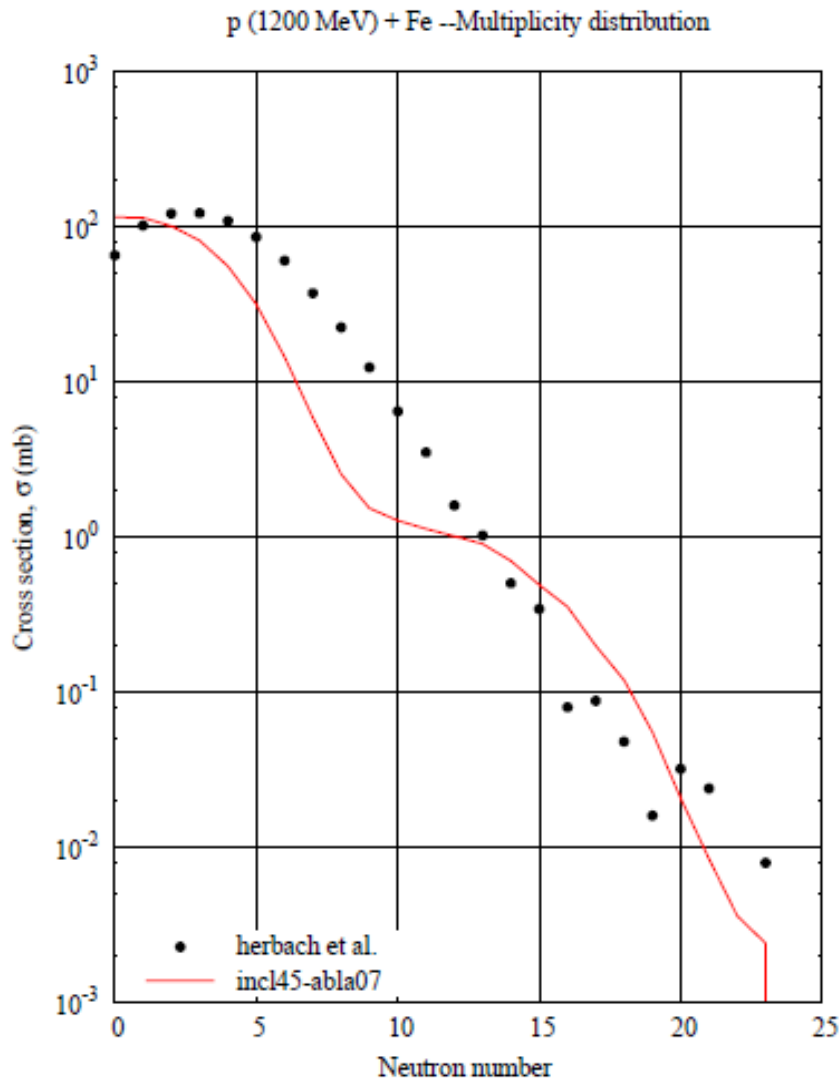


neutron number

# p(1200 MeV) + Fe – Neutron multiplicity distribution

## INCL45-ABLA07

## INCL45-GEMINI++

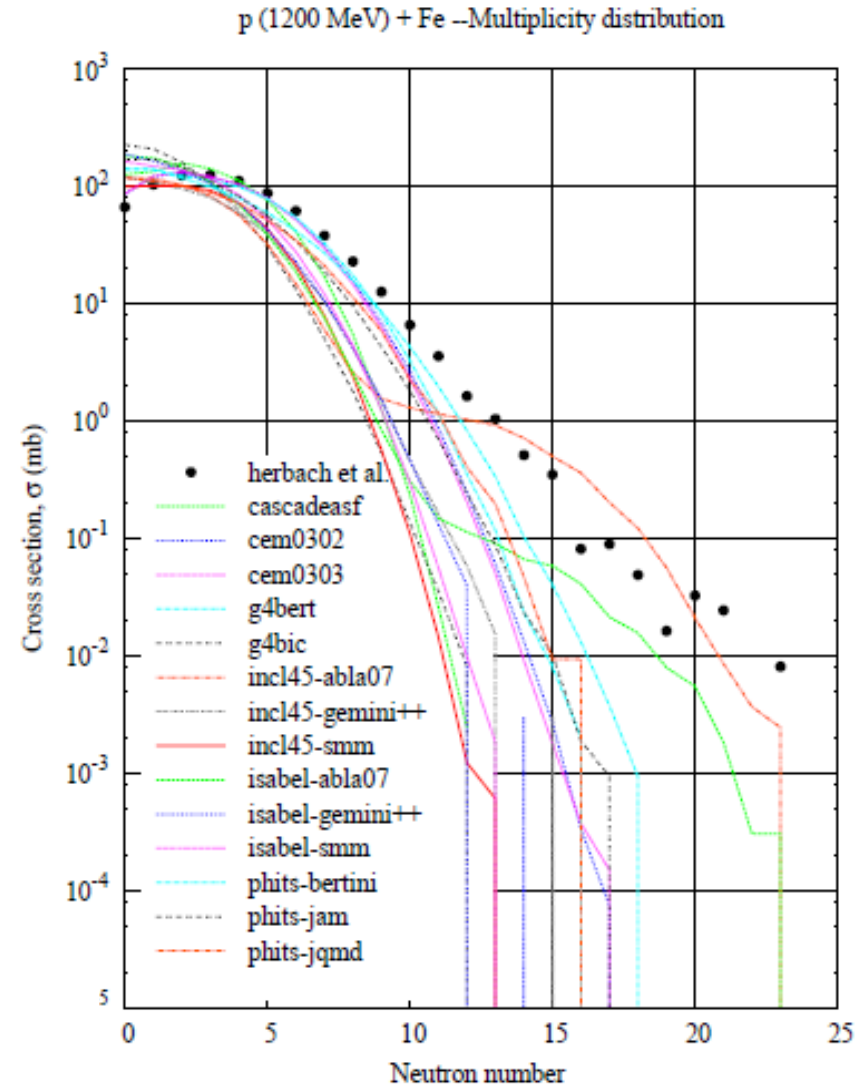
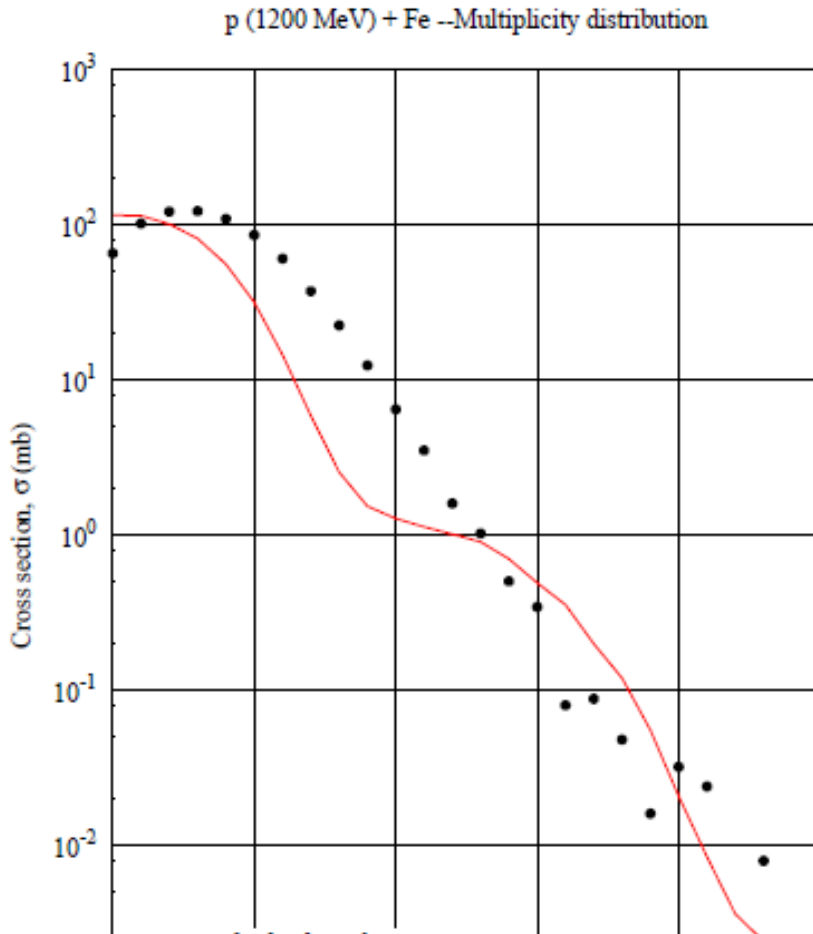


neutron number

# p(1200 MeV) + Fe – Neutron multiplicity distribution

## INCL45-ABLA07

## ALL MODELS



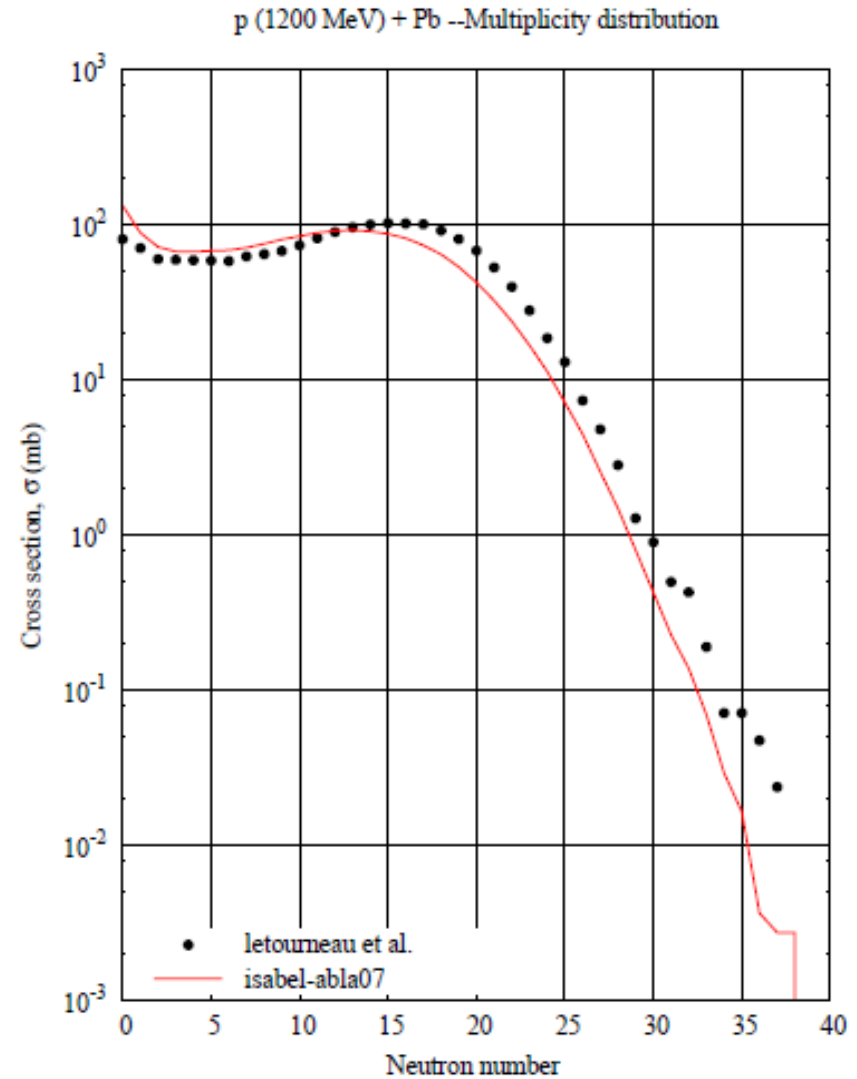
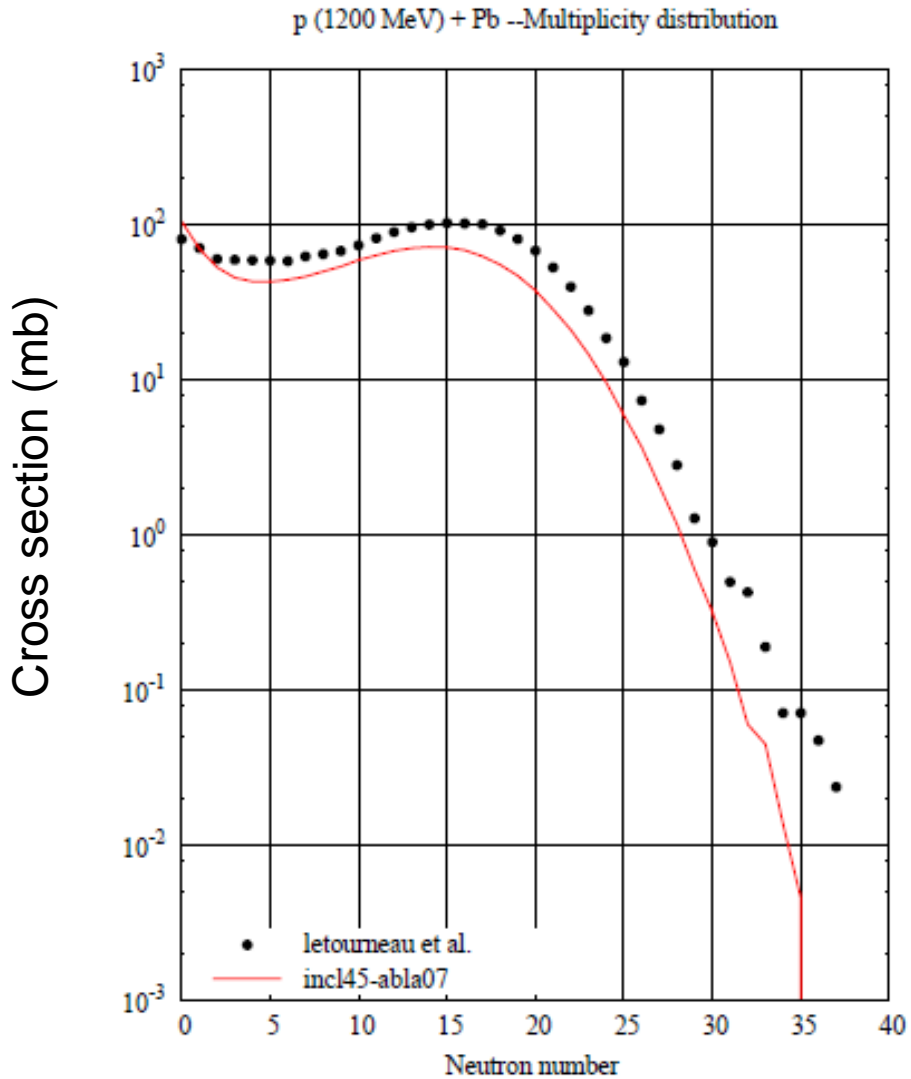
**Strange shape: maybe multifragmentation?**

neutron number

# p(1200 MeV) + Pb – Neutron multiplicity distribution

## INCL45-ABLA07

## ISABEL-ABLA07



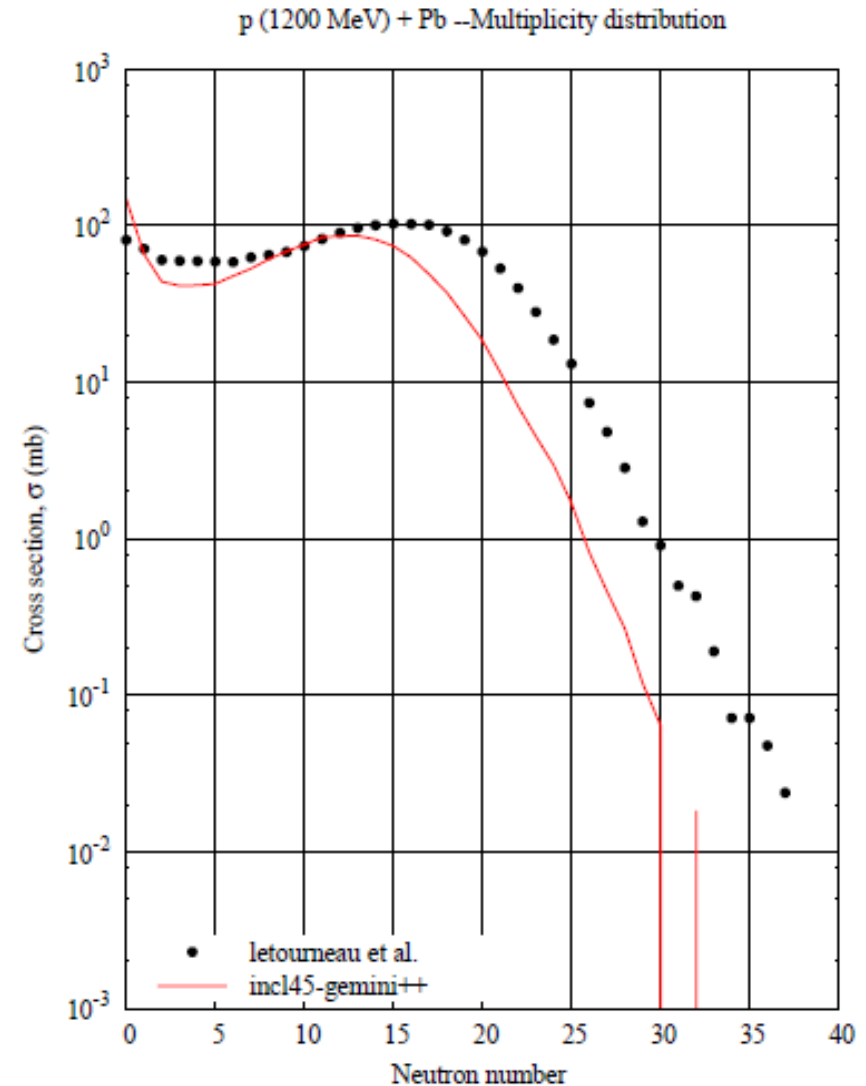
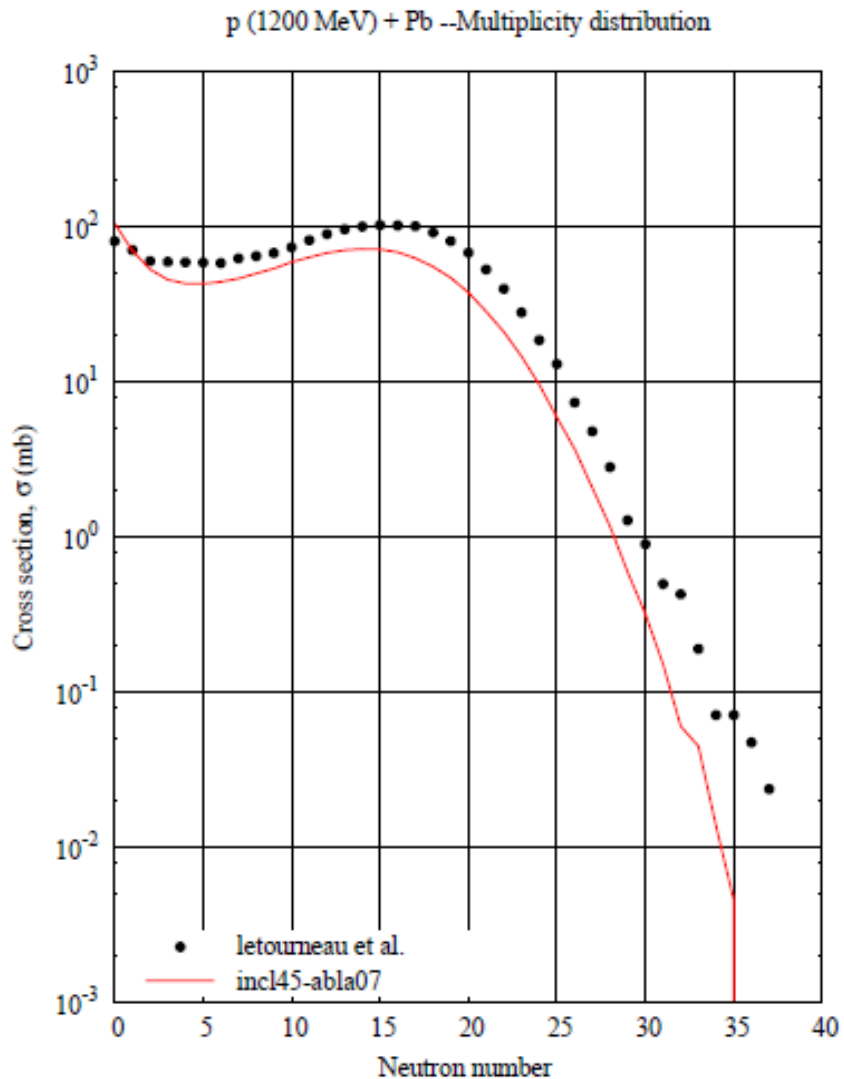
neutron number



# p(1200 MeV) + Pb – Neutron multiplicity distribution

## INCL45-ABLA07

## INCL45-GEMINI++

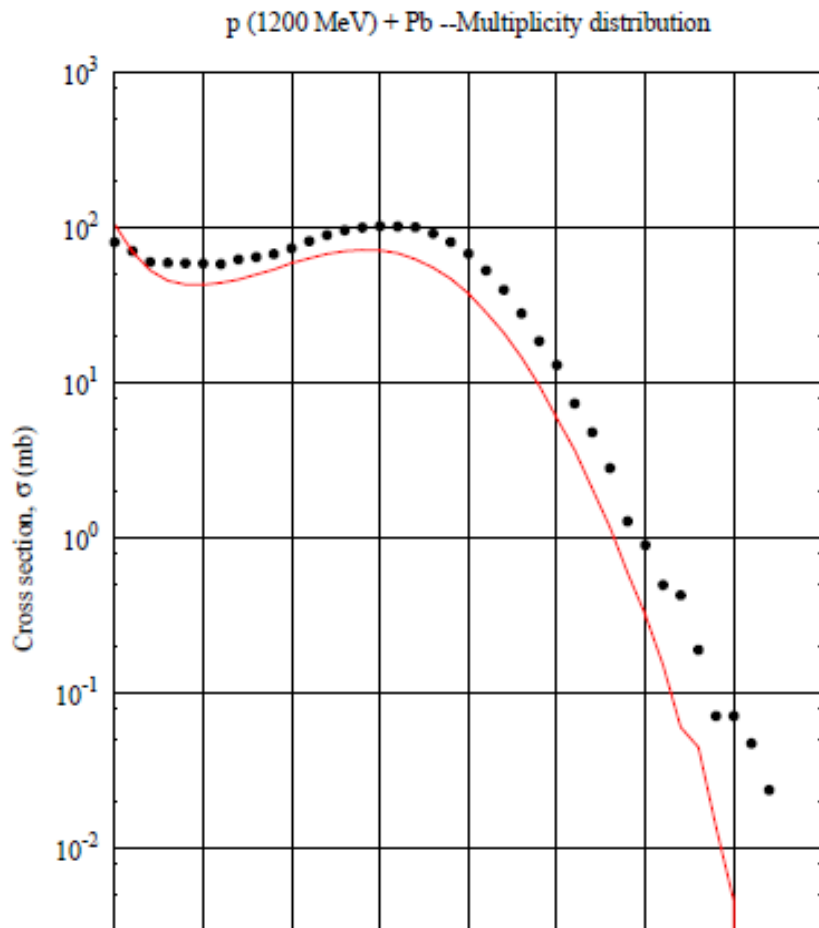


neutron number

# p(1200 MeV) + Pb – Neutron multiplicity distribution

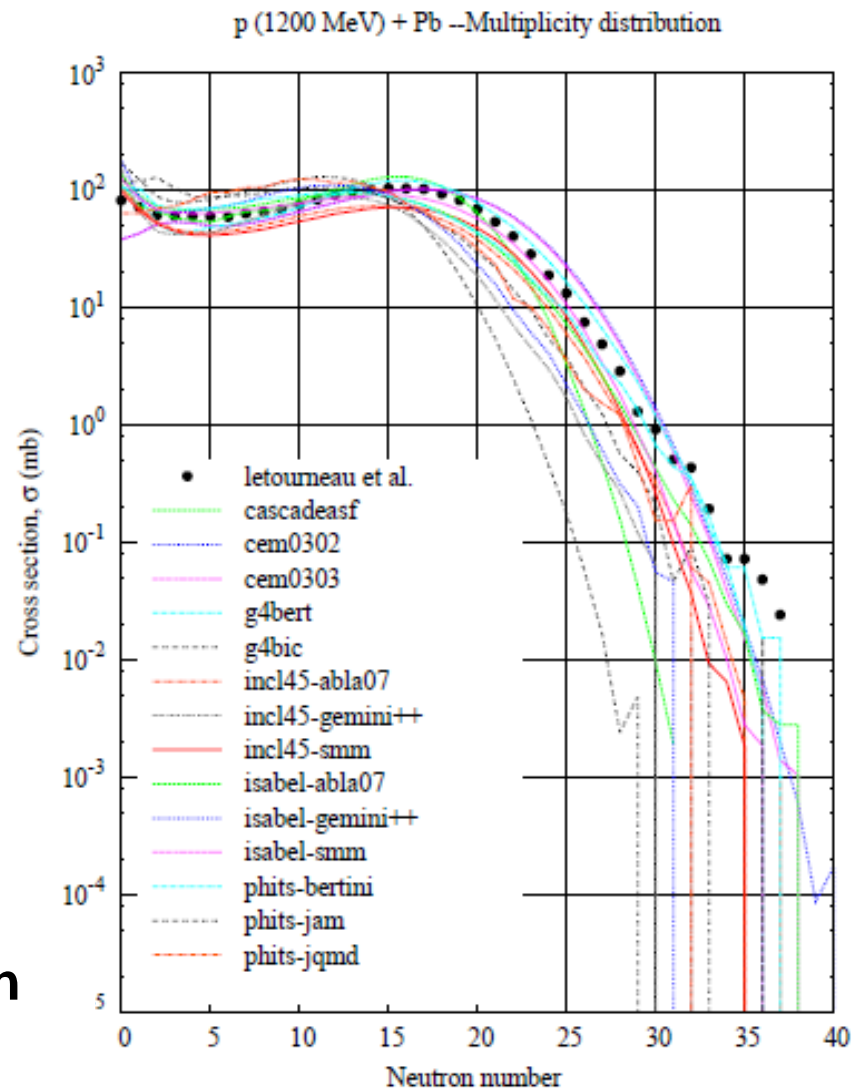
## INCL45-ABLA07

## ALL MODELS



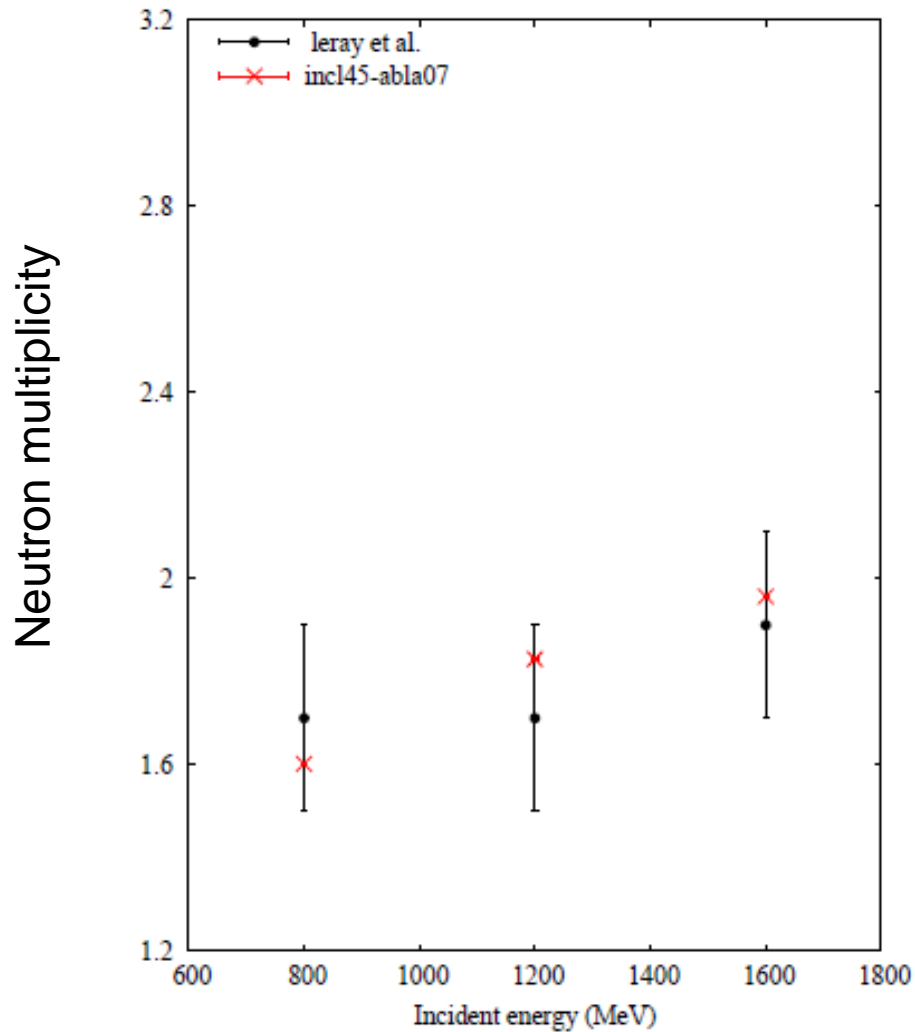
**Practically all models: too low high neutron multiplicities**

**Maybe  $E^*$  too low?**

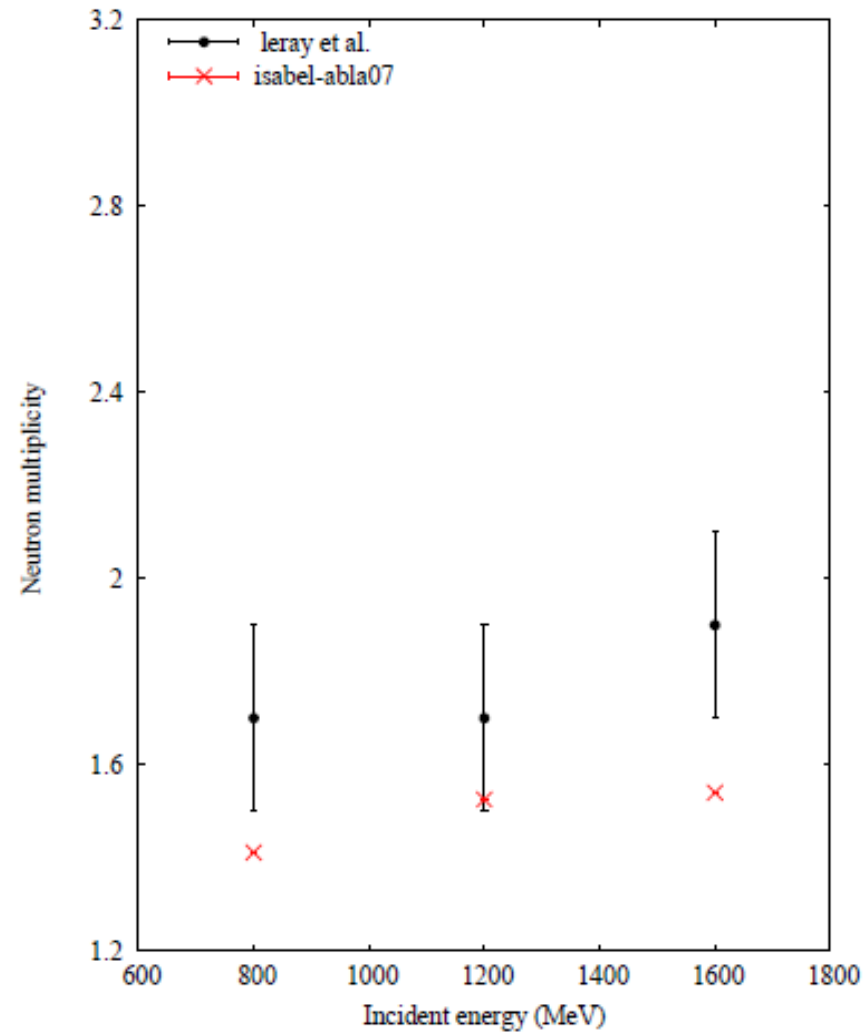


# p + Fe – Average neutron (2-20 MeV) multiplicity

## INCL45-ABLA07



## ISABEL-ABLA07

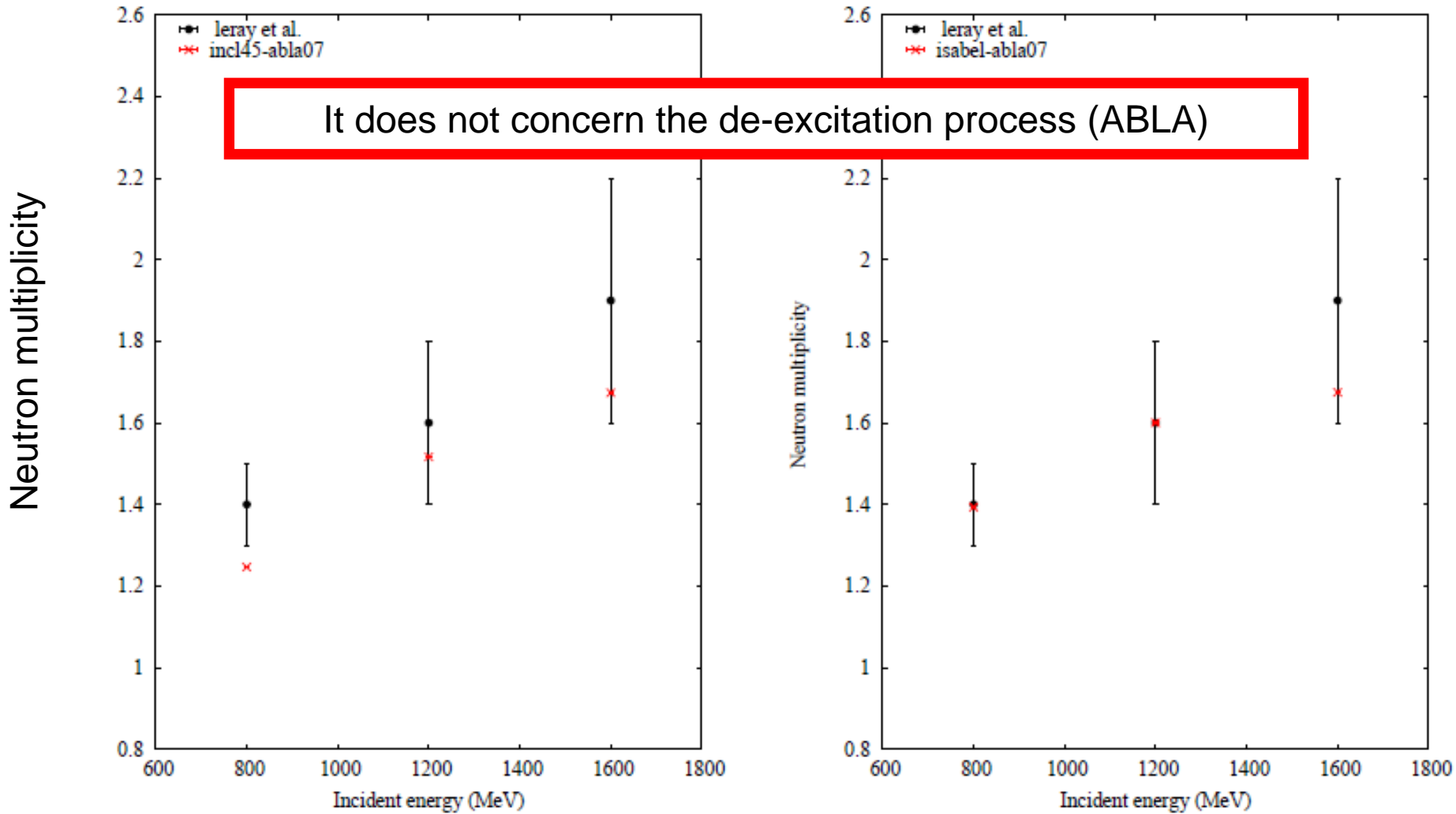


Incident energy (MeV)

# p + Fe – Average neutron (20+ MeV) multiplicity

## INCL45-ABLA07

## ISABEL-ABLA07

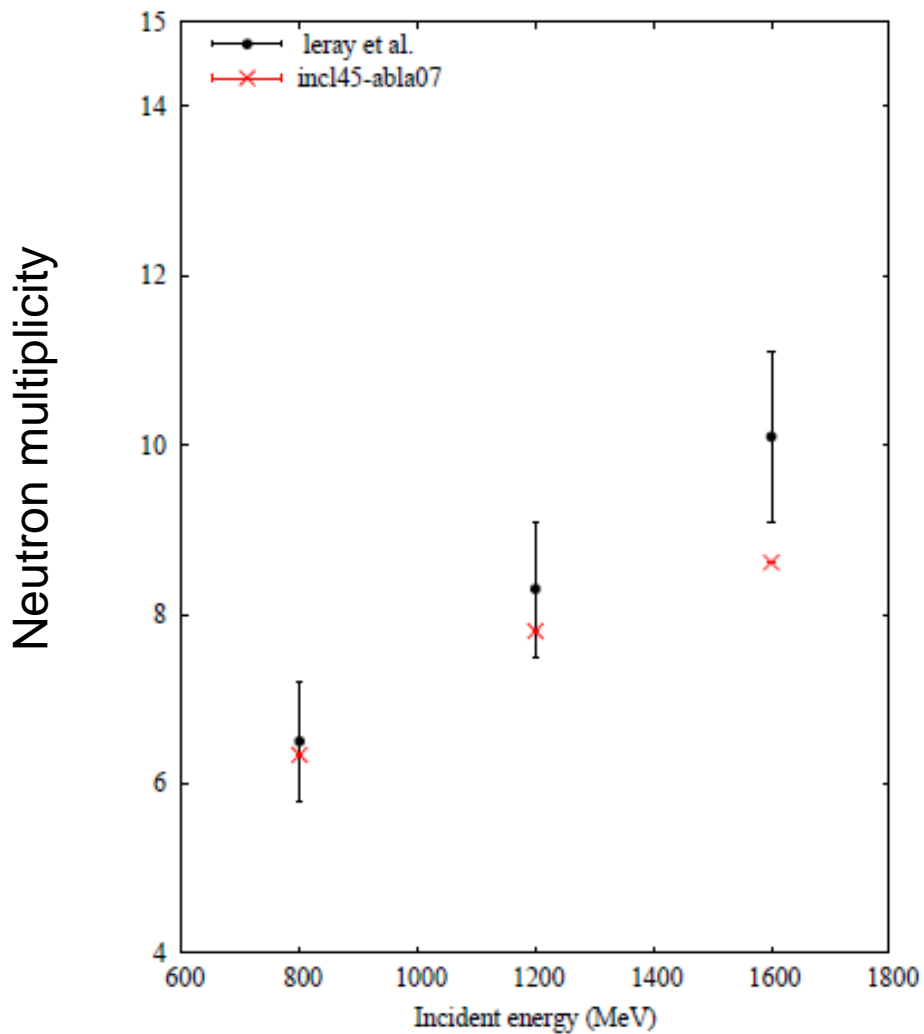


It does not concern the de-excitation process (ABLA)

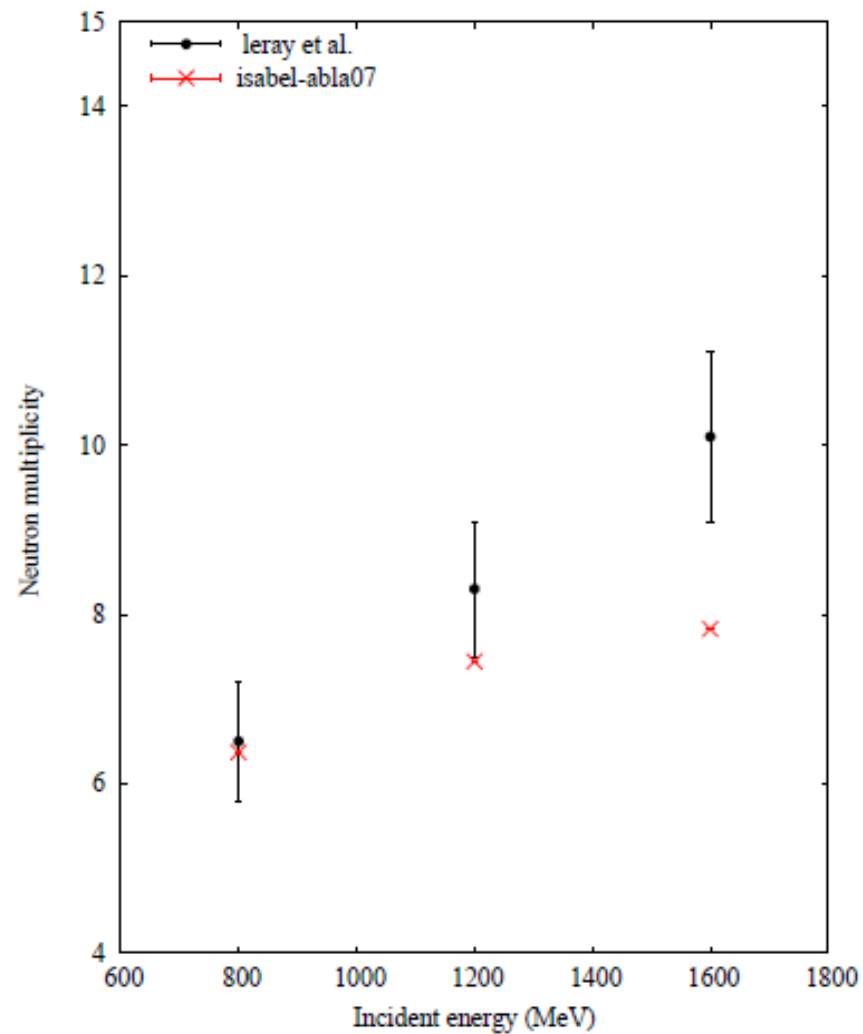
Incident energy (MeV)

# p + Pb – Average neutron (2-20 MeV) multiplicity

## INCL45-ABLA07



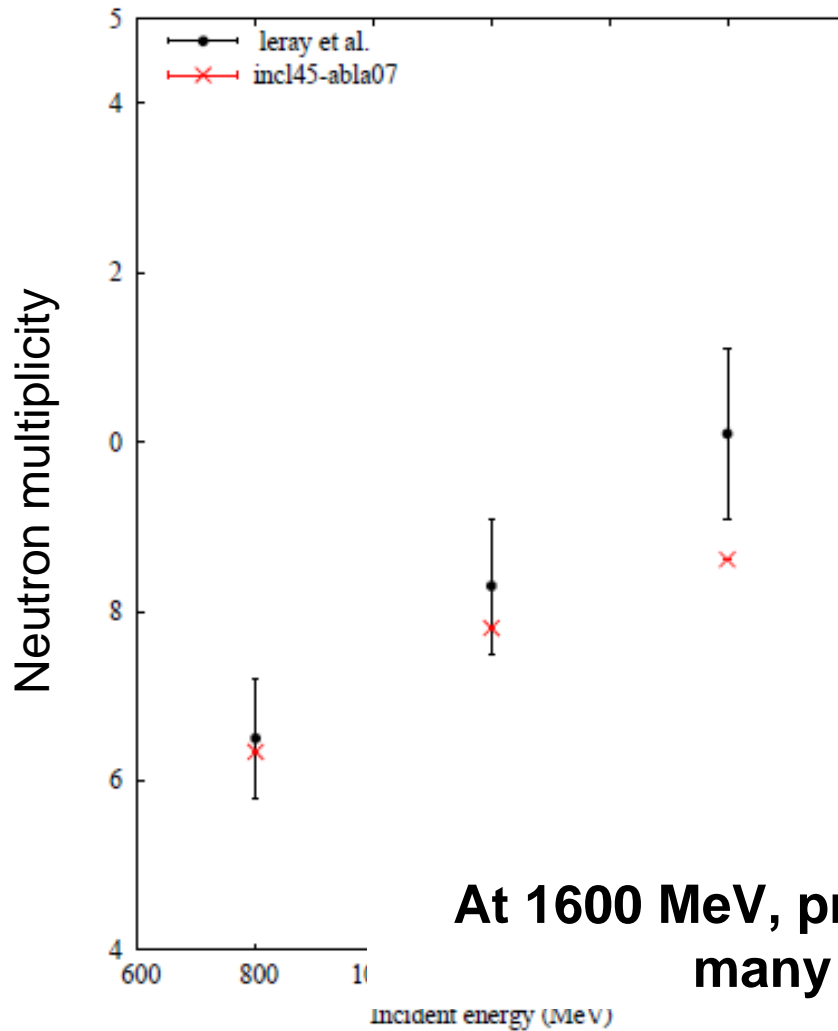
## ISABEL-ABLA07



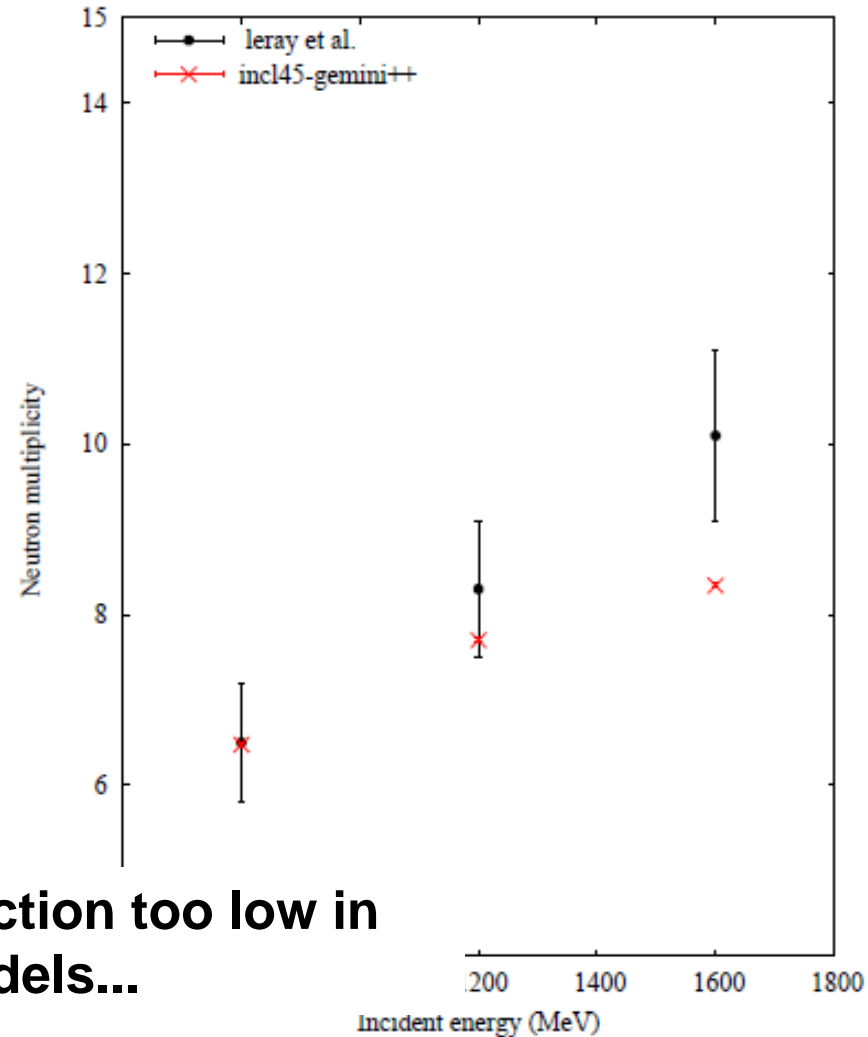
Incident energy (MeV)

# p + Pb – Average neutron (2-20 MeV) multiplicity

## INCL45-ABLA07



## INCL45-GEMINI++



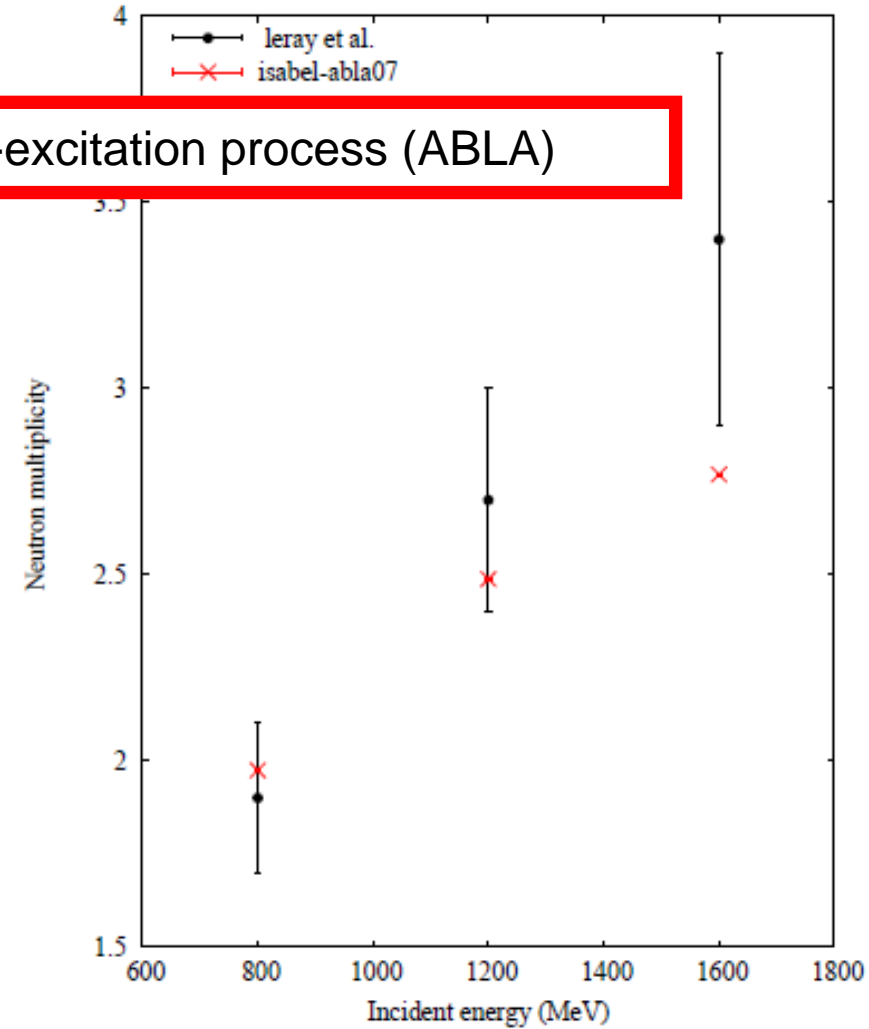
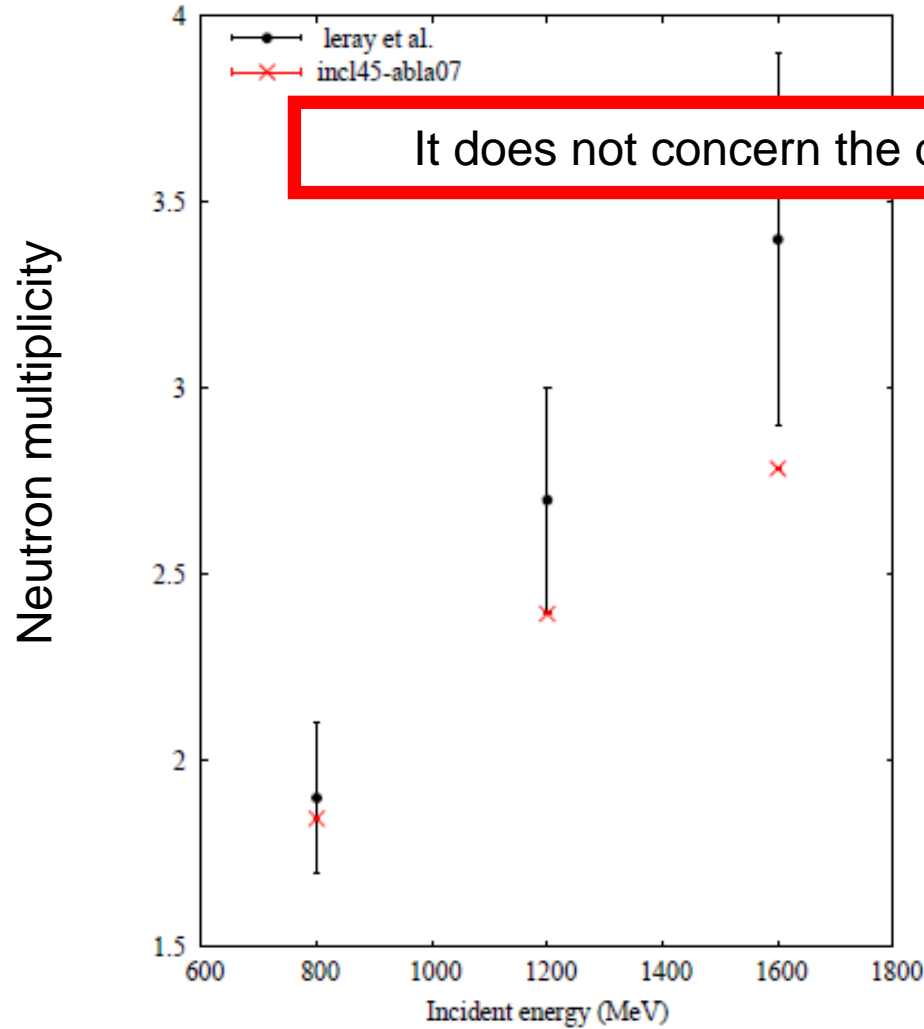
**At 1600 MeV, prediction too low in many models...**

Incident energy (MeV)

# p + Pb – Average neutron (20+ MeV) multiplicity

## INCL45-ABLA07

## ISABEL-ABLA07



Incident energy (MeV)

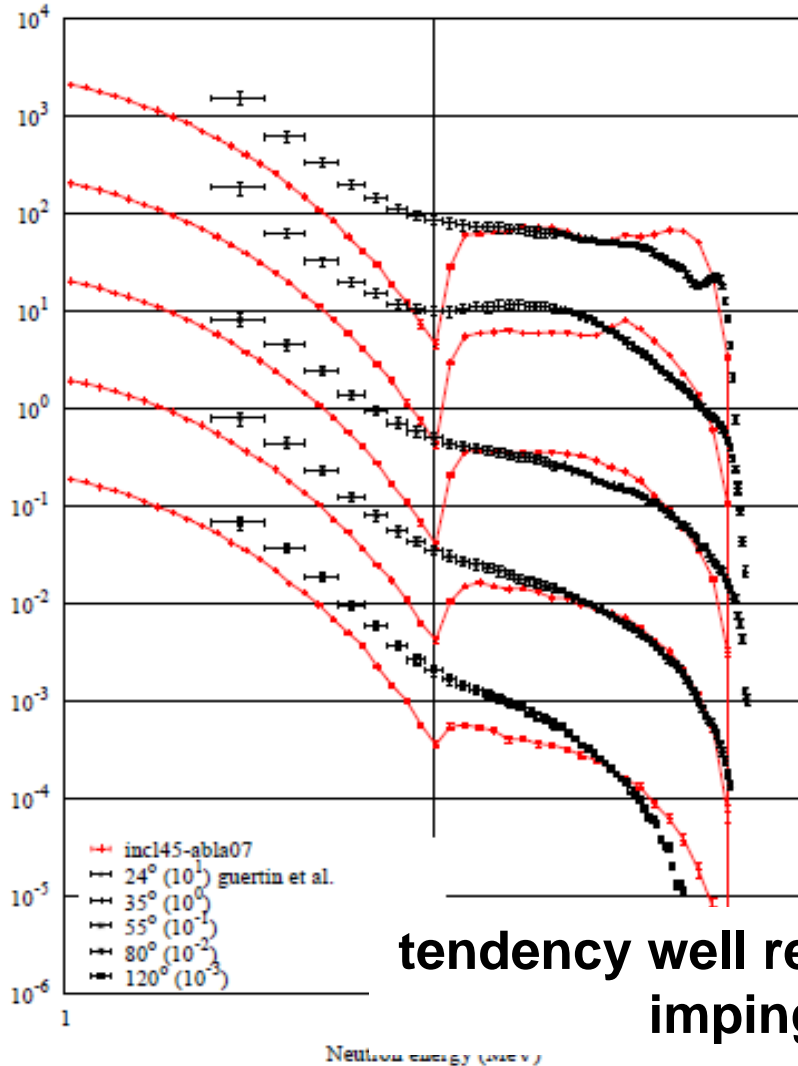
# Neutron spectra



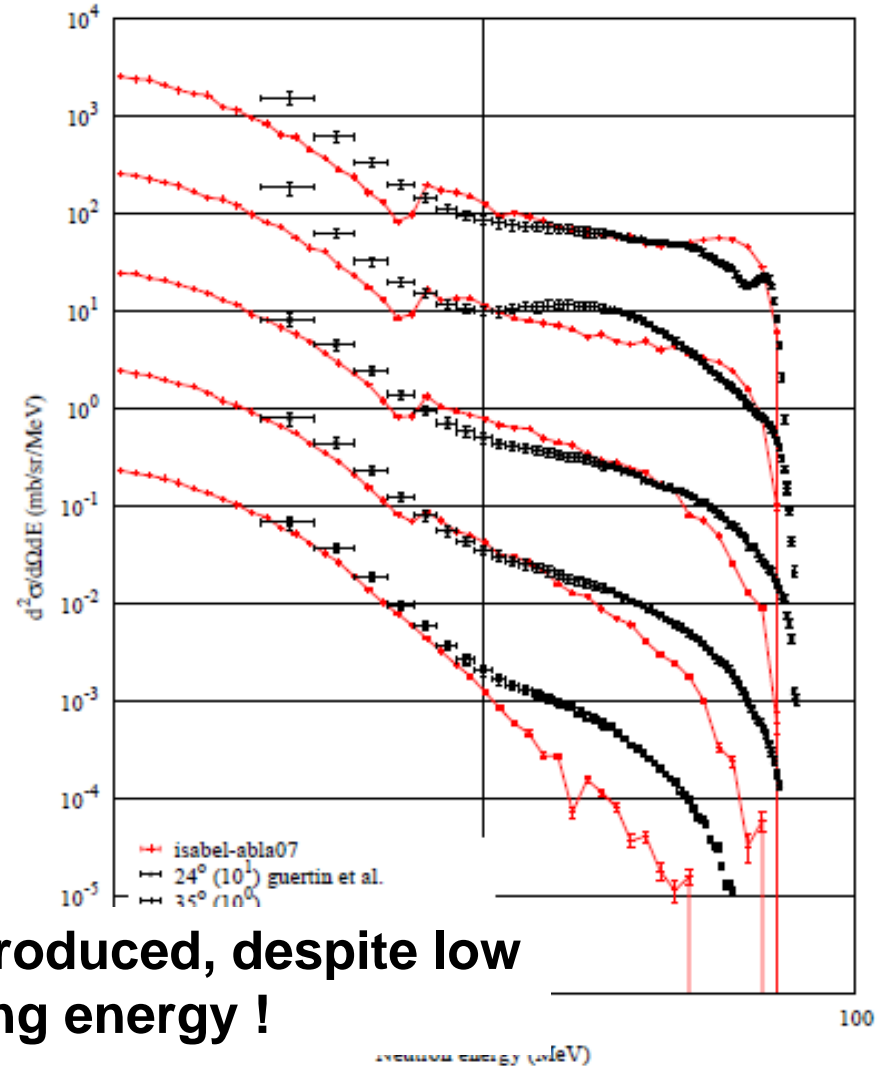
# p(63 MeV) + <sup>208</sup>Pb – Neutron spectrum

## INCL45-ABLA07

double differential cross section (mb/sr/MeV)



## ISABEL-ABLA07



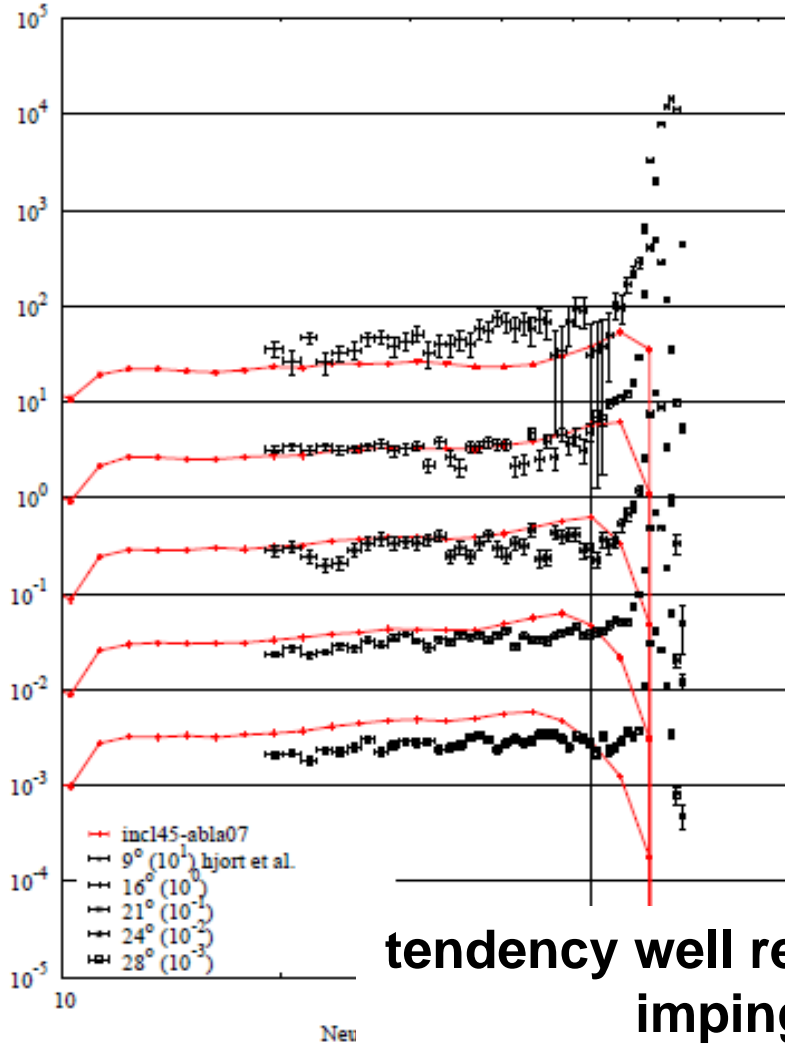
tendency well reproduced, despite low impinging energy !

neutron energy (MeV)

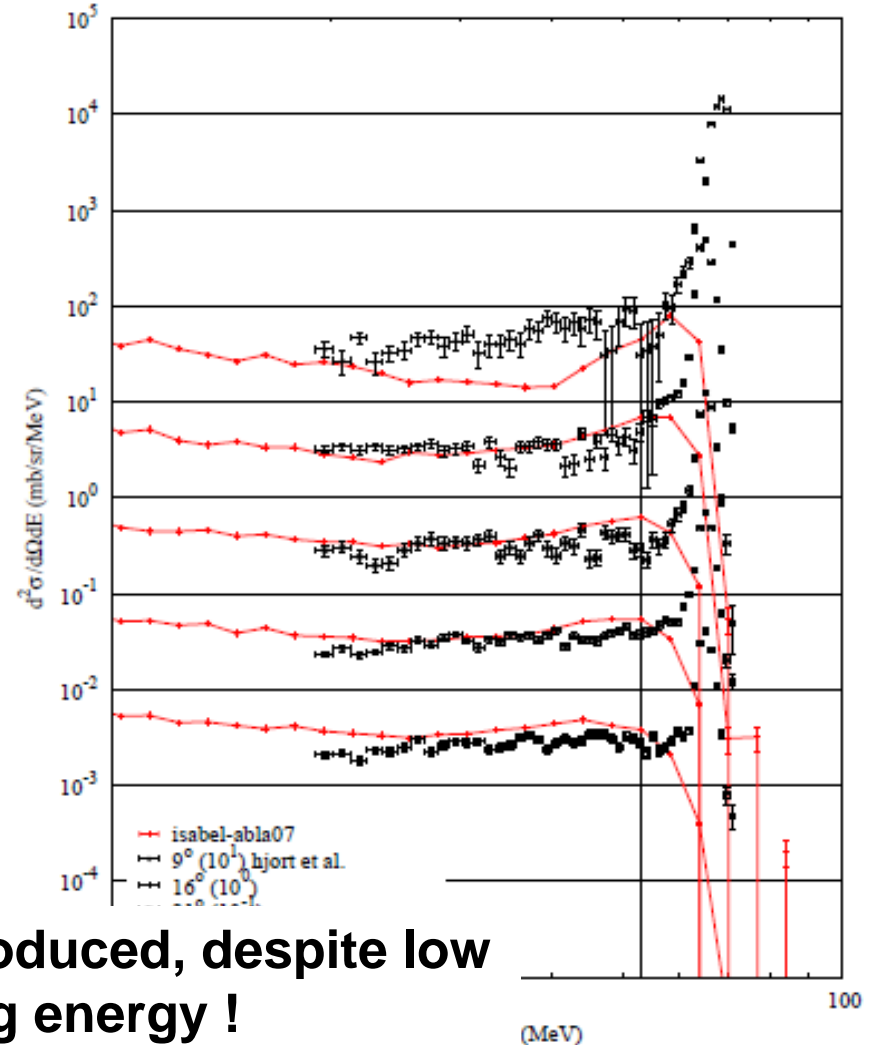
# n(65 MeV) + Fe – Neutron spectrum

## INCL45-ABLA07

double differential cross section (mb/sr/MeV)



## ISABEL-ABLA07

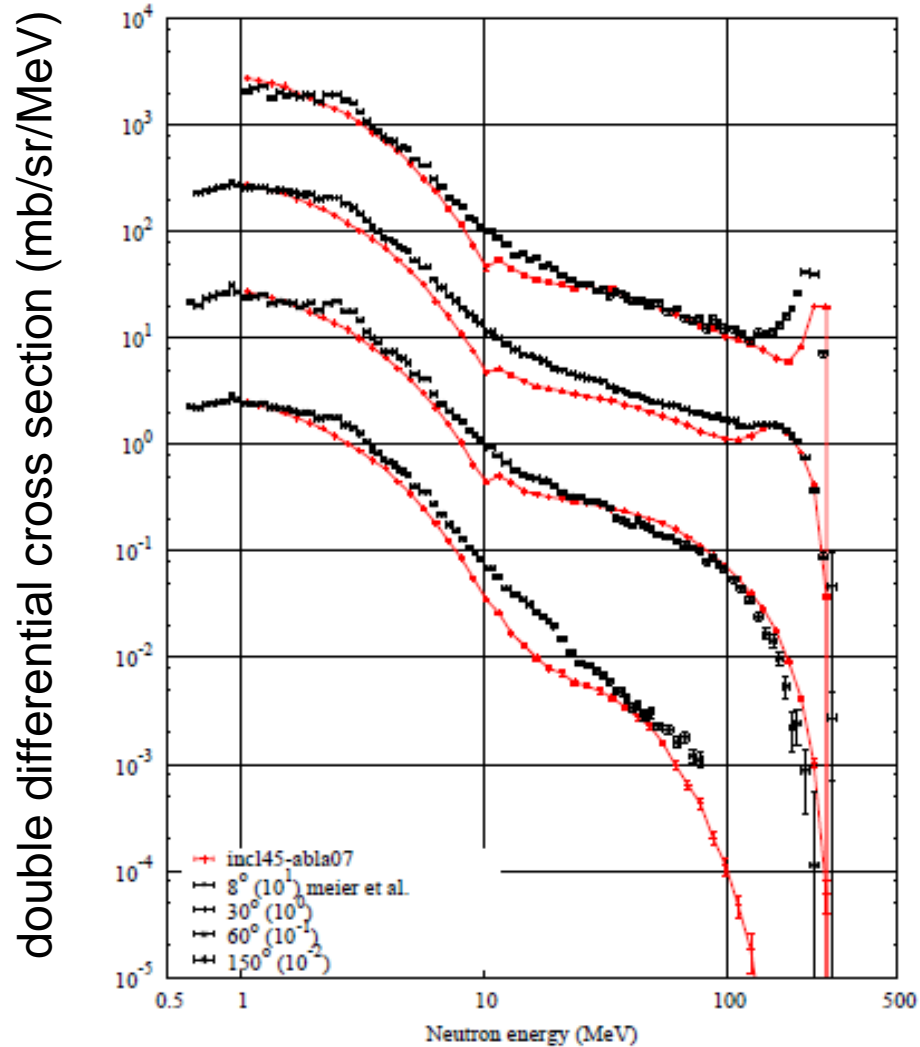


tendency well reproduced, despite low impinging energy !

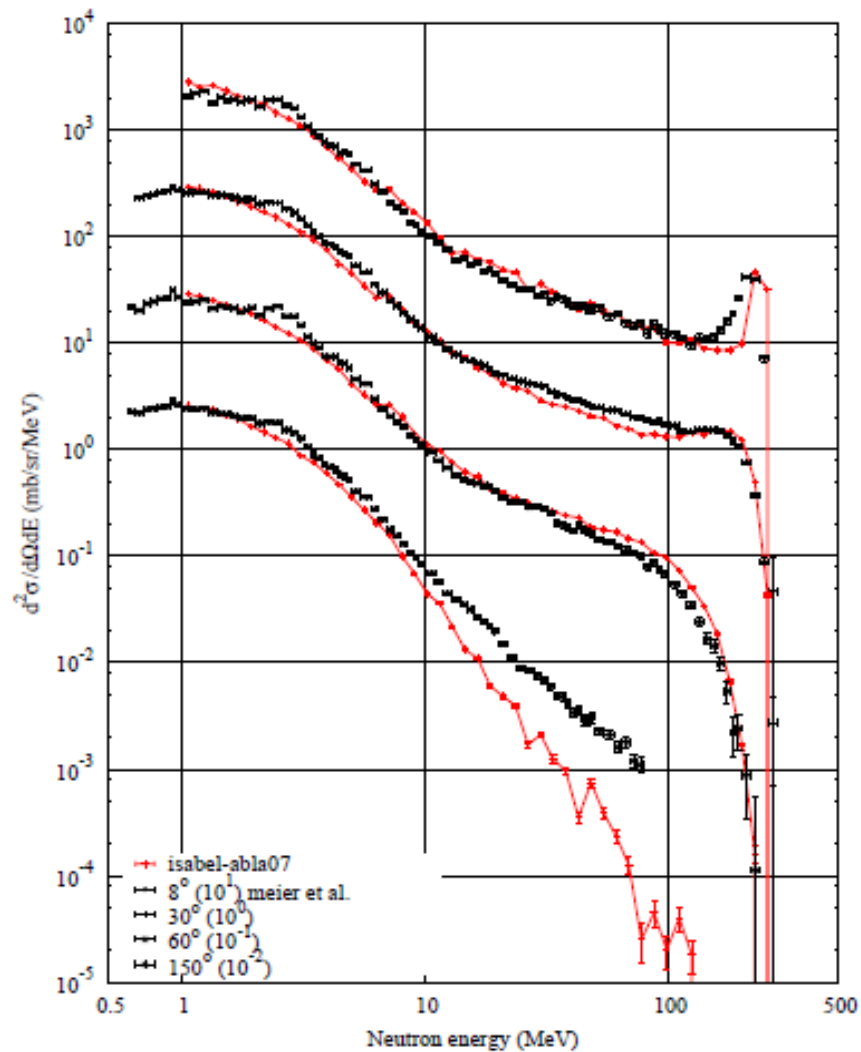
neutron energy (MeV)

# p(256 MeV) + Pb – Neutron spectrum

## INCL45-ABLA07



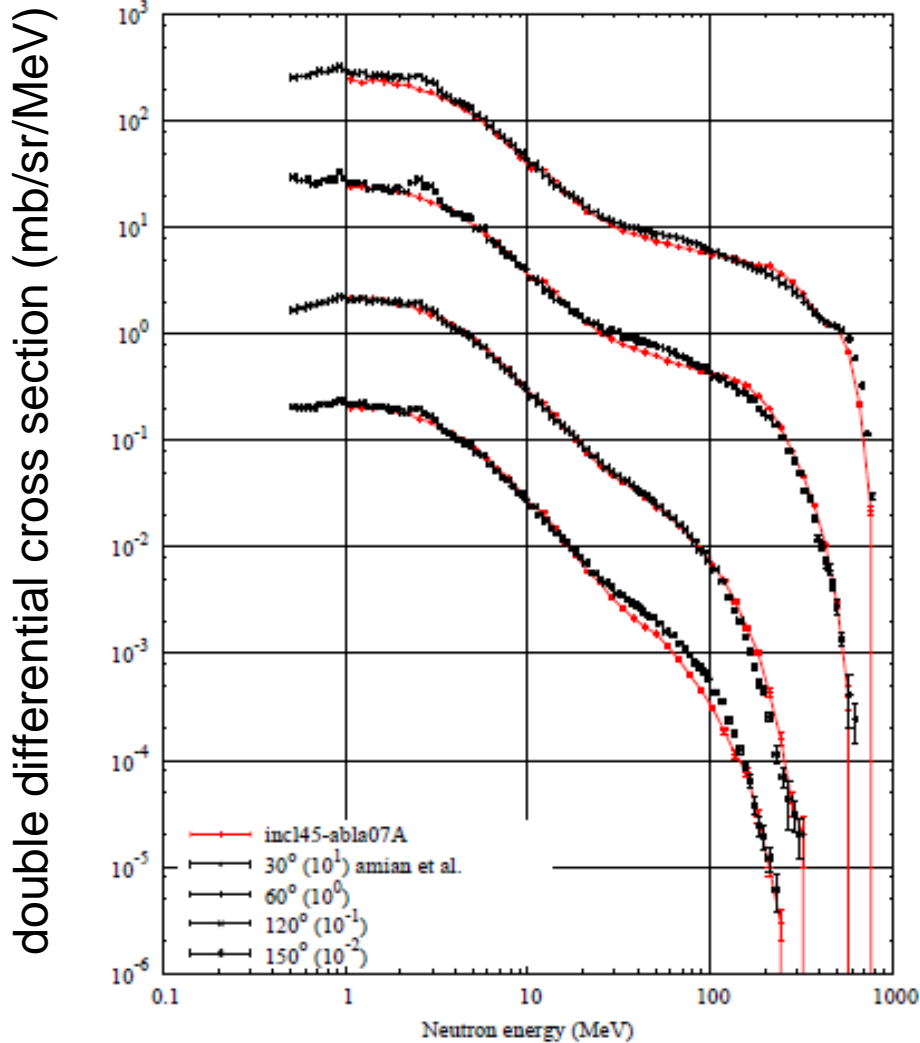
## ISABEL-ABLA07



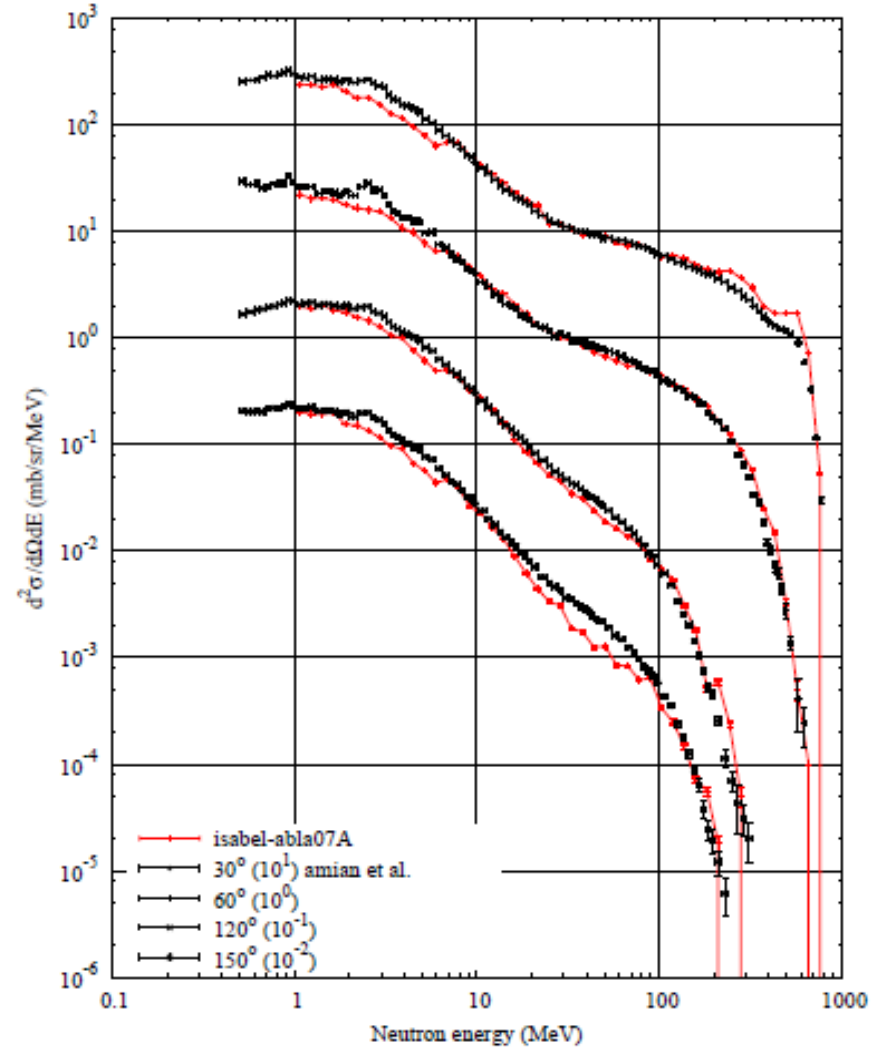
neutron energy (MeV)

# p(800 MeV) + Fe – Neutron spectrum

## INCL45-ABLA07



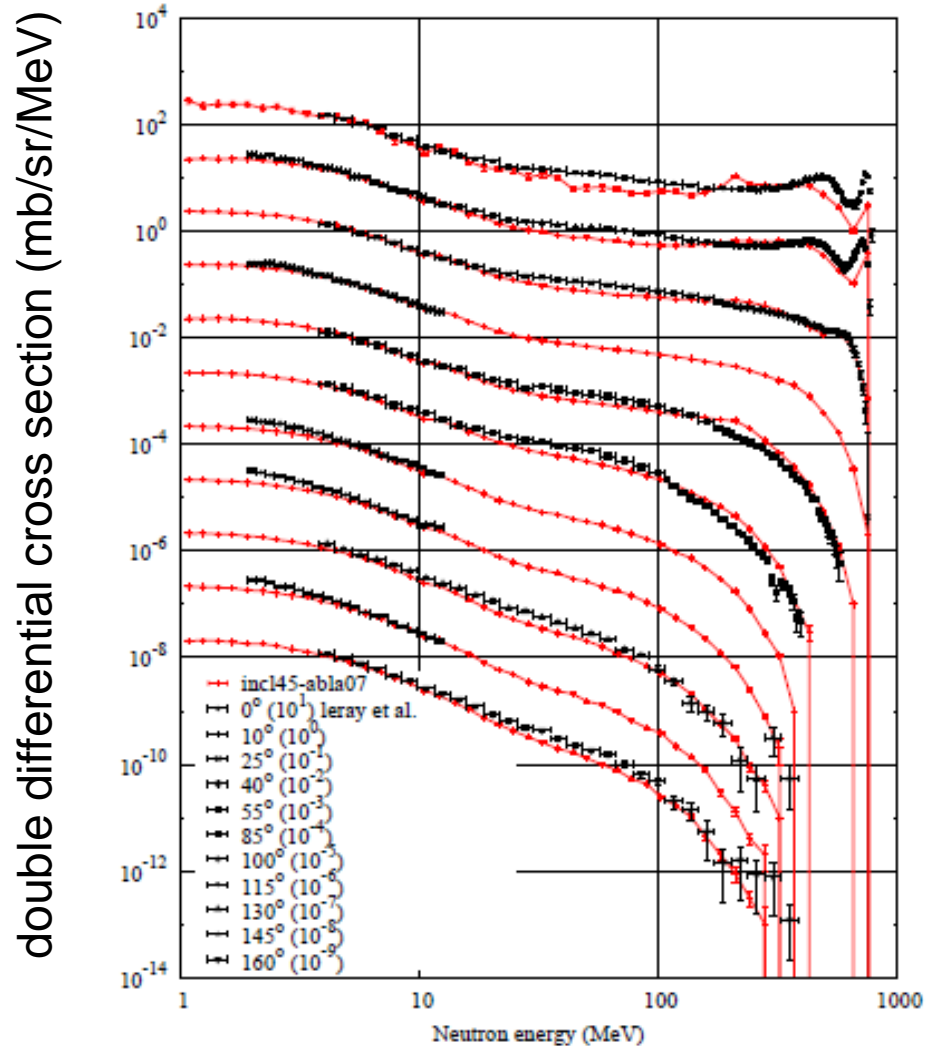
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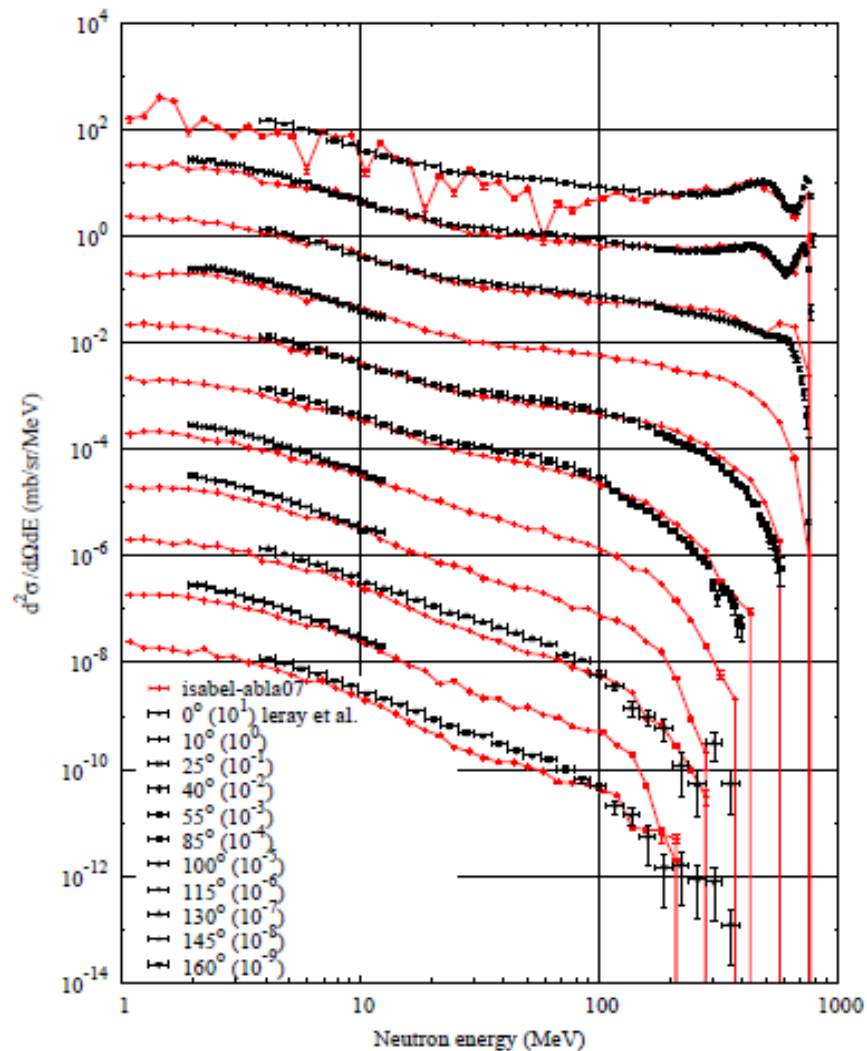
neutron energy (MeV)

# p(800 MeV) + Fe – Neutron spectrum

## INCL45-ABLA07



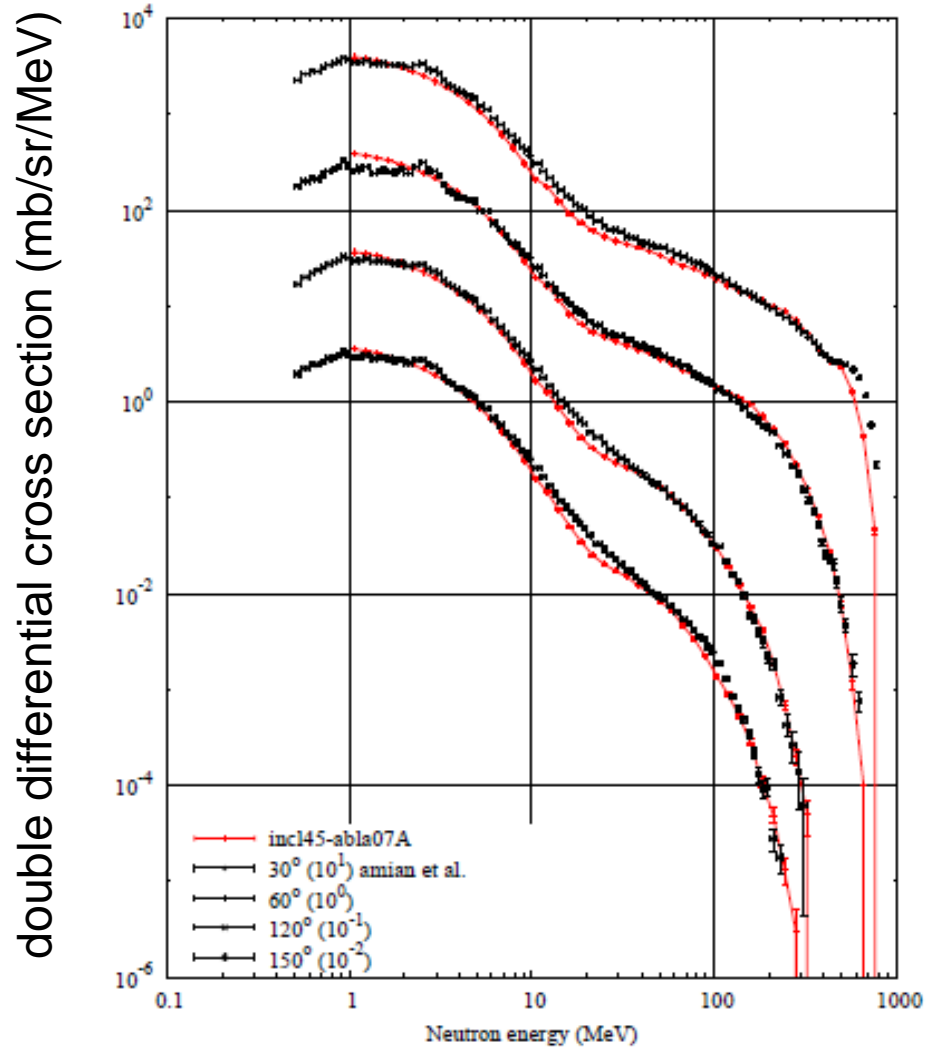
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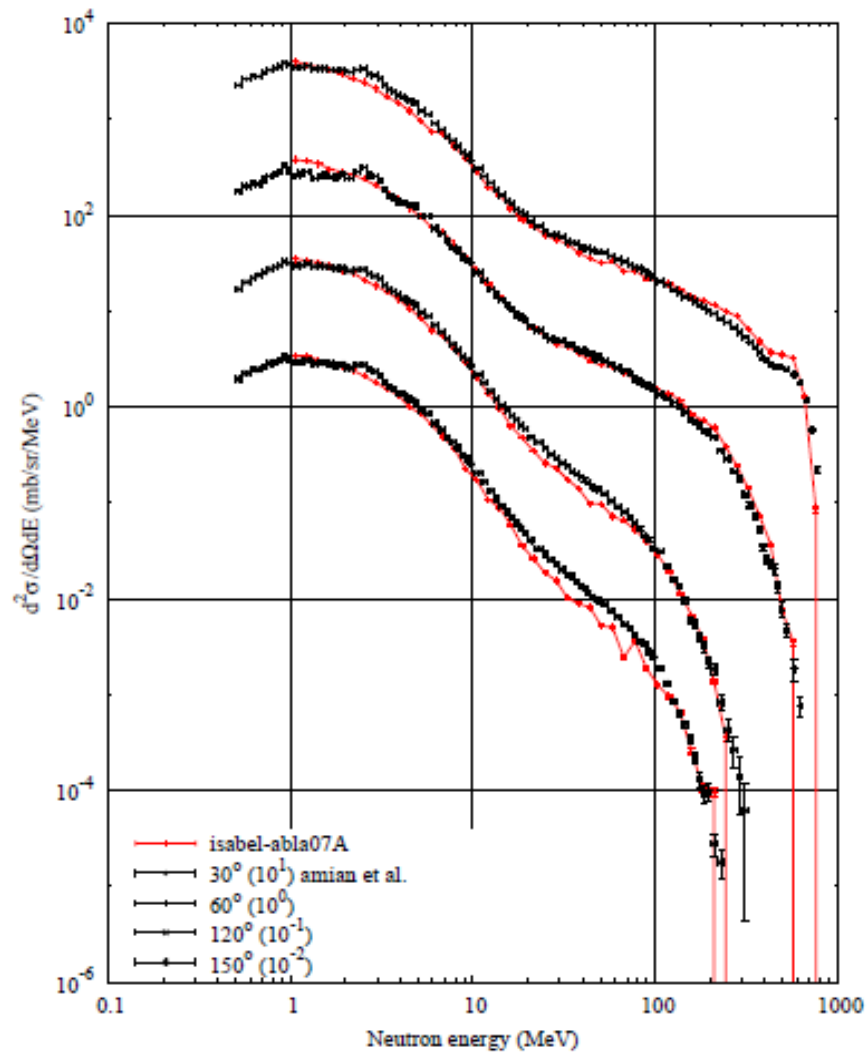
neutron energy (MeV)

# p(800 MeV) + Pb – Neutron spectrum

## INCL45-ABLA07



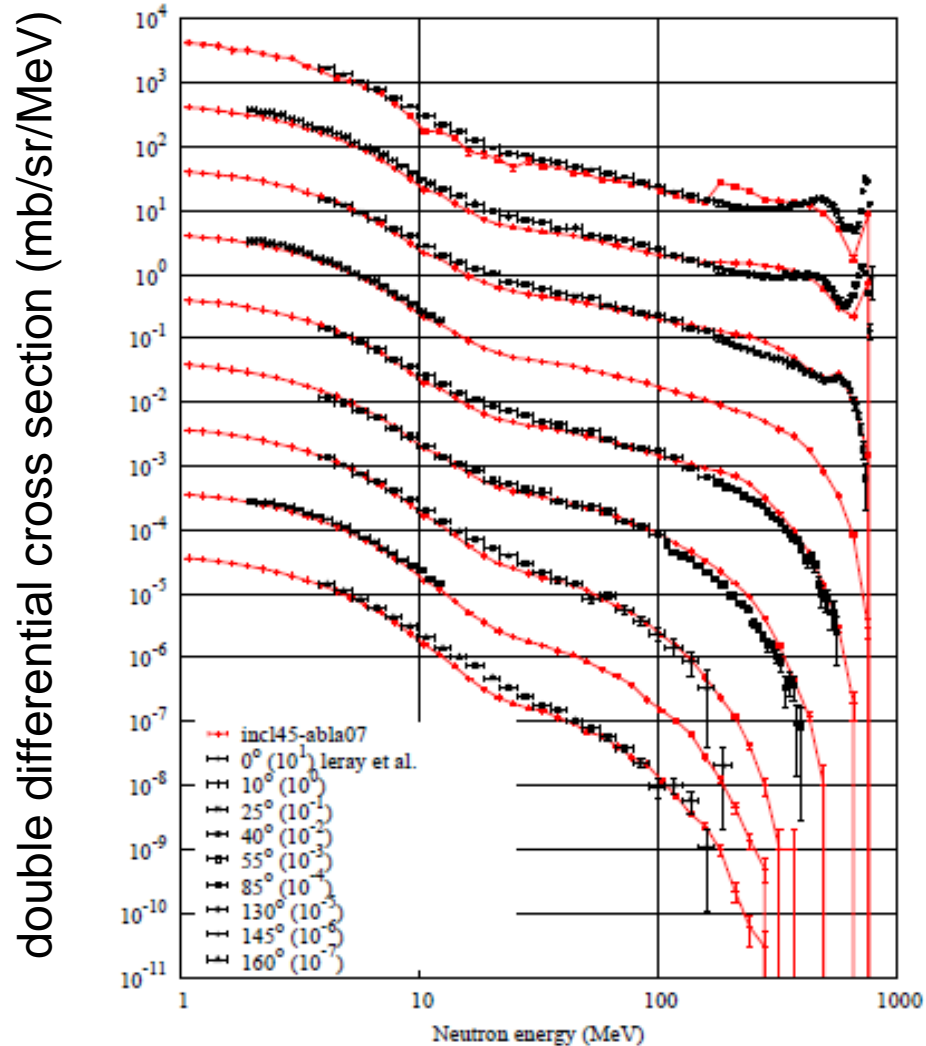
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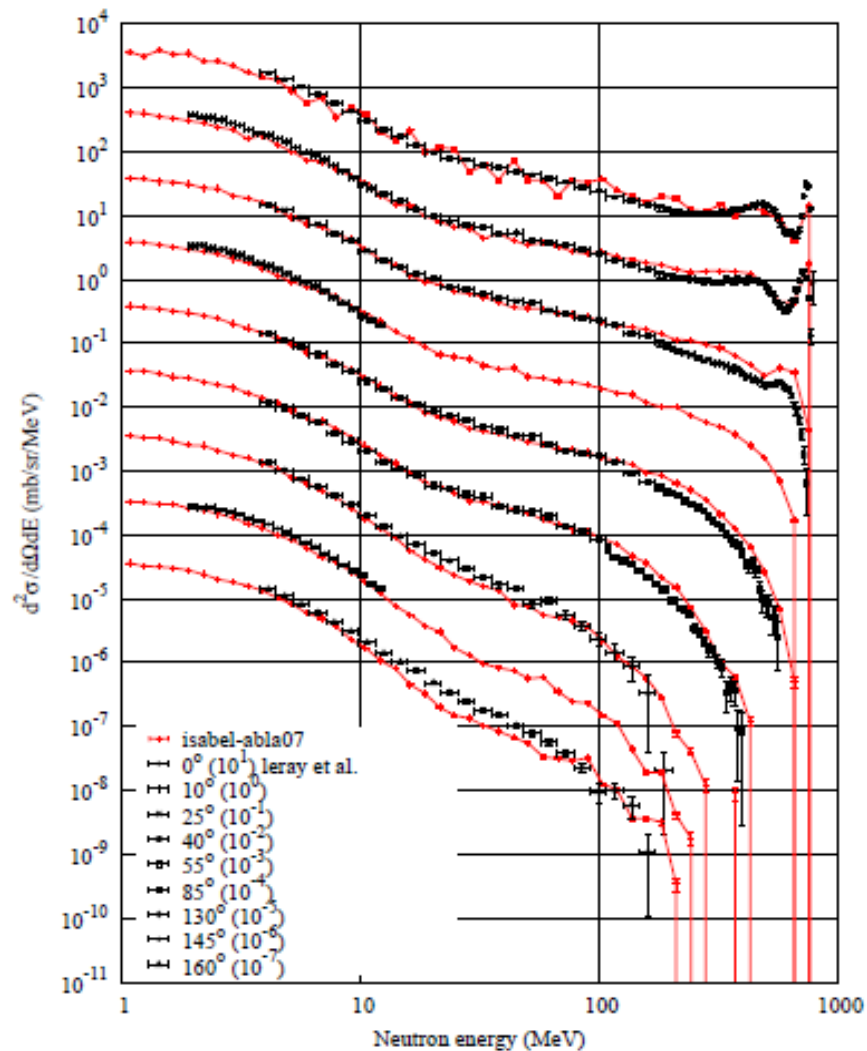
neutron energy (MeV)

# p(800 MeV) + Pb – Neutron spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

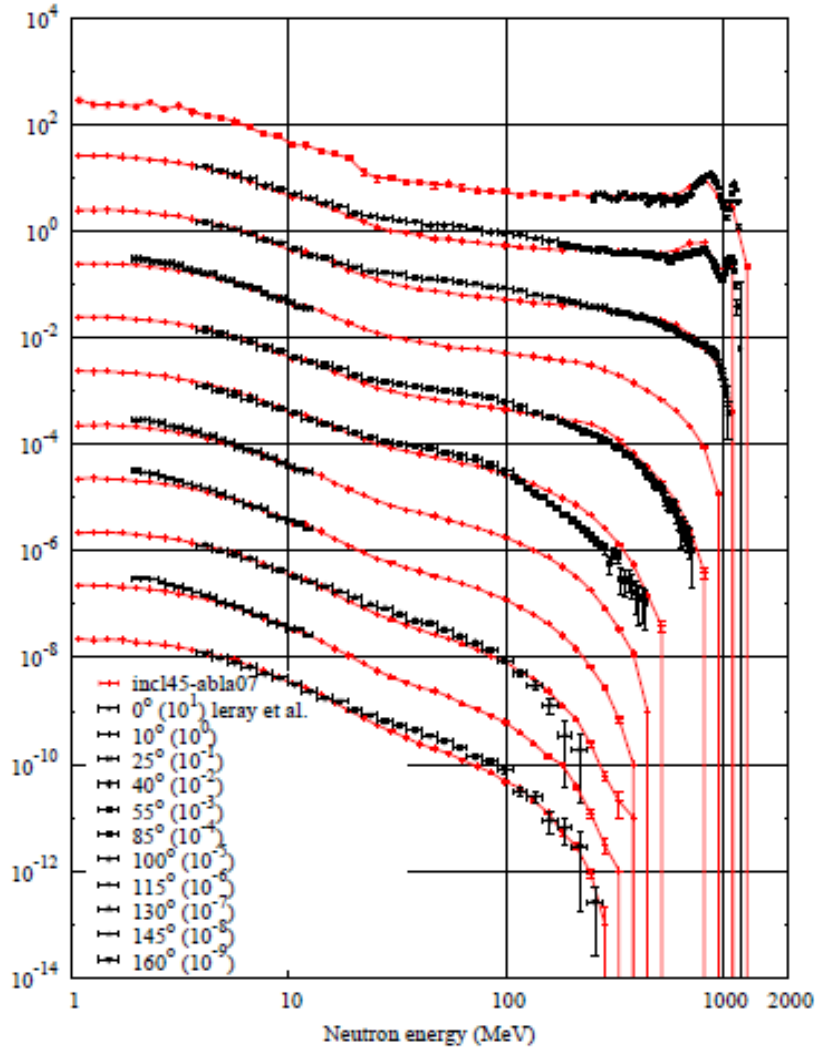


neutron energy (MeV)

# p(1200 MeV) + Fe – Neutron spectrum

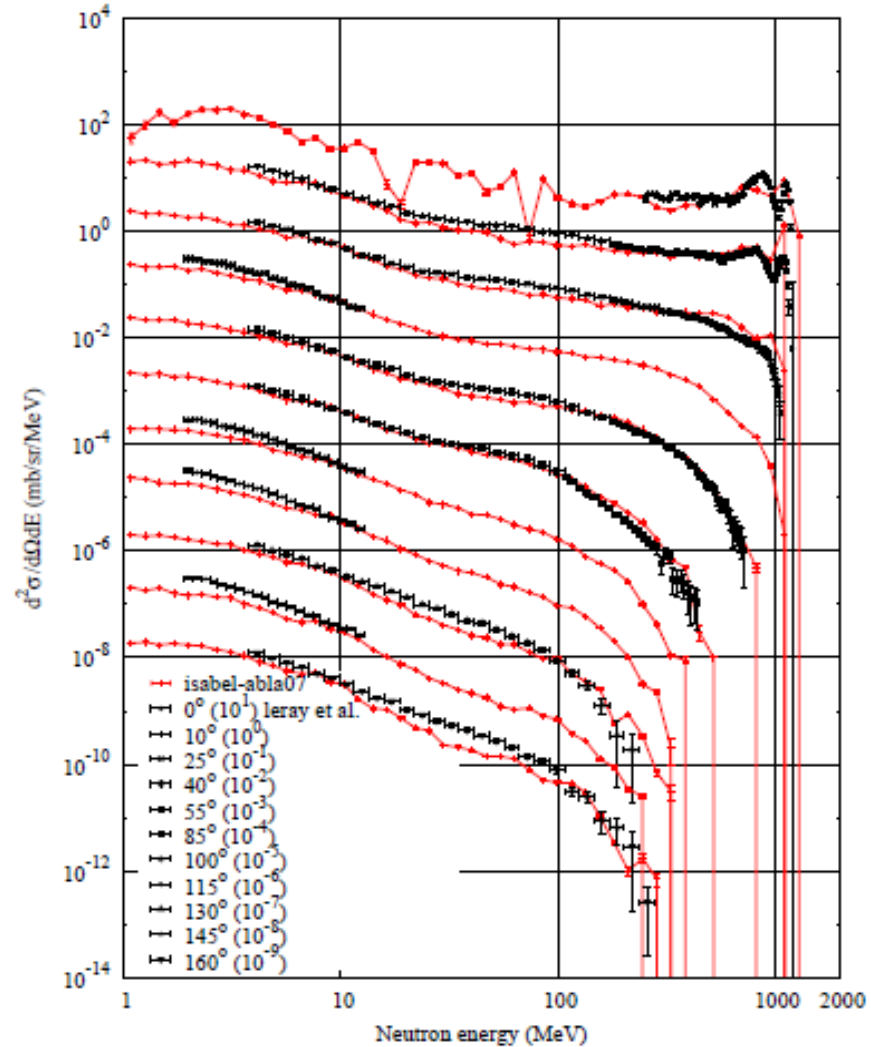
## INCL45-ABLA07

double differential cross section (mb/sr/MeV)



## ISABEL-ABLA07

$d^2\sigma/d\Omega dE$  (mb/sr/MeV)



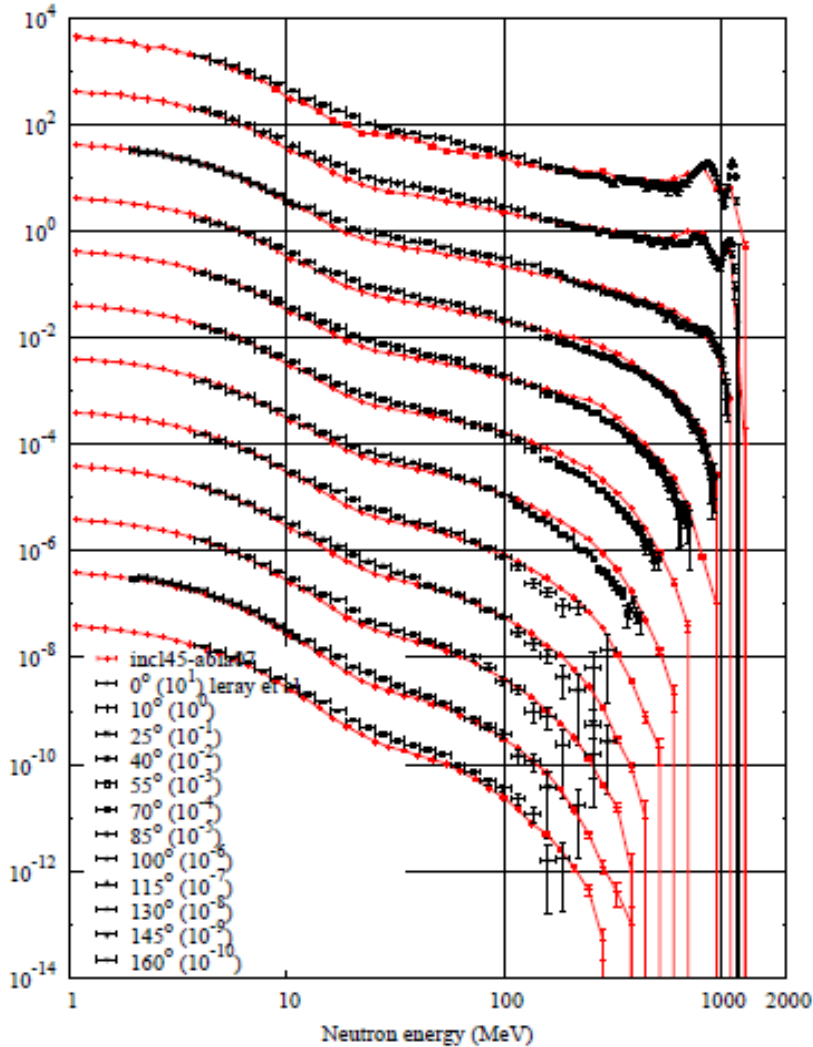
neutron energy (MeV)



# p(1200 MeV) + Pb – Neutron spectrum

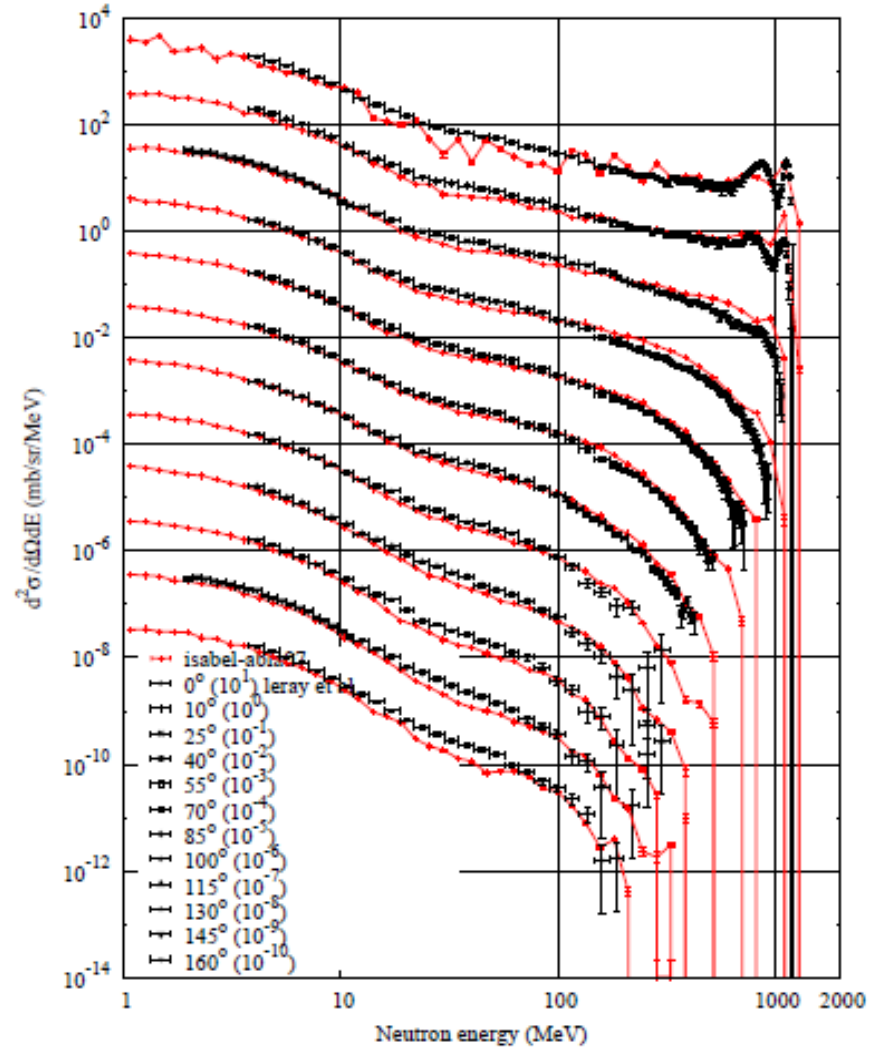
## INCL45-ABLA07

double differential cross section (mb/sr/MeV)



## ISABEL-ABLA07

$d^2\sigma/d\Omega dE$  (mb/sr/MeV)

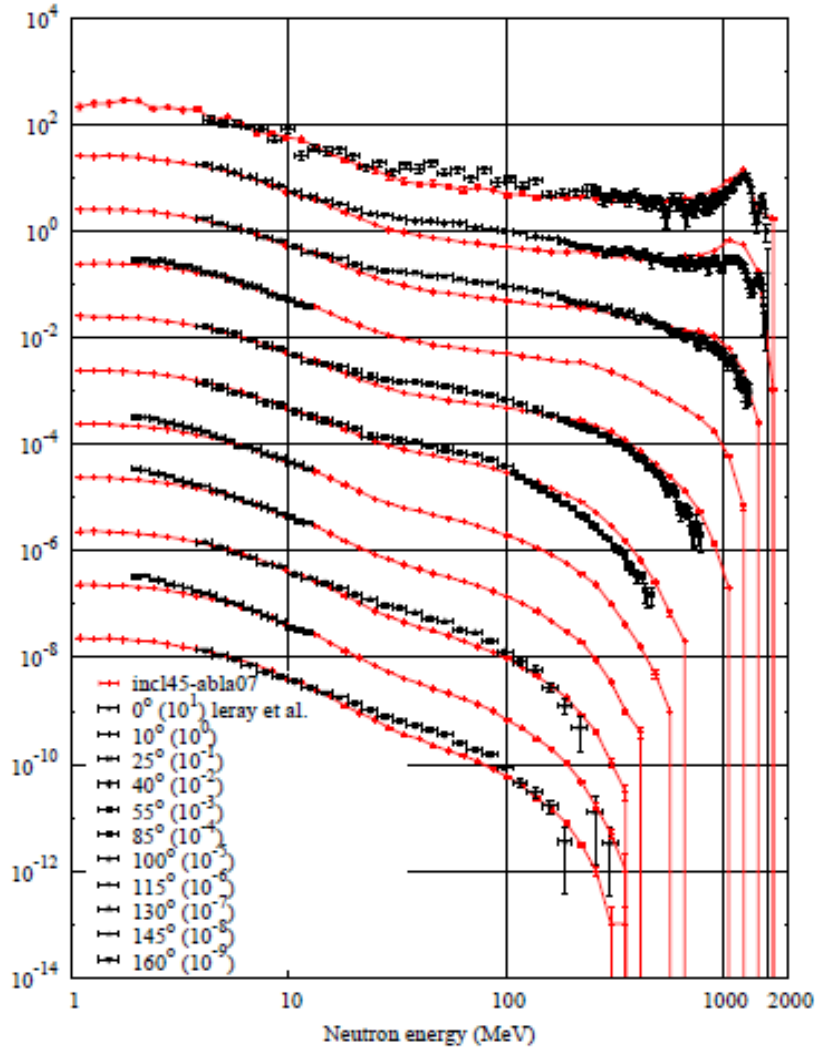


neutron energy (MeV)

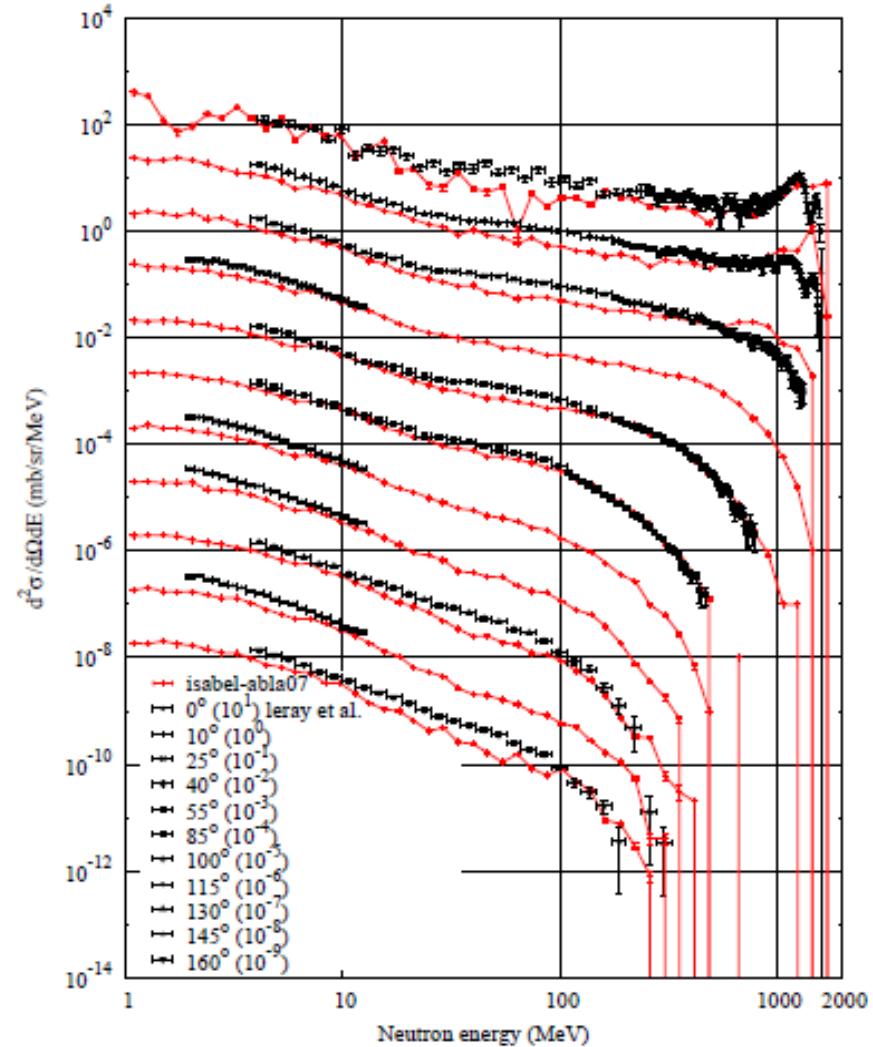
# p(1600 MeV) + Fe – Neutron spectrum

## INCL45-ABLA07

double differential cross section (mb/sr/MeV)



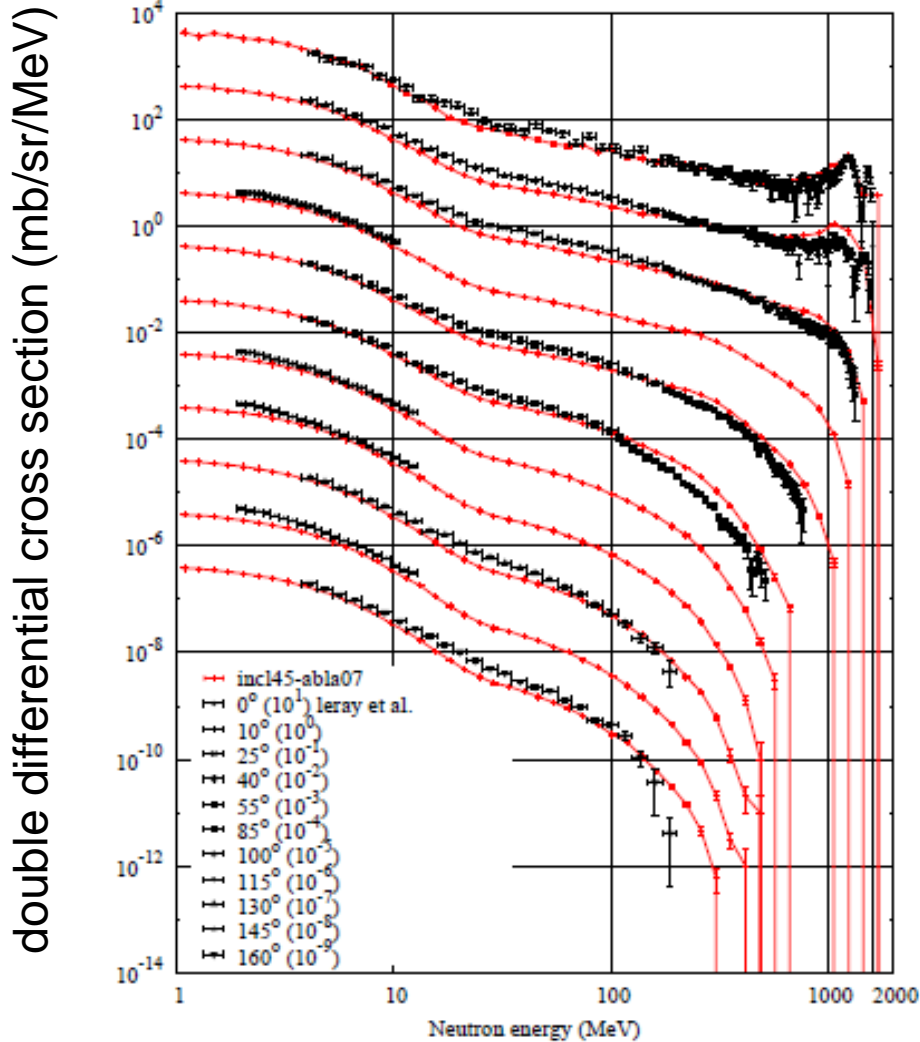
## ISABEL-ABLA07



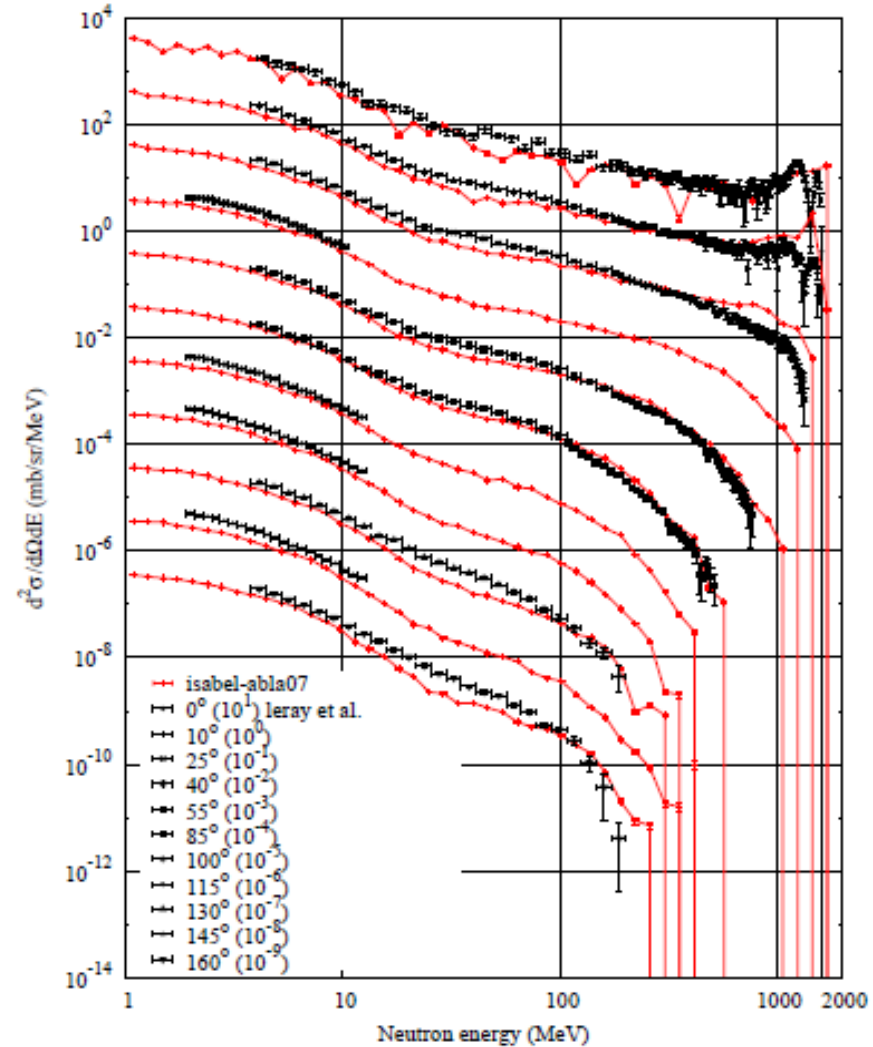
neutron energy (MeV)

# p(1600 MeV) + Pb – Neutron spectrum

## INCL45-ABLA07



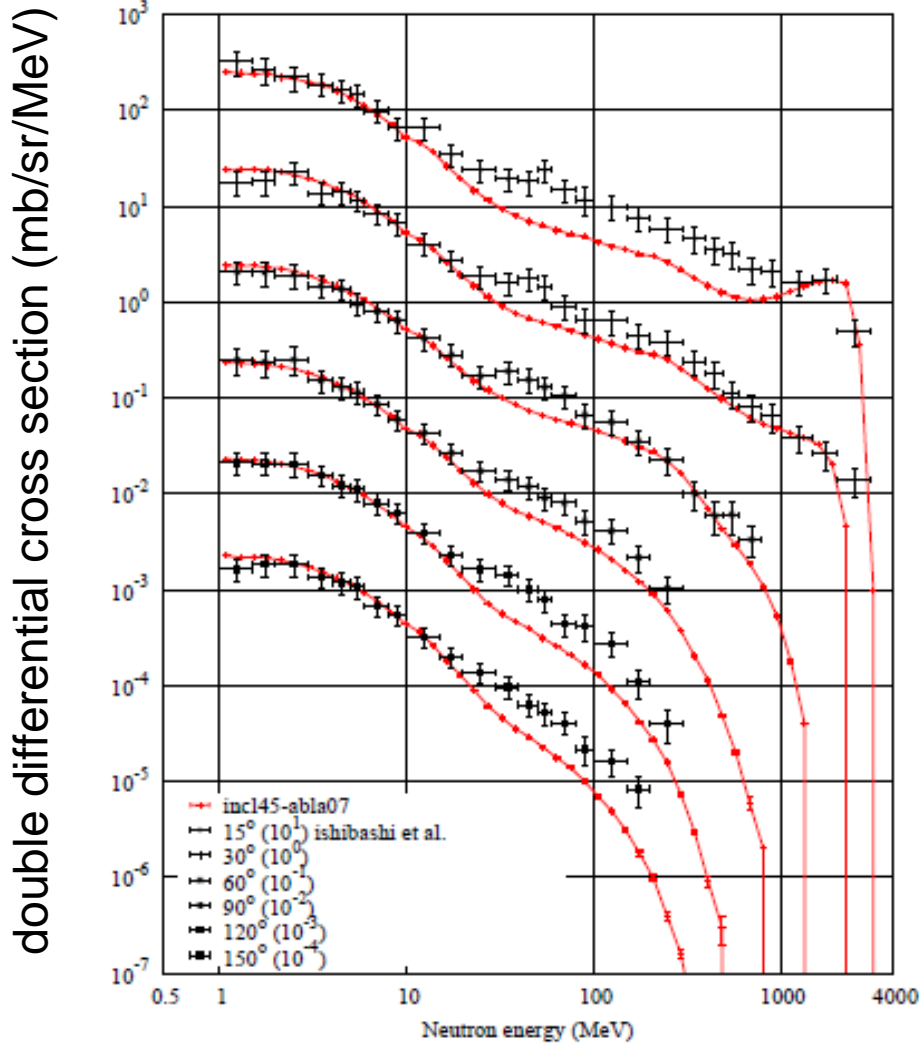
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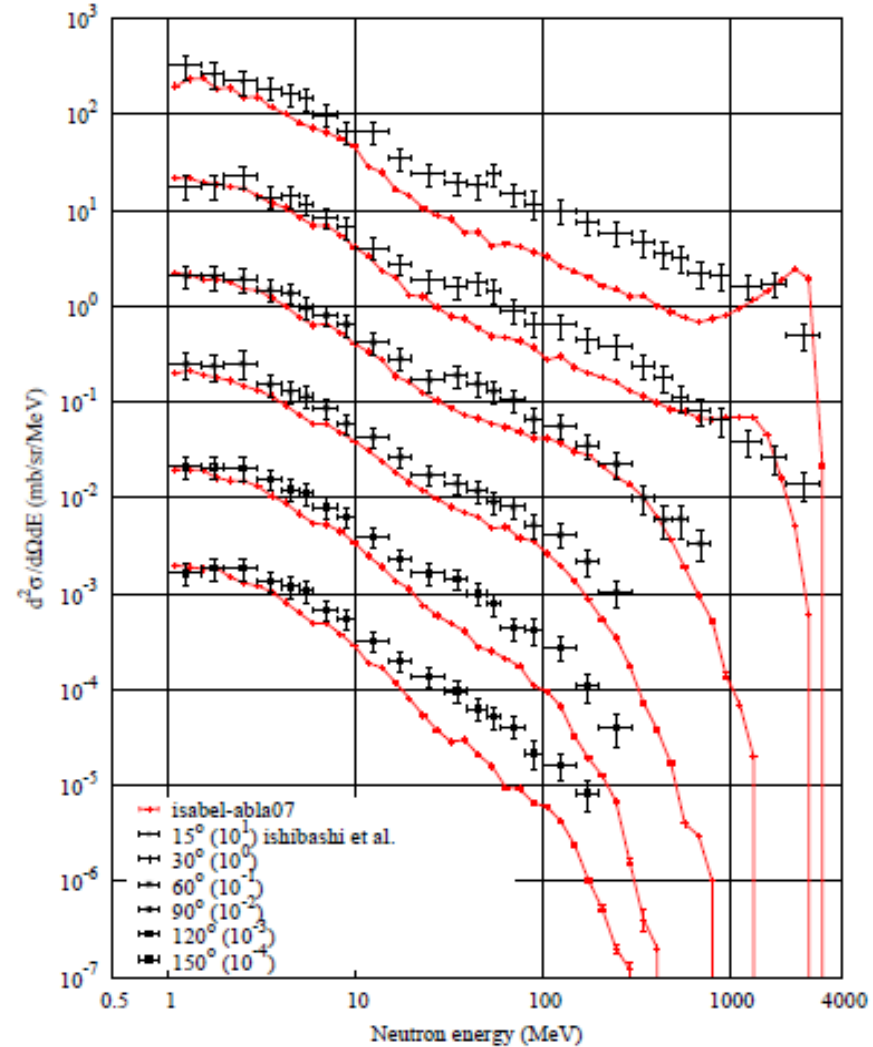
neutron energy (MeV)

# p(3000 MeV) + Fe – Neutron spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

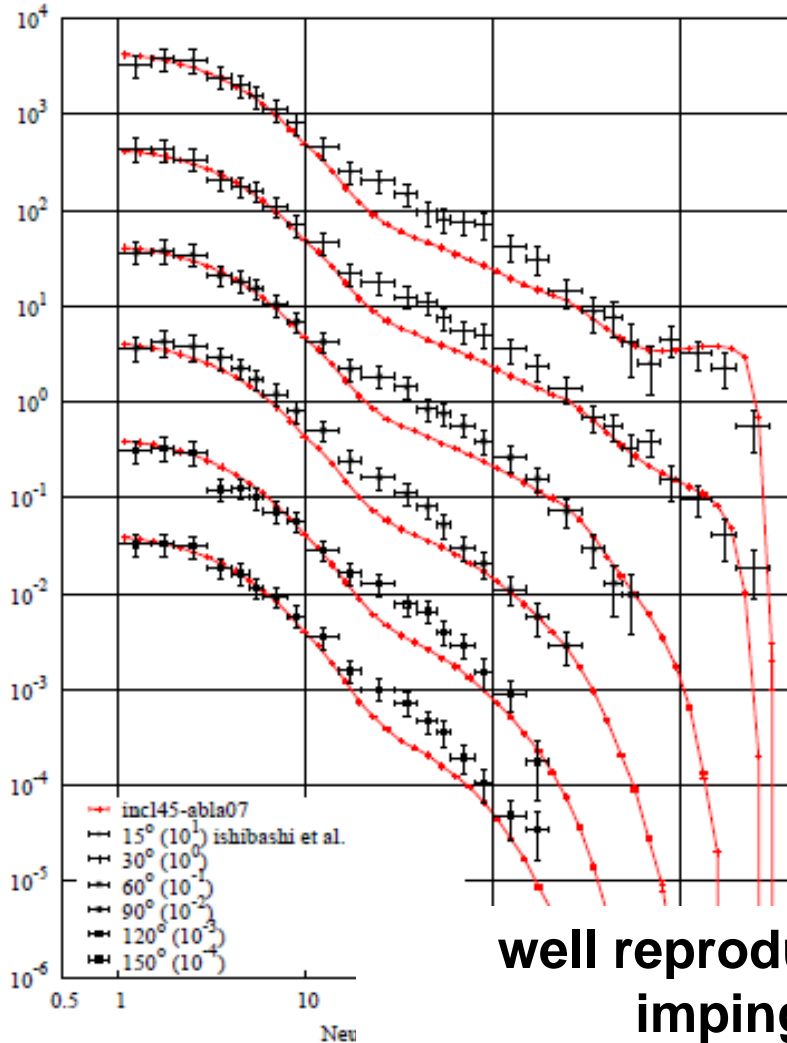


neutron energy (MeV)

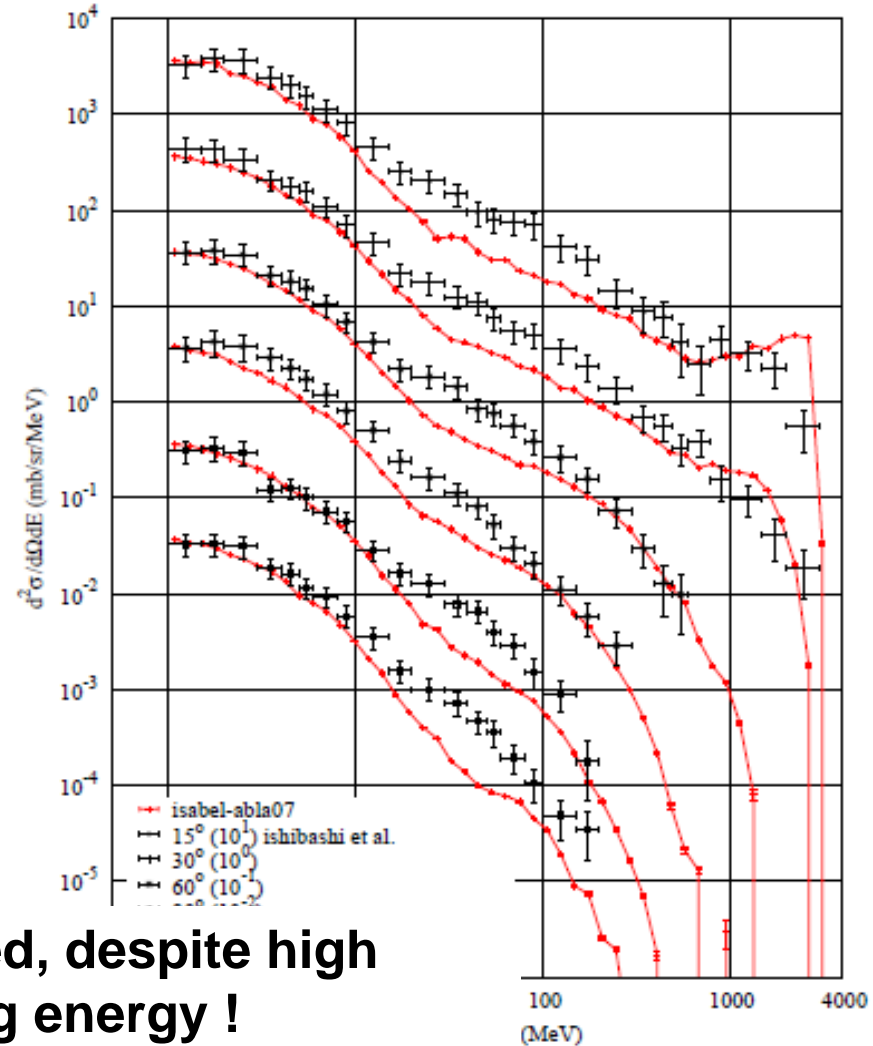
# p(3000 MeV) + Pb – Neutron spectrum

## INCL45-ABLA07

double differential cross section (mb/sr/MeV)



## ISABEL-ABLA07



**well reproduced, despite high  
impinging energy !**

neutron energy (MeV)

# Neutrons

## Neutron average multiplicity

**Status:** Good

**Improvement:**?

## Neutron multiplicity distribution

**Status:**

- INCL45+ any de-excitation and ISABEL+ any de-excitation: All models are too low at high neutron multiplicities (check  $E^*$  coming from INC models?)
- p(1200 MeV) + Fe: strange shape (only with ABLA07); maybe break-up?

**Improvement:** Test break-up contribution. Test shape of spectra with modified initial  $E^*$  distribution. Other?

## Neutron double differential cross sections (starting from for $E_{neutron} = 256$ MeV)

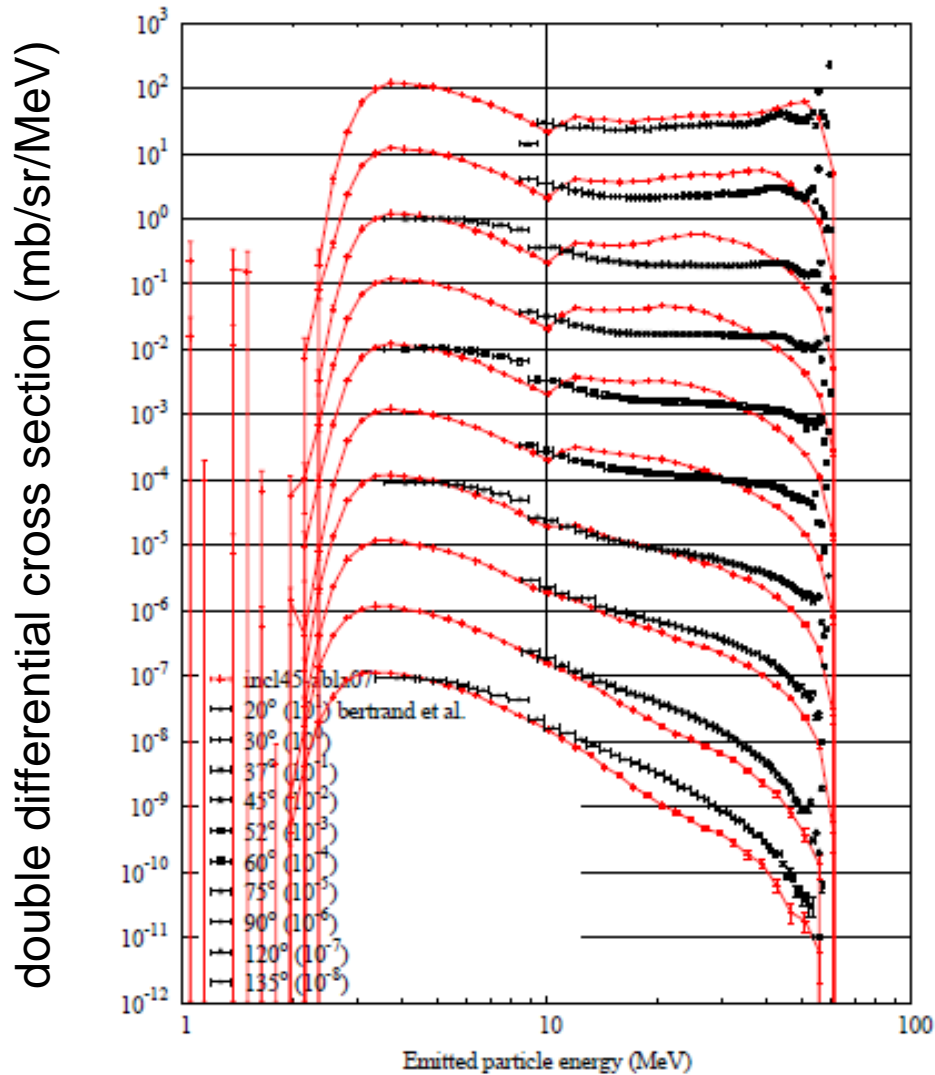
**Status:** Good

**Improvement:** INC models?

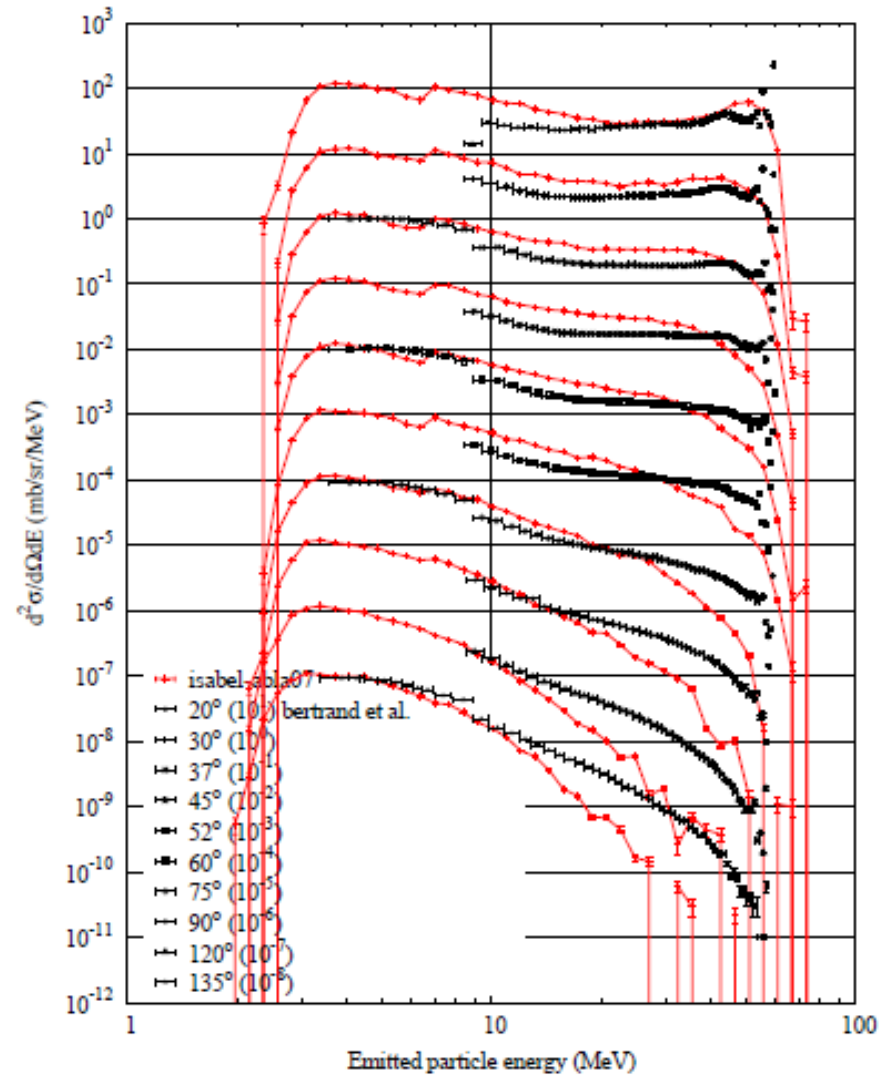
# Proton spectra

# p(62 MeV) + $^{56}\text{Fe}$ – Proton spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

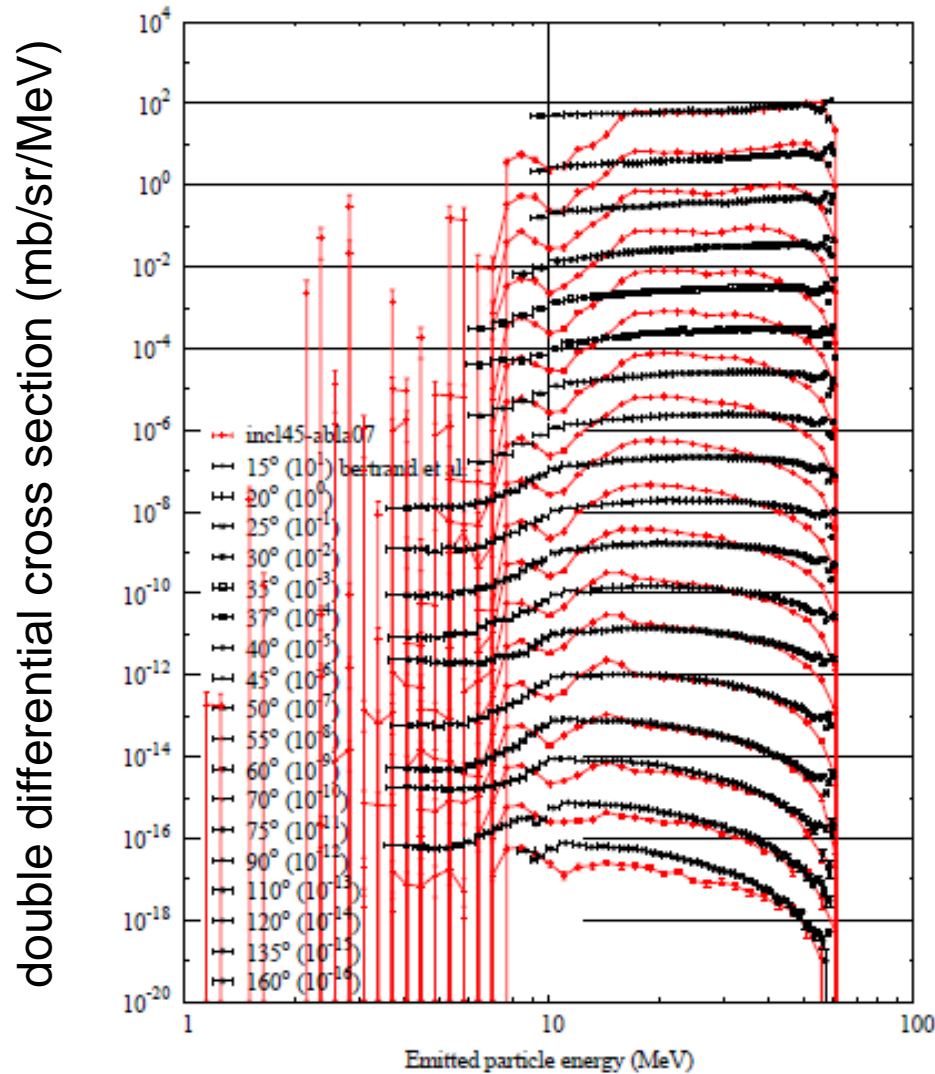


emitted-particle energy (MeV)

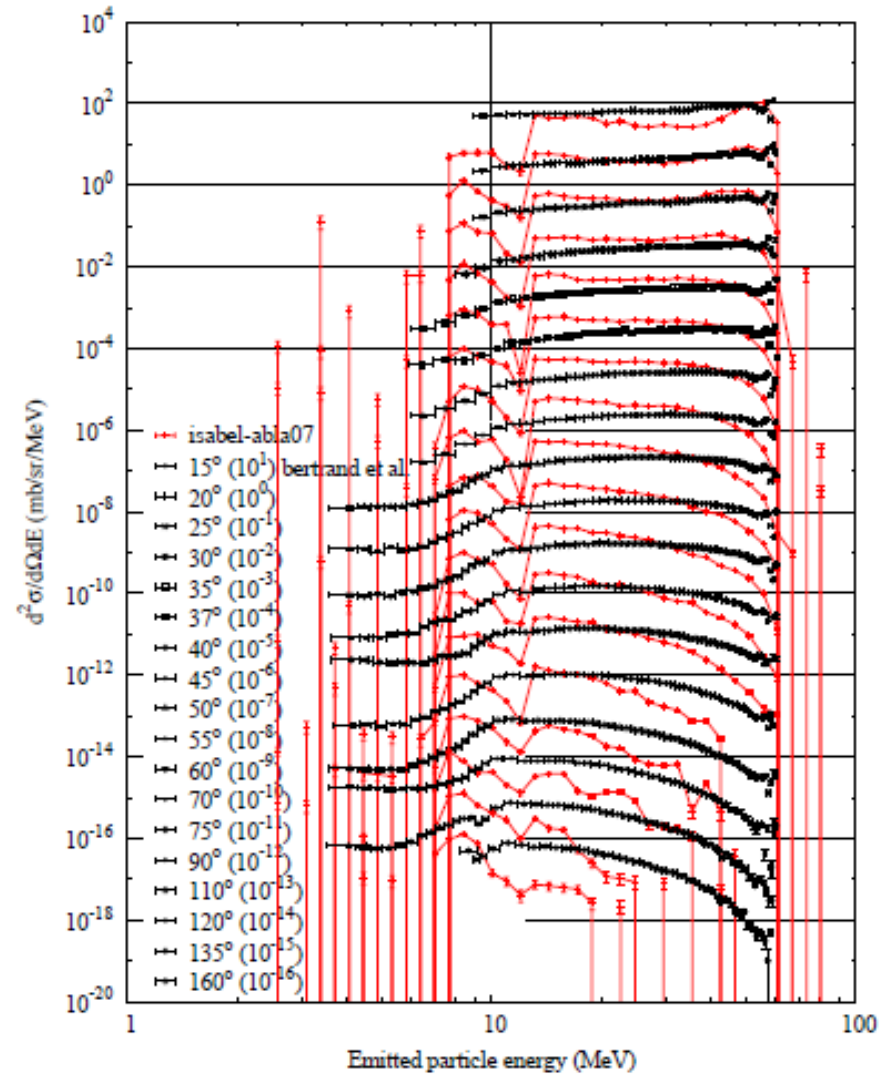


# p(62 MeV) + Bi – Proton spectrum

## INCL45-ABLA07



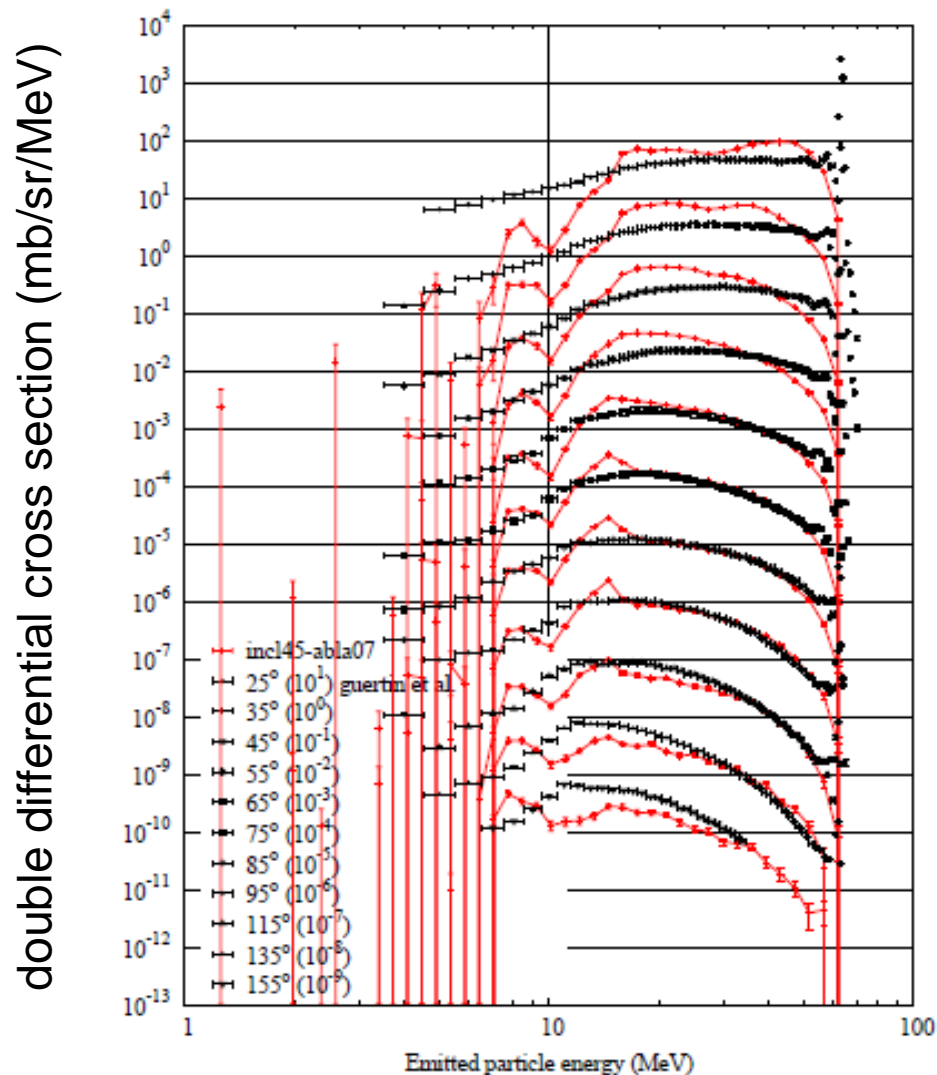
## ISABEL-ABLA07



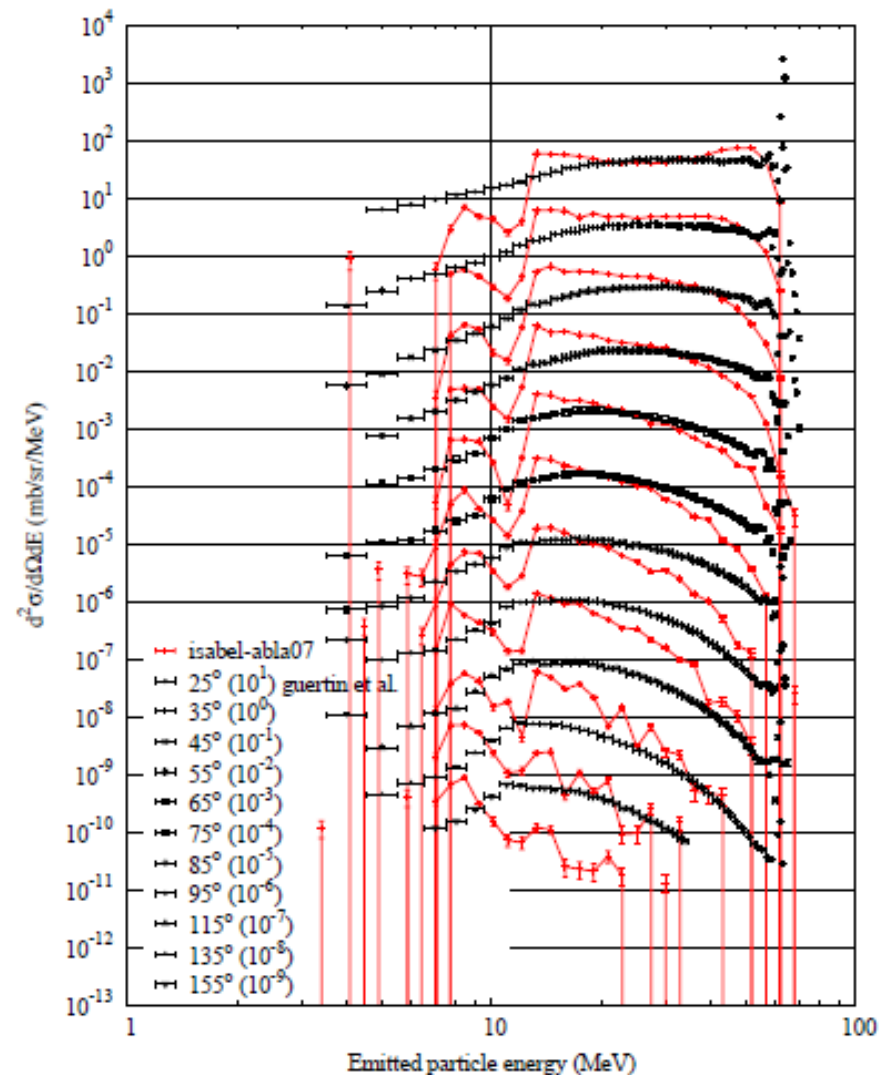
emitted-particle energy (MeV)

# p(63 MeV) + <sup>208</sup>Pb – Proton spectrum

## INCL45-ABLA07



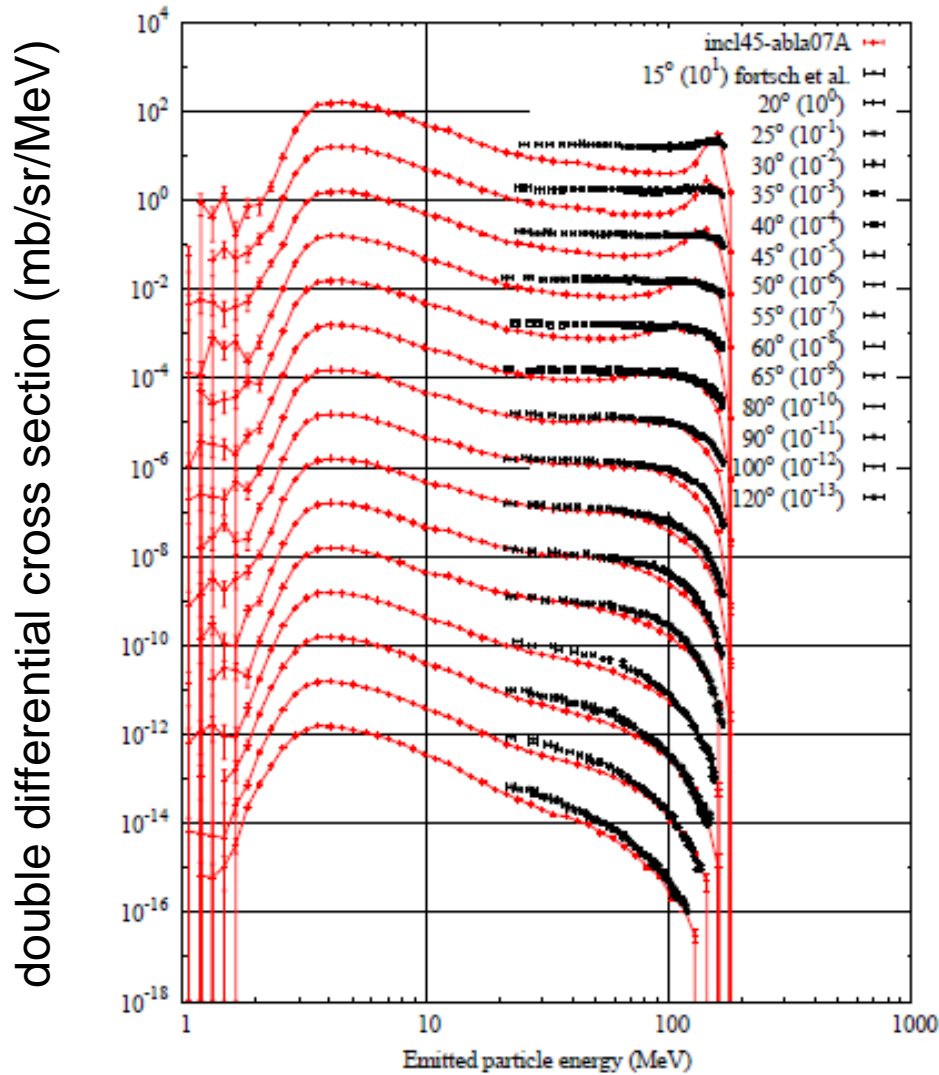
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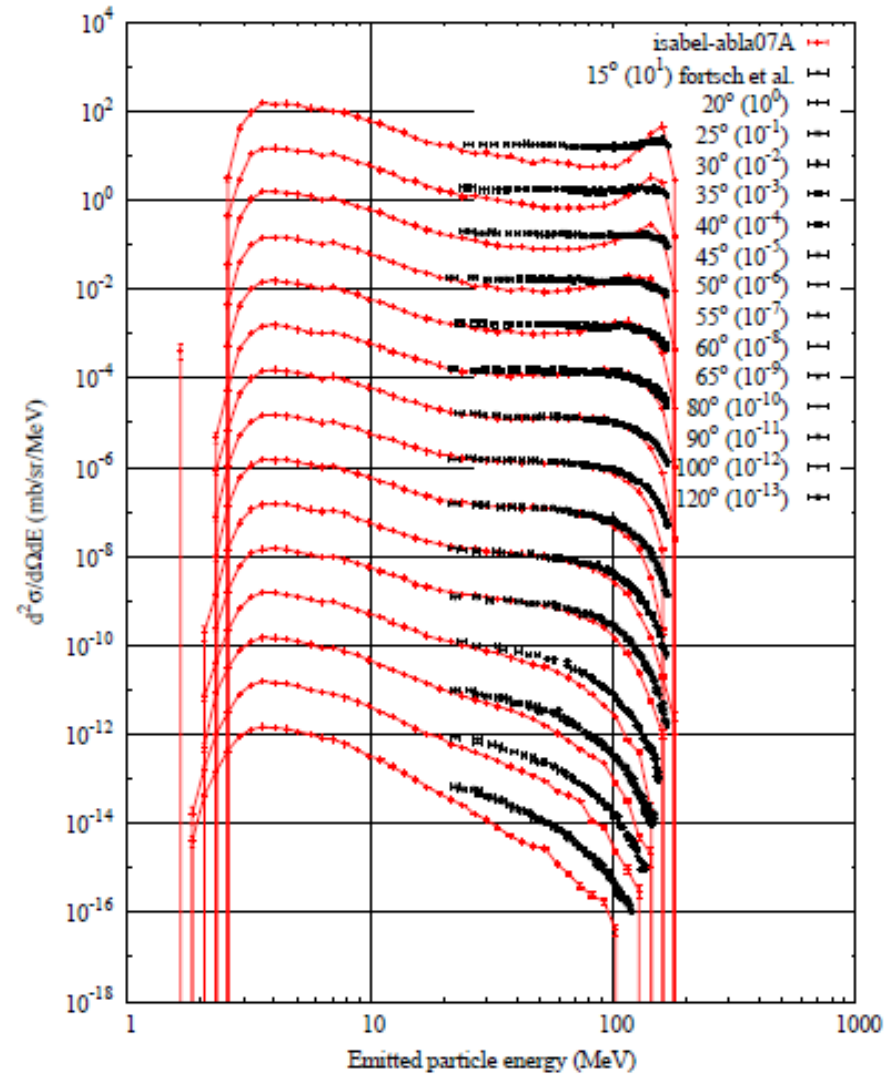
emitted-particle energy (MeV)

# p(175 MeV) + Ni – Proton spectrum

## INCL45-ABLA07



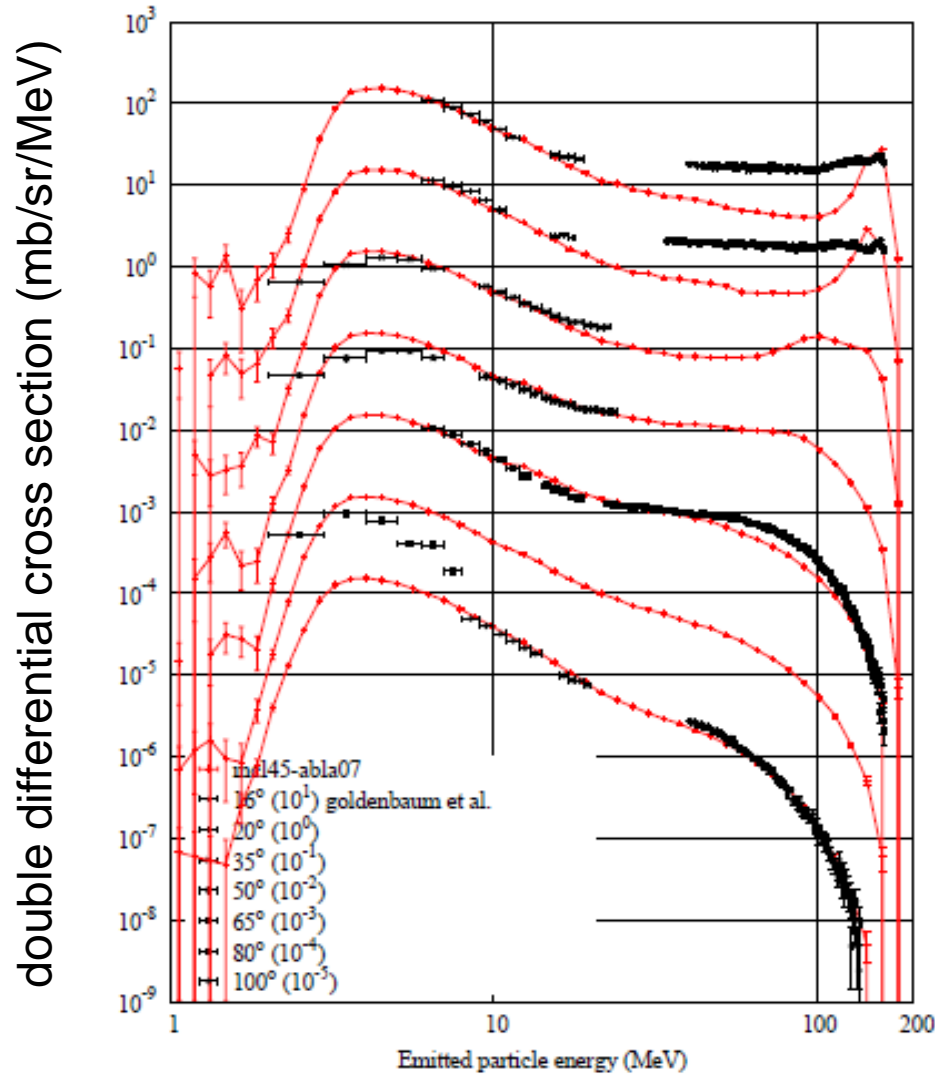
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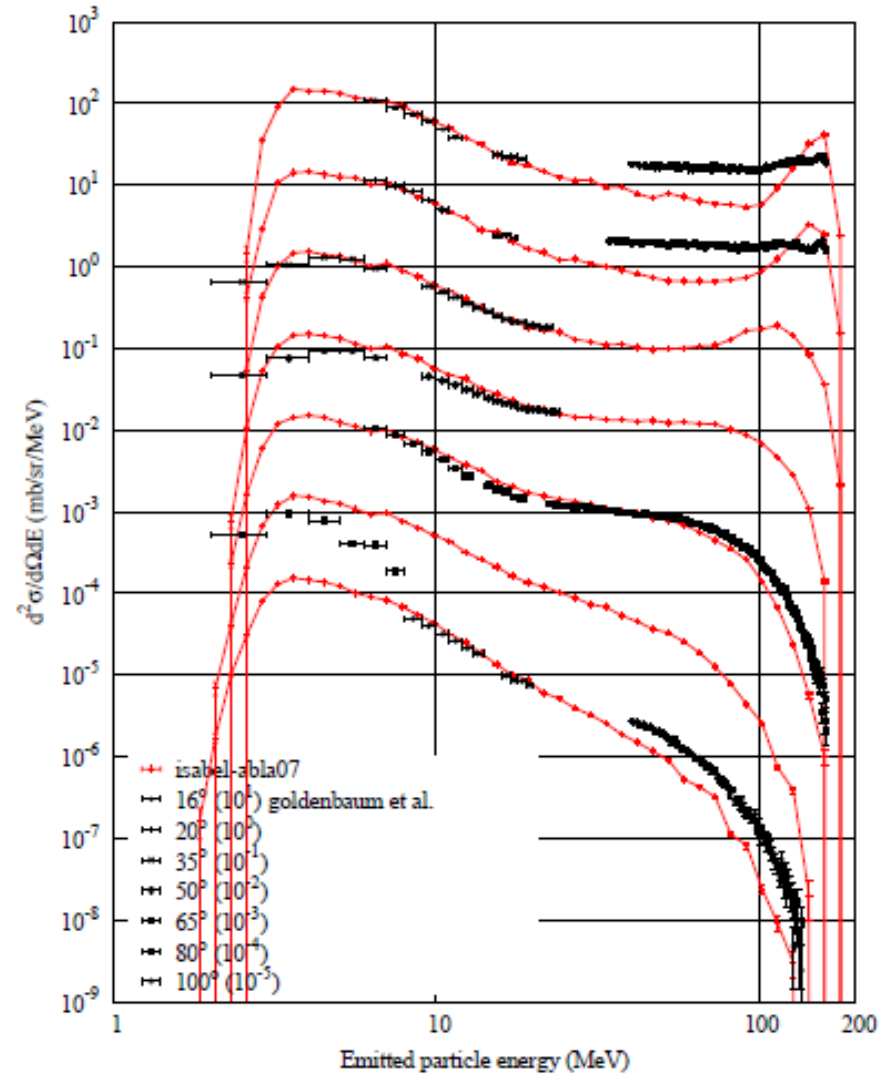
emitted-particle energy (MeV)

# p(175 MeV) + Ni – Proton spectrum

## INCL45-ABLA07



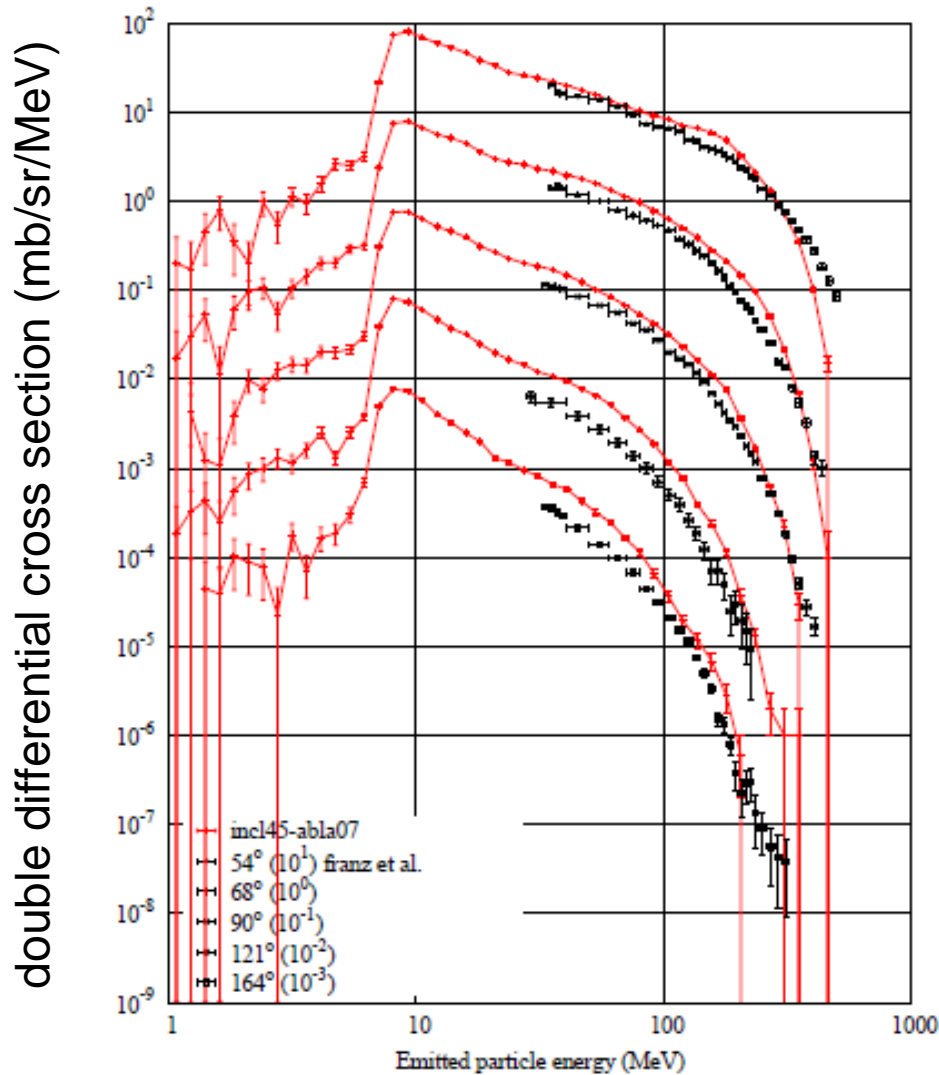
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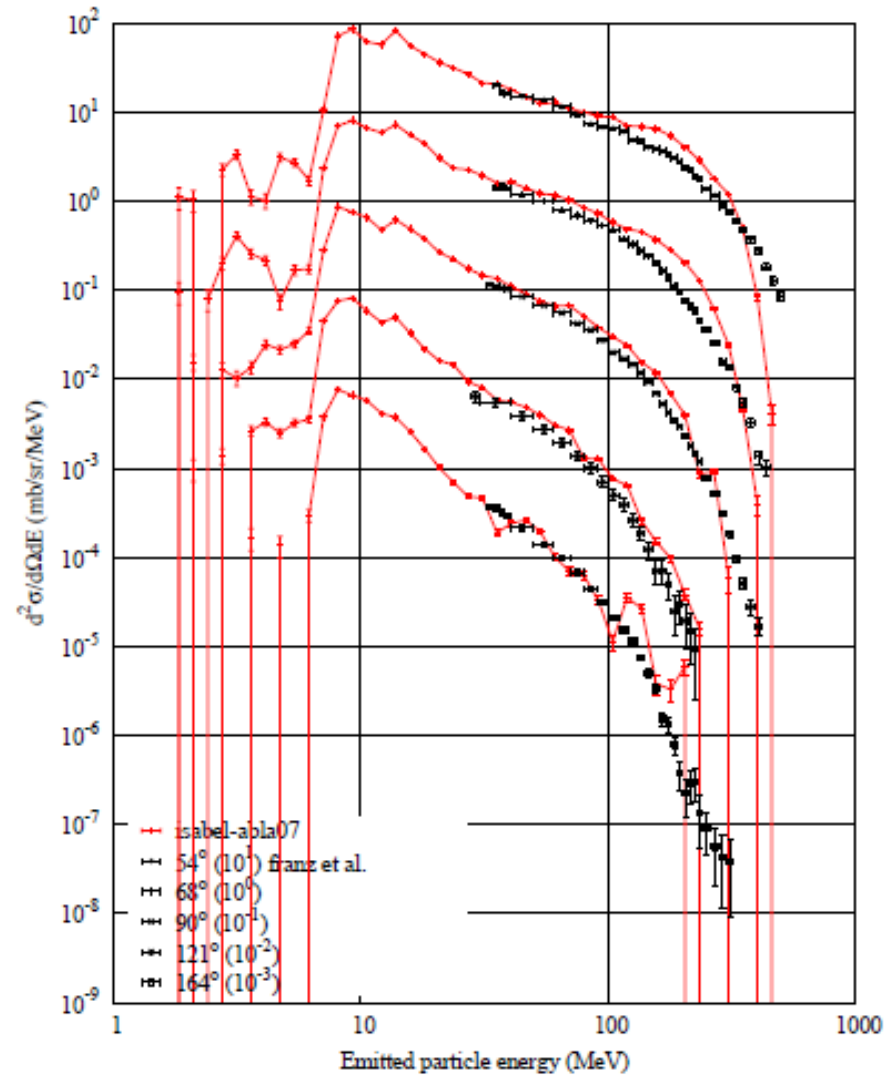
emitted-particle energy (MeV)

# n(542 MeV) + Bi – Proton spectrum

## INCL45-ABLA07



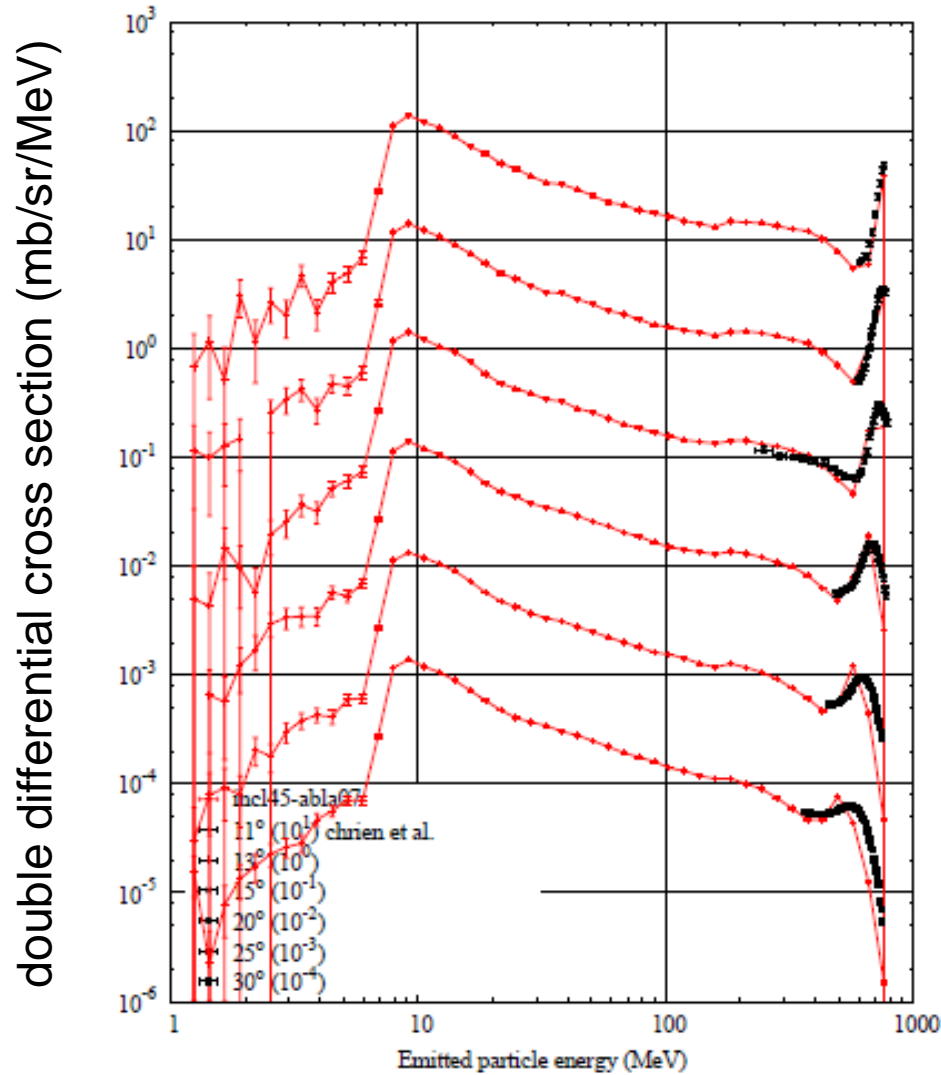
## ISABEL-ABLA07



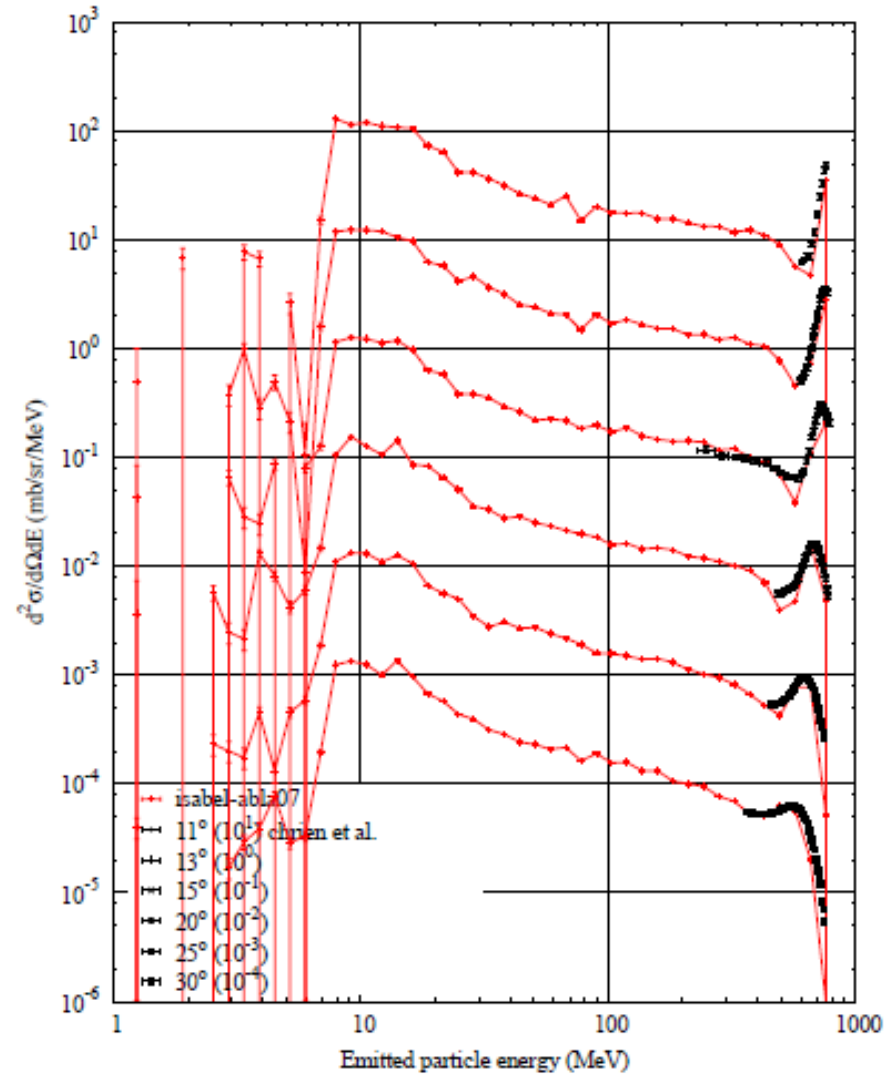
emitted-particle energy (MeV)

# p(800 MeV) + <sup>208</sup>Pb – Proton spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

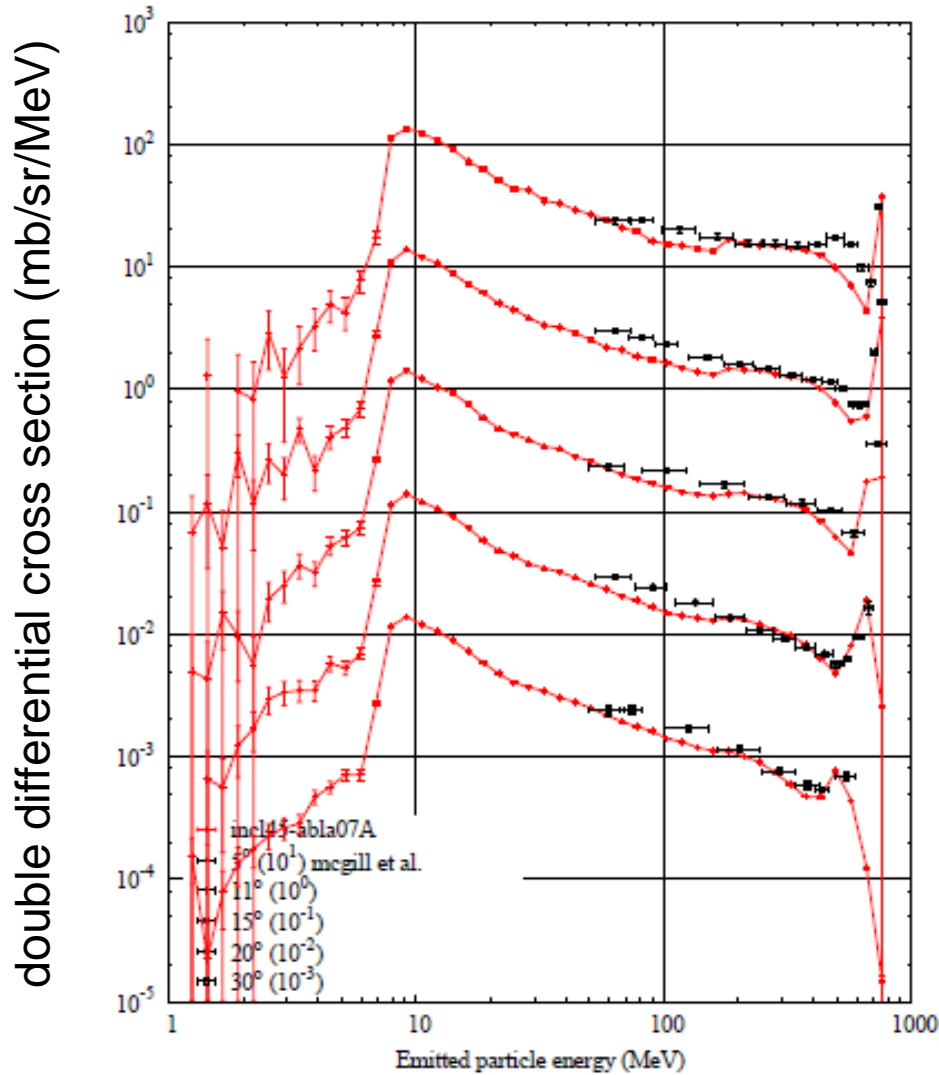


emitted-particle energy (MeV)

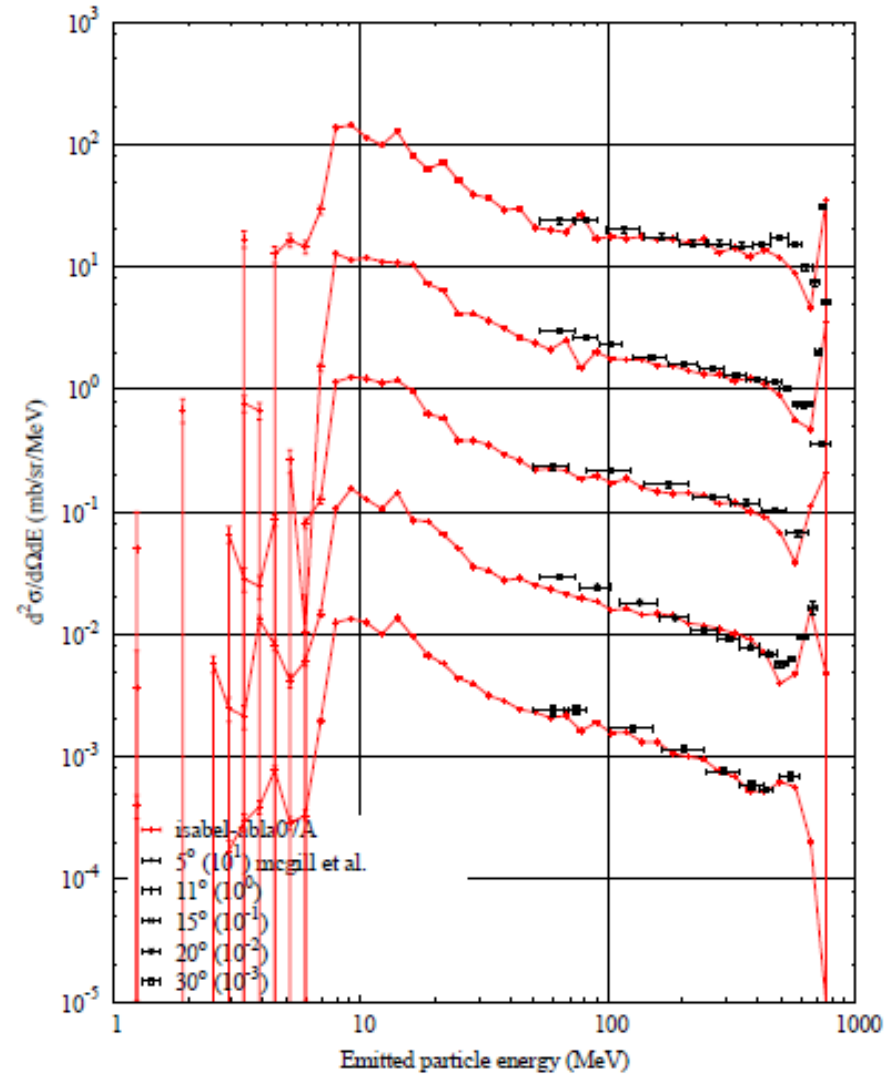


# p(800 MeV) + $^{208}\text{Pb}$ – Proton spectrum

## INCL45-ABLA07



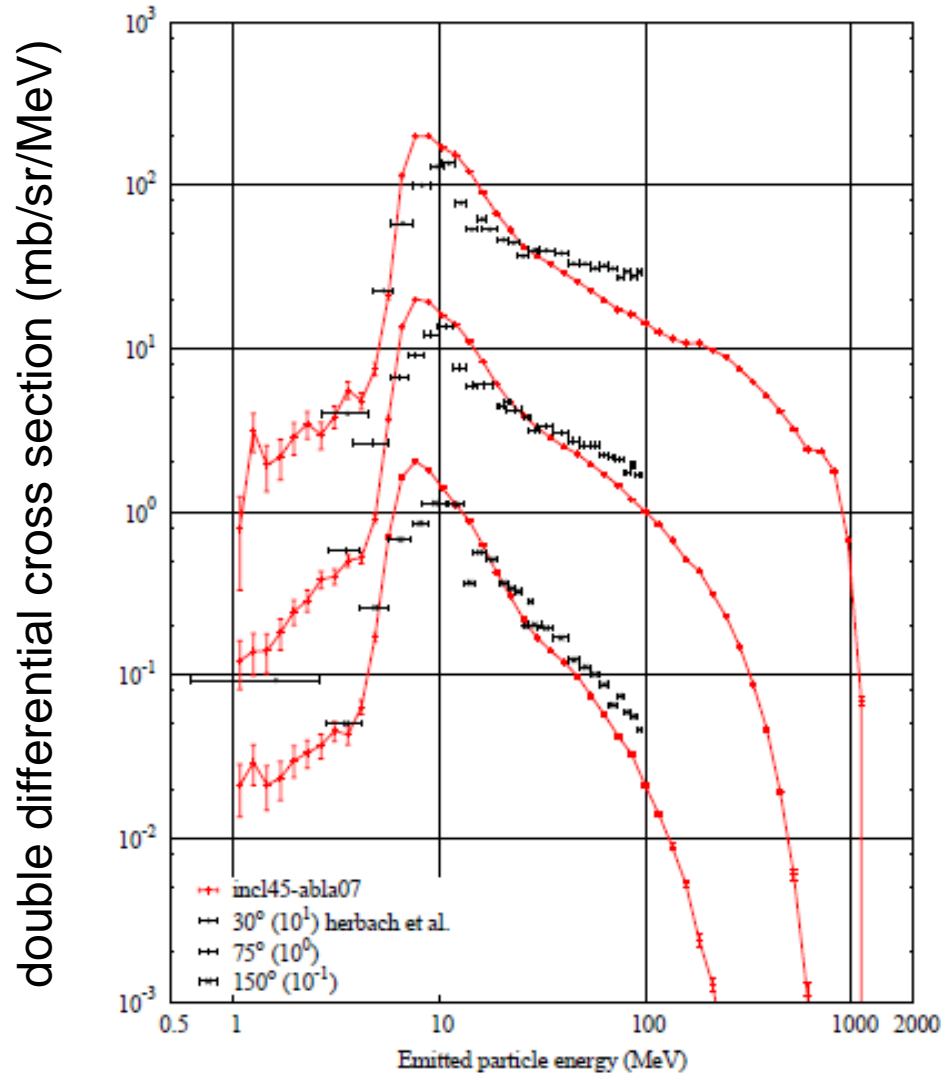
## ISABEL-ABLA07



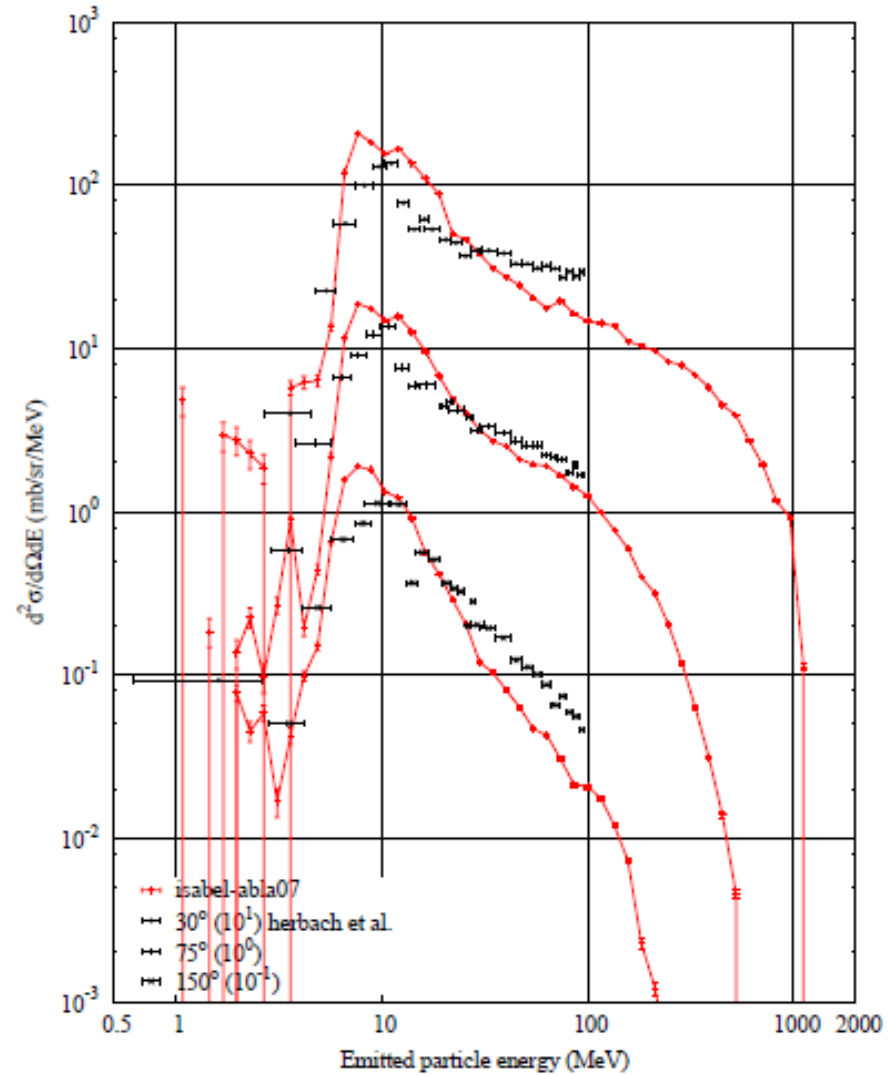
emitted-particle energy (MeV)

# p(1200 MeV) + Ta – Proton spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

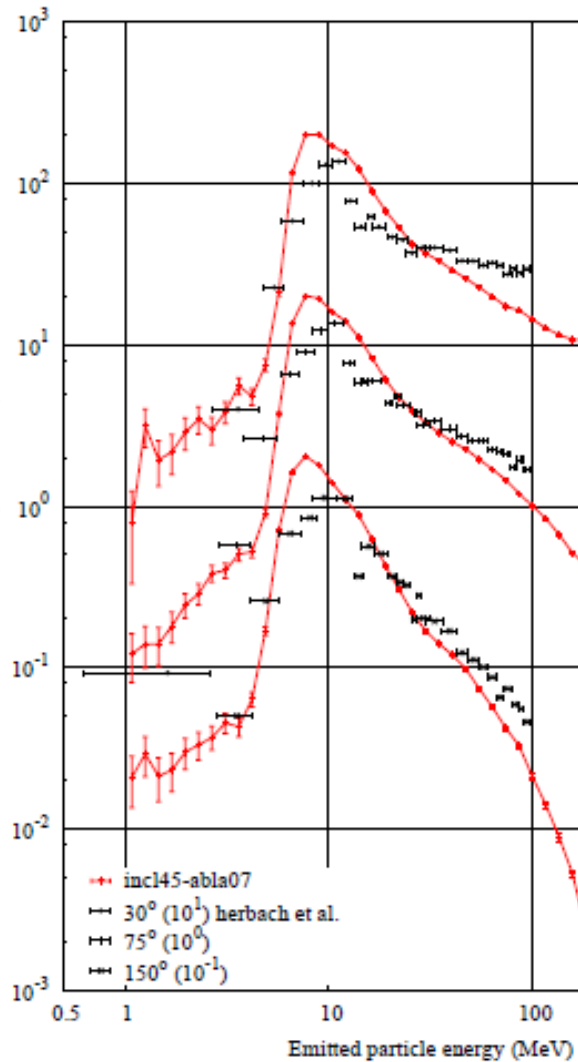


emitted-particle energy (MeV)

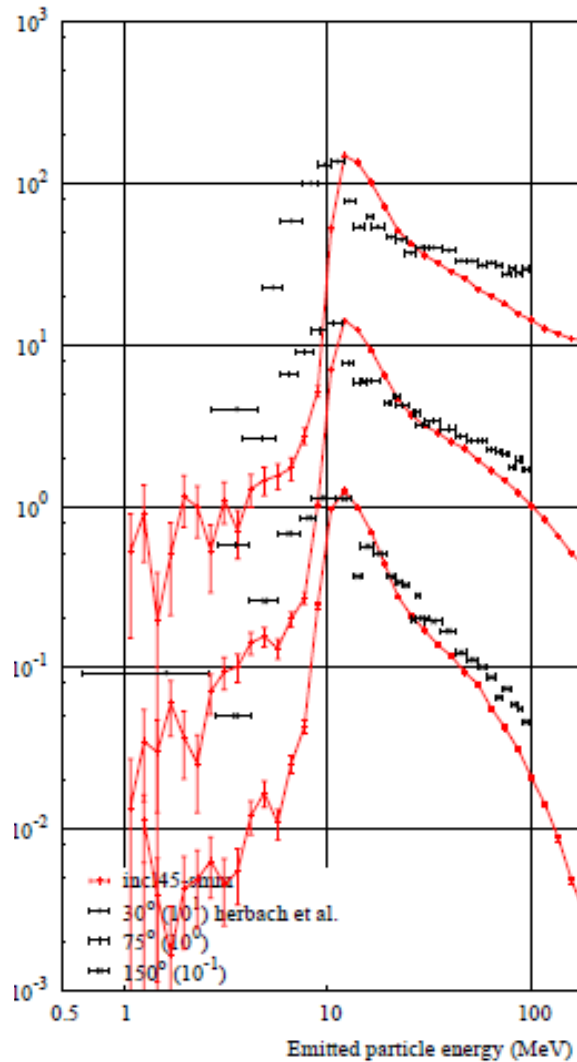


# p(1200 MeV) + Ta – Proton spectrum

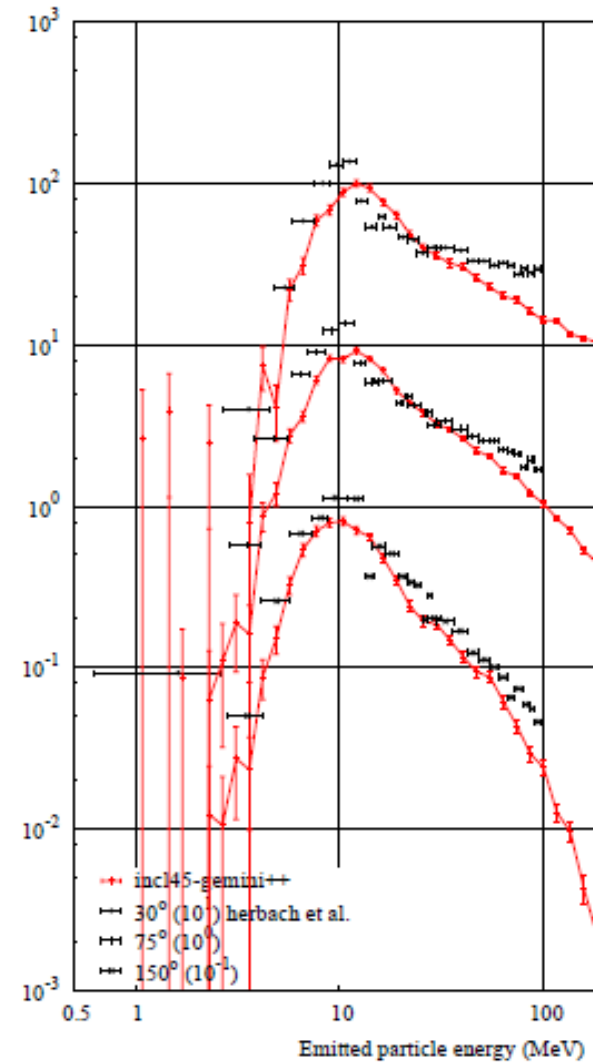
## INCL45-ABLA07



## INCL45-SMM



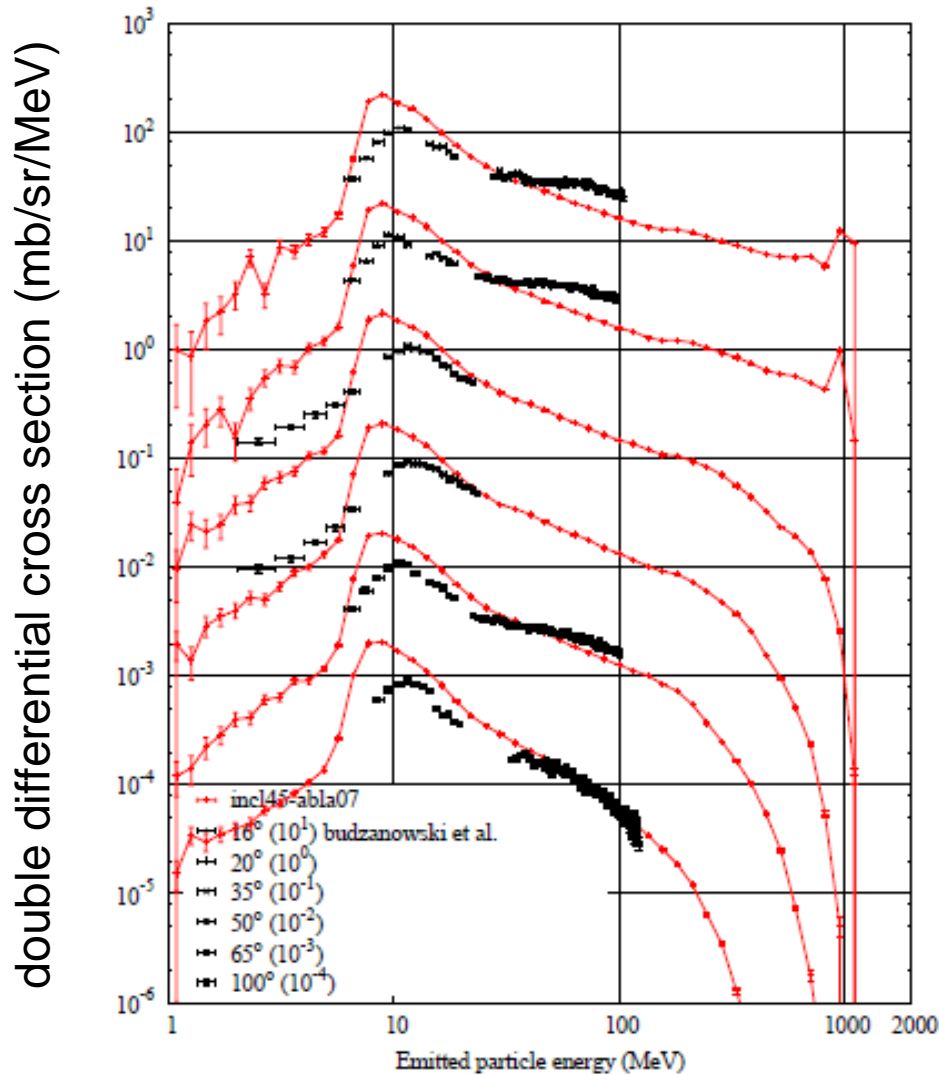
## INCL45-GEMINI++



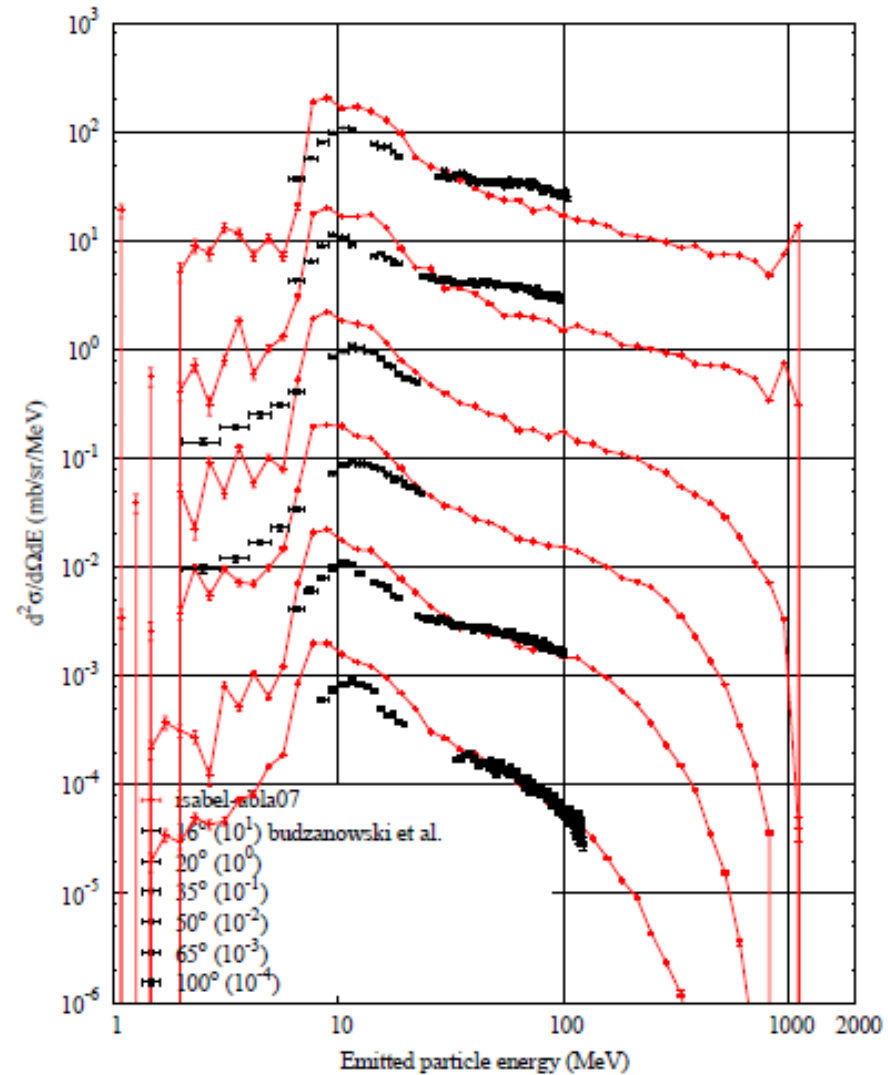
emitted-particle energy (MeV)

# p(1200 MeV) + Au – Proton spectrum

## INCL45-ABLA07



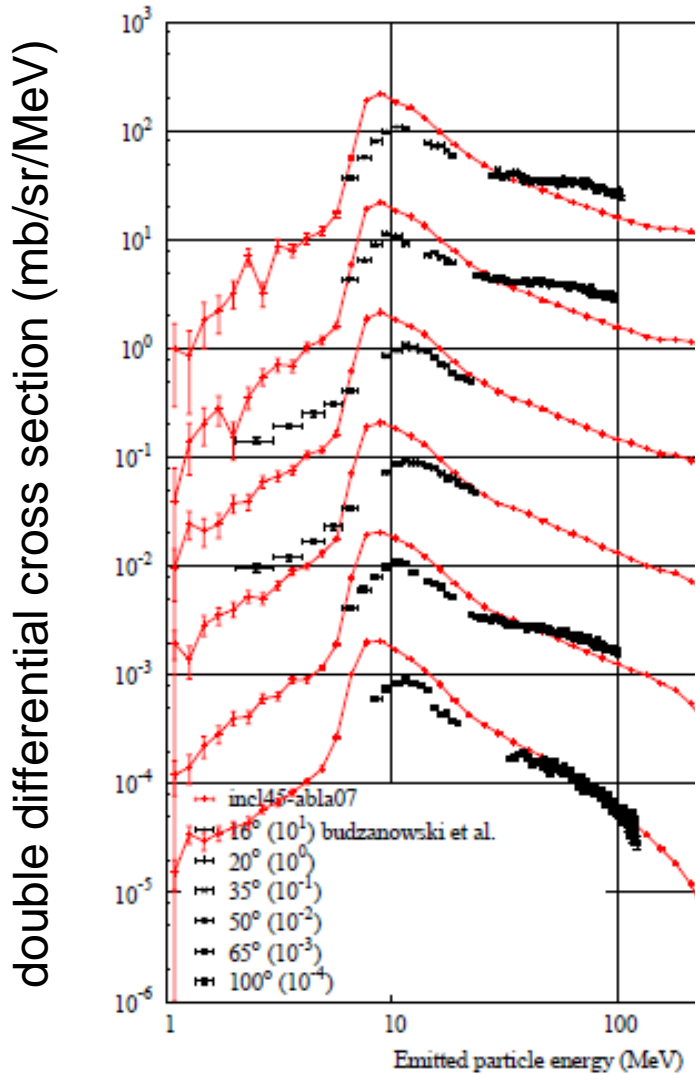
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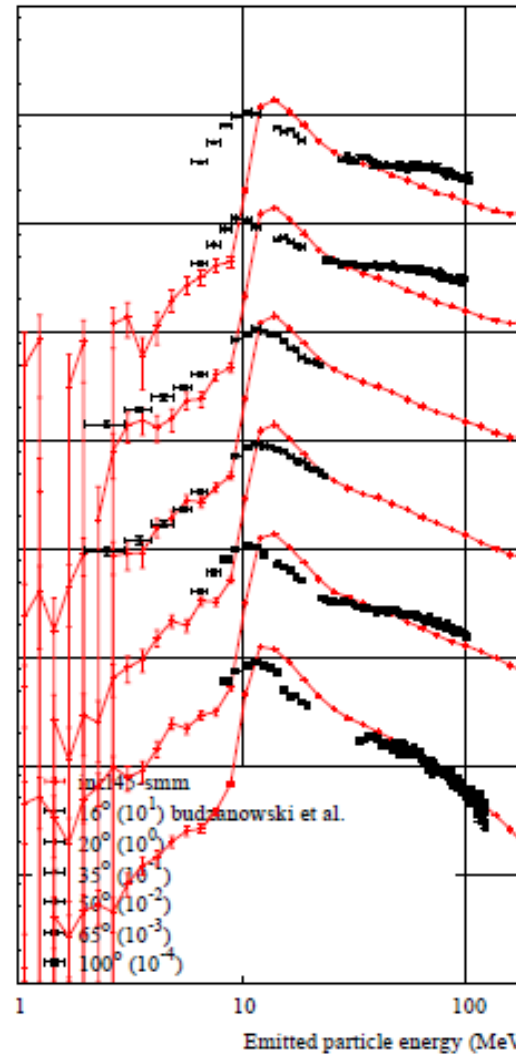
emitted-particle energy (MeV)

# p(1200 MeV) + Au – Proton spectrum

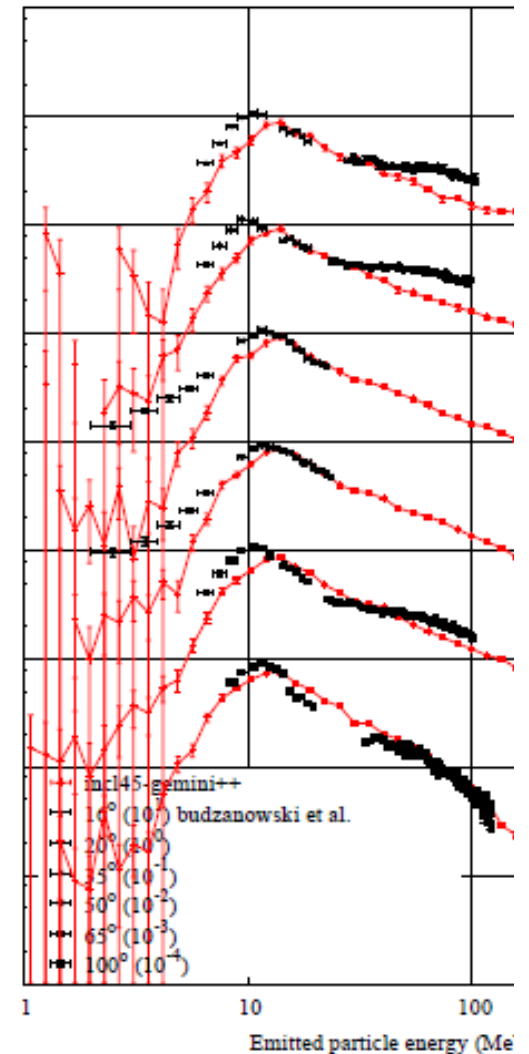
## INCL45-ABLA07



## INCL45-SMM



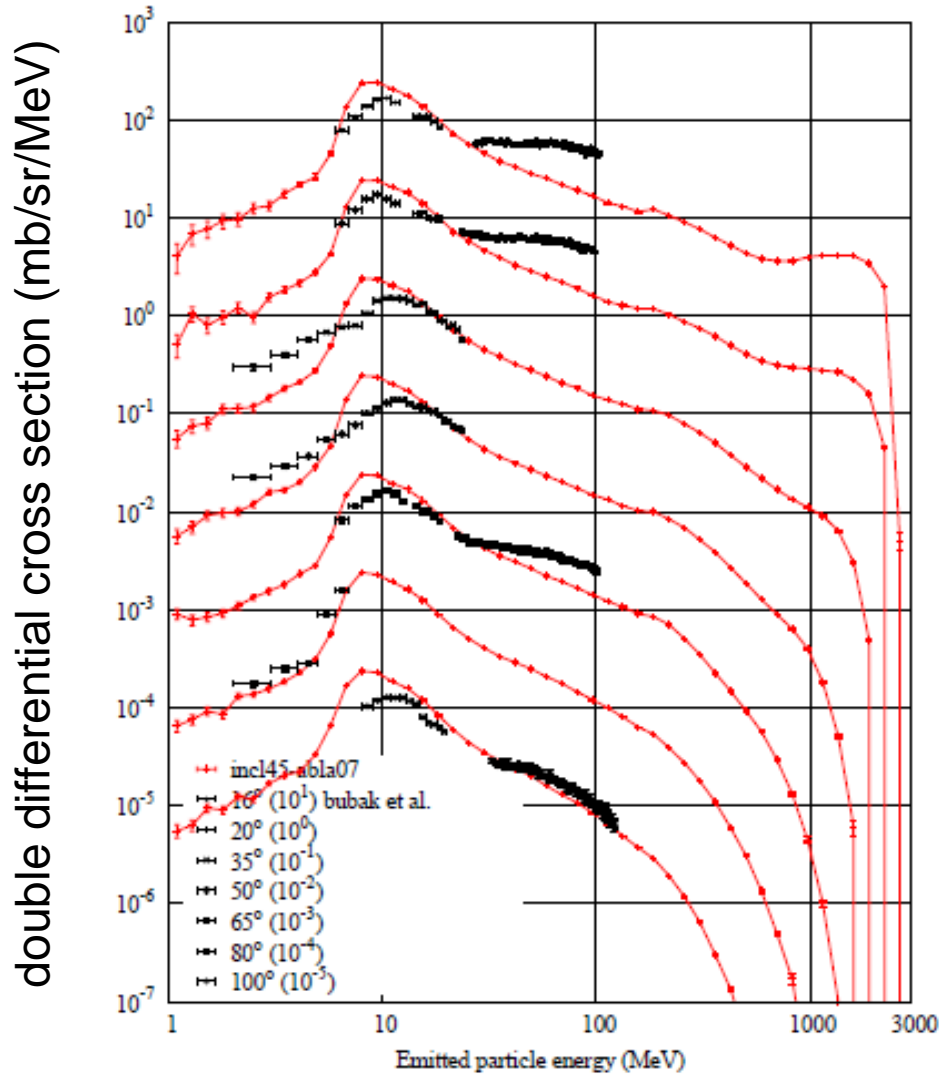
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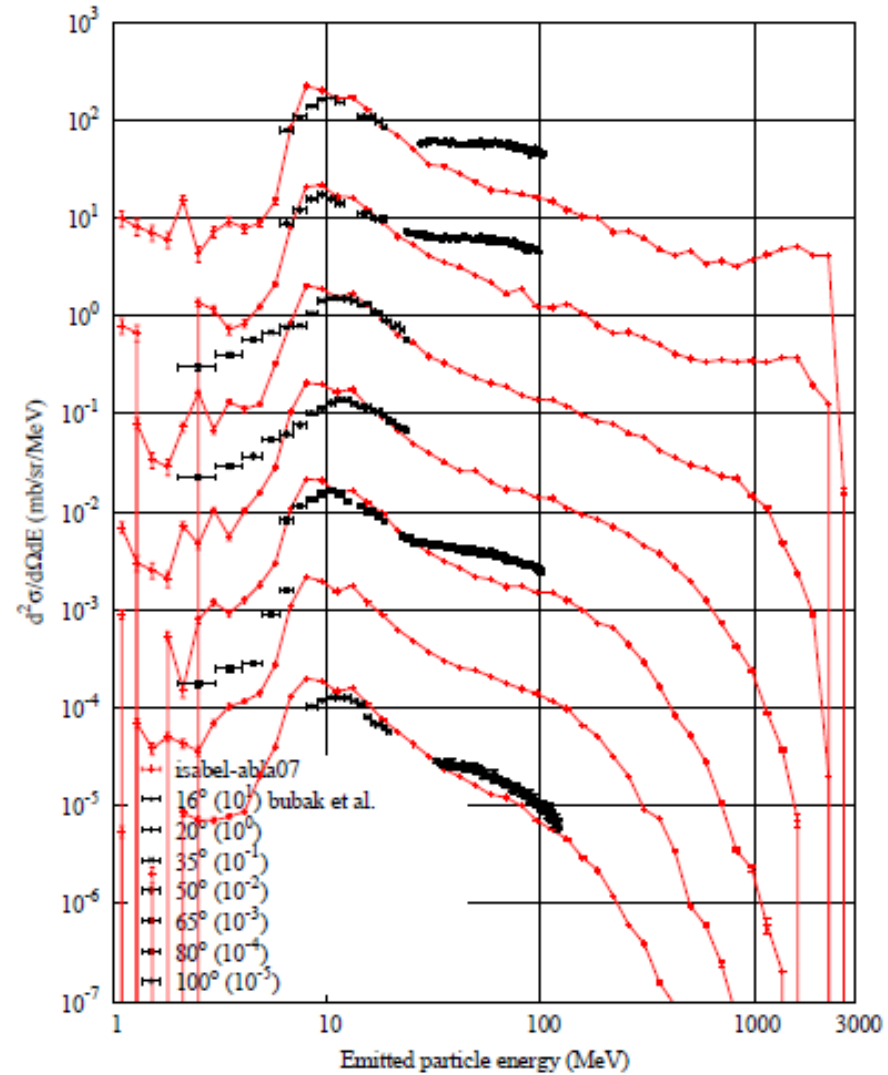
emitted-particle energy (MeV)

# p(2500 MeV) + Au – Proton spectrum

## INCL45-ABLA07



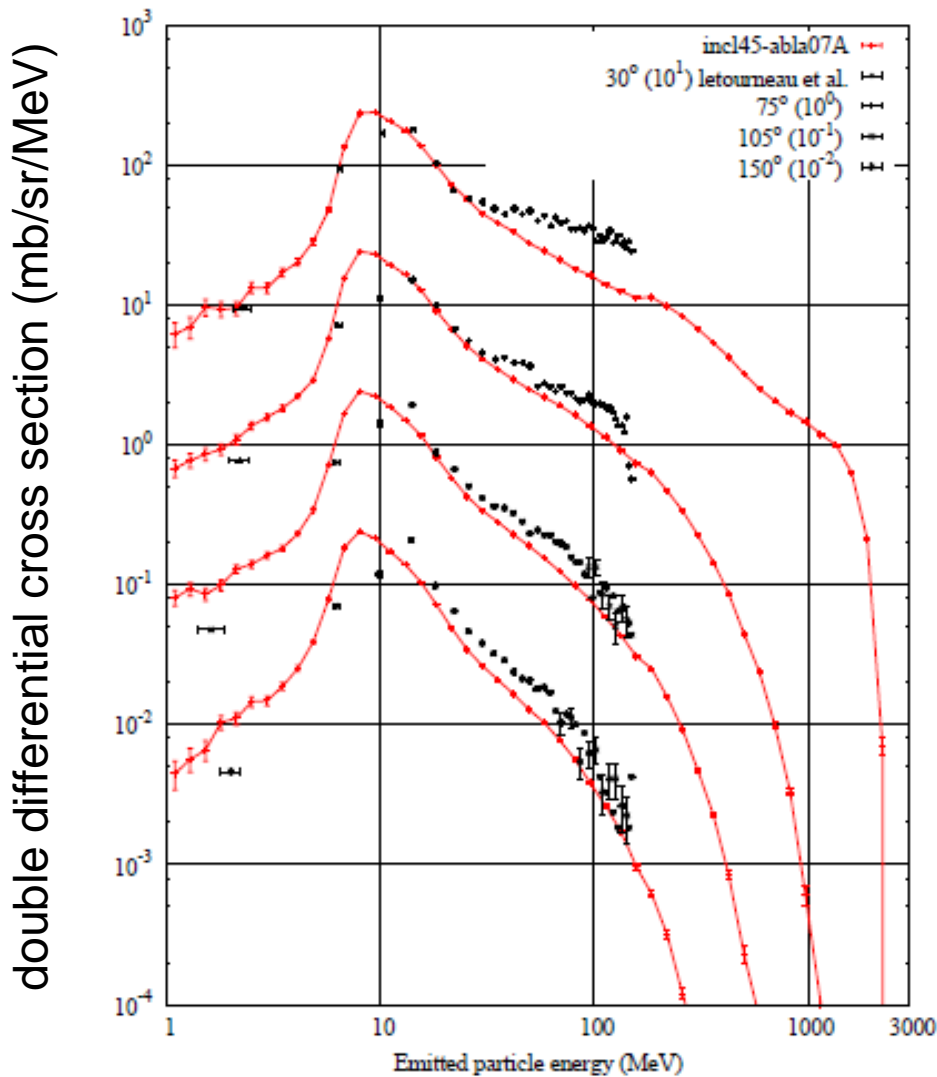
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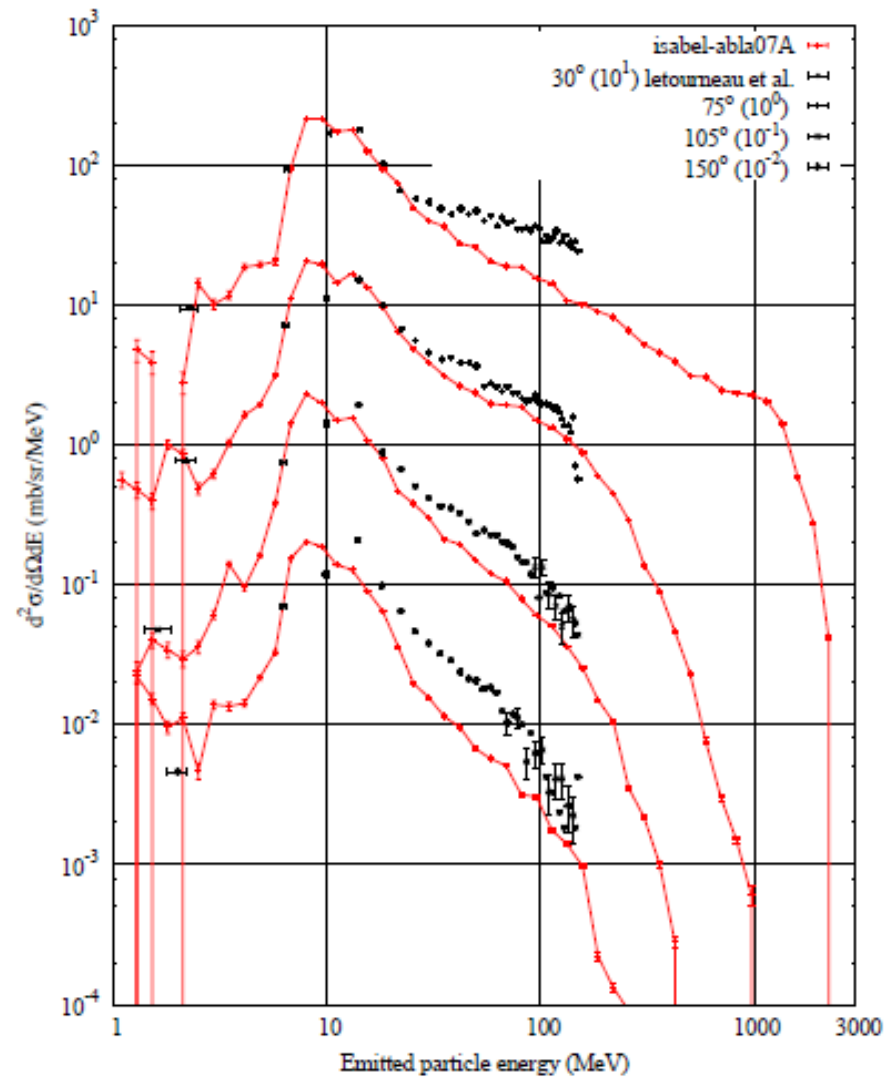
emitted-particle energy (MeV)

# p(2500 MeV) + Au – Proton spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

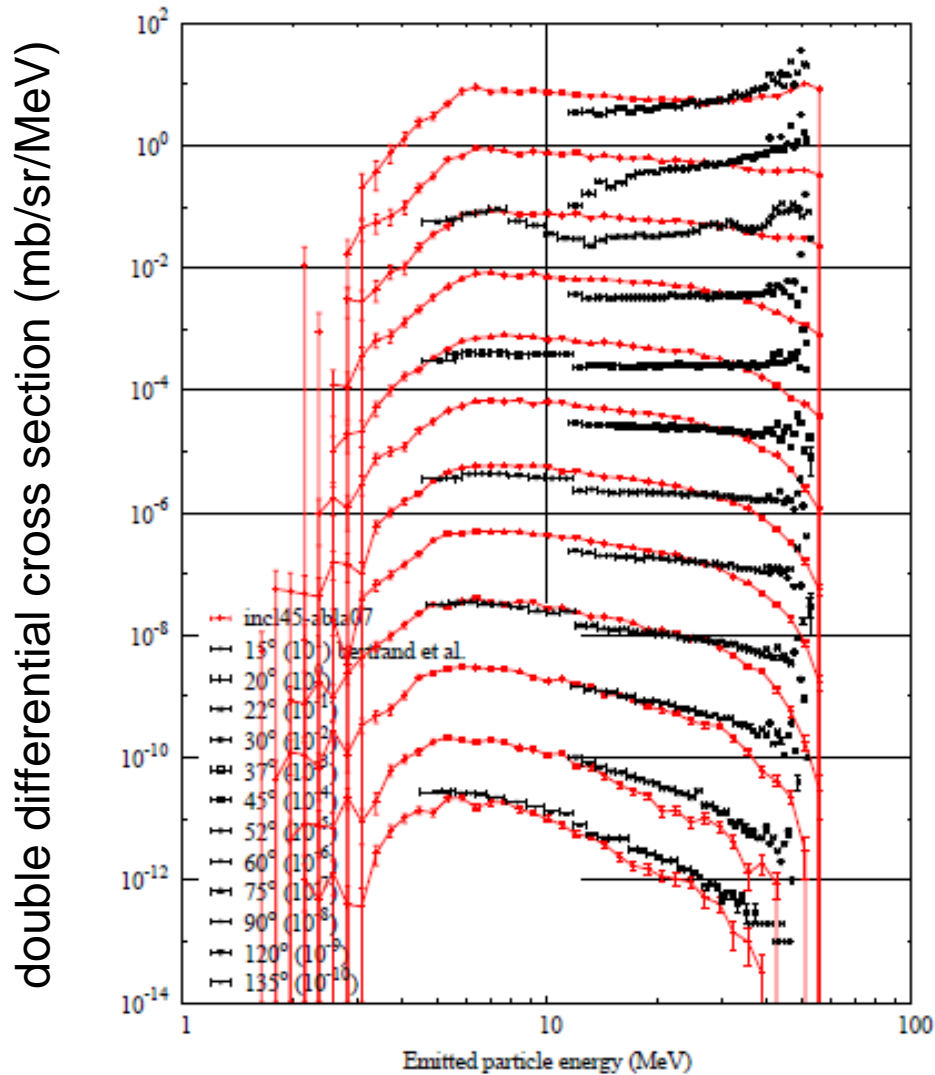


emitted-particle energy (MeV)

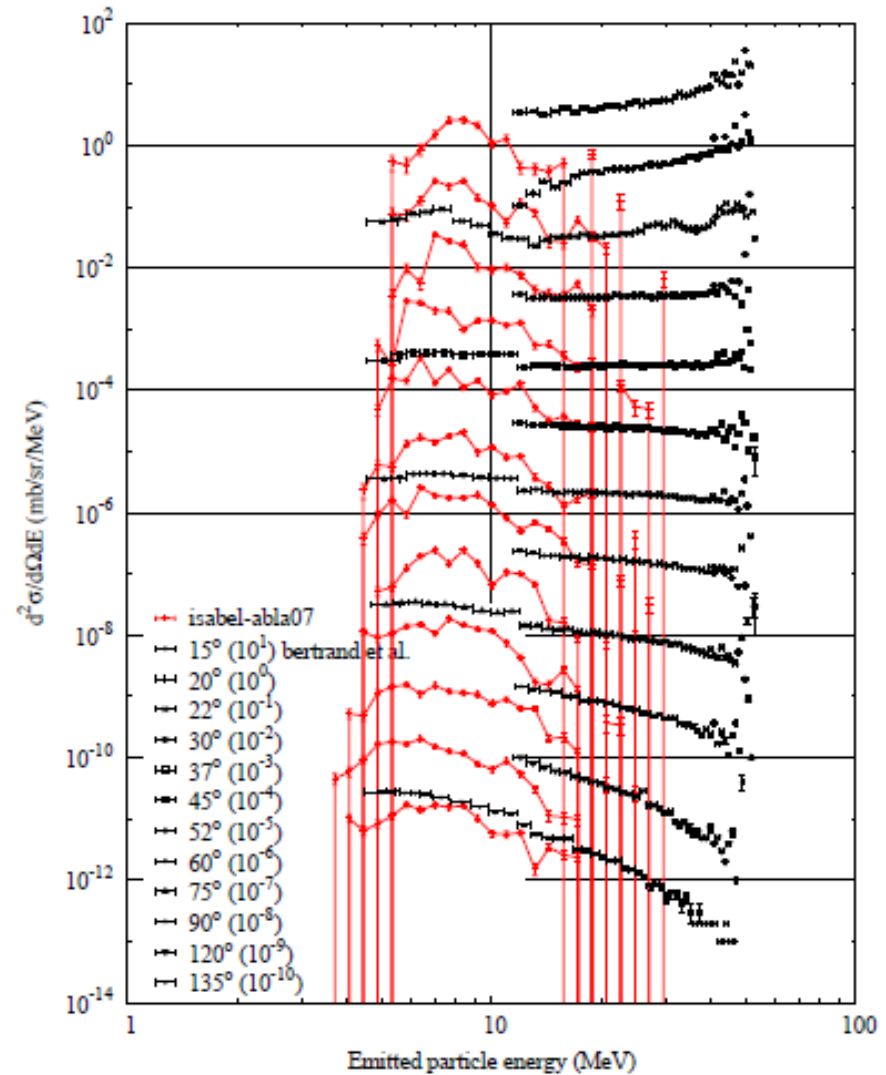
# Deuteron spectra

# p(62 MeV) + <sup>56</sup>Fe – Deuteron spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

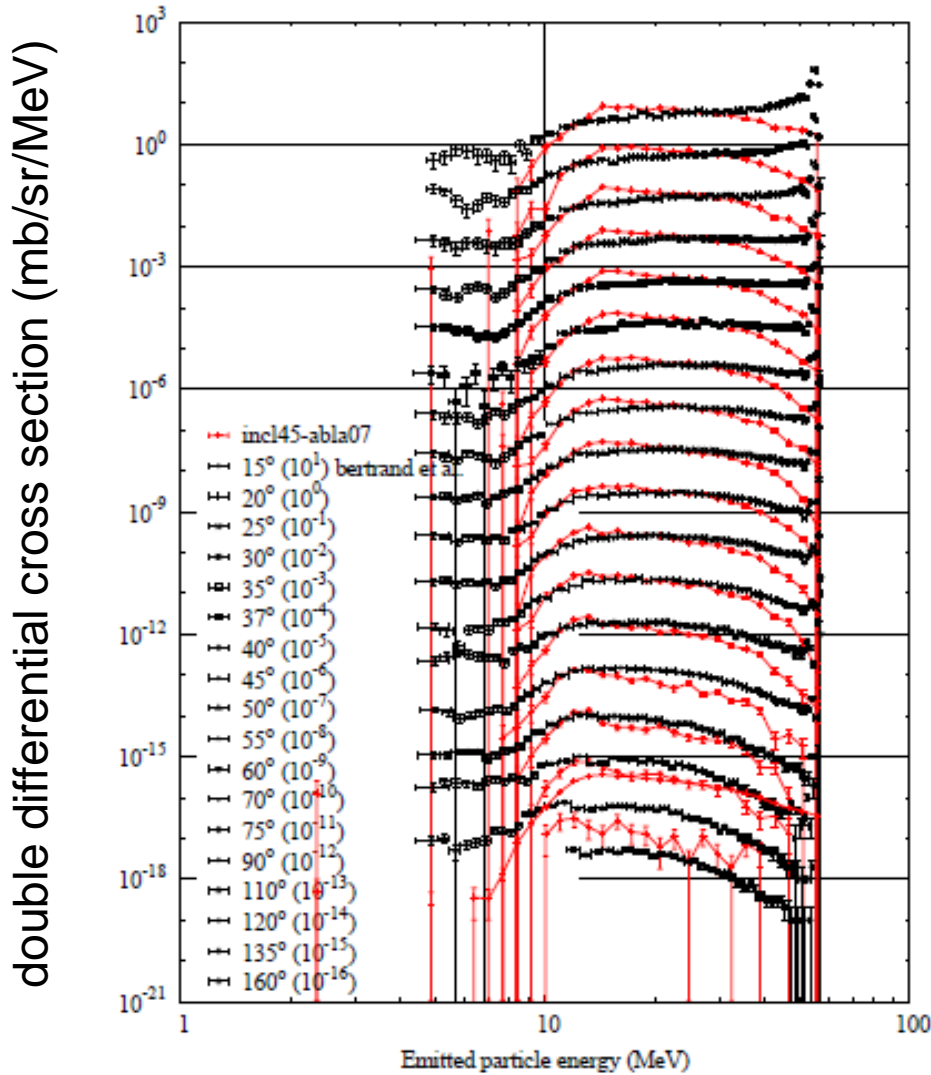


emitted-particle energy (MeV)

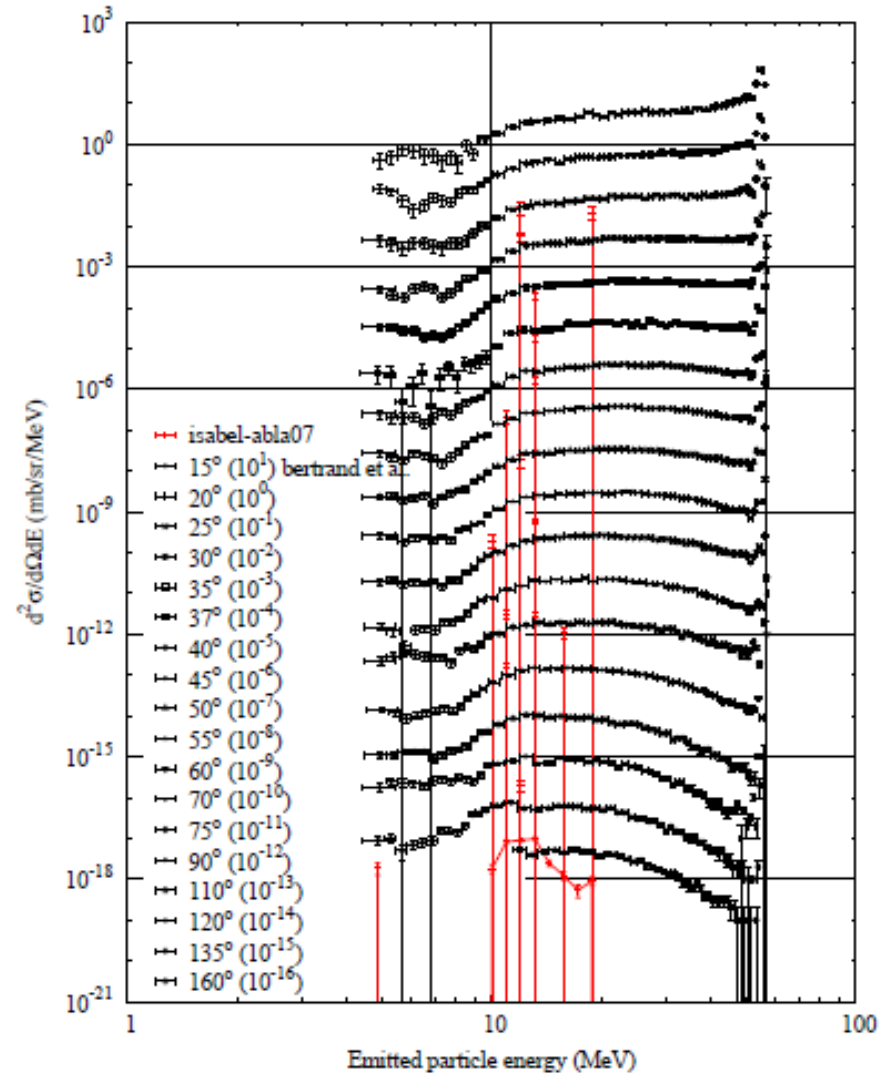


# p(62 MeV) + Bi – Deuteron spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

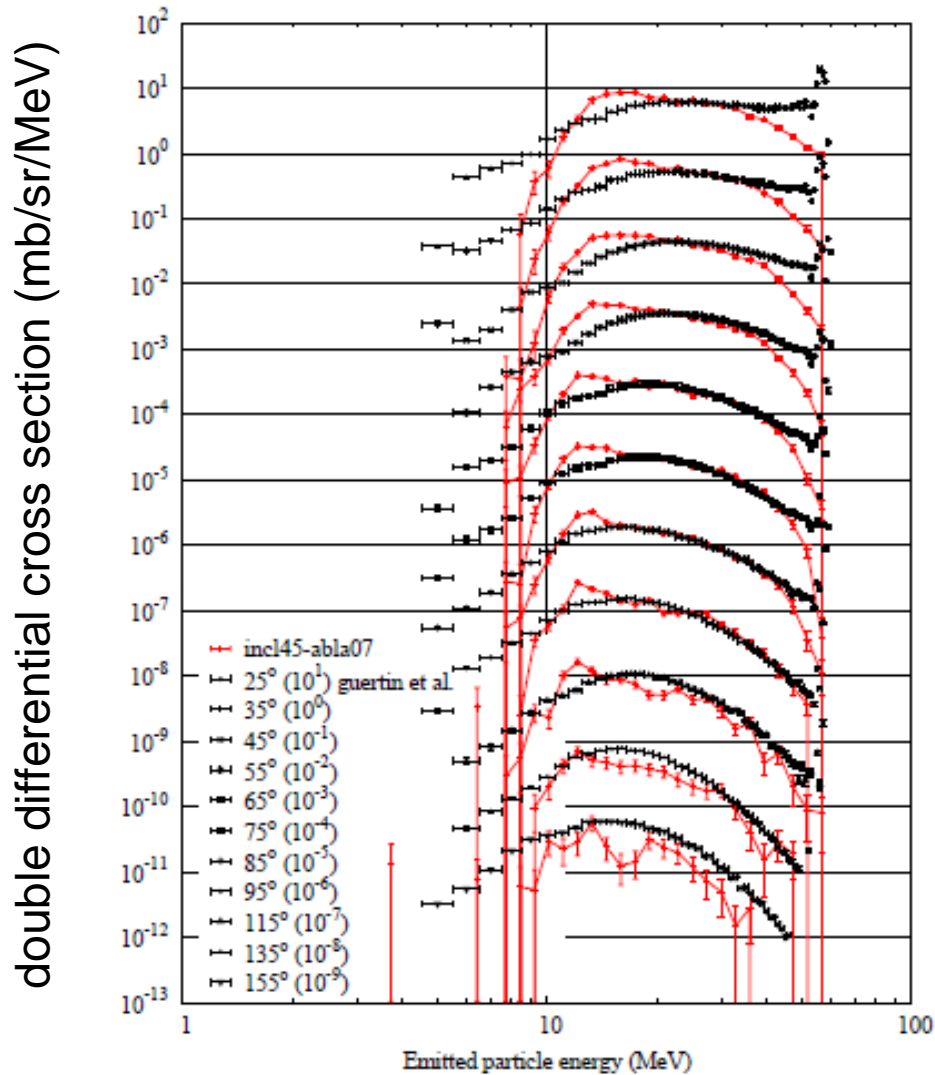


emitted-particle energy (MeV)

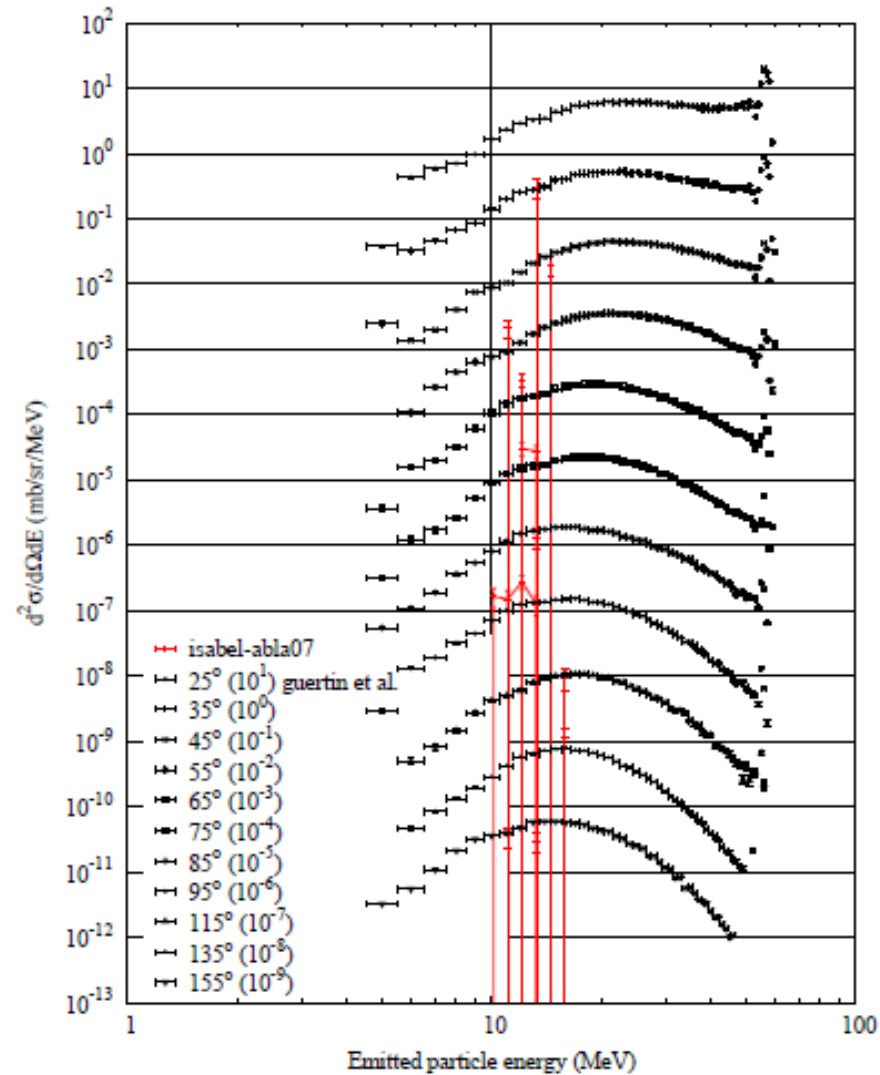


# p(63 MeV) + <sup>208</sup>Pb – Deuteron spectrum

## INCL45-ABLA07



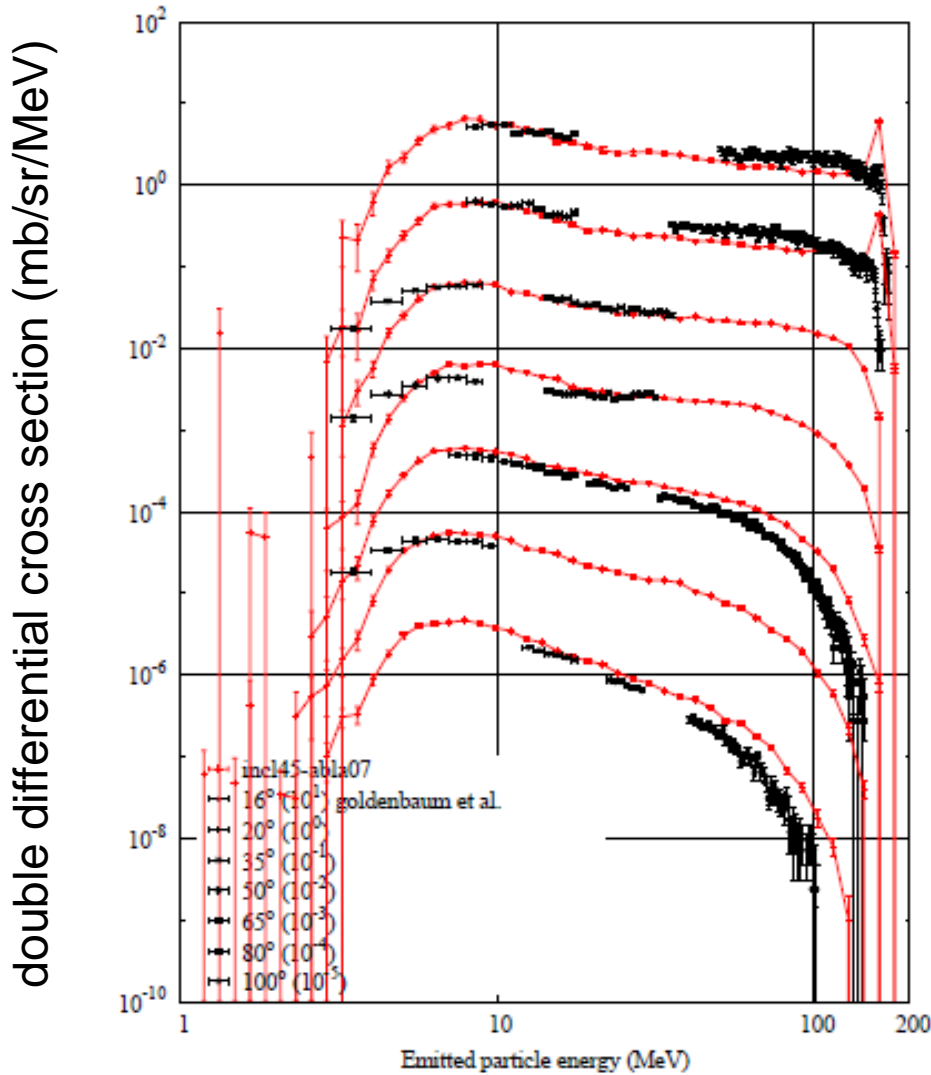
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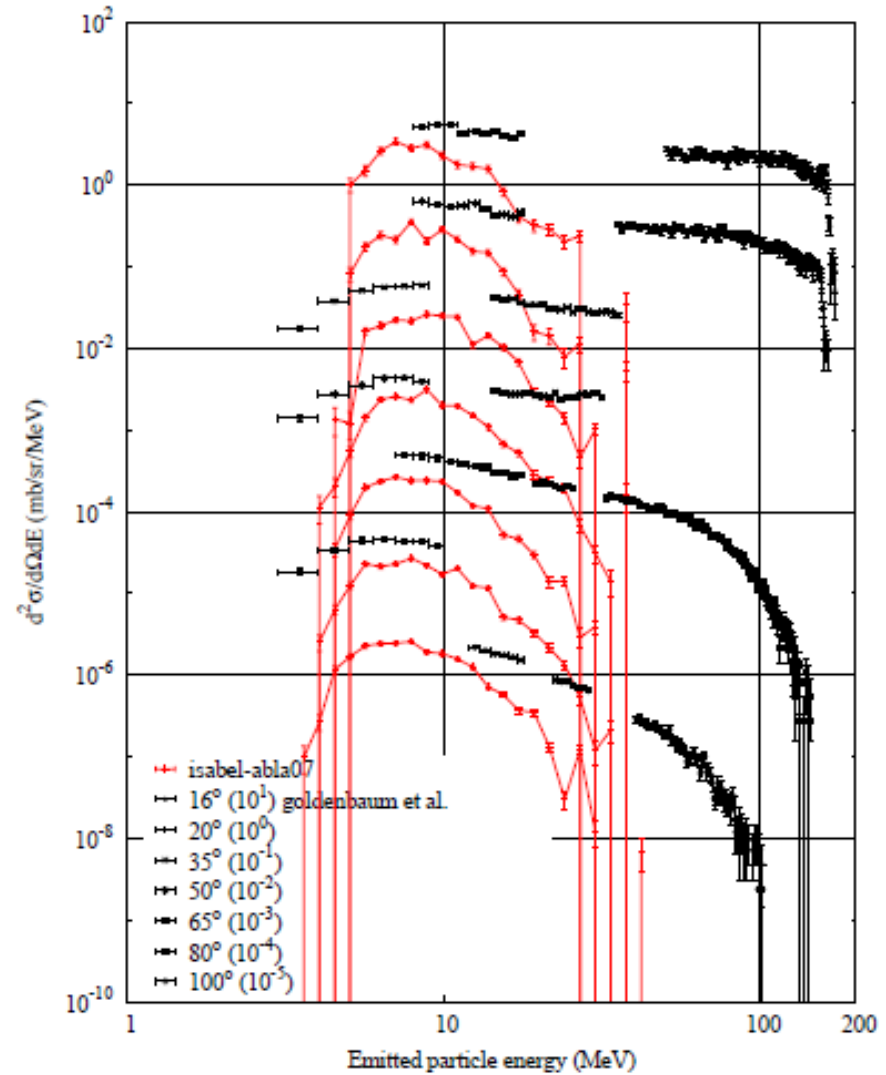
emitted-particle energy (MeV)

# p(175 MeV) + Ni – Deuteron spectrum

## INCL45-ABLA07



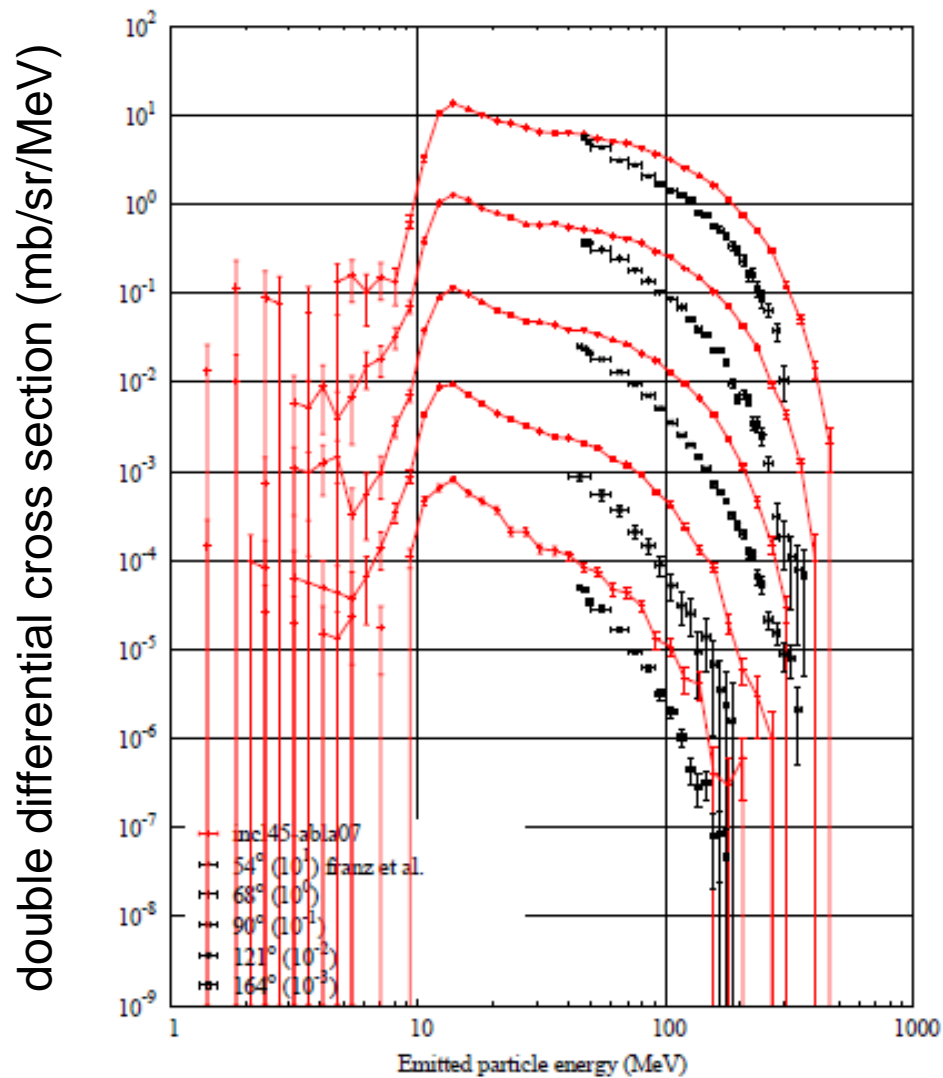
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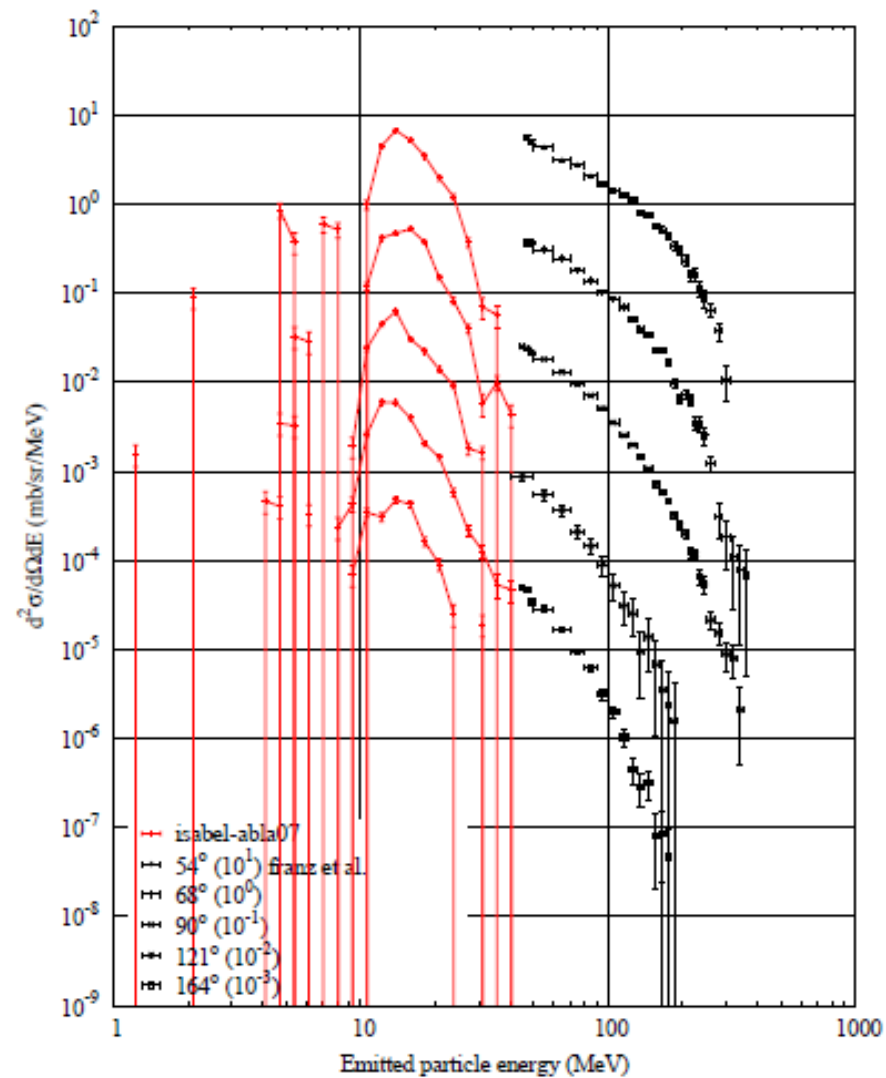
emitted-particle energy (MeV)

# n(542 MeV) + Bi – Deuteron spectrum

## INCL45-ABLA07



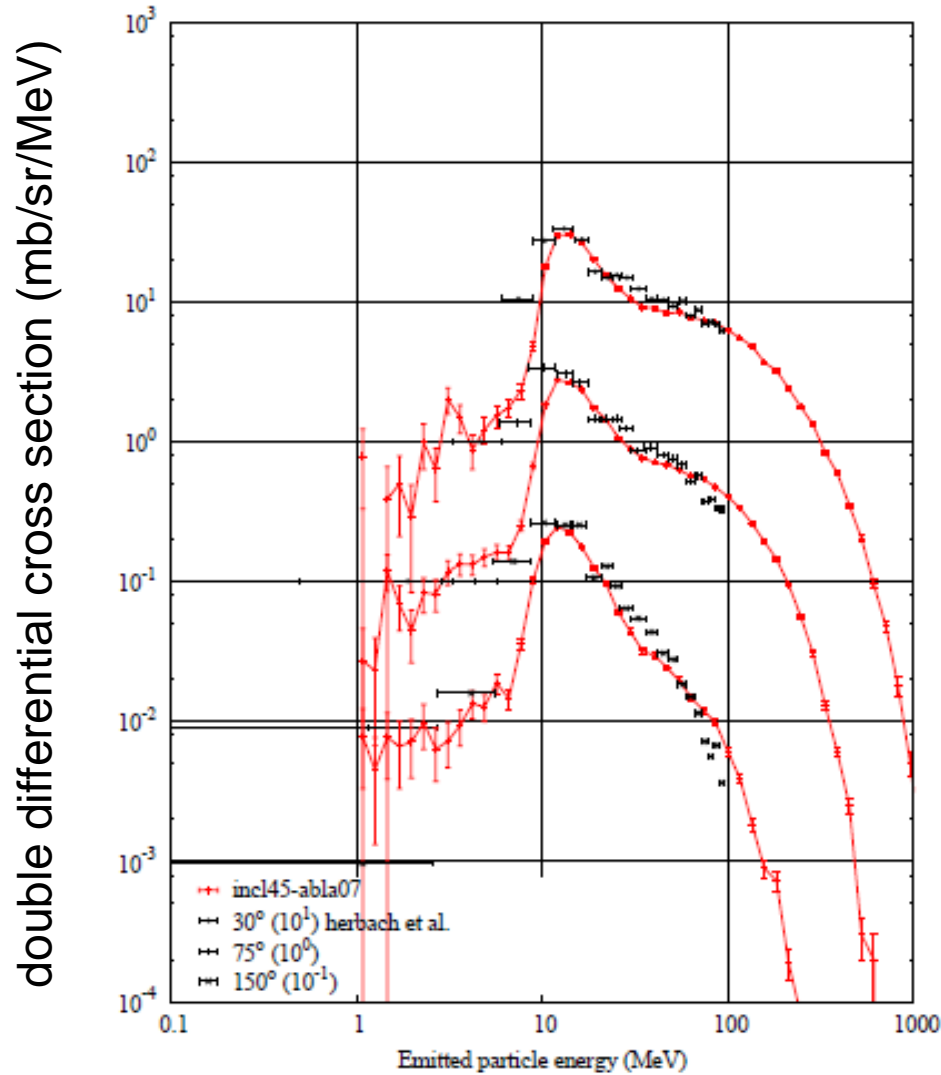
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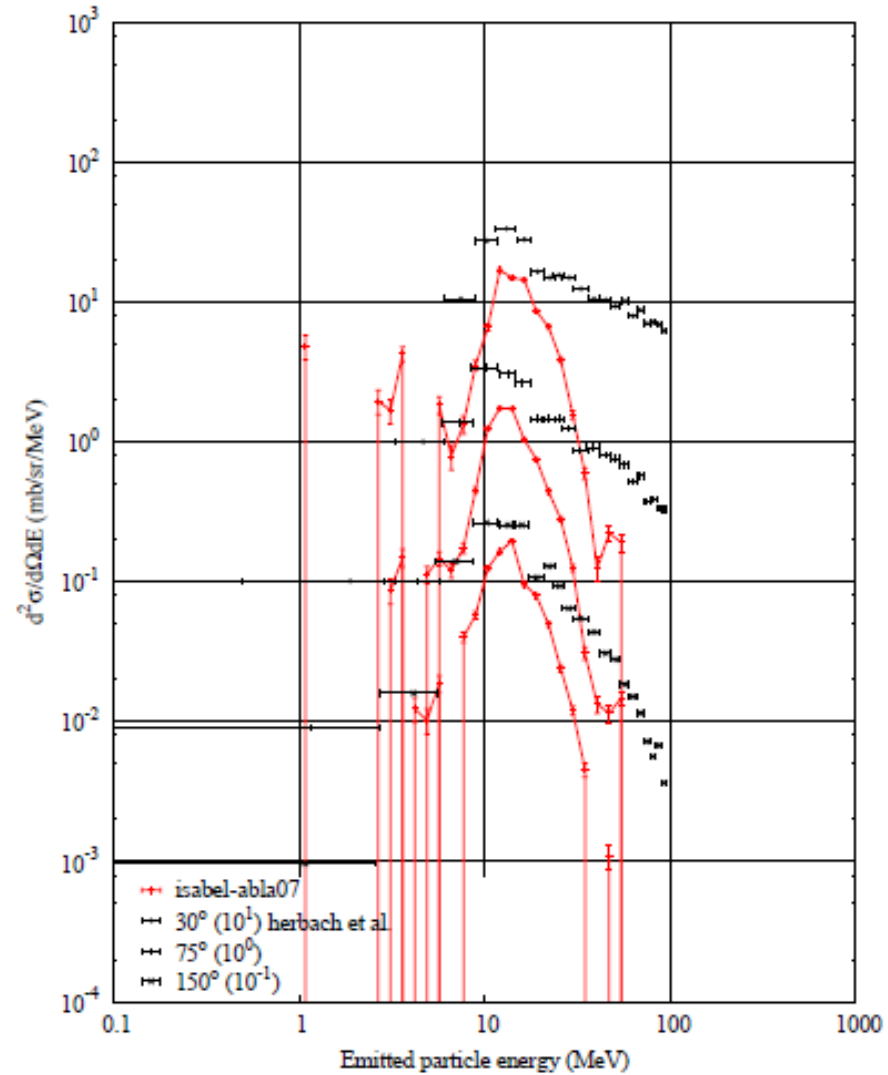
emitted-particle energy (MeV)

# p(1200 MeV) + Ta – Deuteron spectrum

## INCL45-ABLA07



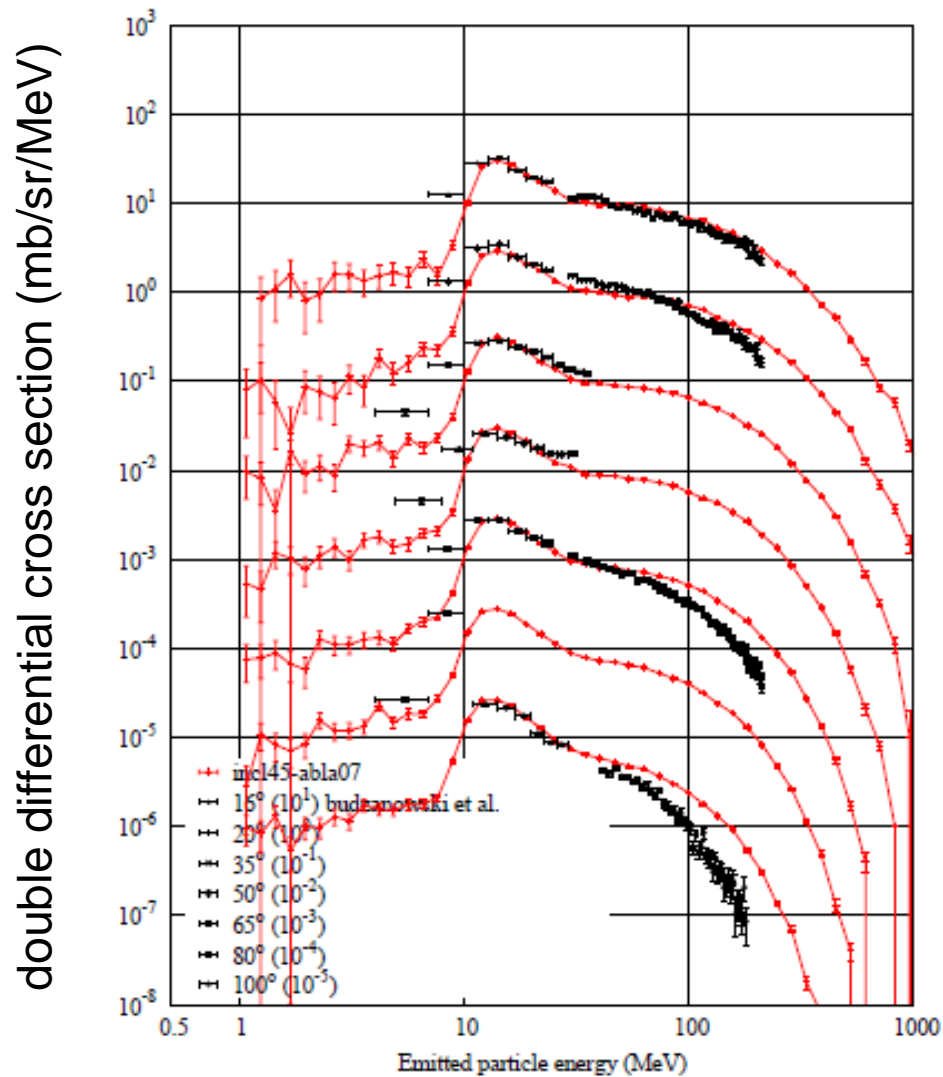
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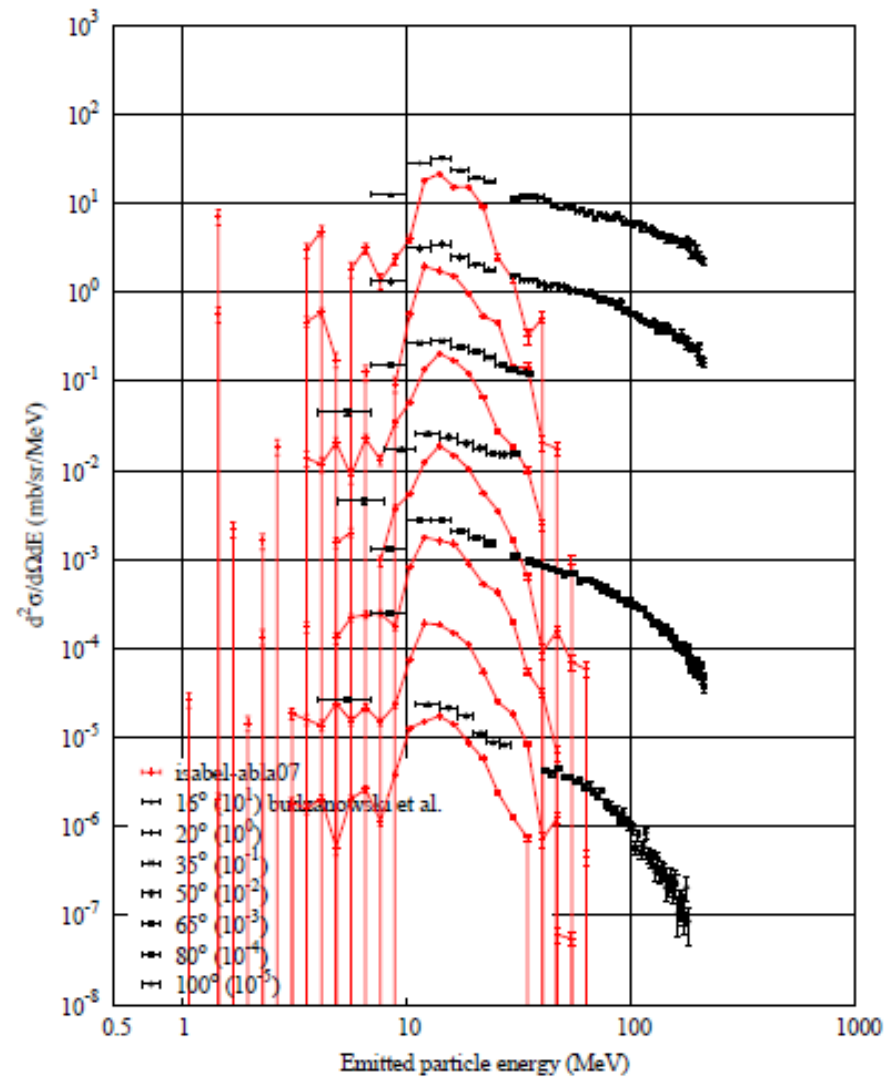
emitted-particle energy (MeV)

# p(1200 MeV) + Au – Deuteron spectrum

## INCL45-ABLA07



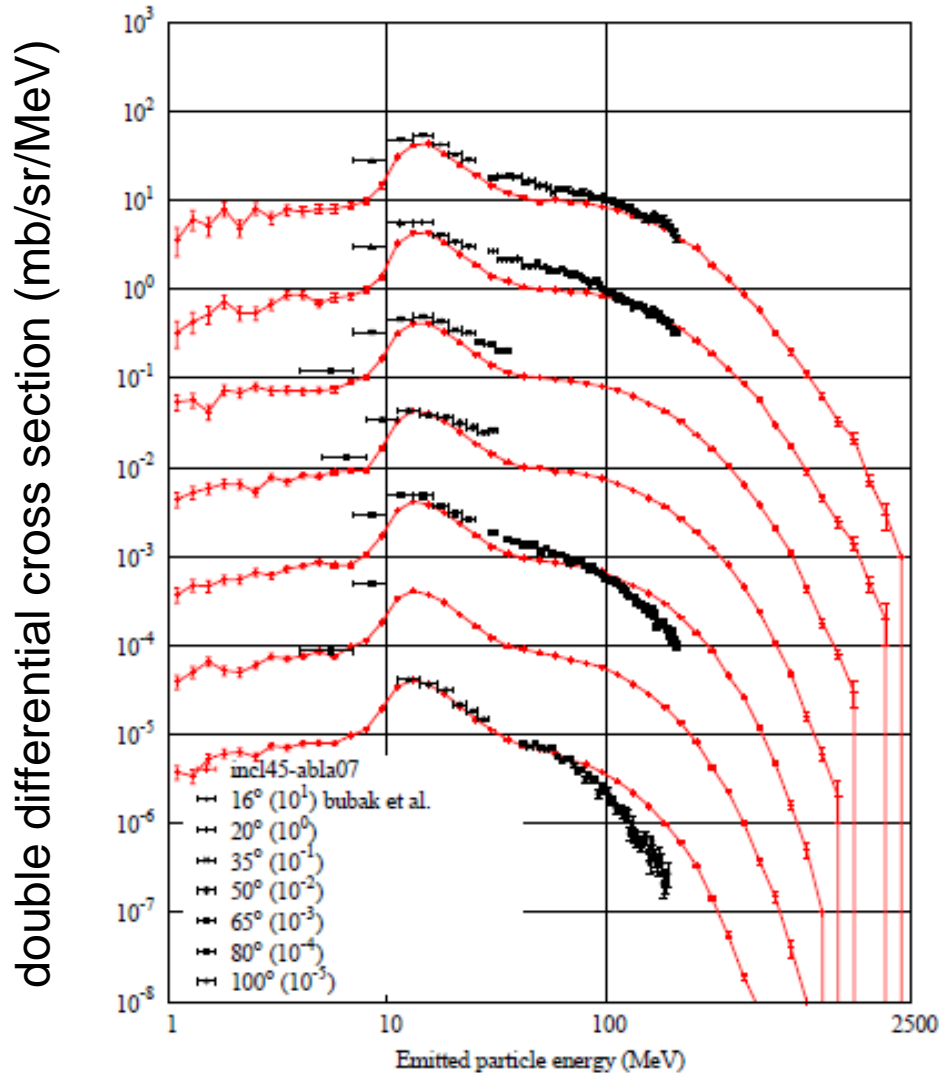
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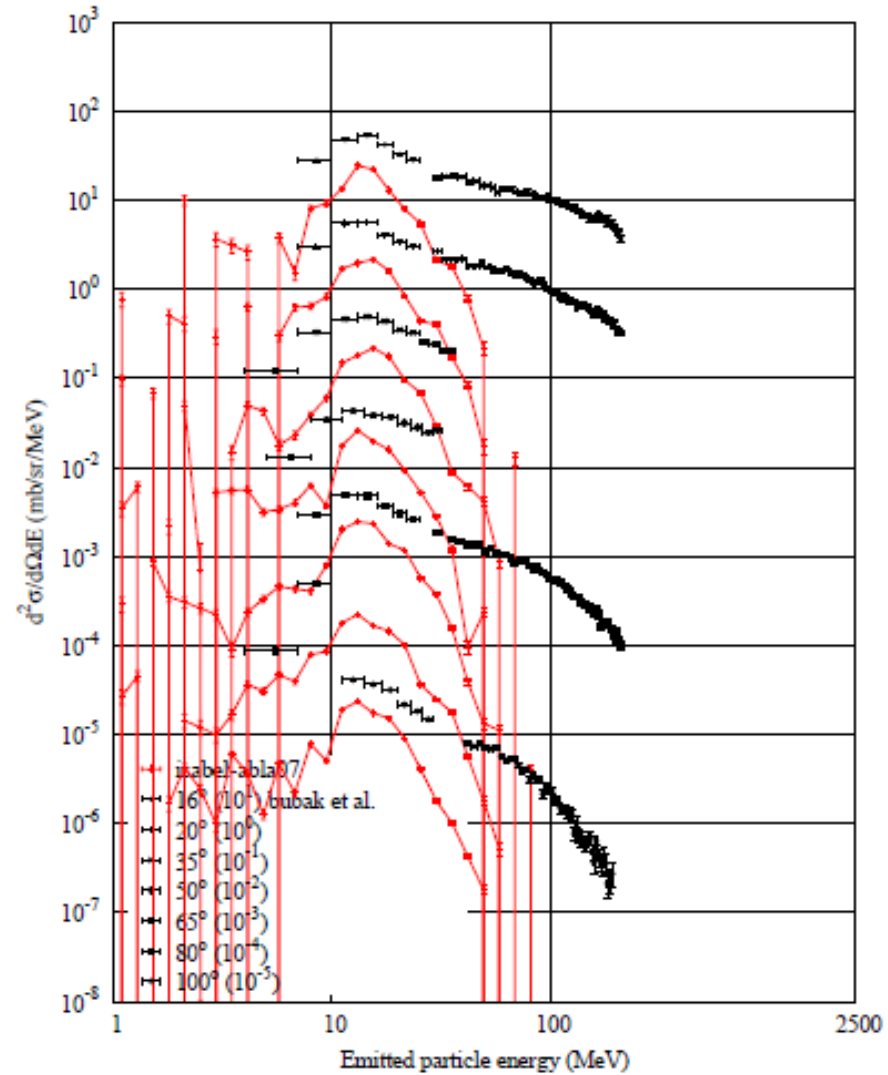
emitted-particle energy (MeV)

# p(2500 MeV) + Au – Deuteron spectrum

## INCL45-ABLA07



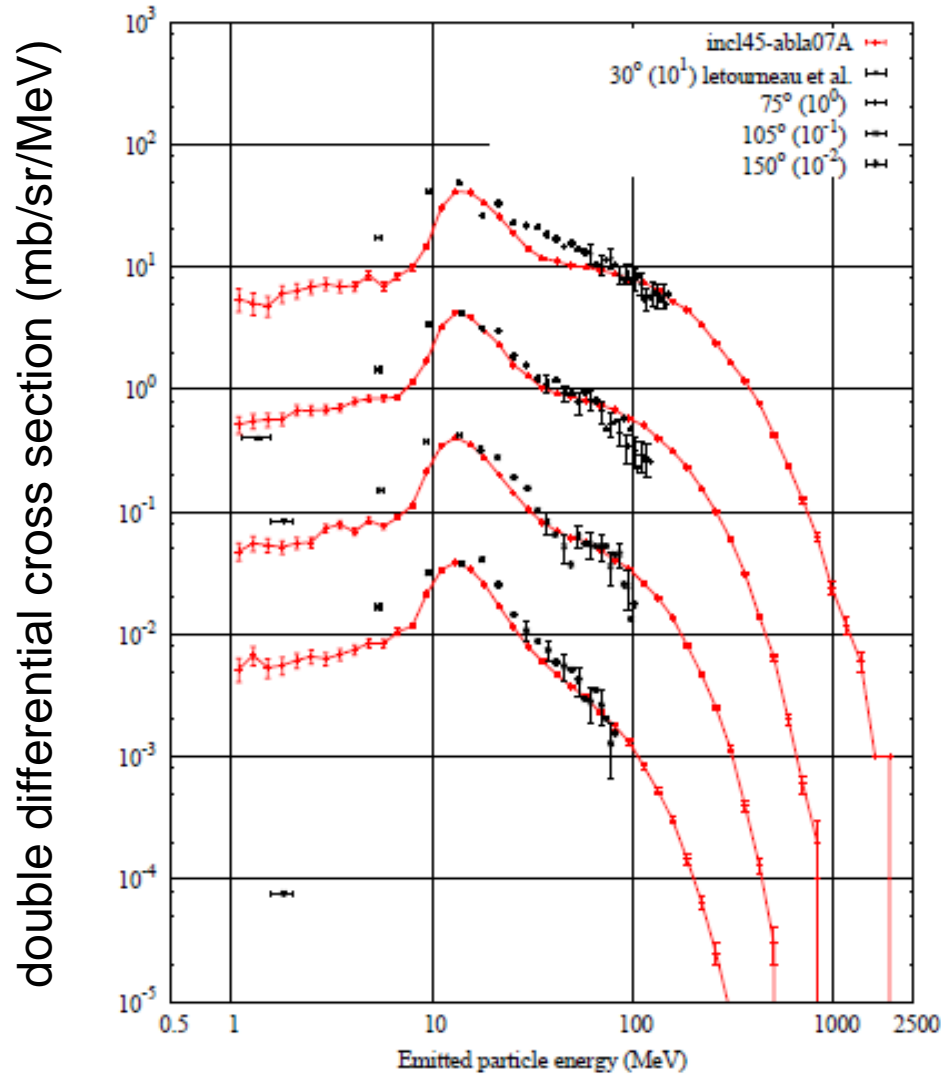
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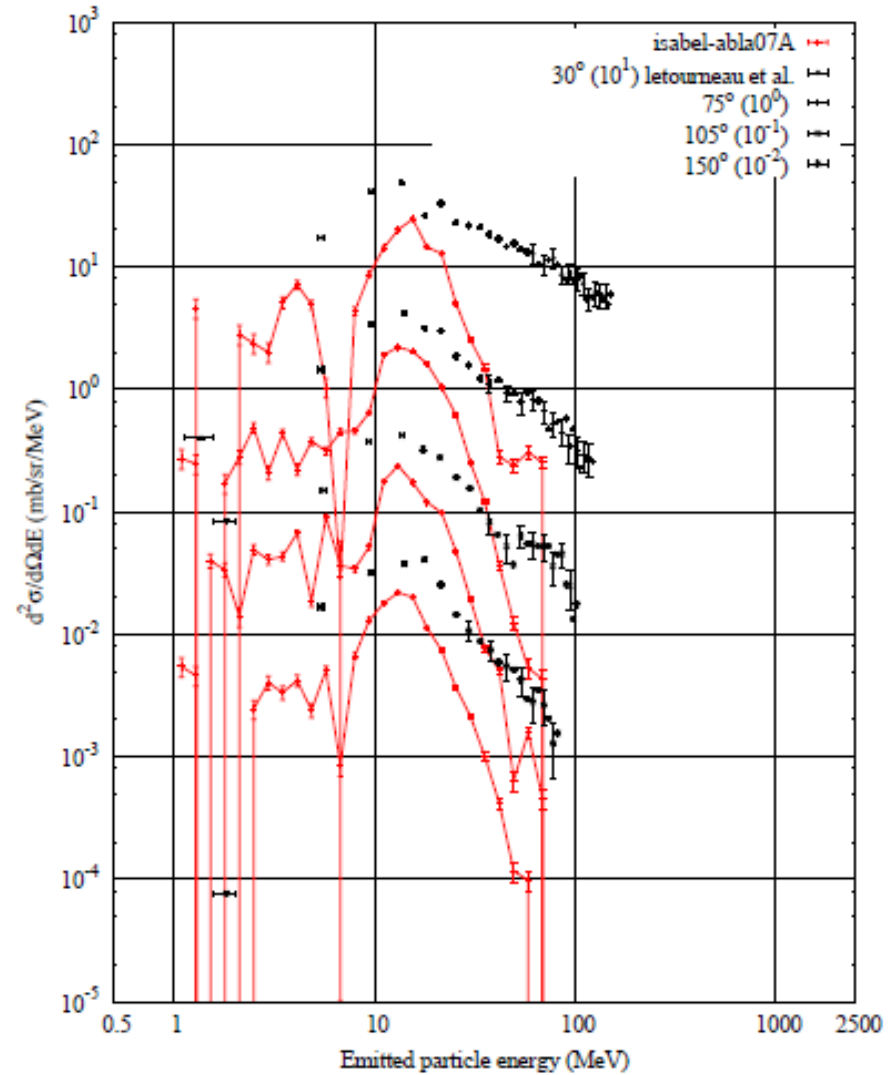
emitted-particle energy (MeV)

# p(2500 MeV) + Au – Deuteron spectrum

## INCL45-ABLA07



## ISABEL-ABLA07



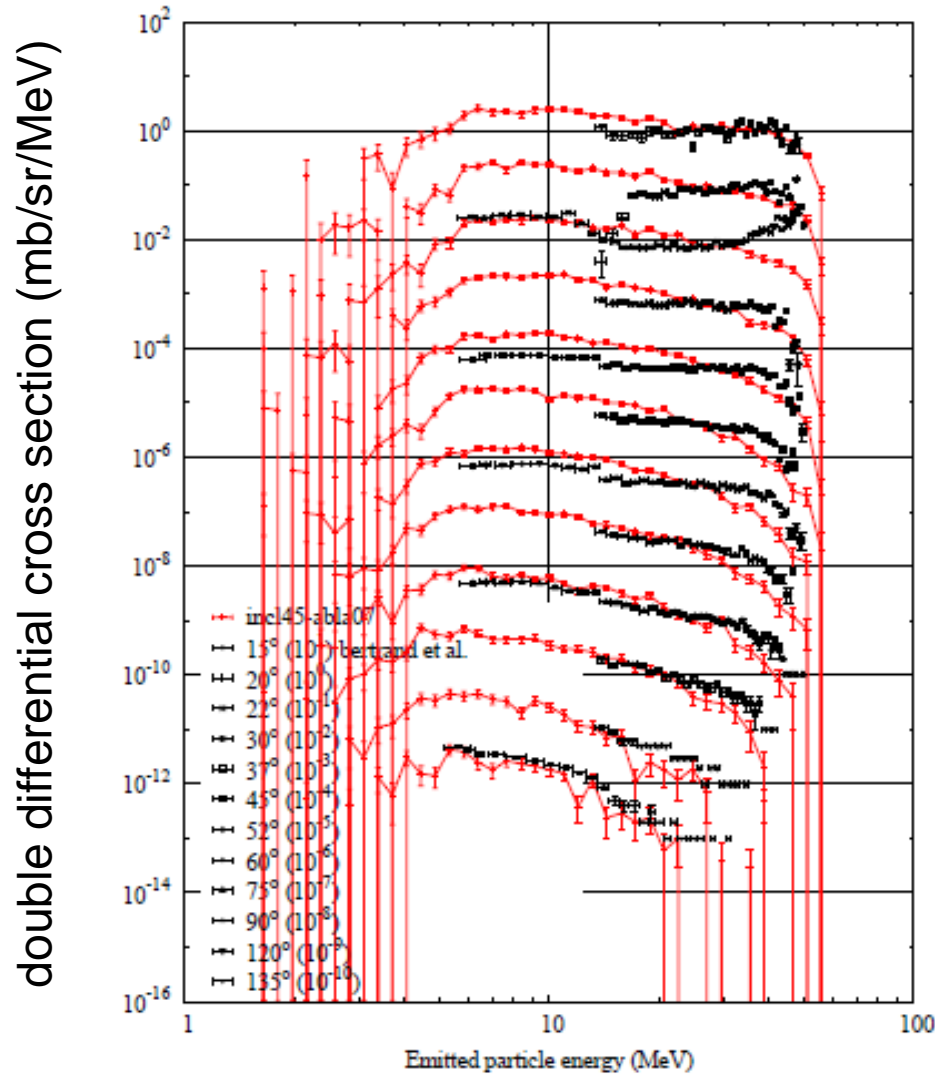
emitted-particle energy (MeV)

# Tritium spectra

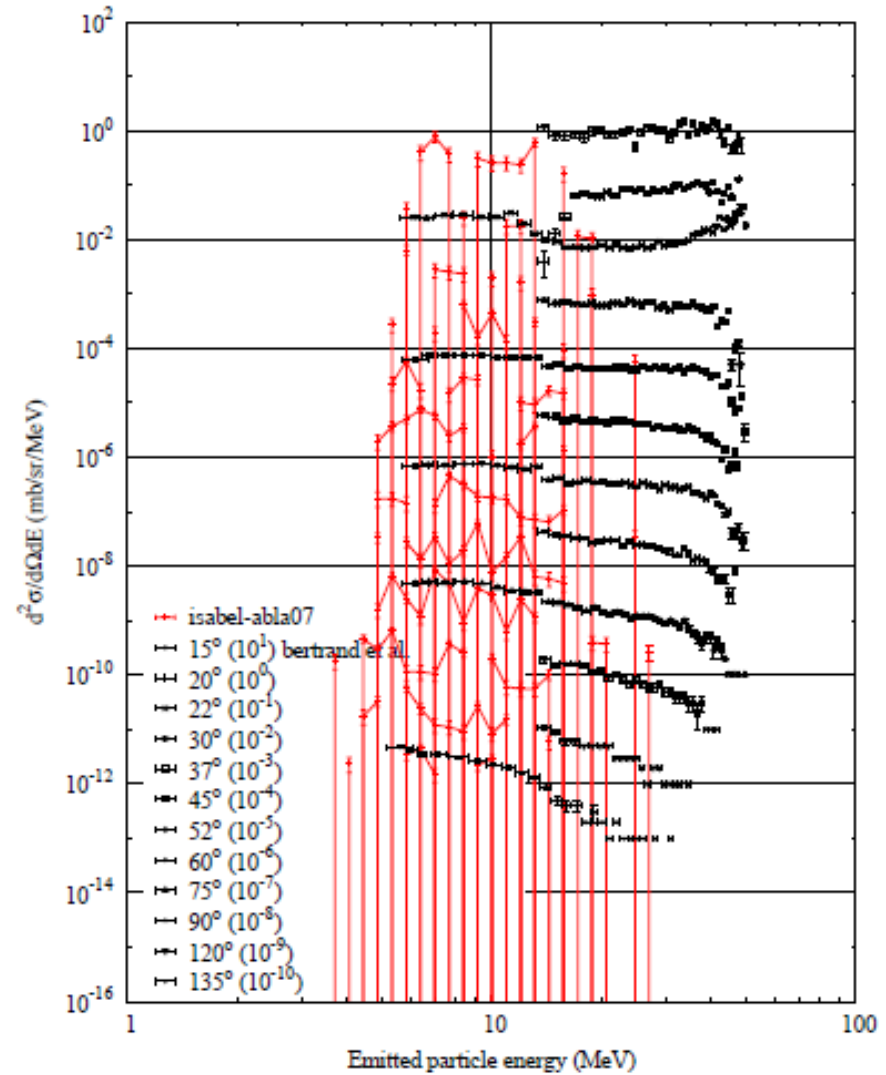


# p(62 MeV) + $^{56}\text{Fe}$ – Tritium spectrum

## INCL45-ABLA07



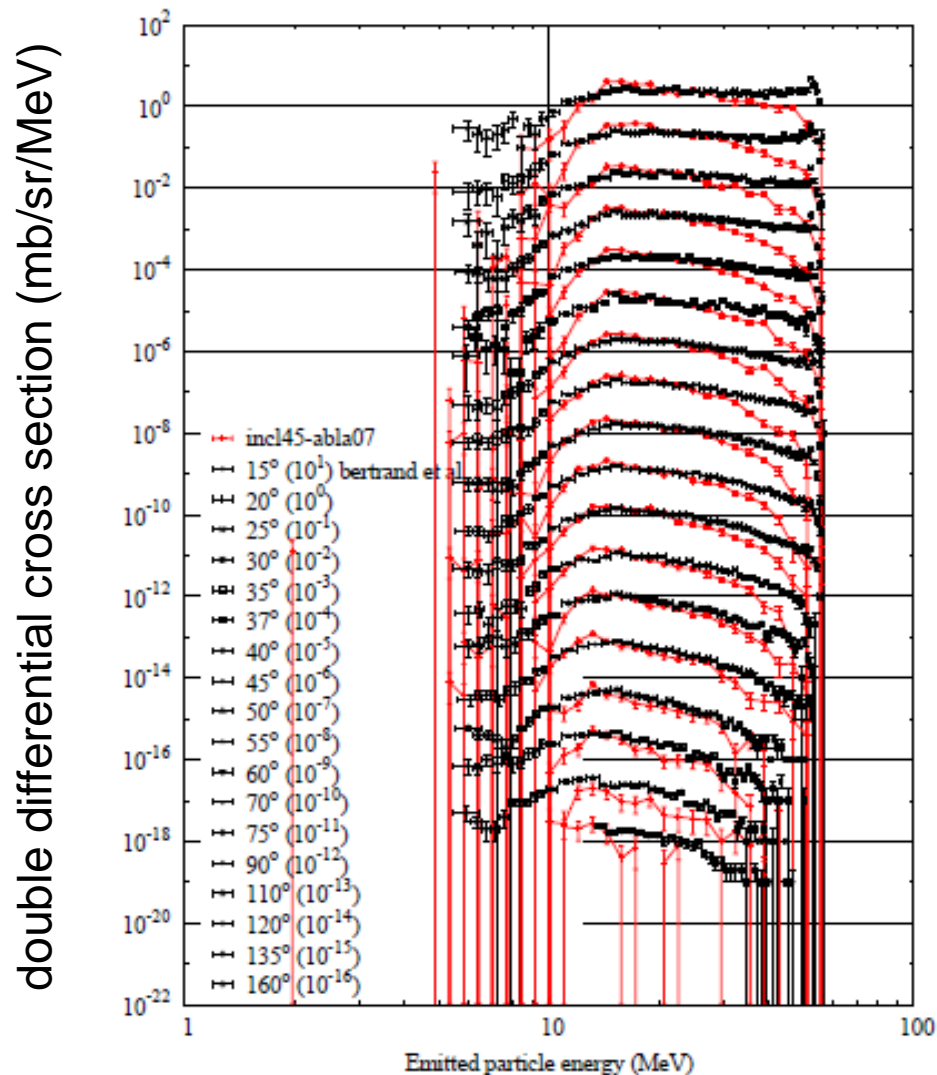
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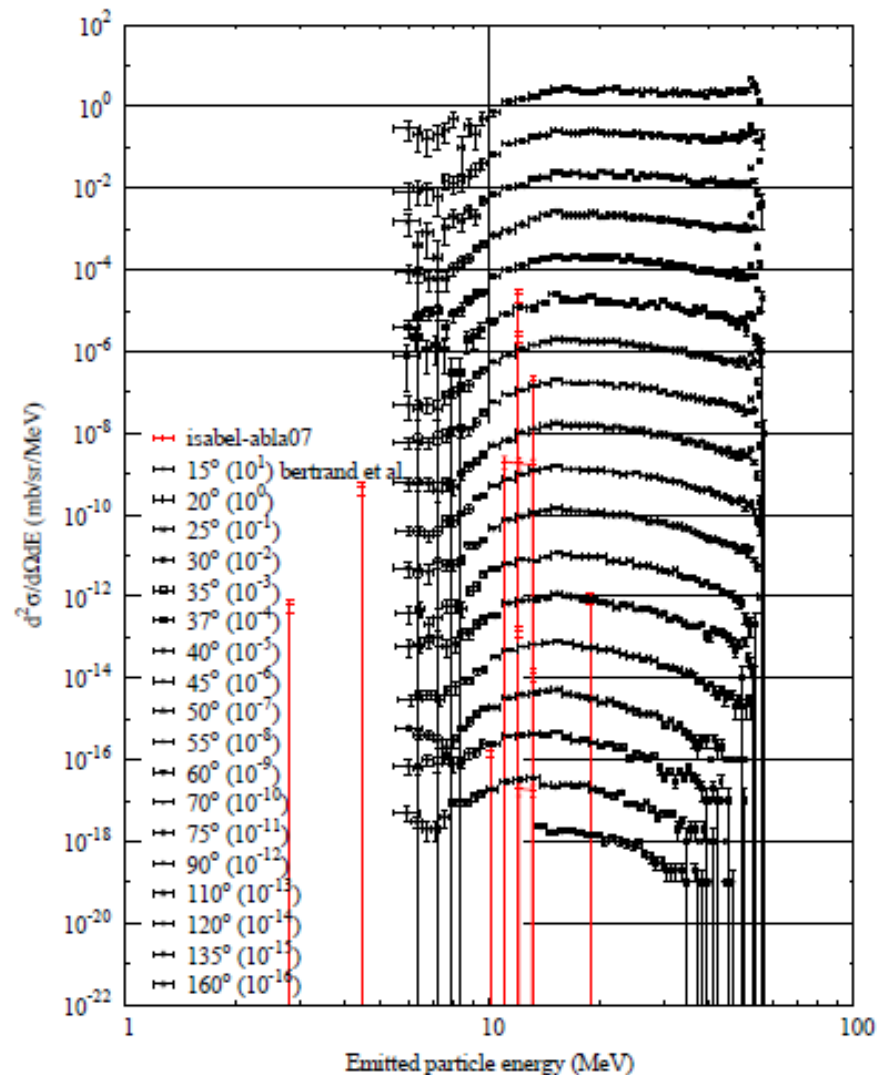
emitted-particle energy (MeV)

# p(62 MeV) + Bi – Tritium spectrum

## INCL45-ABLA07



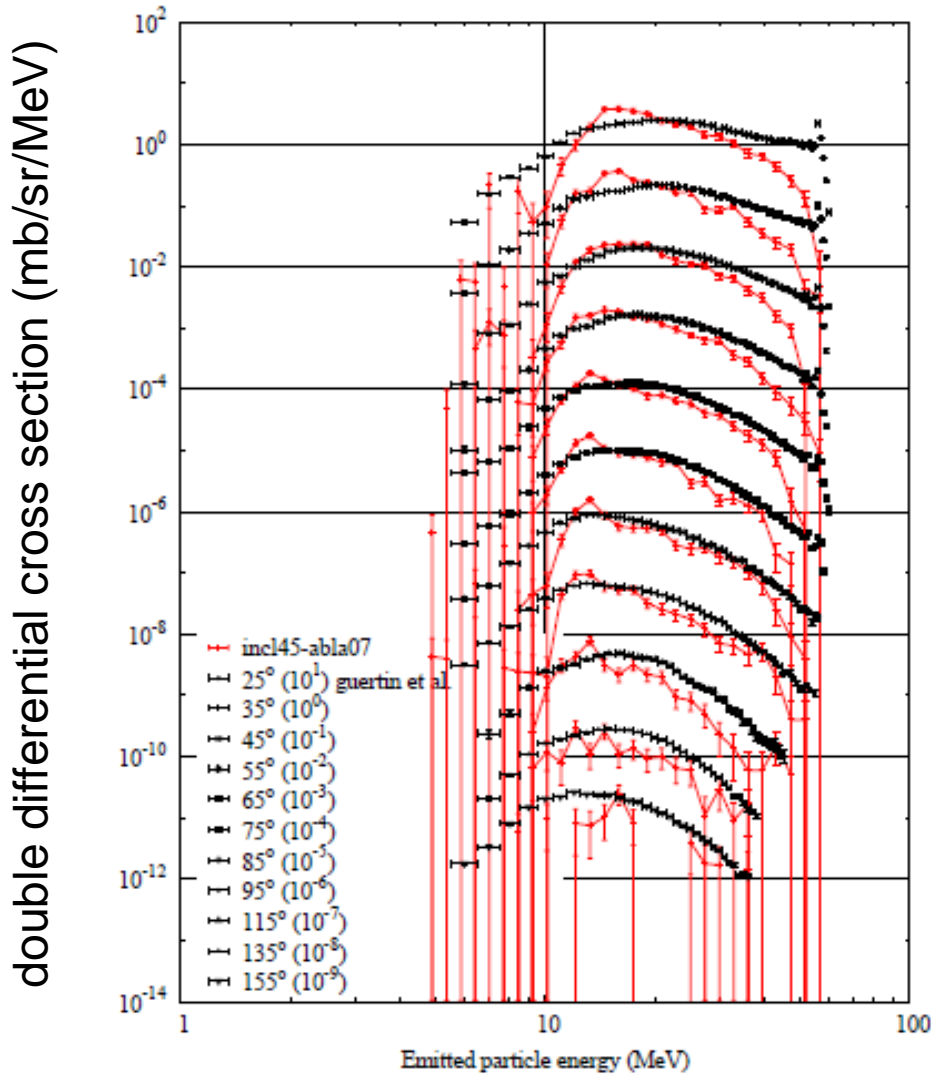
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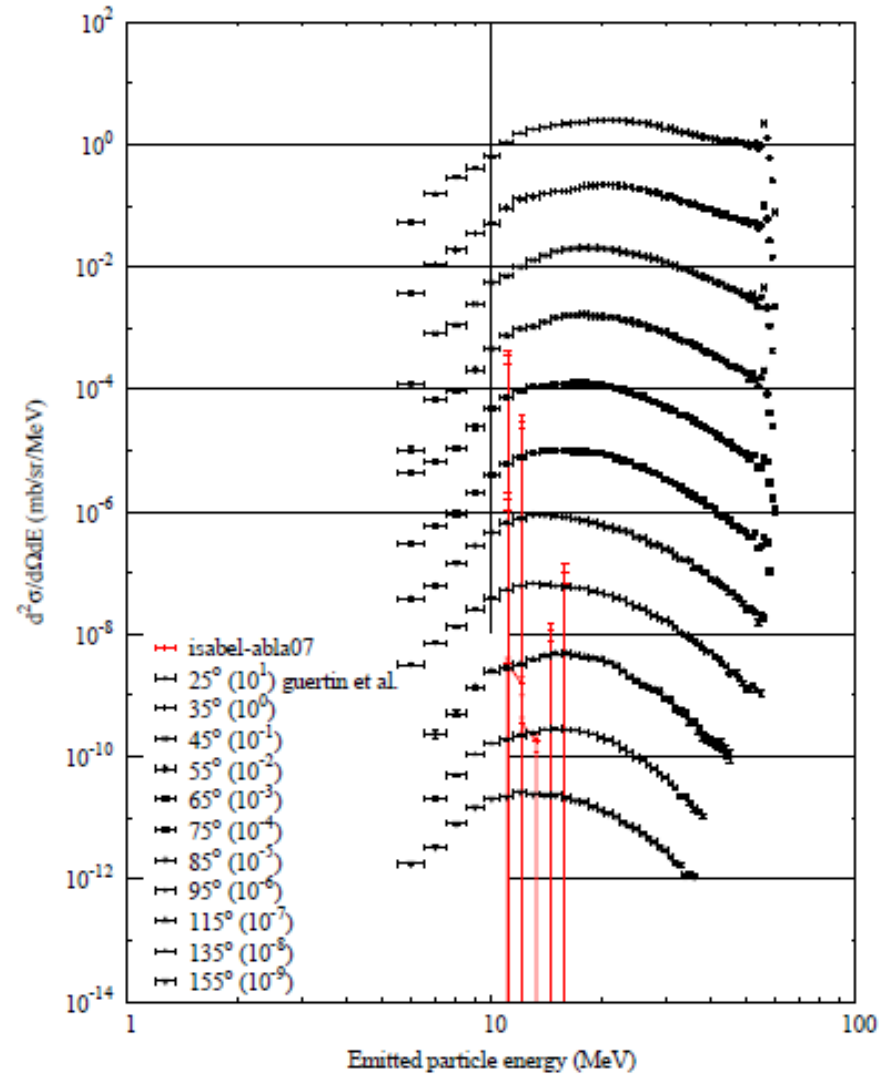
emitted-particle energy (MeV)

# p(63 MeV) + $^{208}\text{Pb}$ – Tritium spectrum

## INCL45-ABLA07



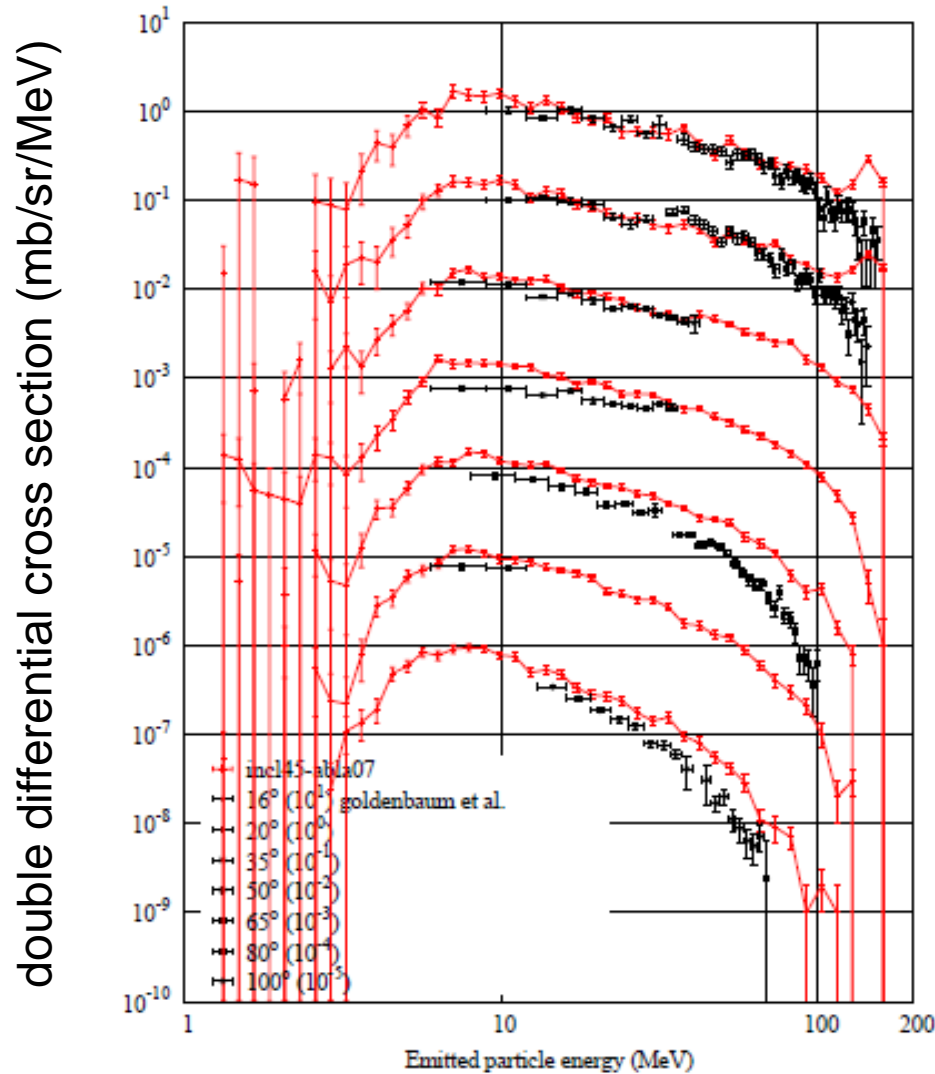
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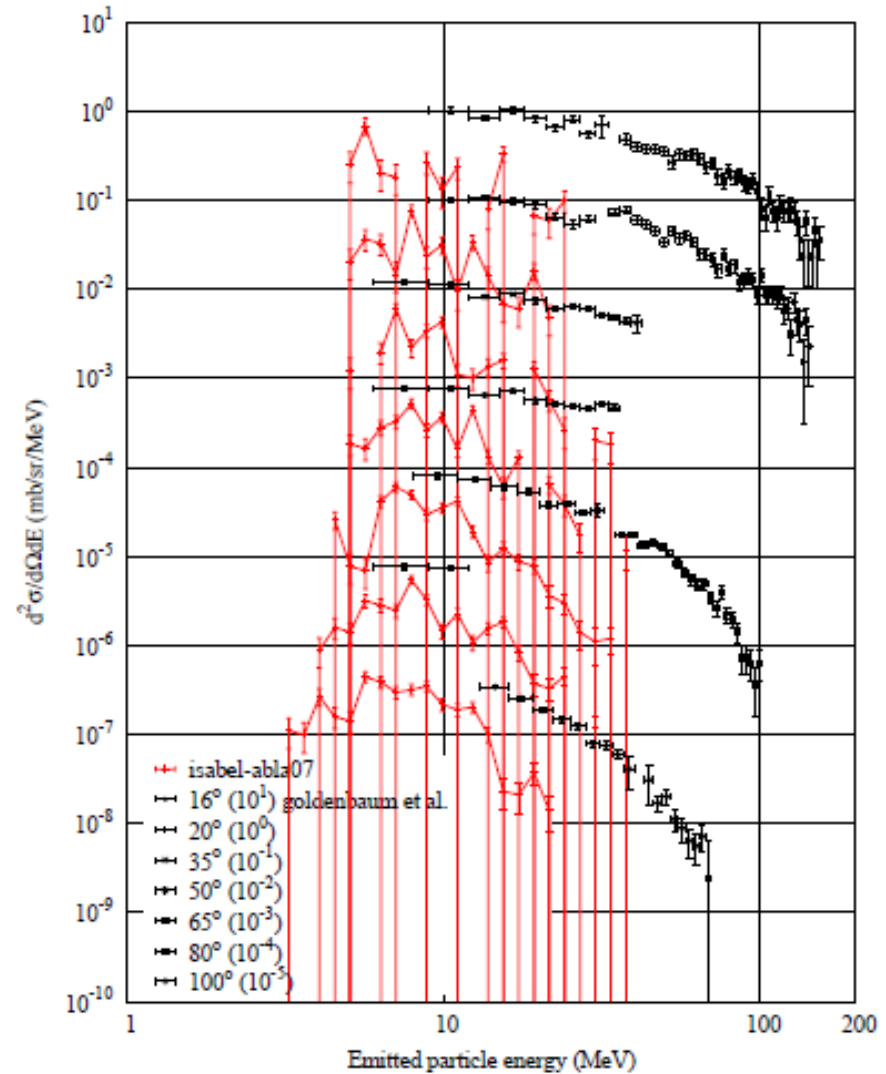
emitted-particle energy (MeV)

# p(175 MeV) + Ni – Tritium spectrum

## INCL45-ABLA07



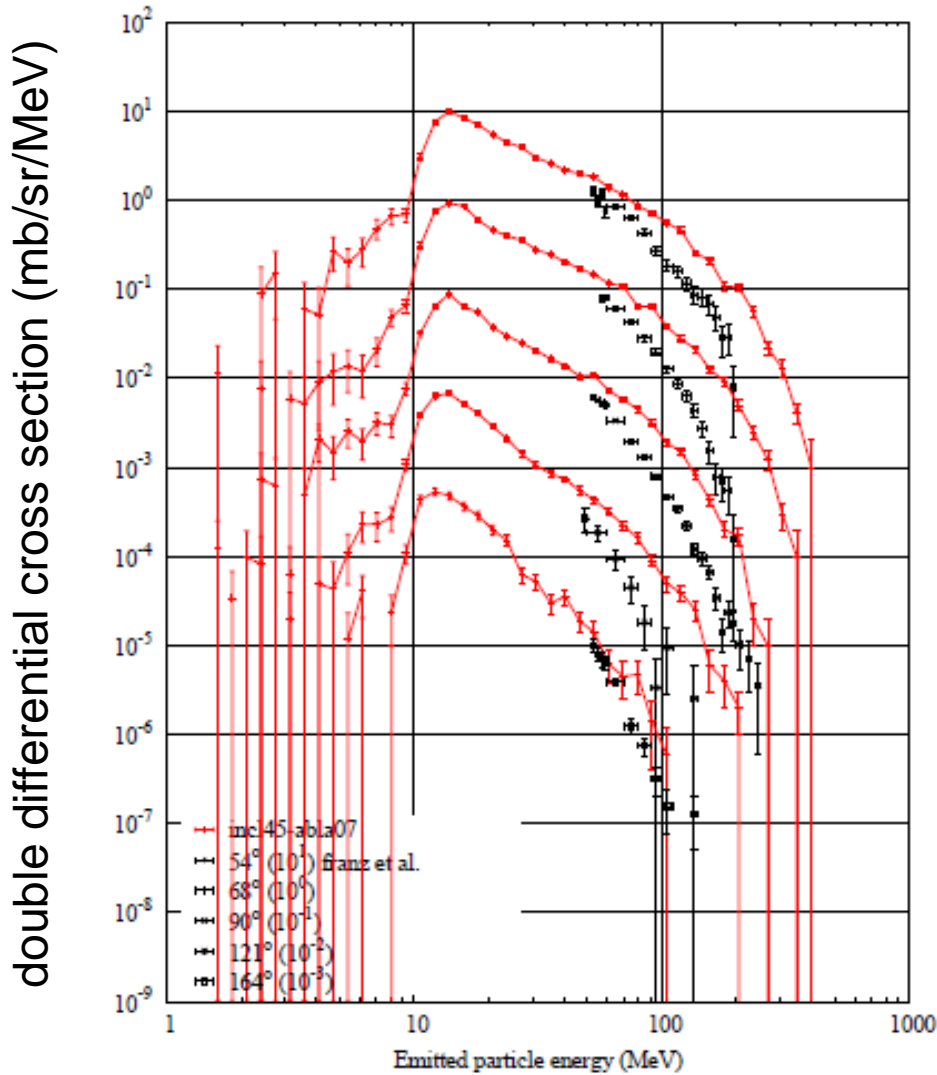
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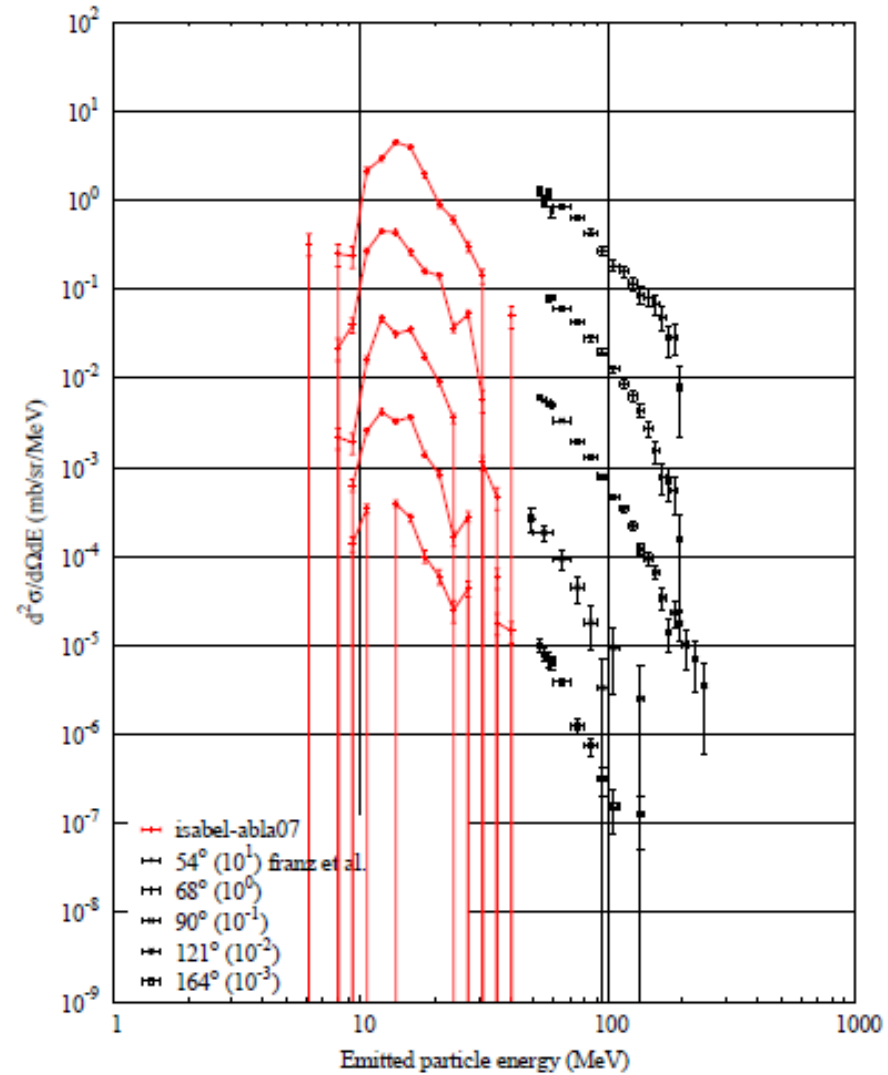
emitted-particle energy (MeV)

# n(542 MeV) + Bi – Tritium spectrum

## INCL45-ABLA07



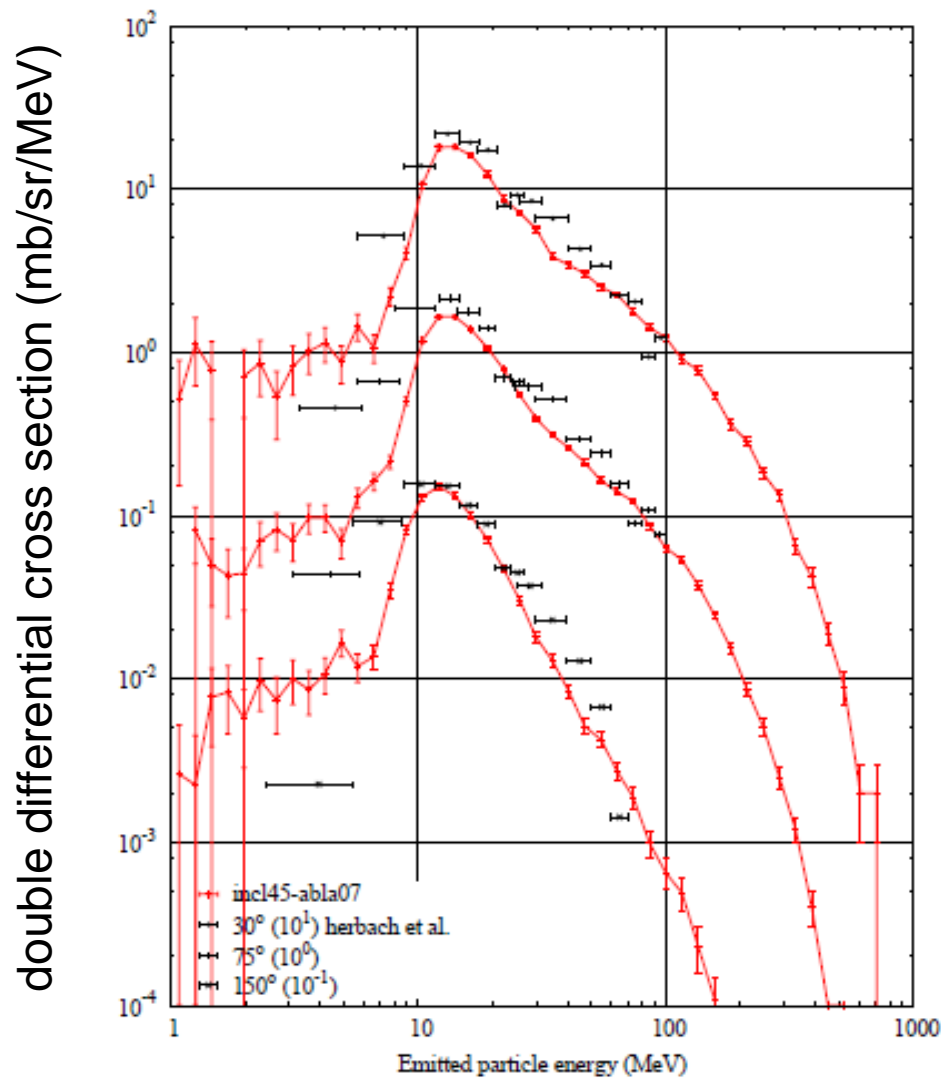
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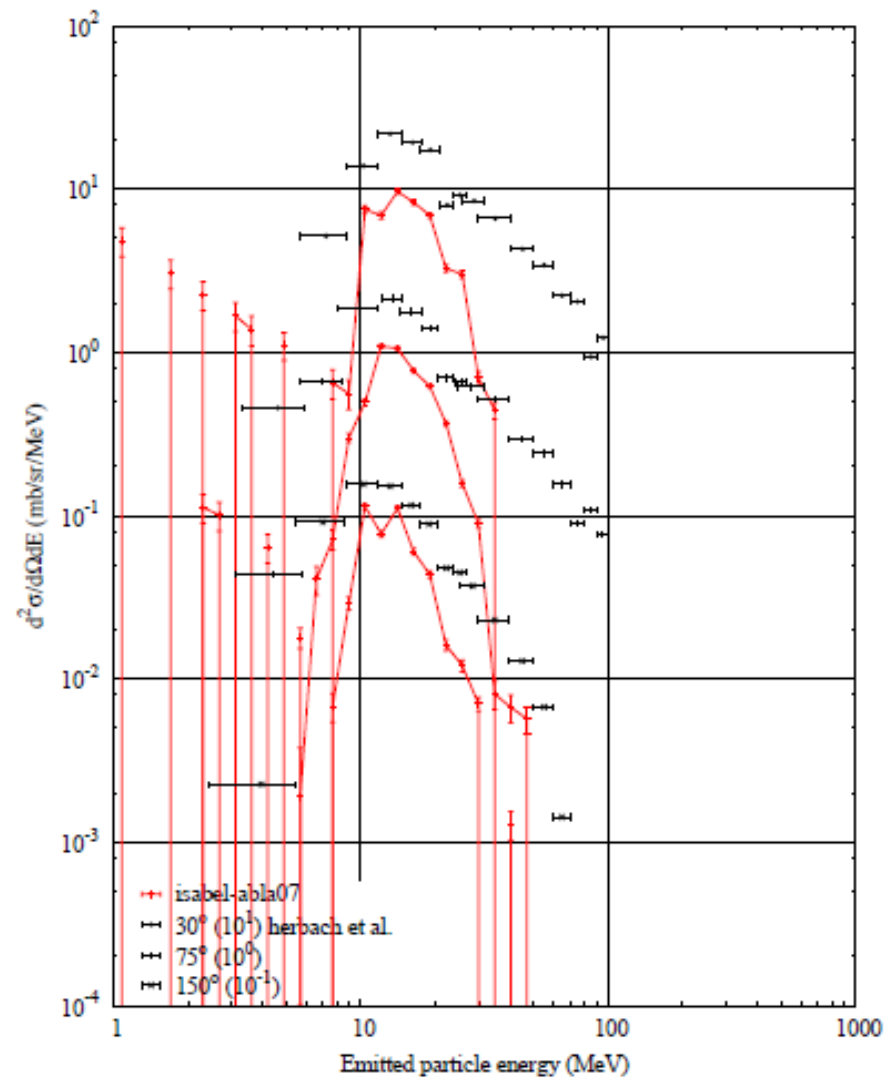
emitted-particle energy (MeV)

# p(1200 MeV) + Ta – Tritium spectrum

## INCL45-ABLA07



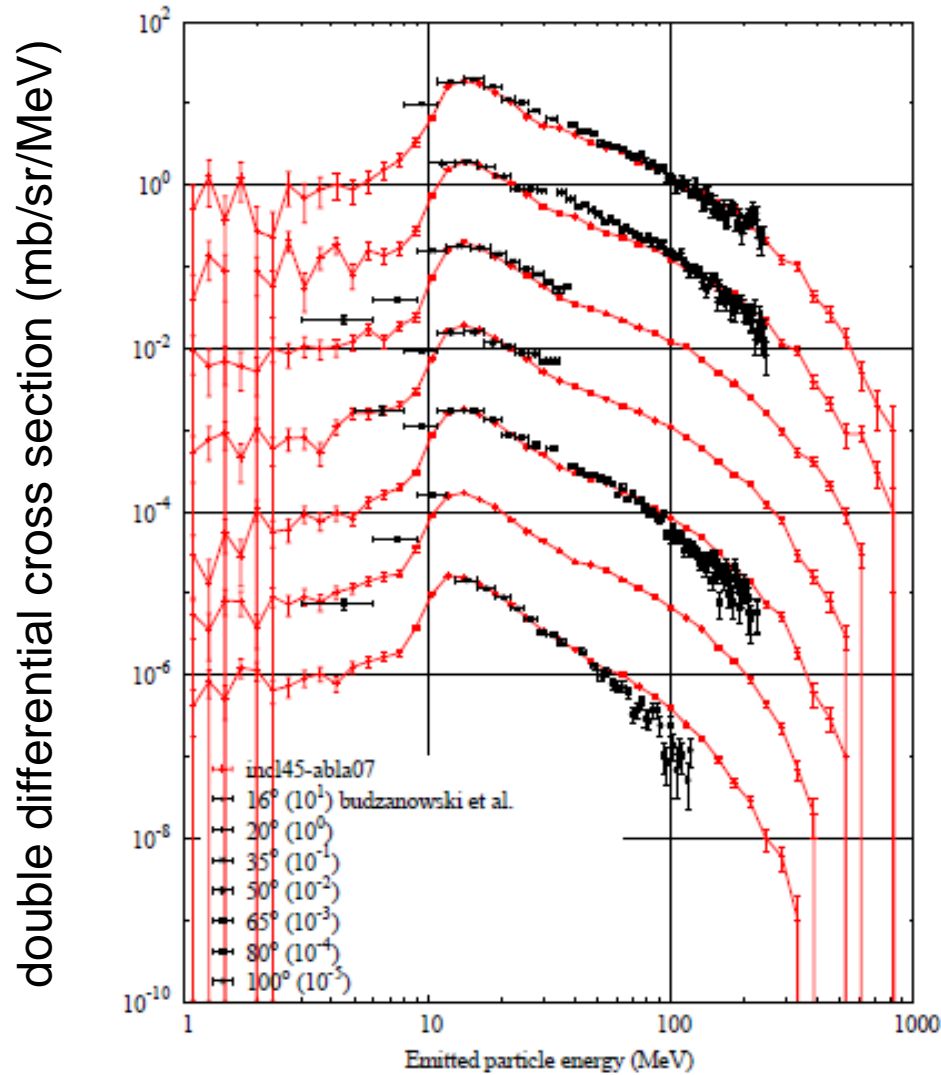
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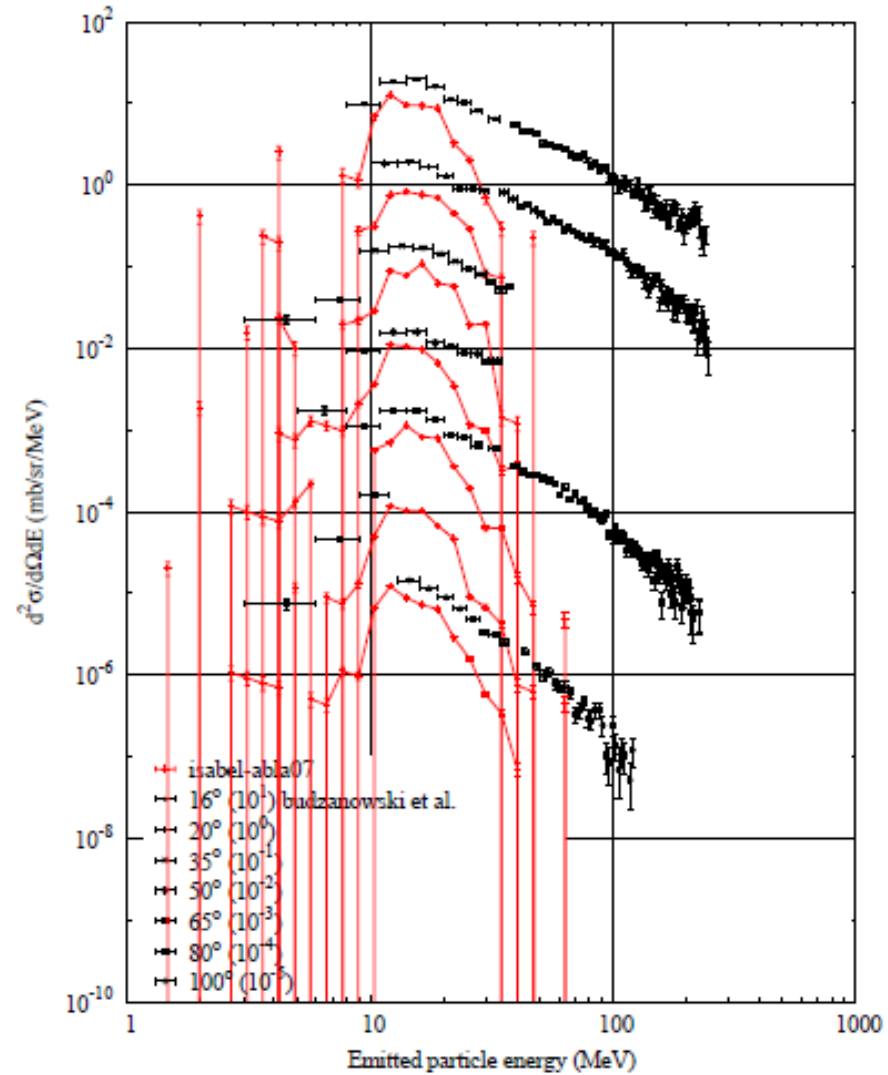
emitted-particle energy (MeV)

# p(1200 MeV) + Au – Tritium spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

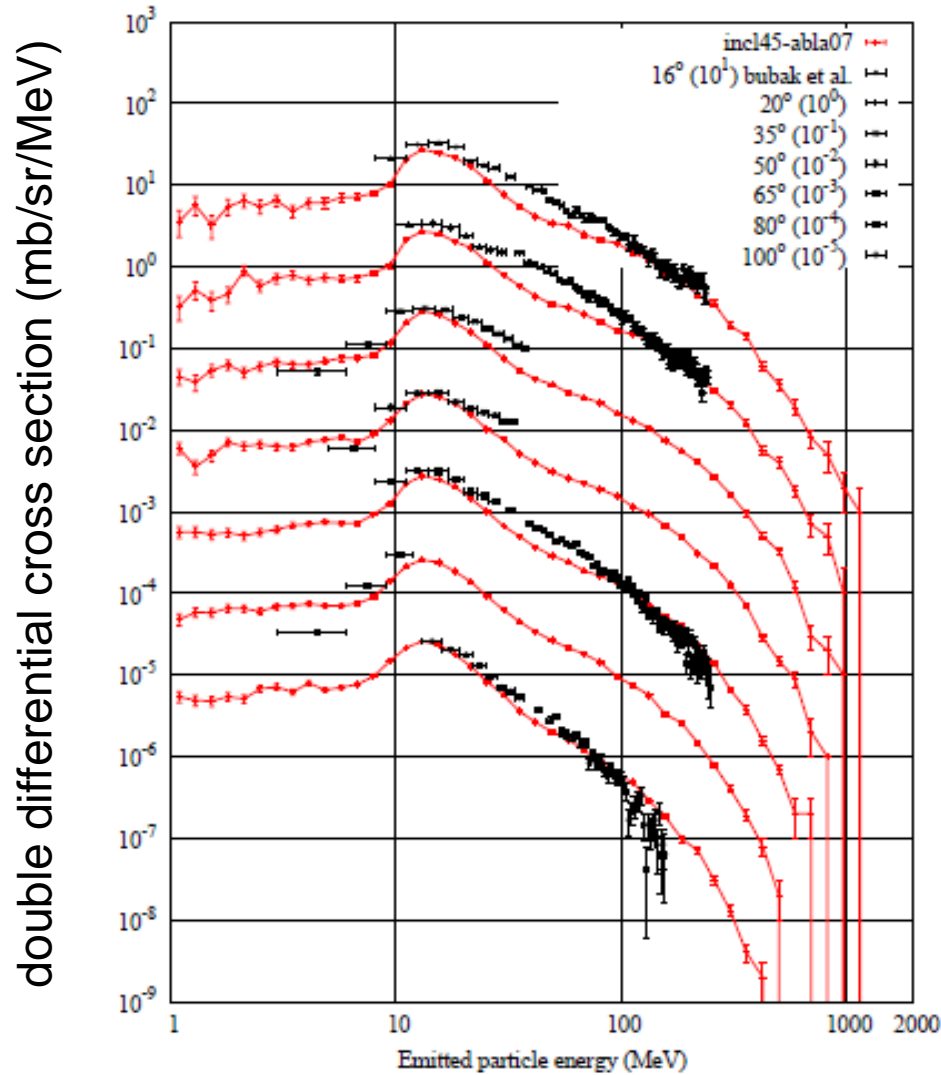


emitted-particle energy (MeV)

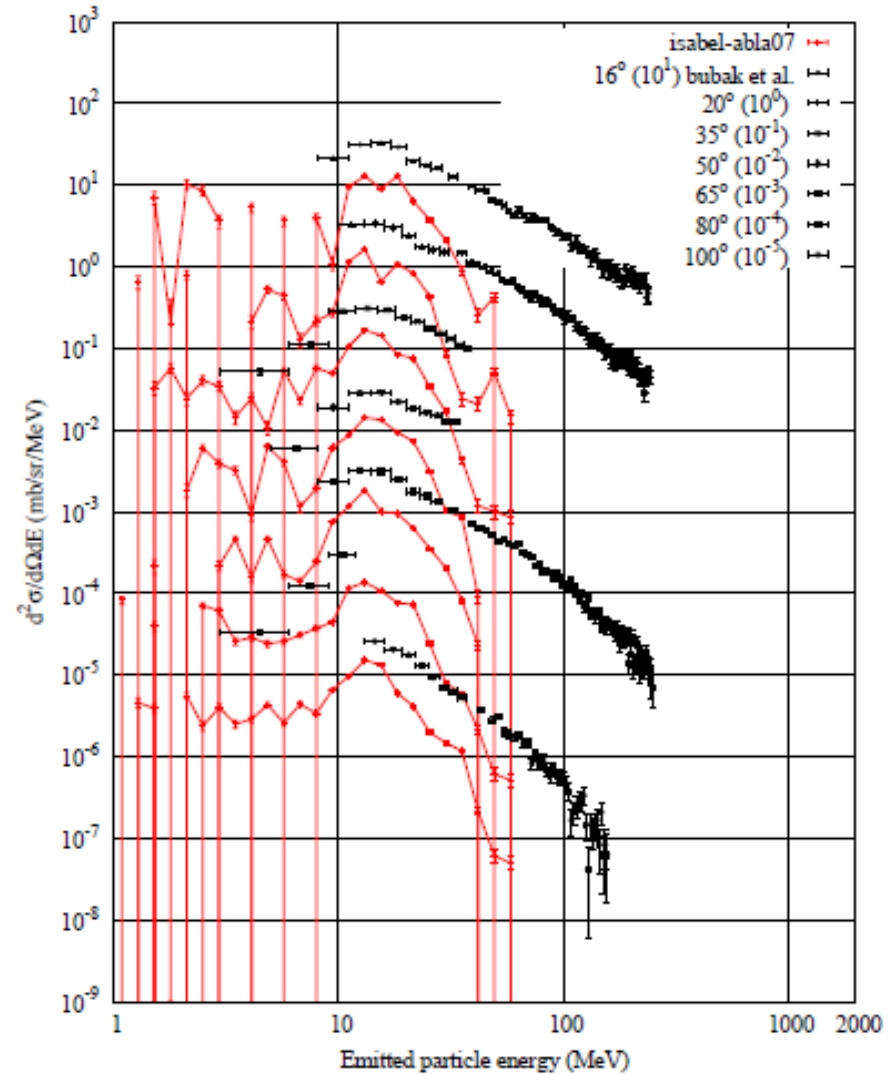


# p(2500 MeV) + Au – Tritium spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

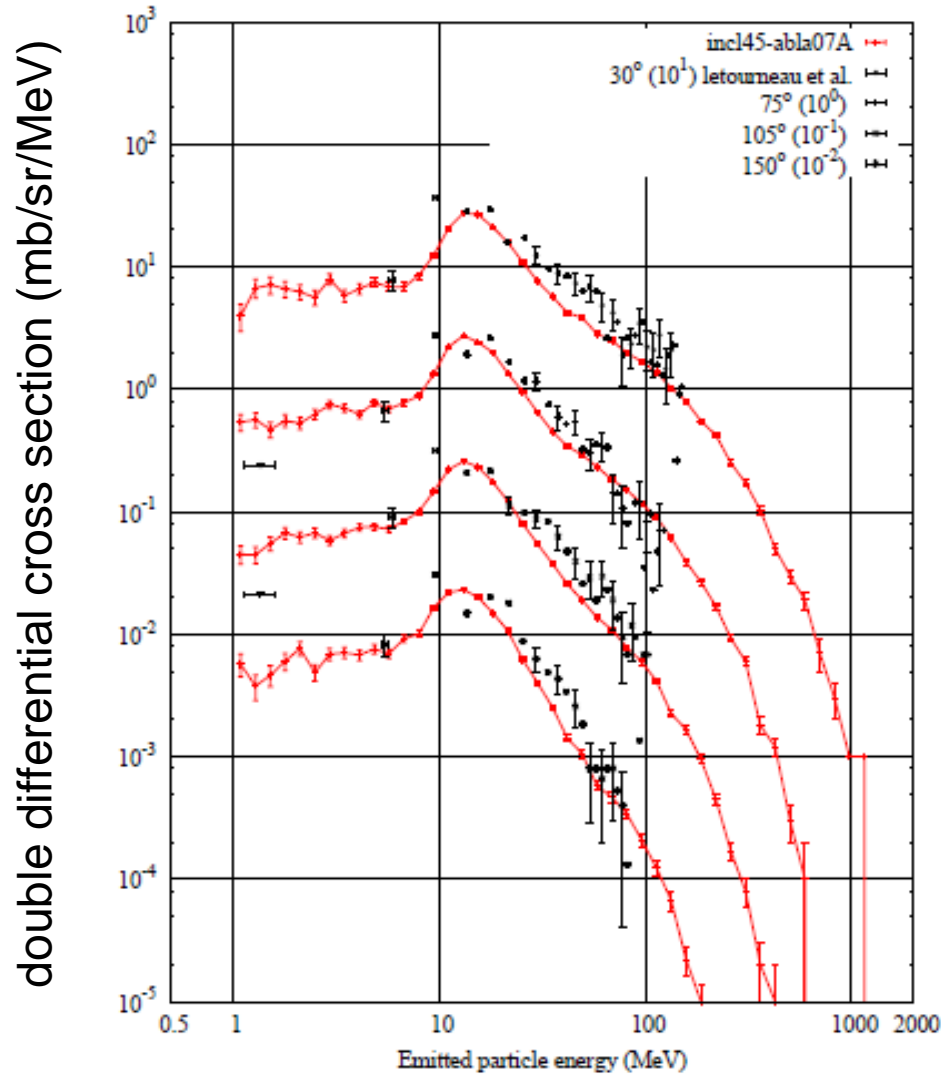


emitted-particle energy (MeV)

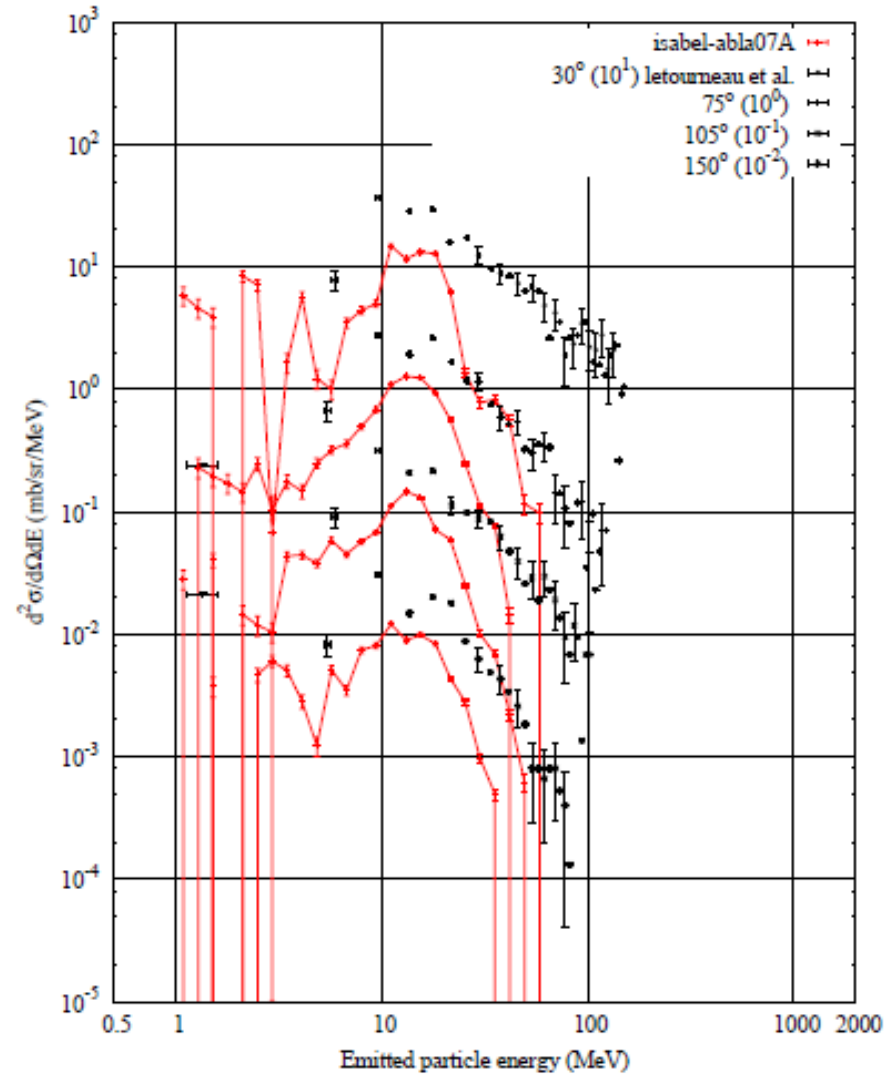


# p(2500 MeV) + Au – Tritium spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

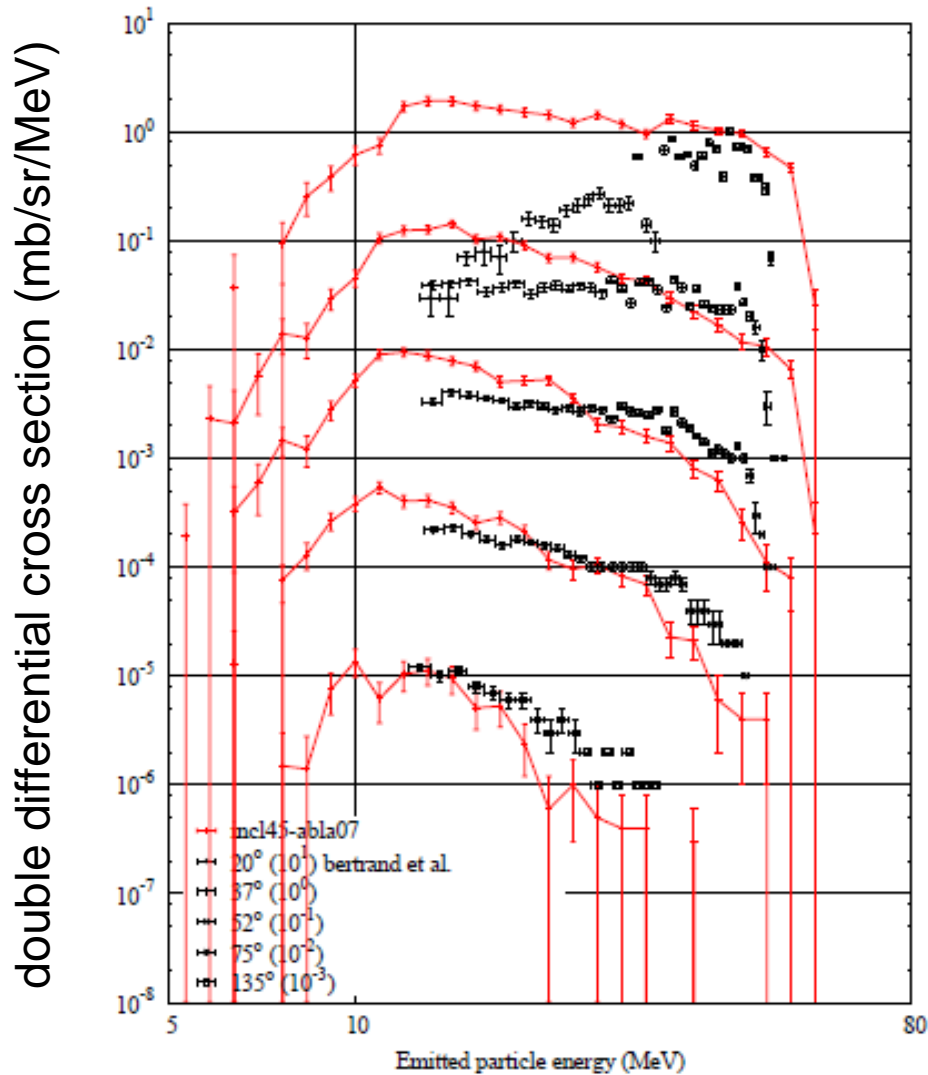


emitted-particle energy (MeV)

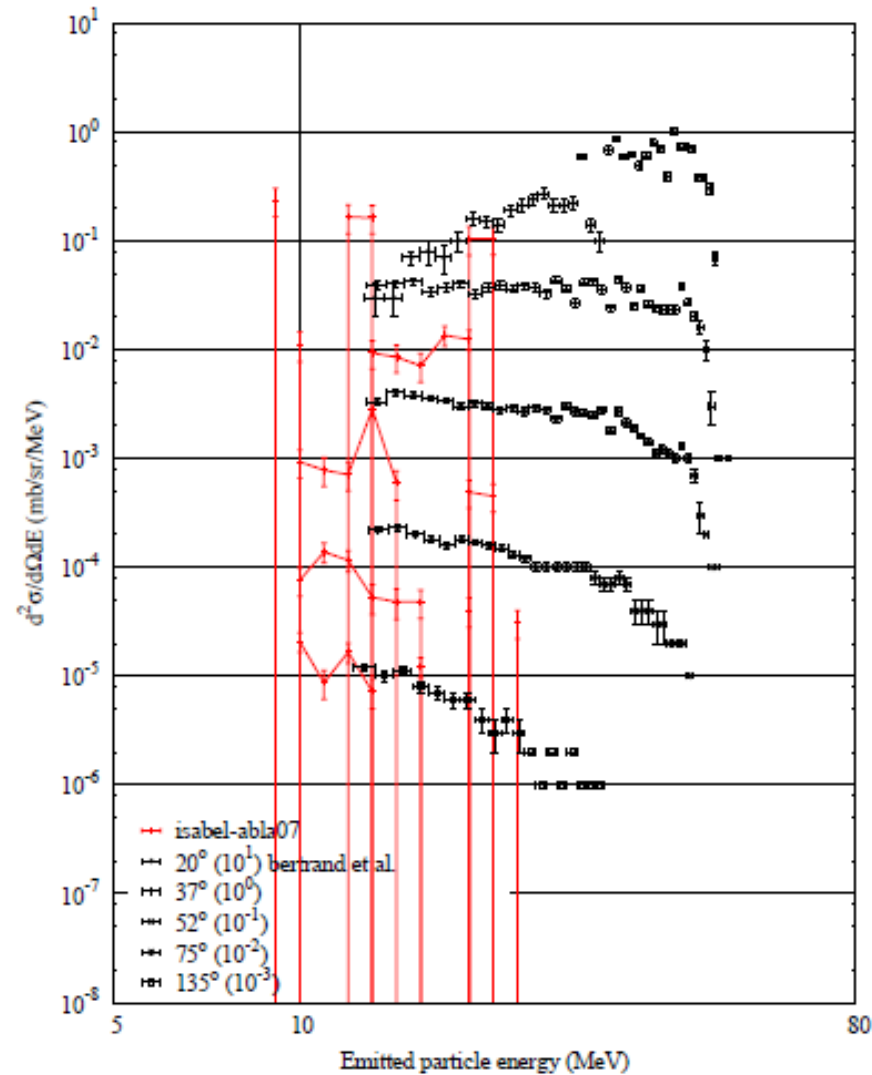
**$^3\text{He}$  spectra**

# p(62 MeV) + Bi – $^3\text{He}$ spectrum

## INCL45-ABLA07



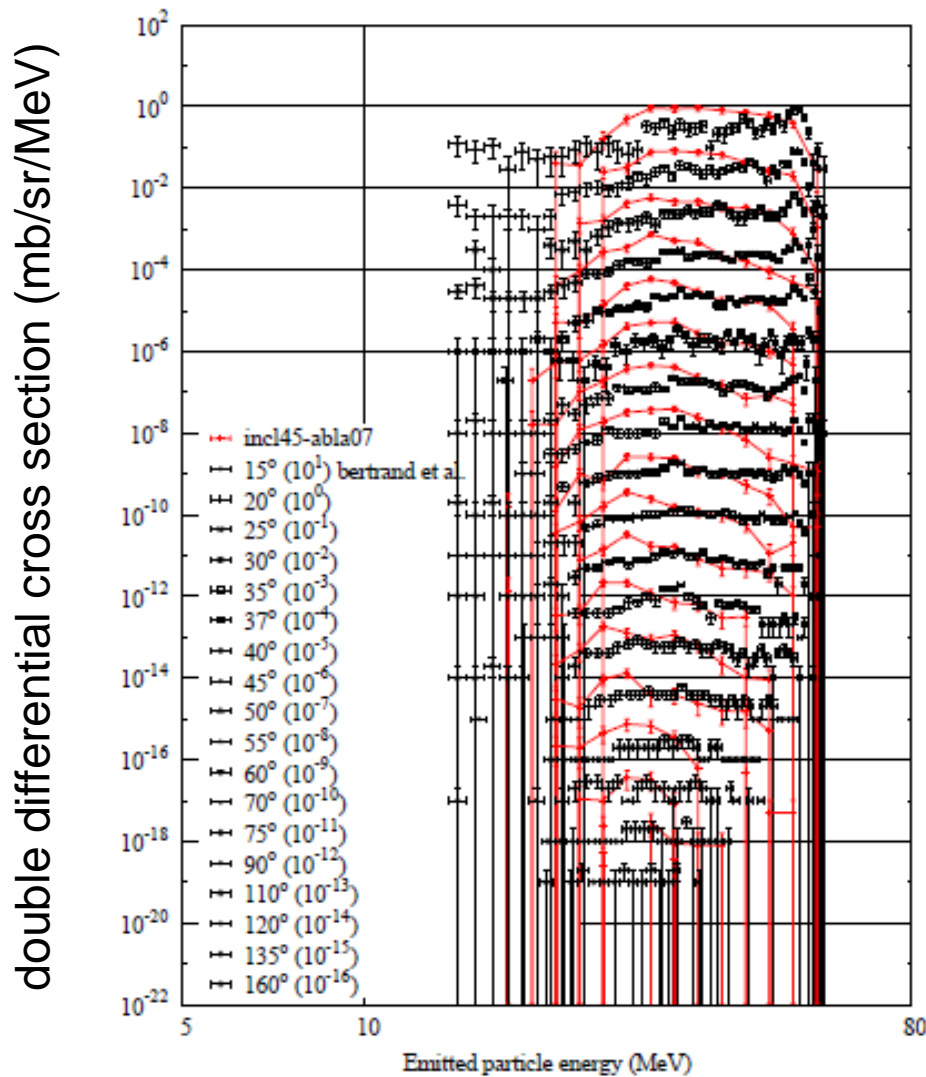
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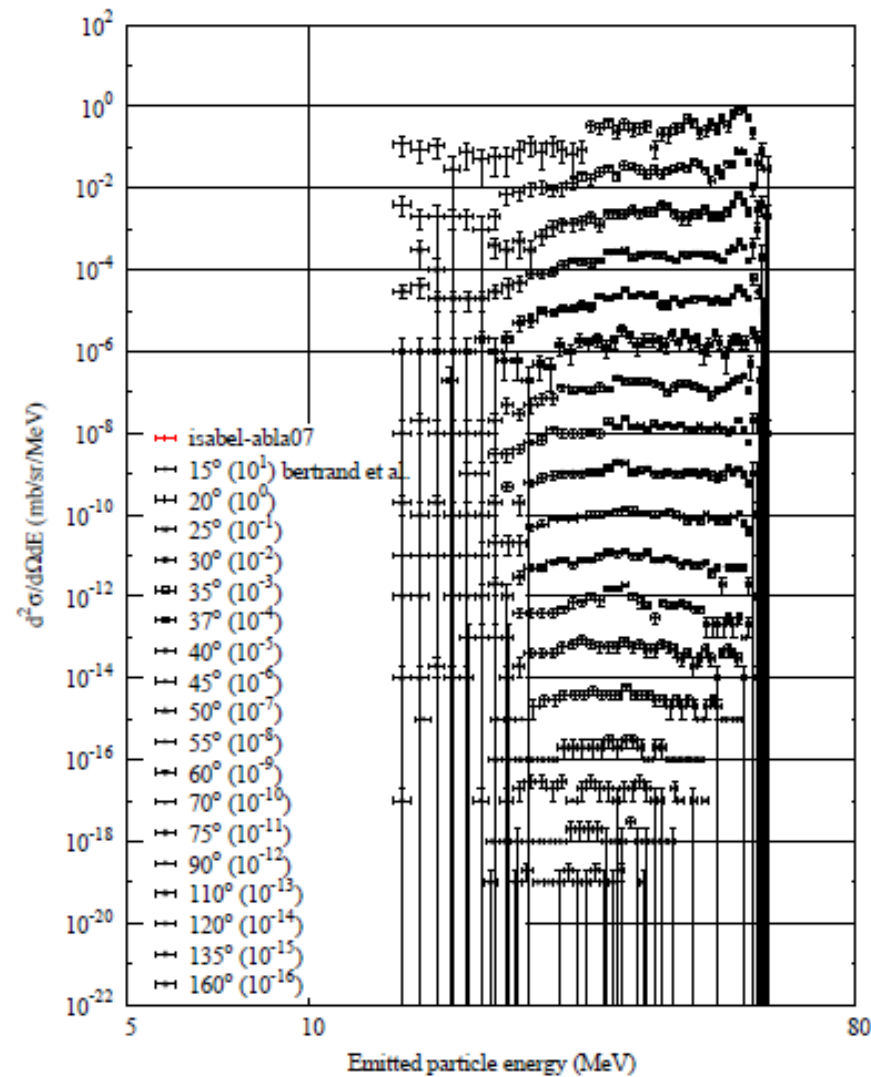
emitted-particle energy (MeV)

# p(62 MeV) + Bi – $^3\text{He}$ spectrum

## INCL45-ABLA07



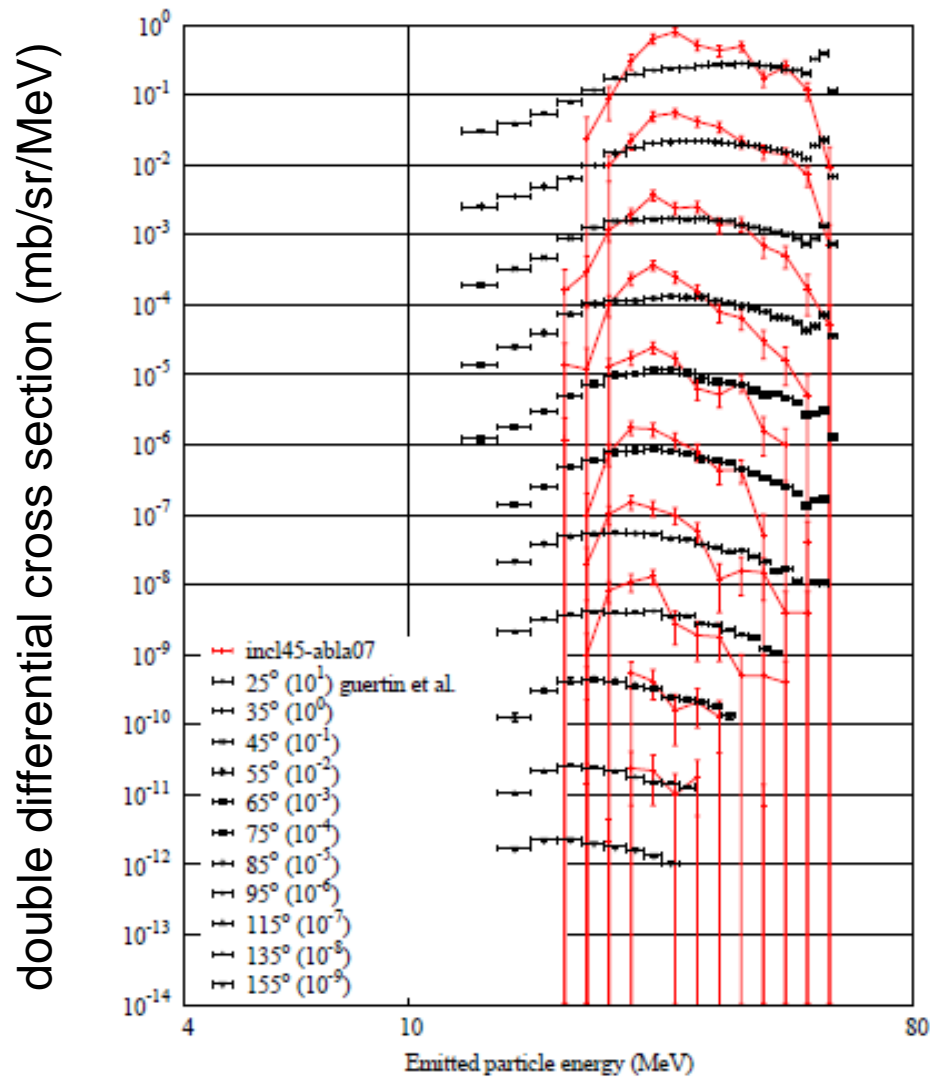
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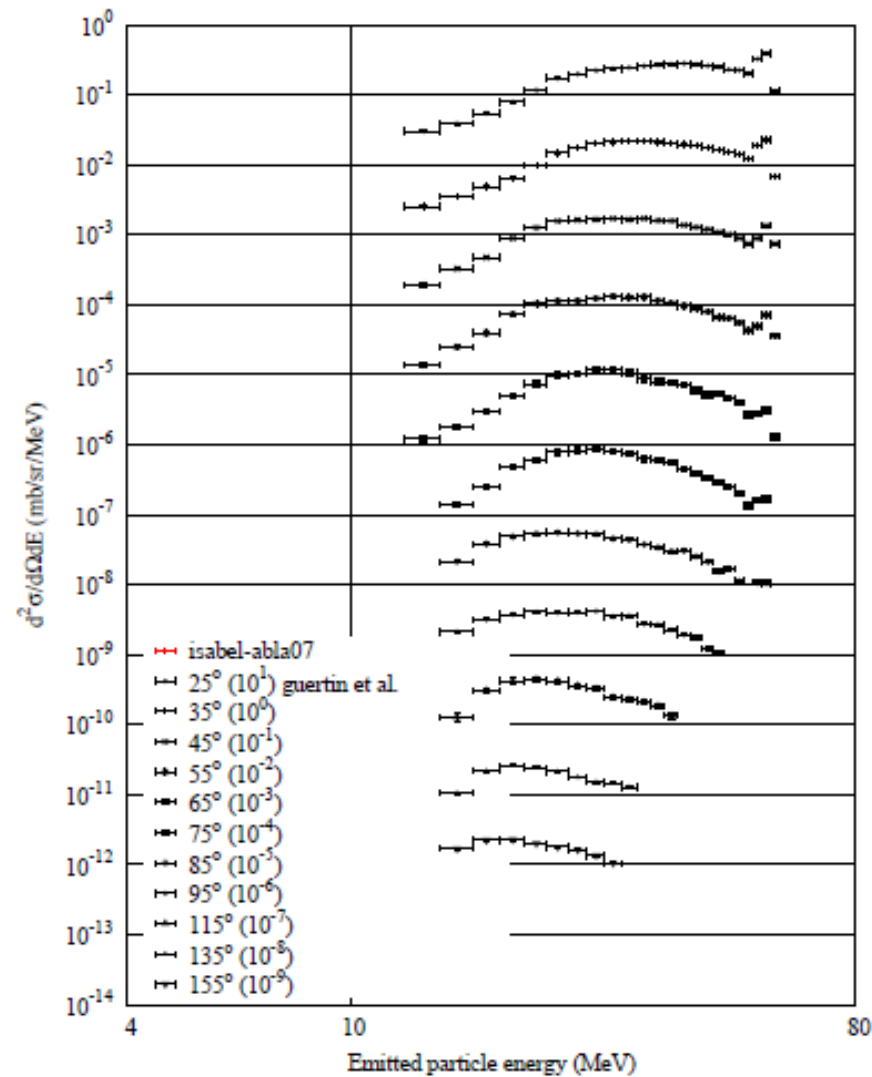
emitted-particle energy (MeV)

# p(63 MeV) + $^{208}\text{Pb}$ – $^3\text{He}$ spectrum

## INCL45-ABLA07



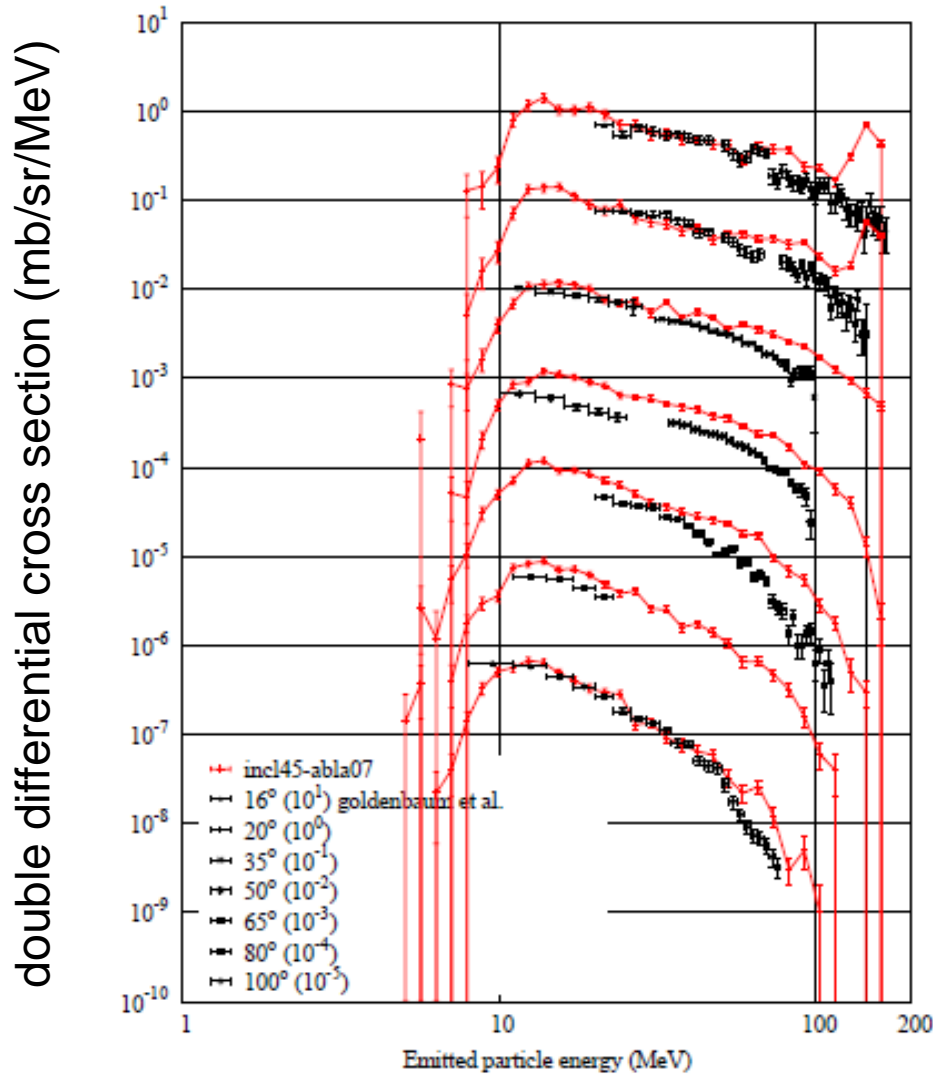
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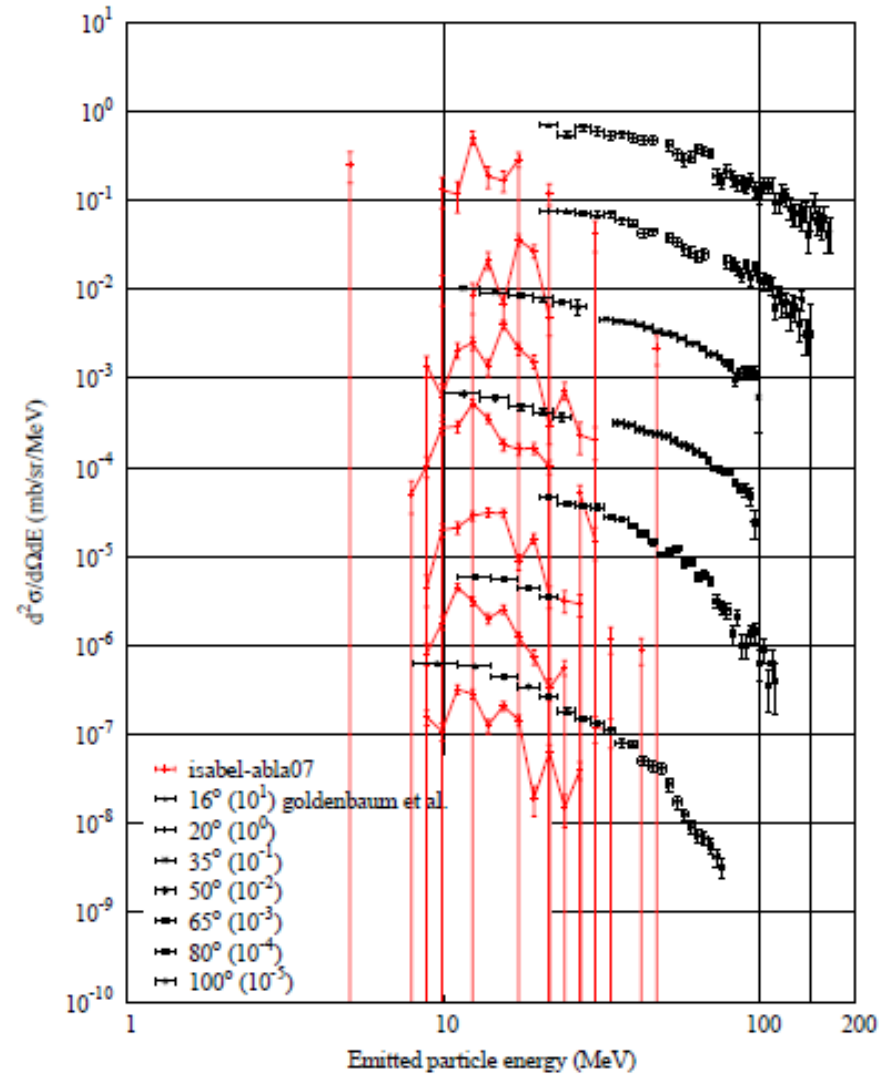
emitted-particle energy (MeV)

# p(175 MeV) + Ni – $^3\text{He}$ spectrum

## INCL45-ABLA07



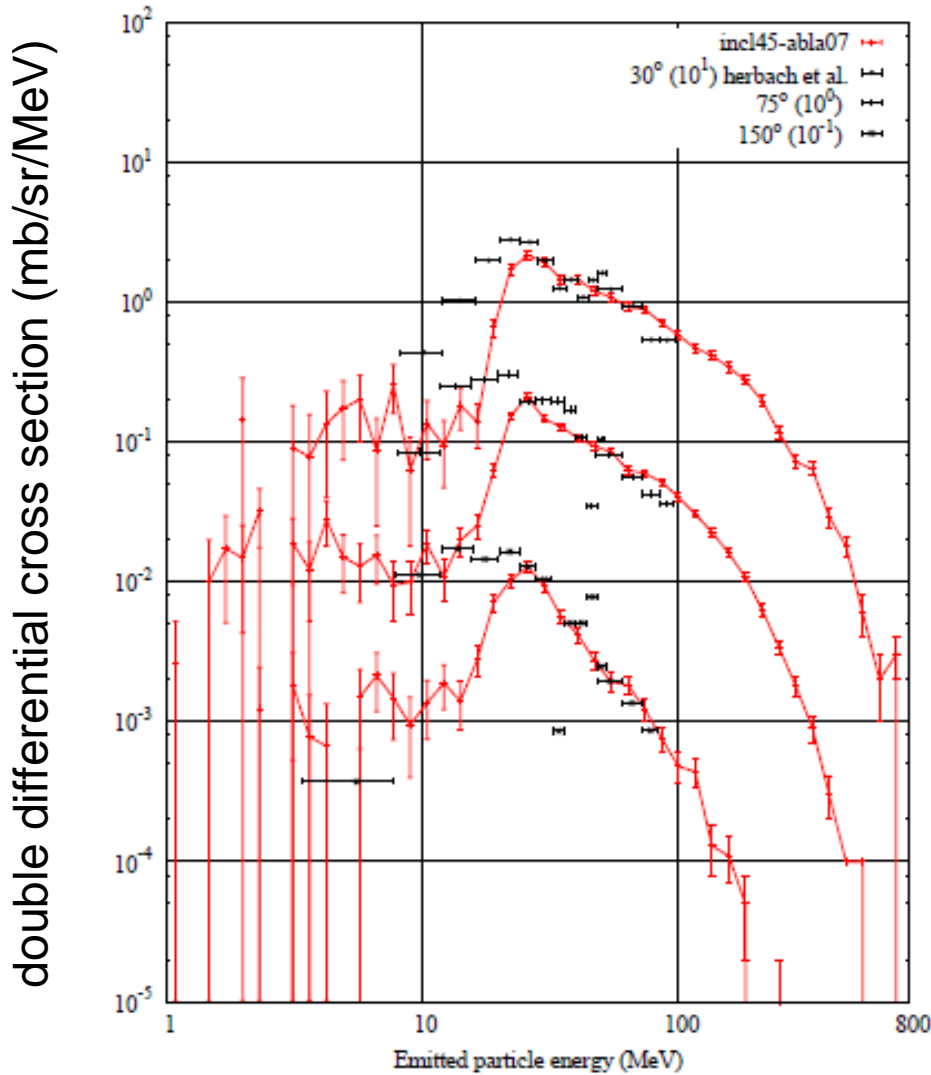
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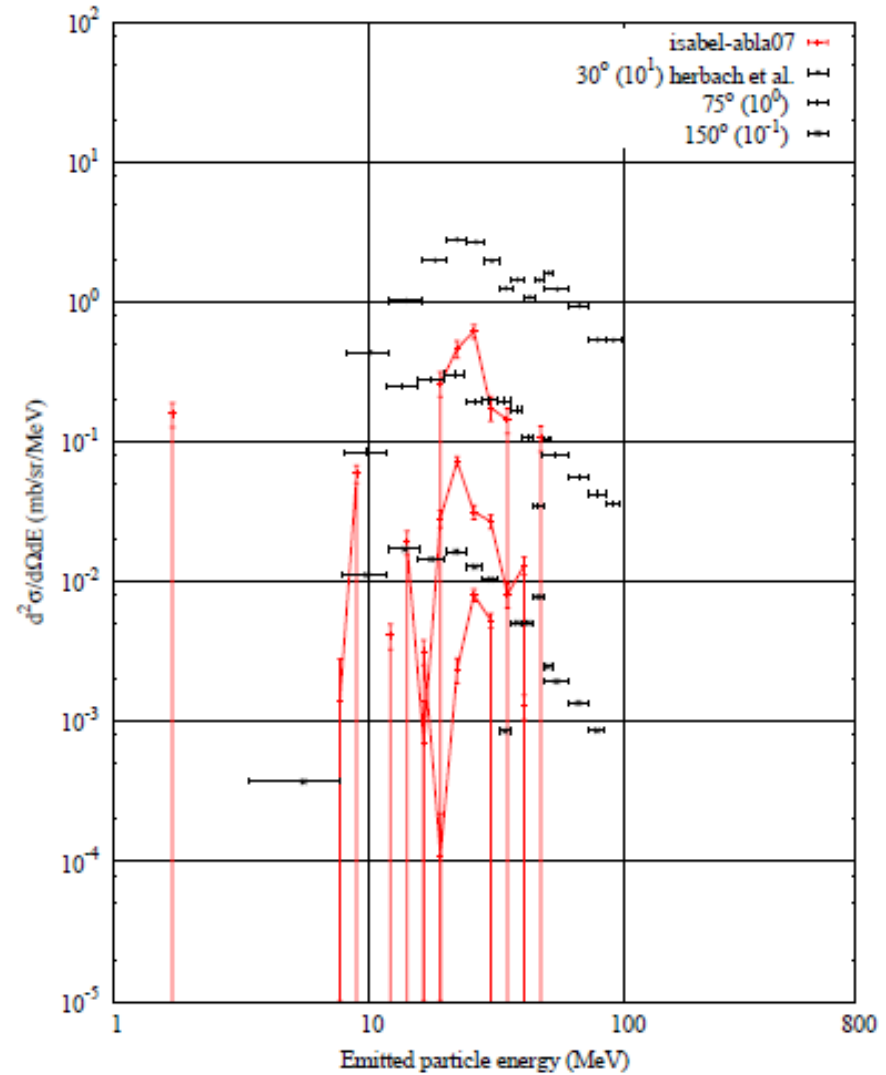
emitted-particle energy (MeV)

# p(1200 MeV) + Ta – $^3\text{He}$ spectrum

## INCL45-ABLA07



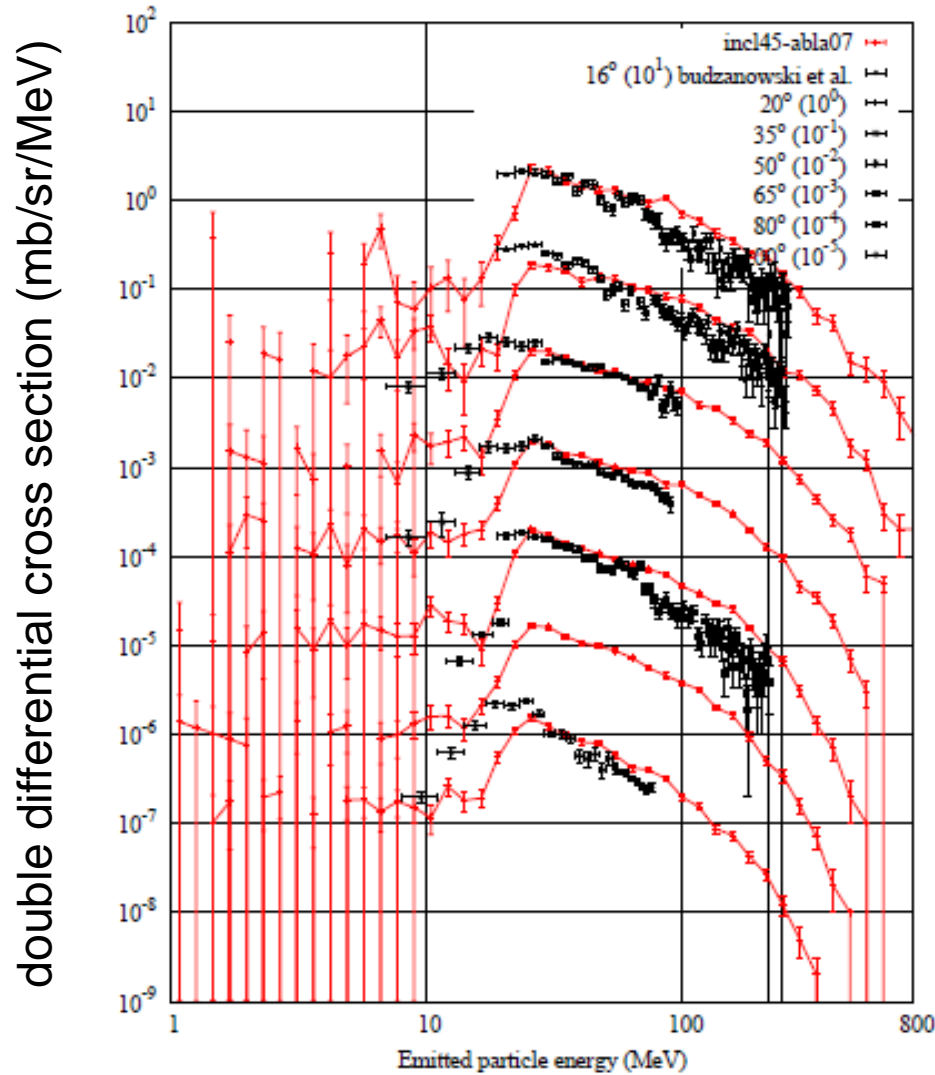
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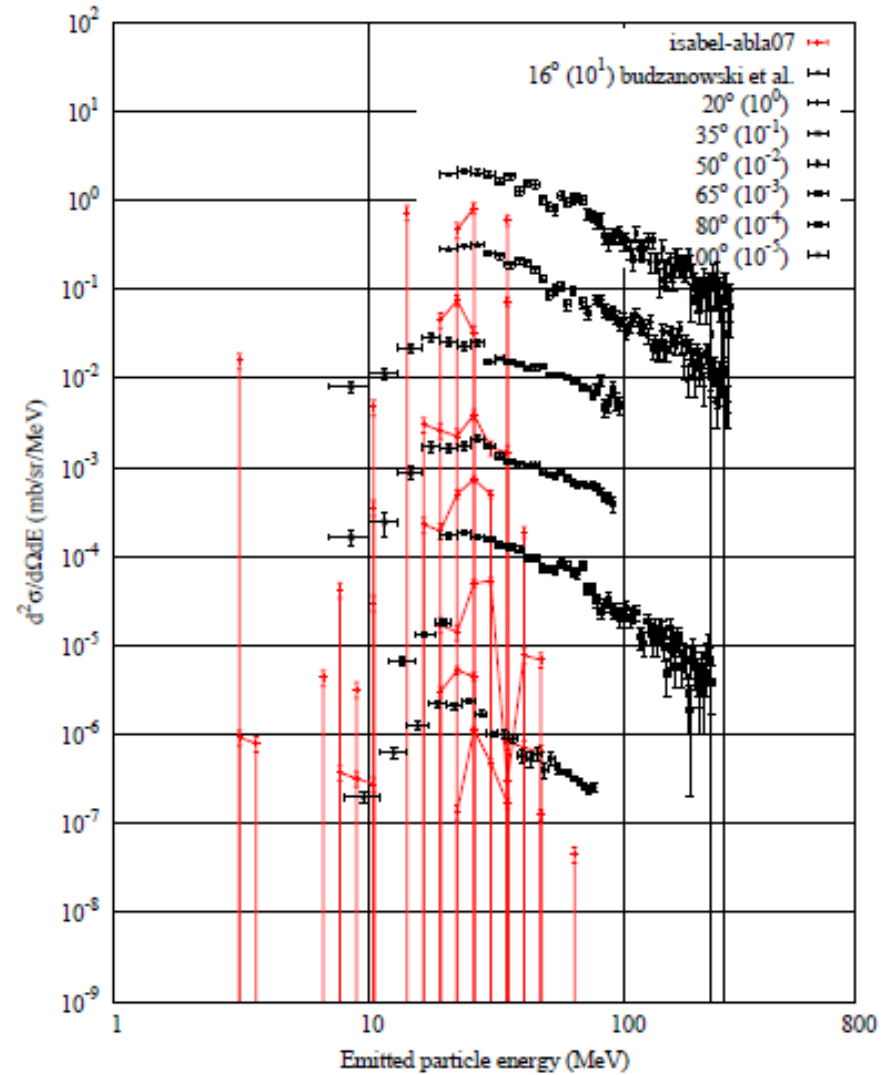
emitted-particle energy (MeV)

# p(1200 MeV) + Au – $^3\text{He}$ spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

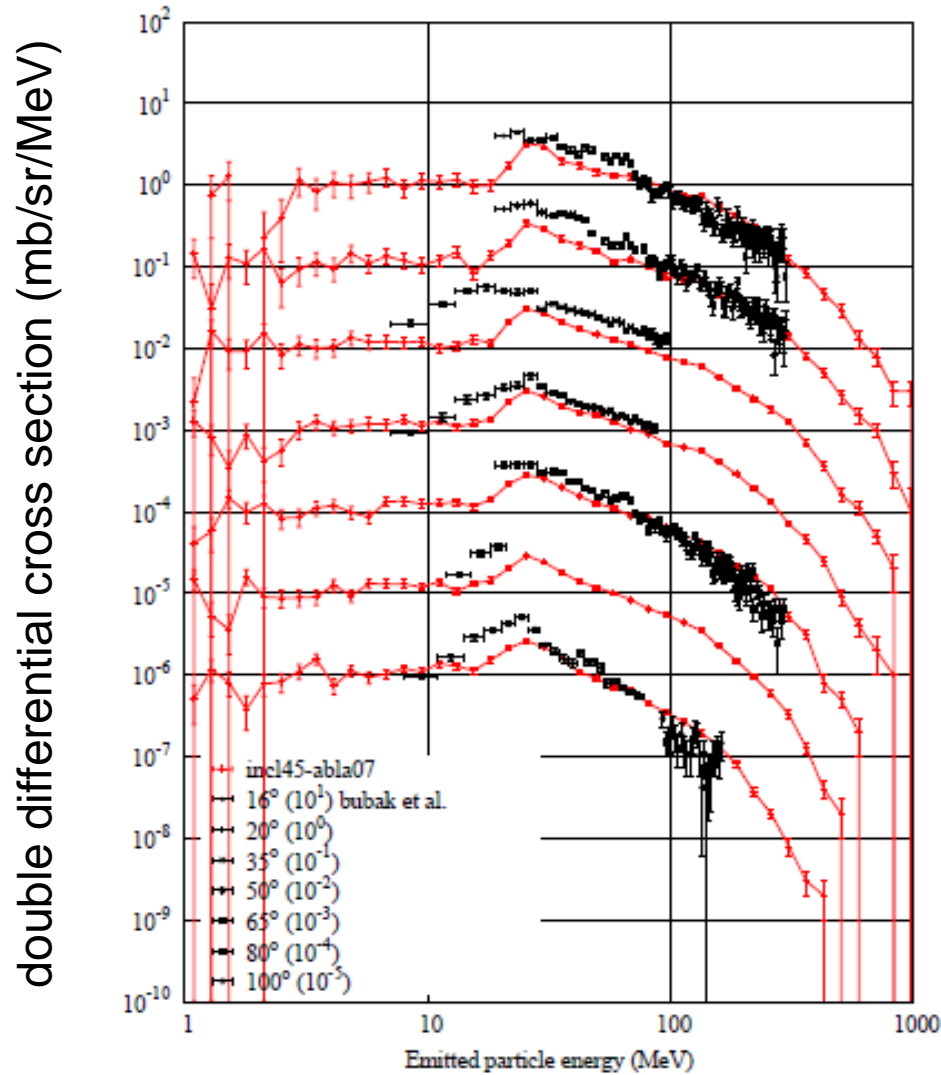


emitted-particle energy (MeV)

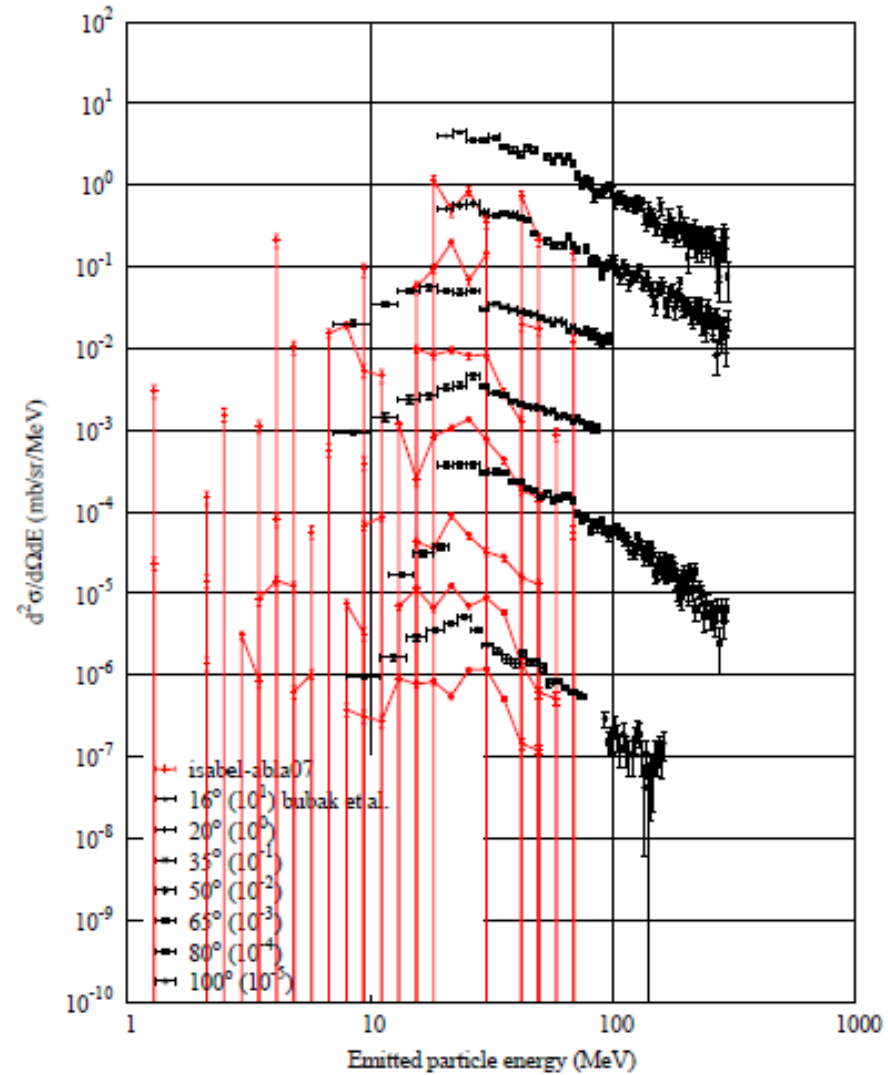


# p(2500 MeV) + Au – $^3\text{He}$ spectrum

## INCL45-ABLA07



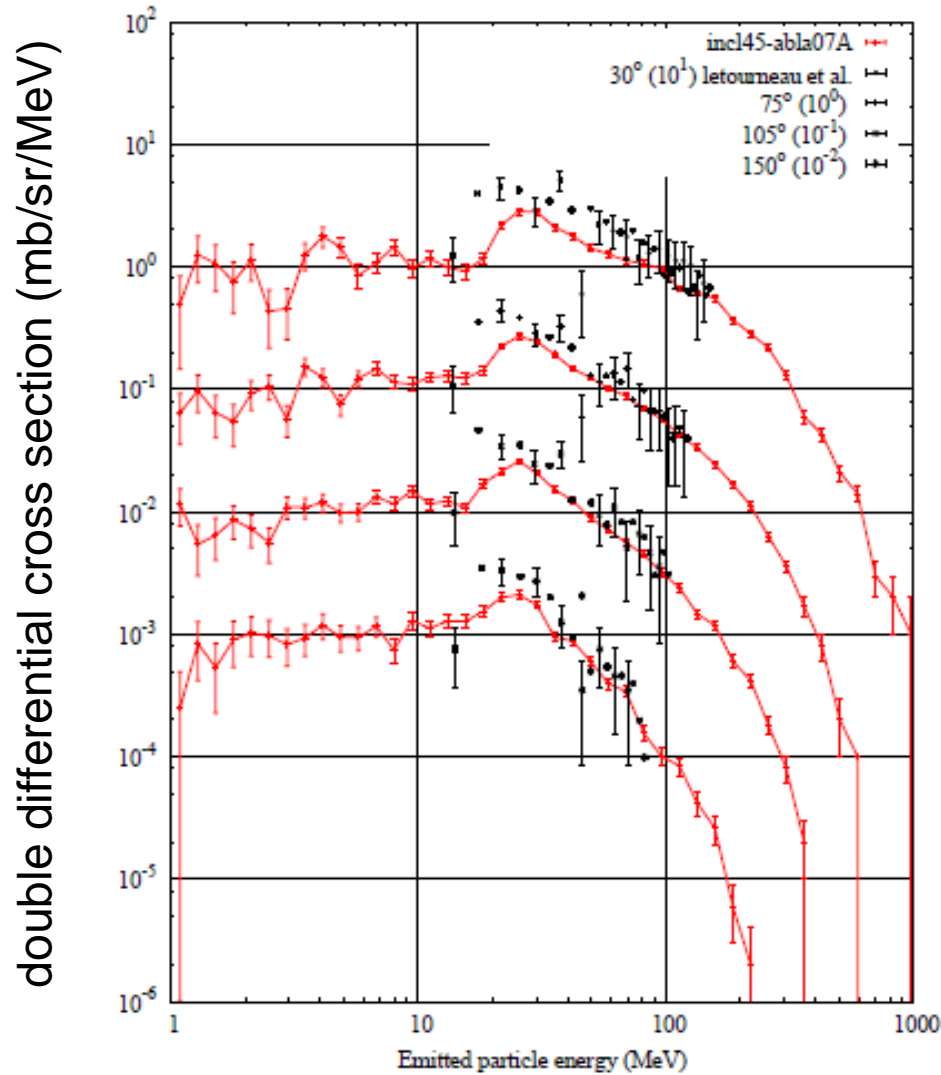
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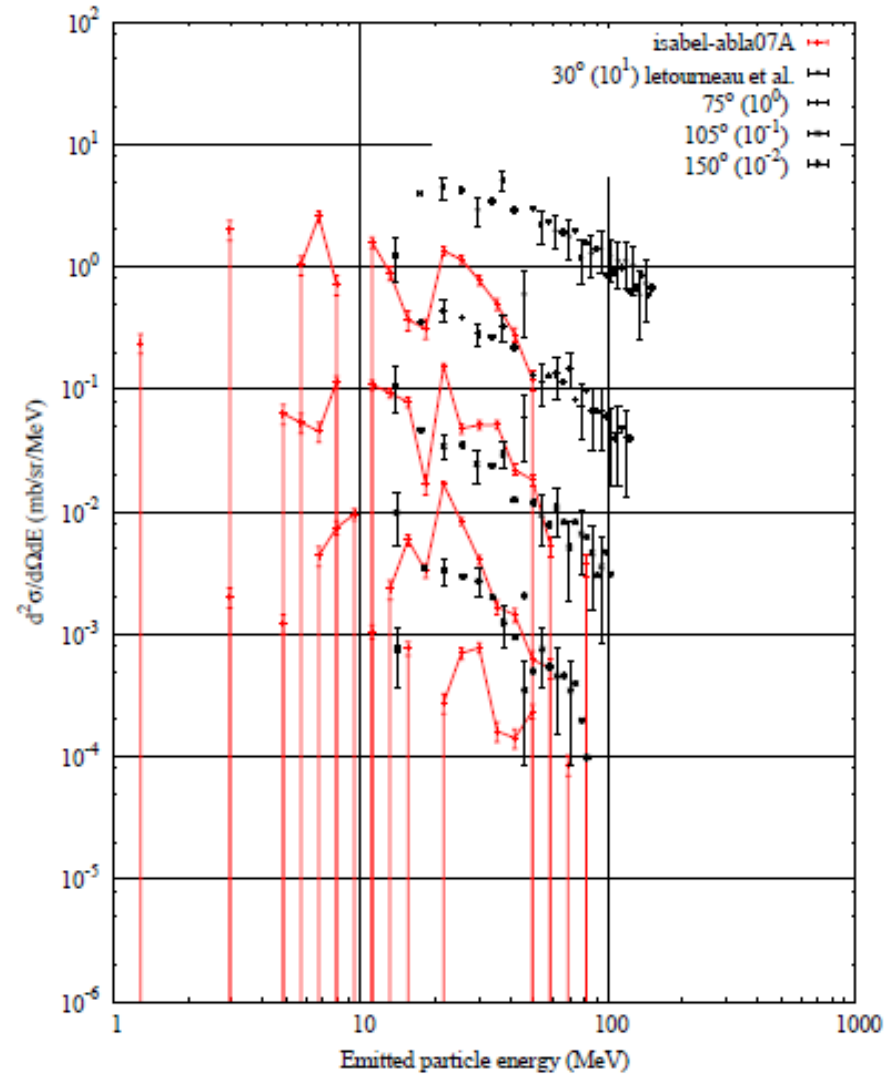
emitted-particle energy (MeV)

# p(2500 MeV) + Au – $^3\text{He}$ spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

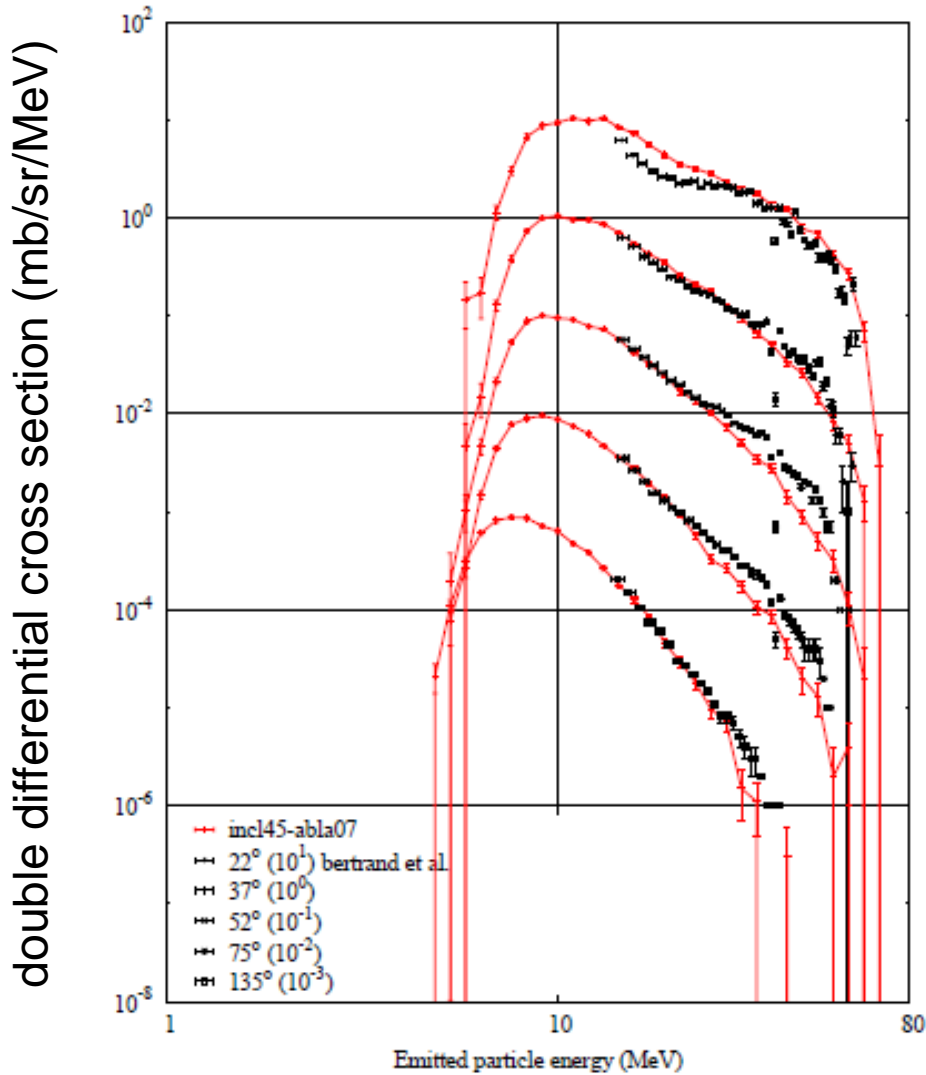


emitted-particle energy (MeV)

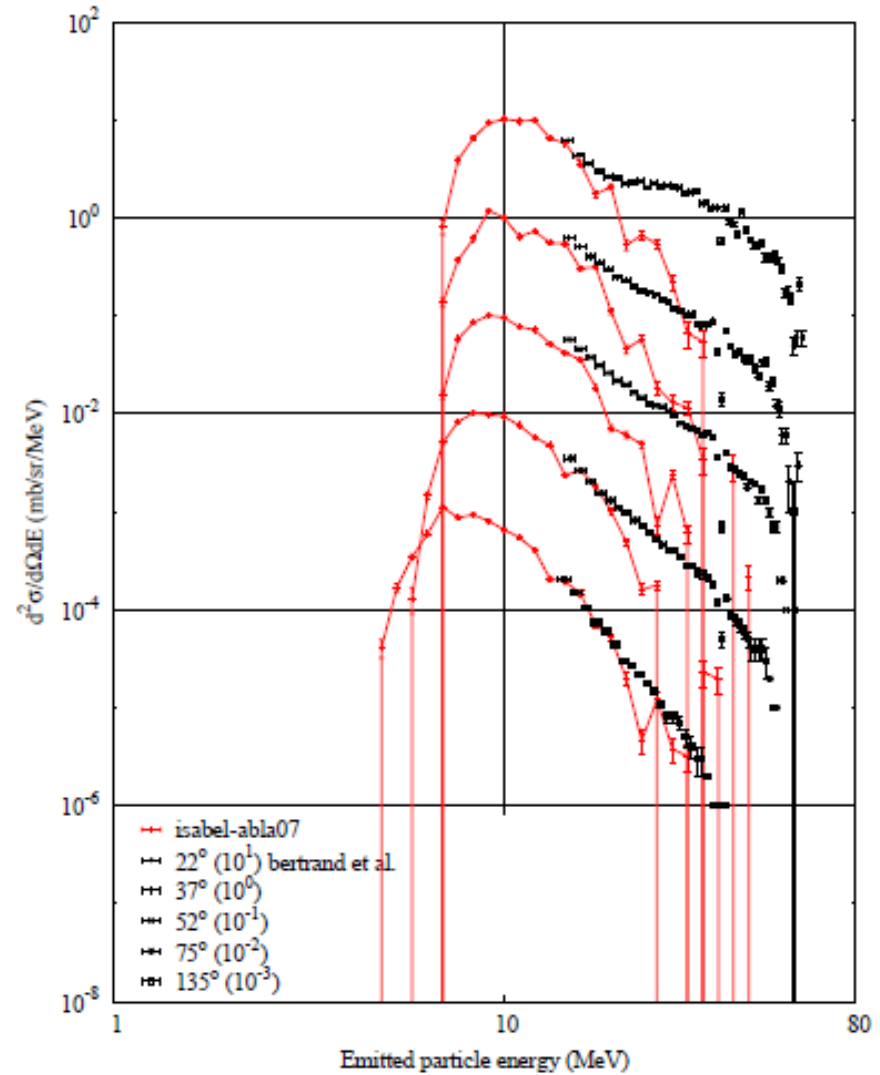
**$^4\text{He}$  spectra**

# p(62 MeV) + $^{56}\text{Fe}$ – $^4\text{He}$ spectrum

## INCL45-ABLA07



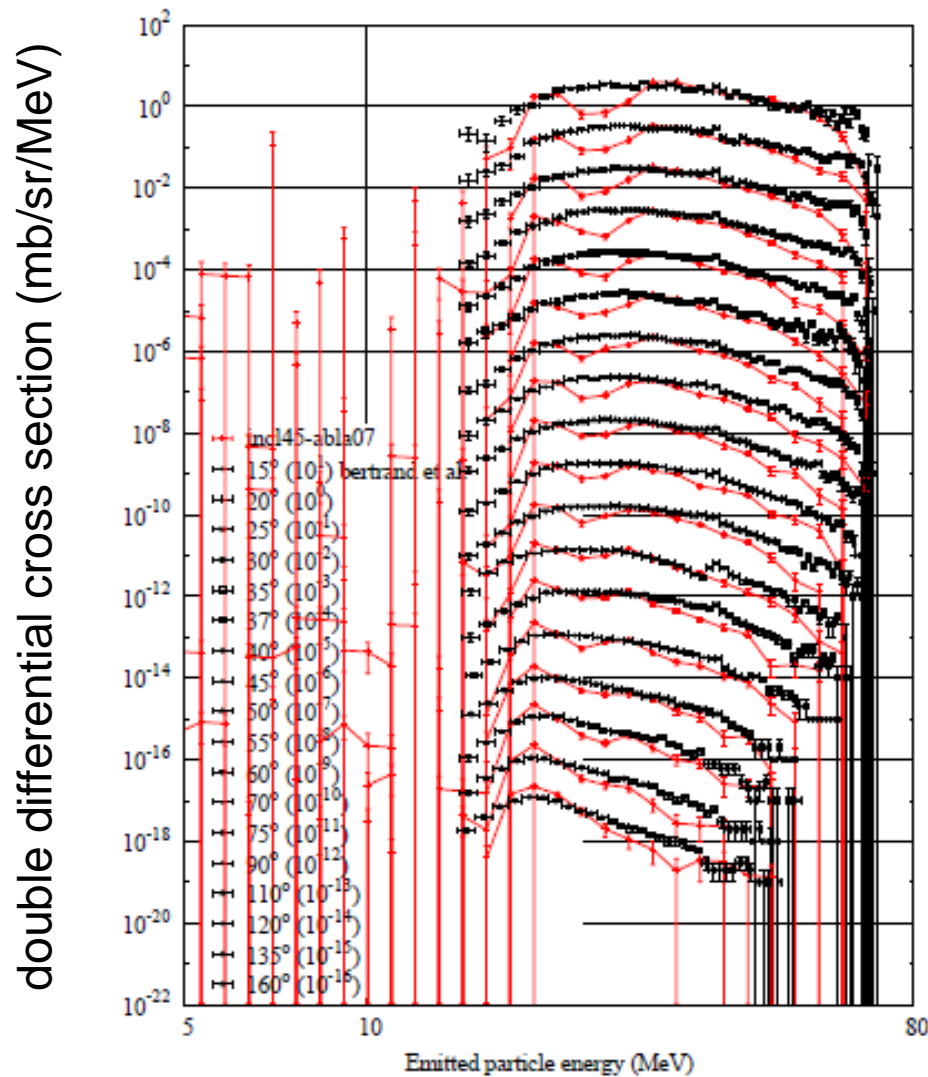
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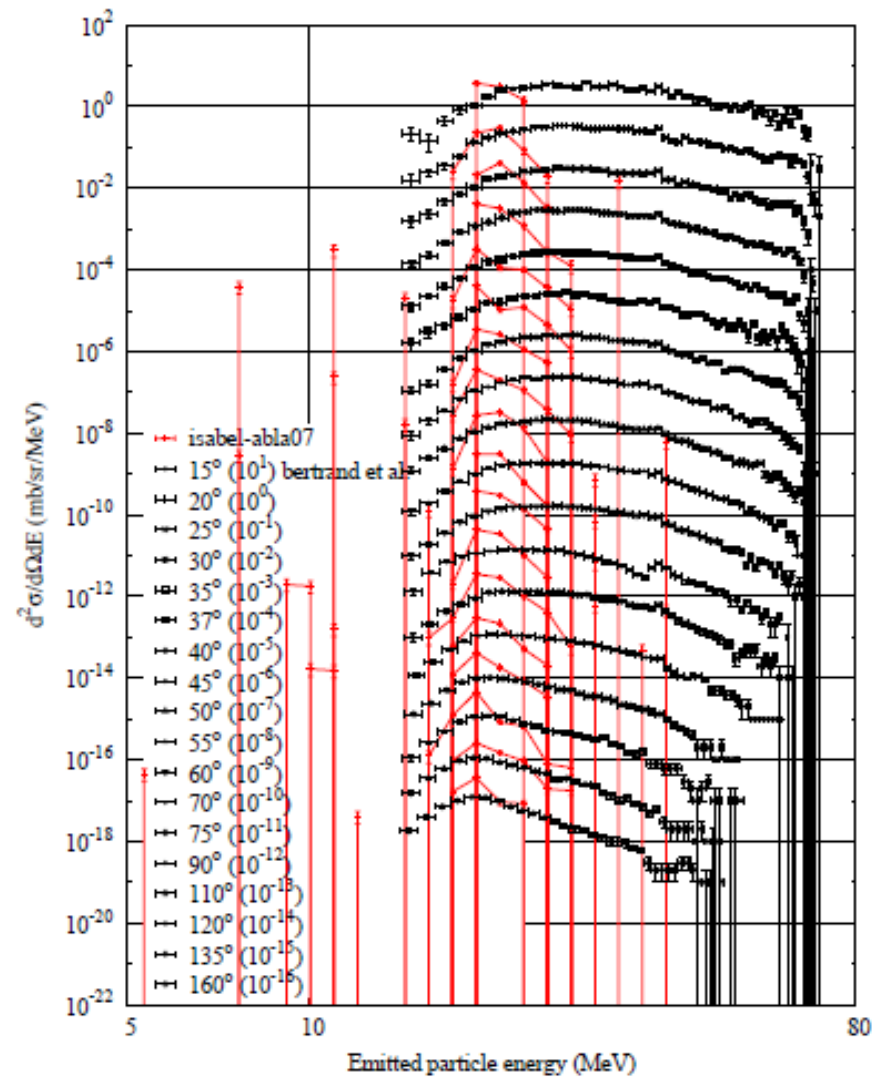
emitted-particle energy (MeV)

# p(62 MeV) + Bi – $^4\text{He}$ spectrum

## INCL45-ABLA07



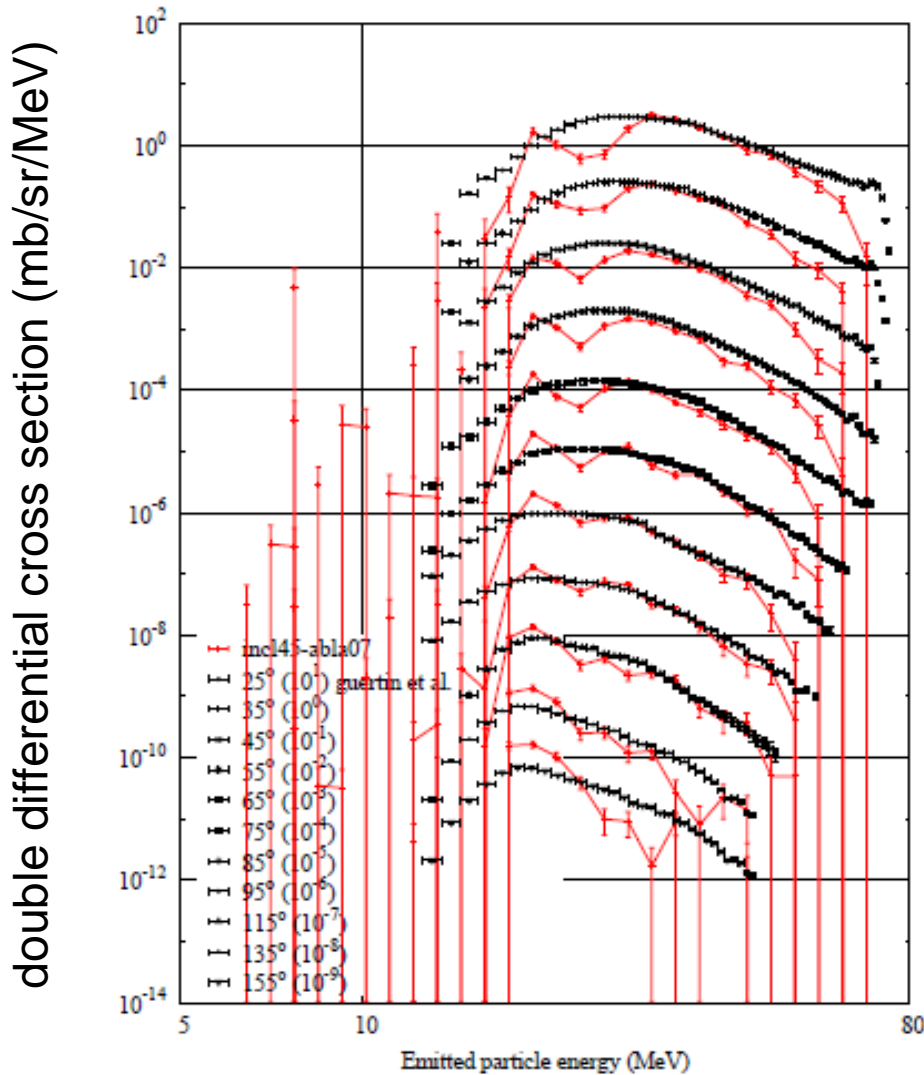
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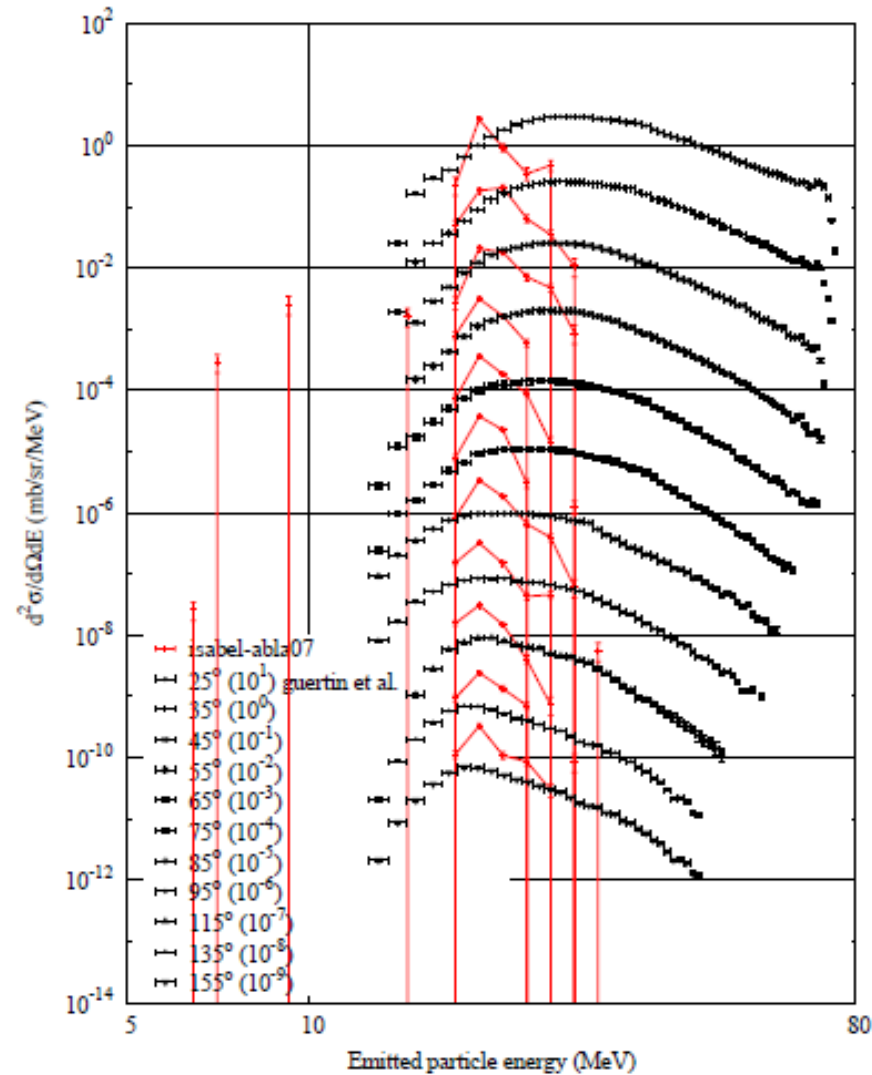
emitted-particle energy (MeV)

# p(63 MeV) + $^{208}\text{Pb}$ – $^4\text{He}$ spectrum

## INCL45-ABLA07



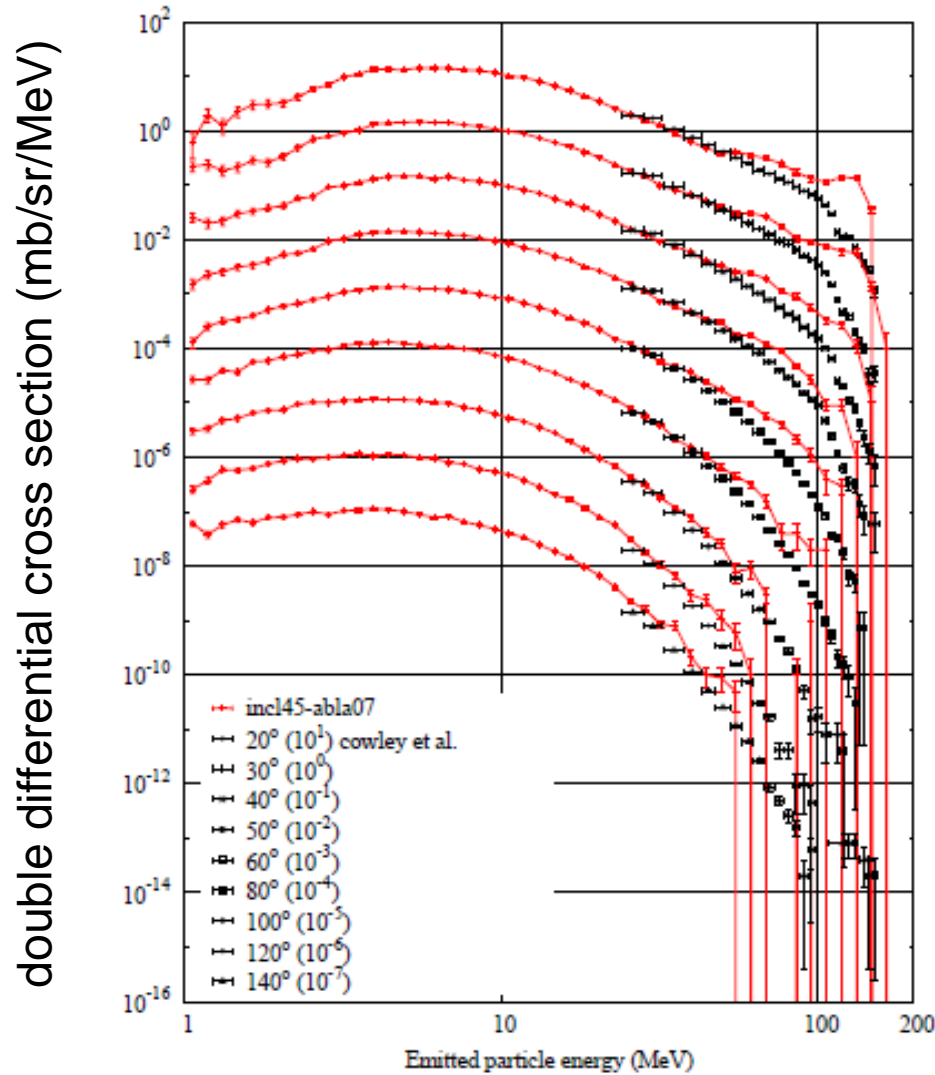
## ISABEL-ABLA07



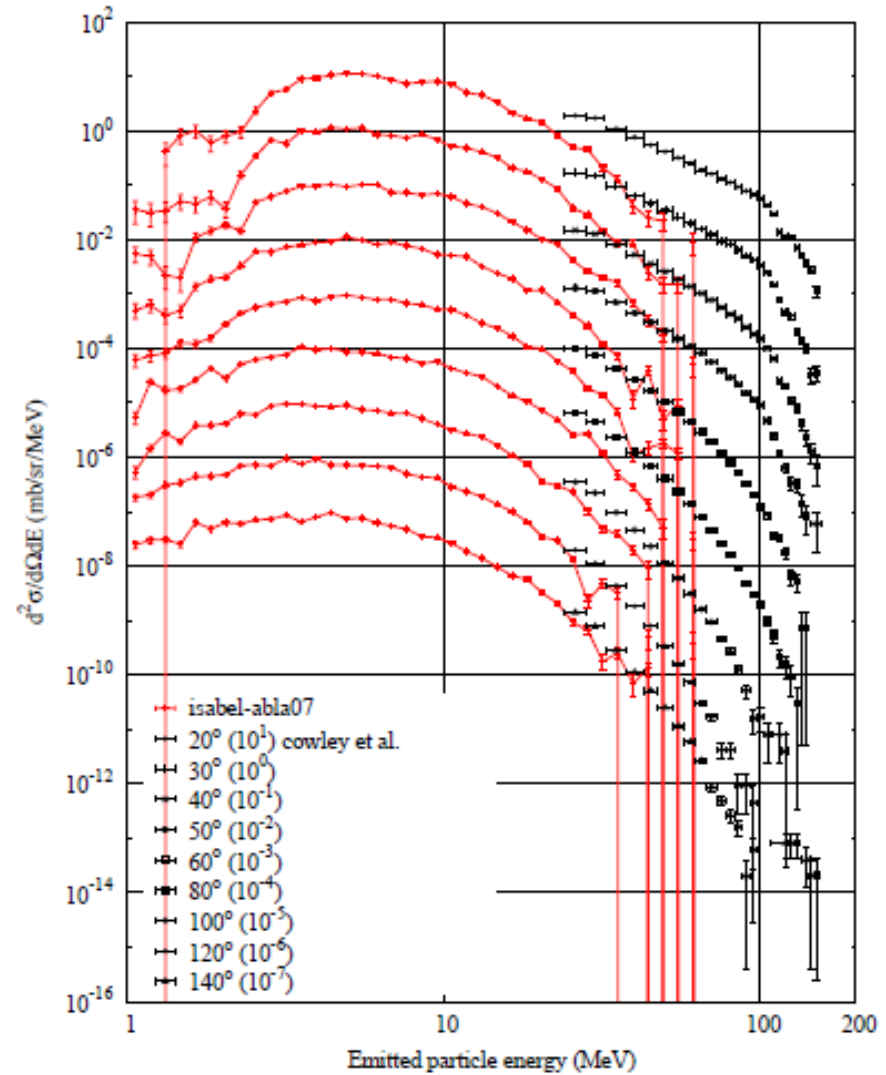
emitted-particle energy (MeV)

# p(160 MeV) + Al – $^4\text{He}$ spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

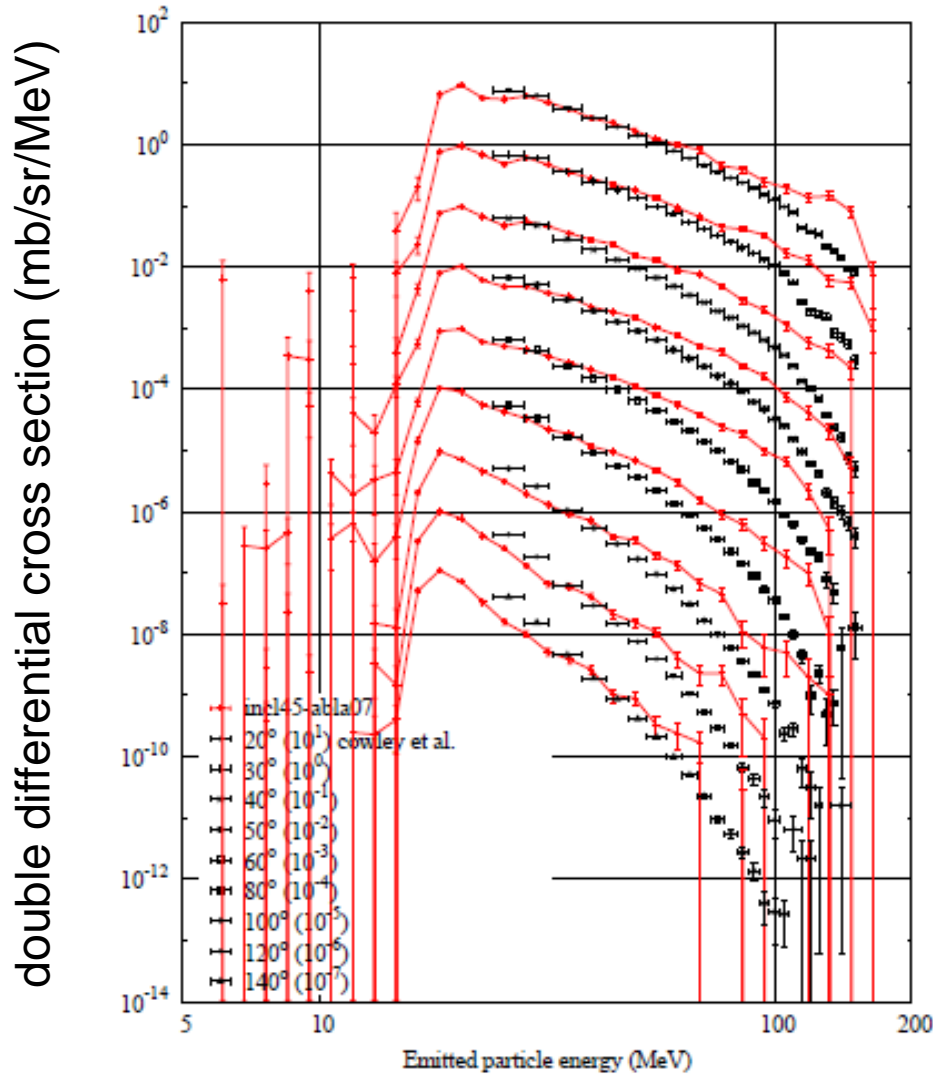


emitted-particle energy (MeV)

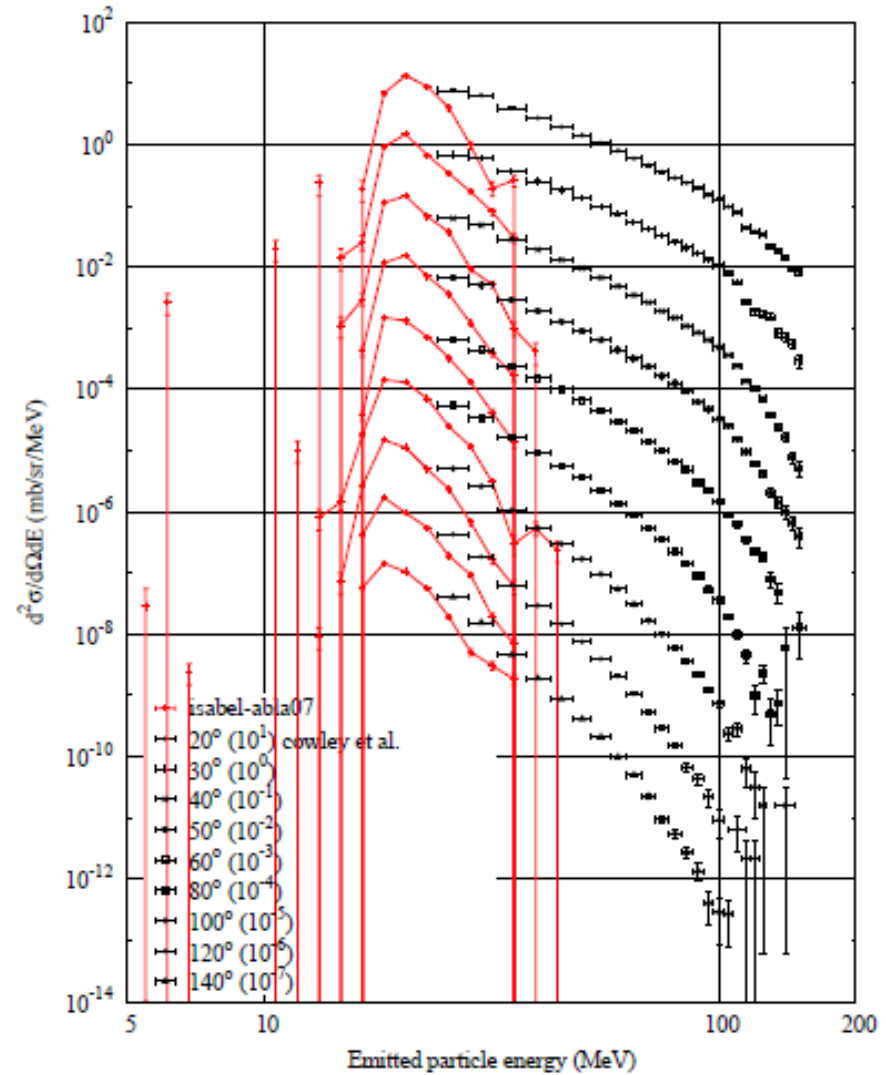


# p(160 MeV) + Au – $^4\text{He}$ spectrum

## INCL45-ABLA07



## ISABEL-ABLA07

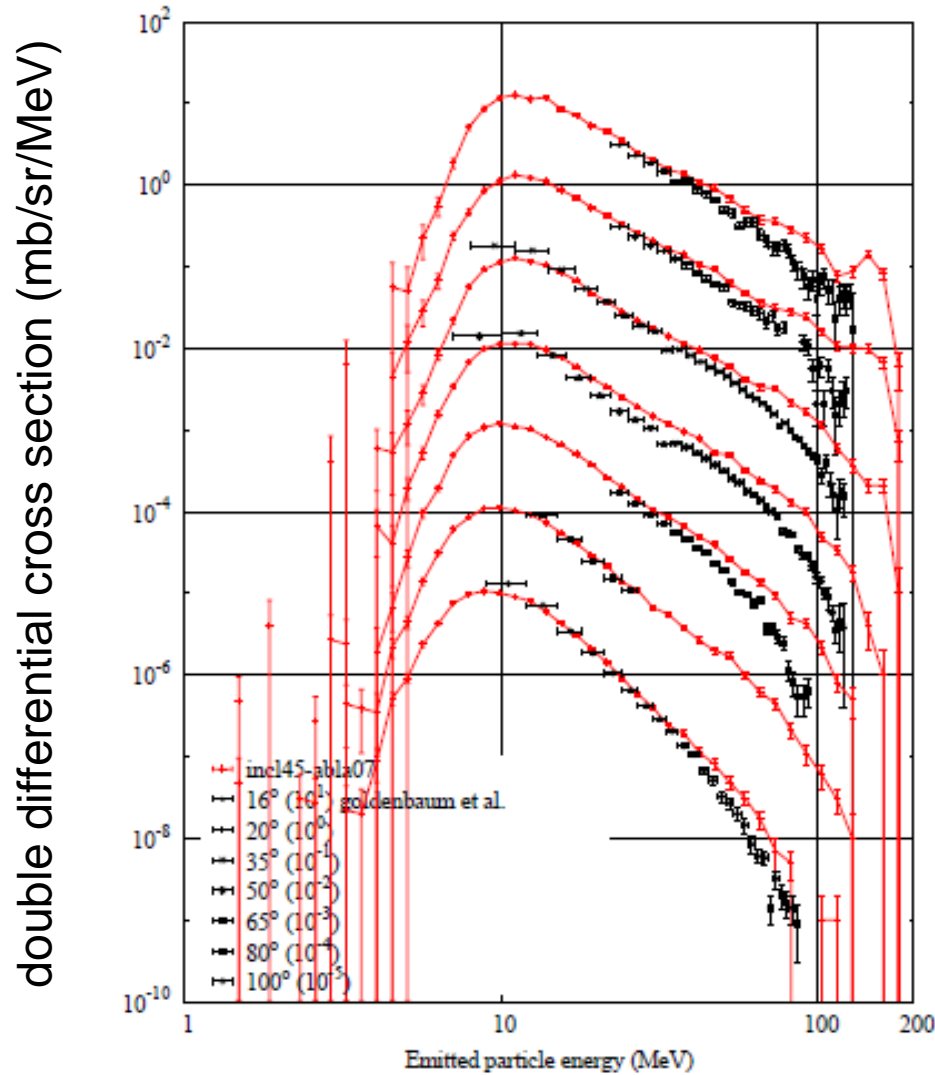


emitted-particle energy (MeV)

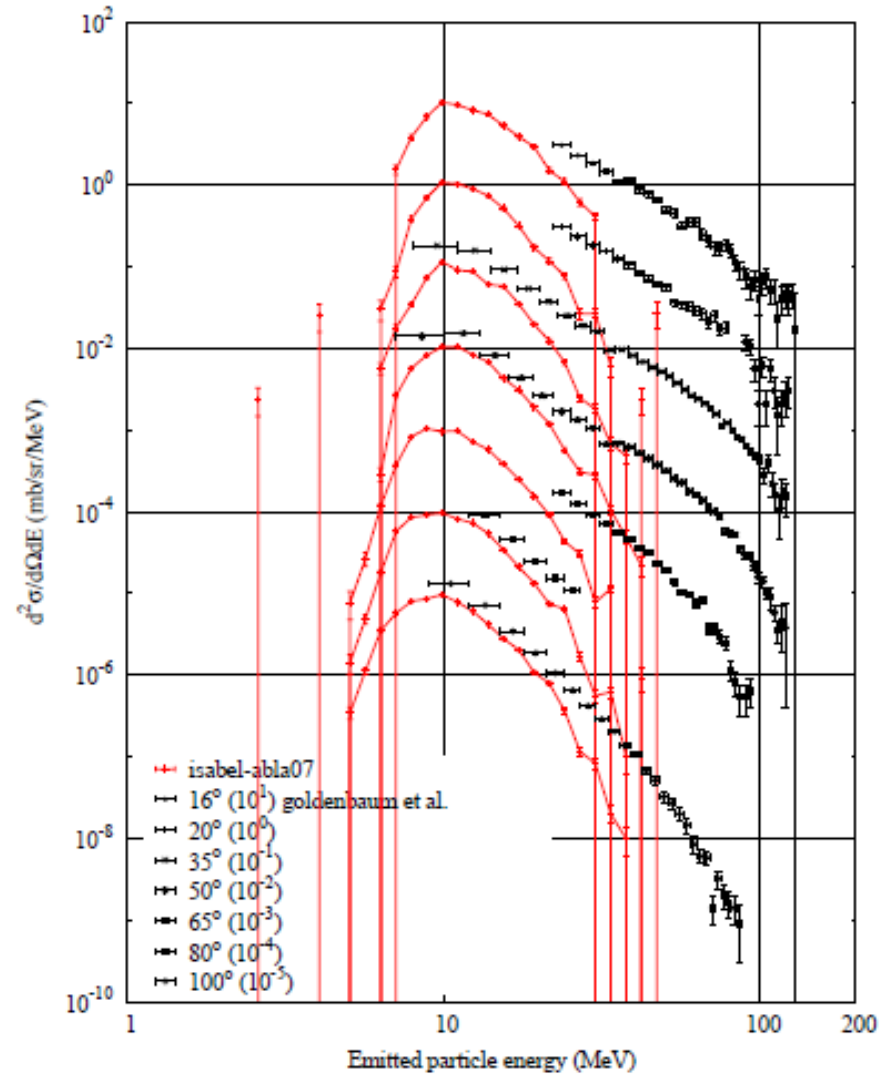


# p(175 MeV) + Ni – <sup>4</sup>He spectrum

## INCL45-ABLA07



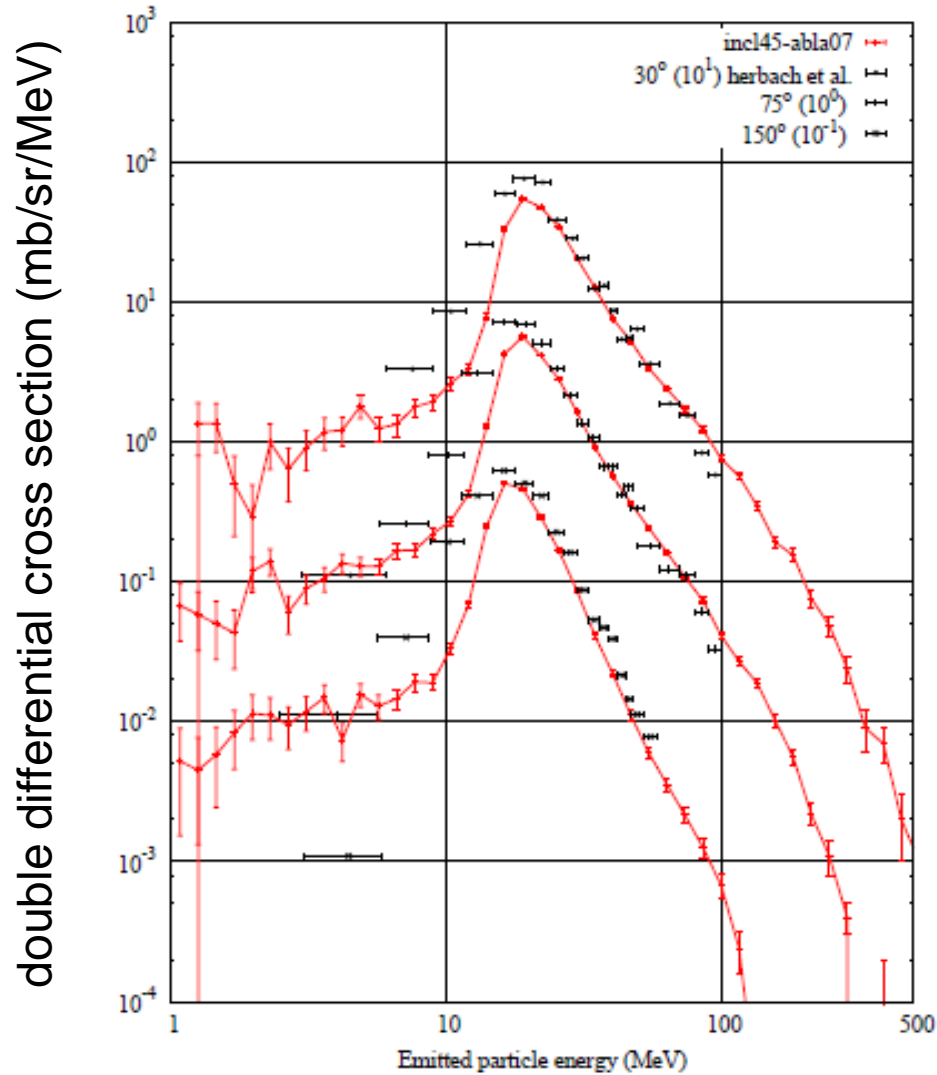
## ISABEL-ABLA07



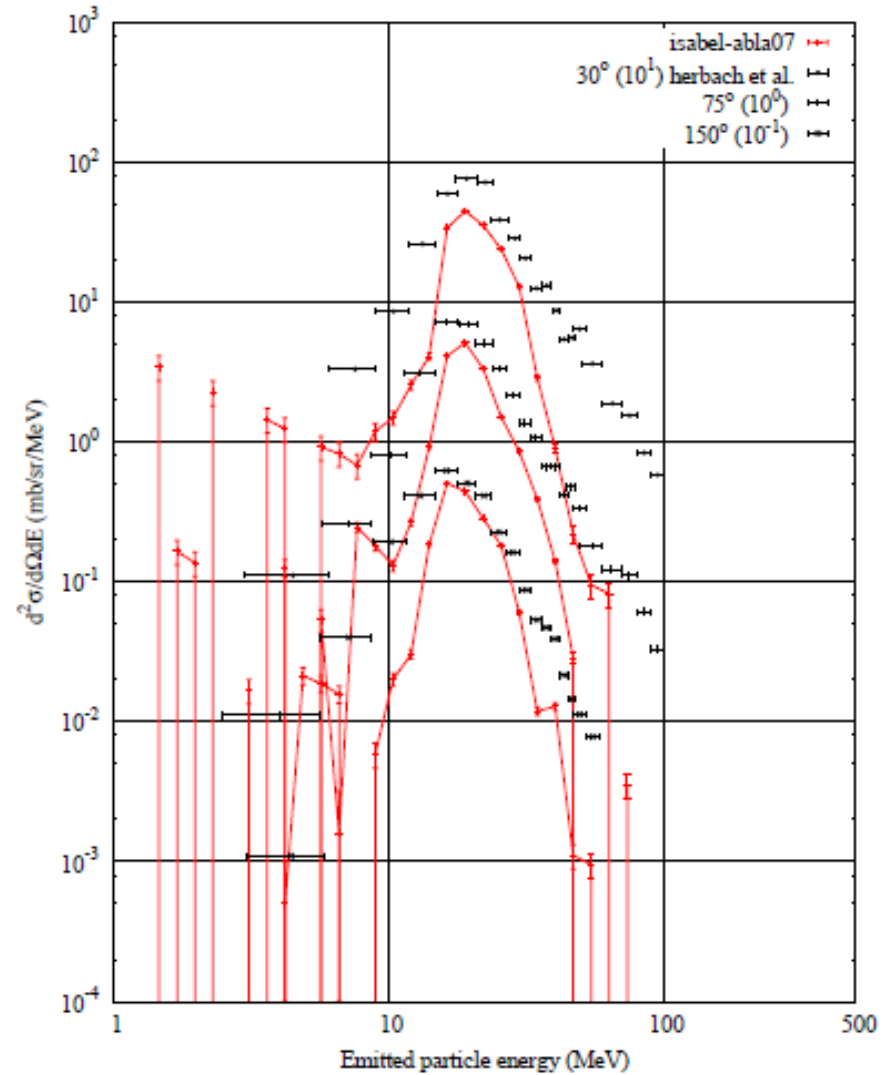
emitted-particle energy (MeV)

# p(1200 MeV) + Ta – $^4\text{He}$ spectrum

## INCL45-ABLA07



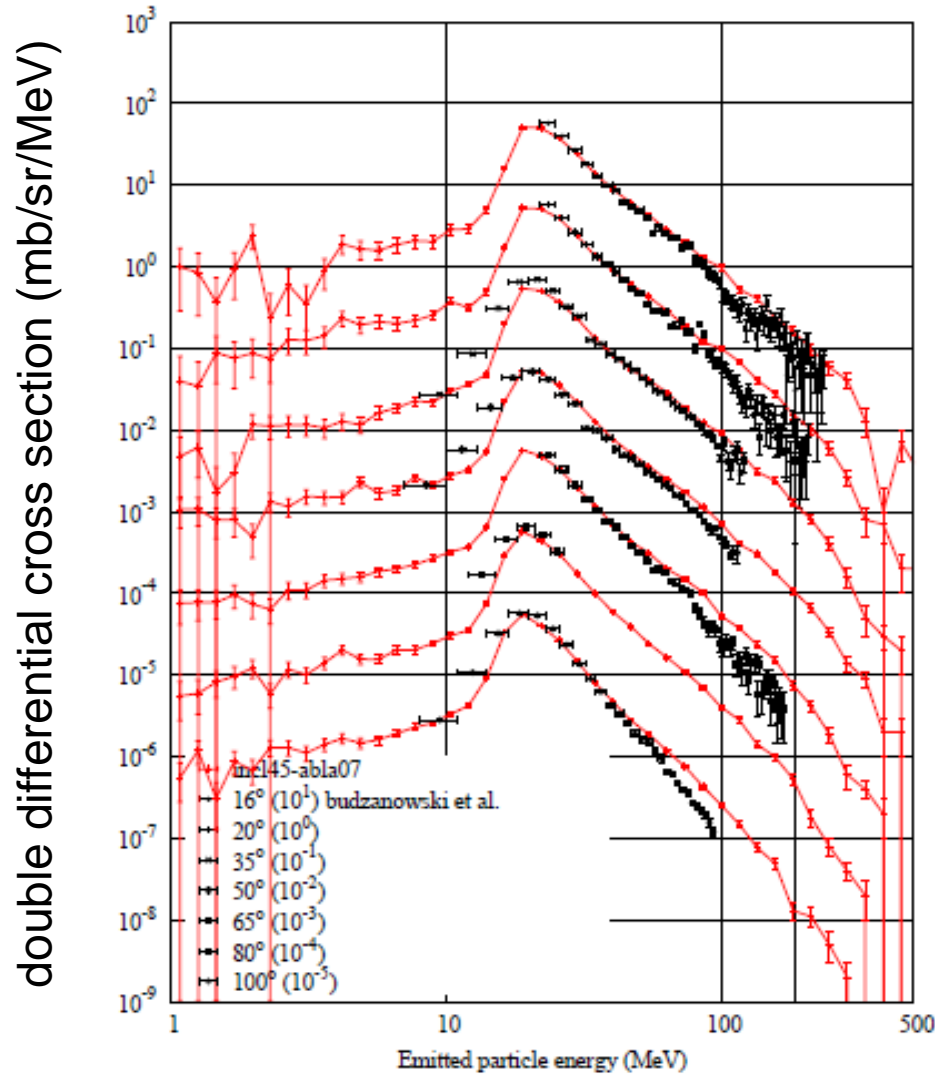
## ISABEL-ABLA07



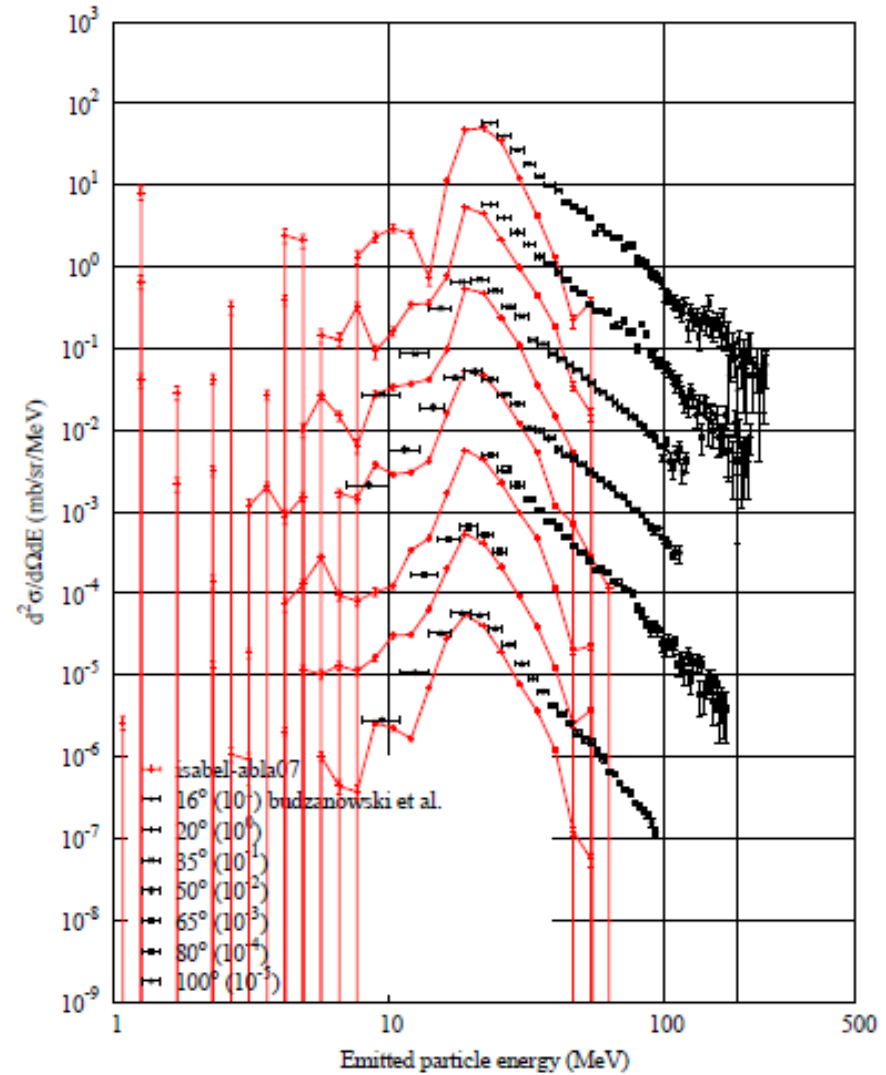
emitted-particle energy (MeV)

# p(1200 MeV) + Au – $^4\text{He}$ spectrum

## INCL45-ABLA07



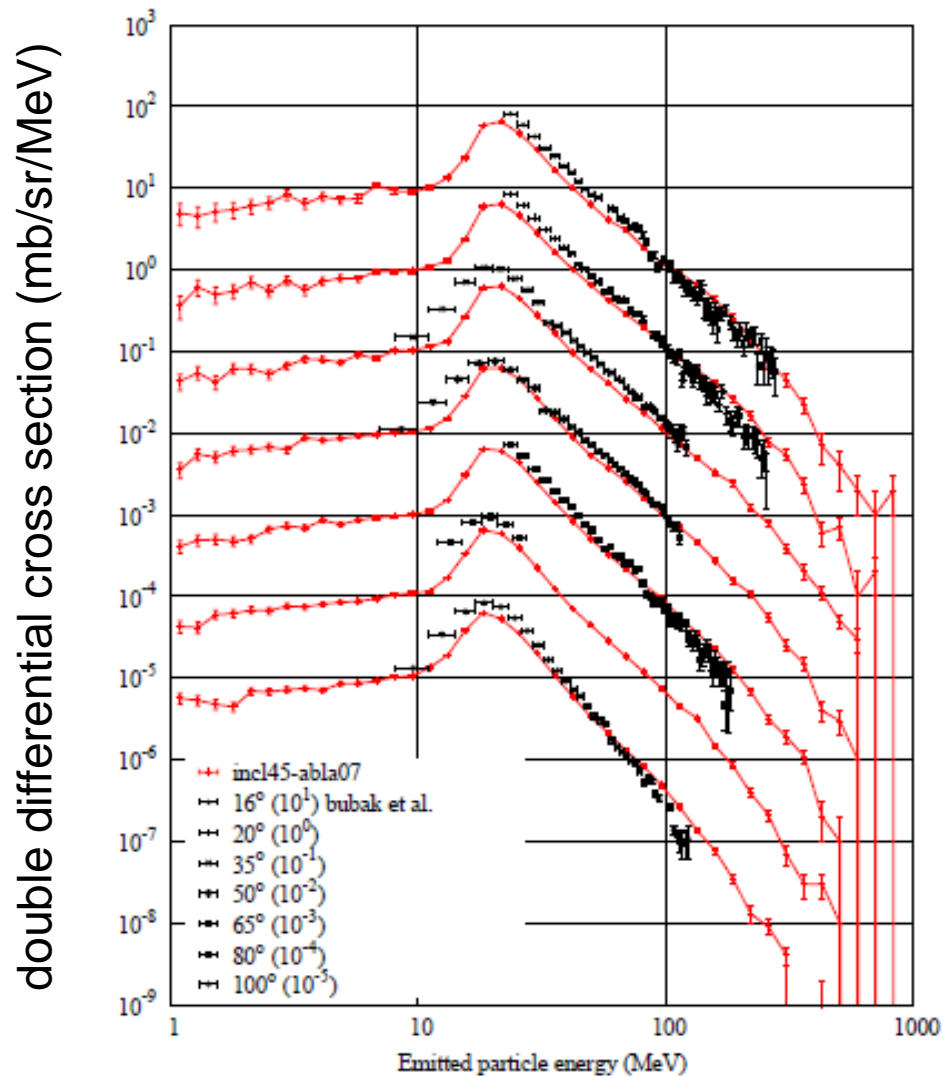
## ISABEL-ABLA07



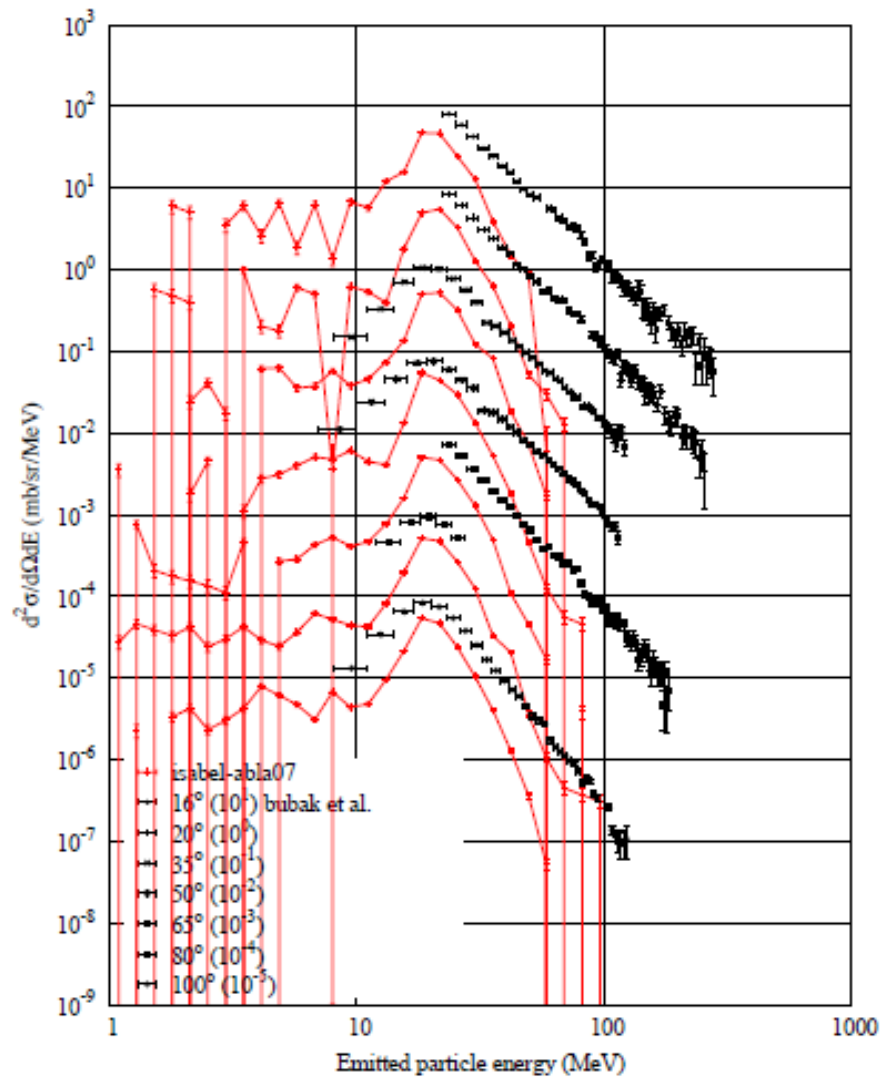
emitted-particle energy (MeV)

# p(2500 MeV) + Au – $^4\text{He}$ spectrum

## INCL45-ABLA07



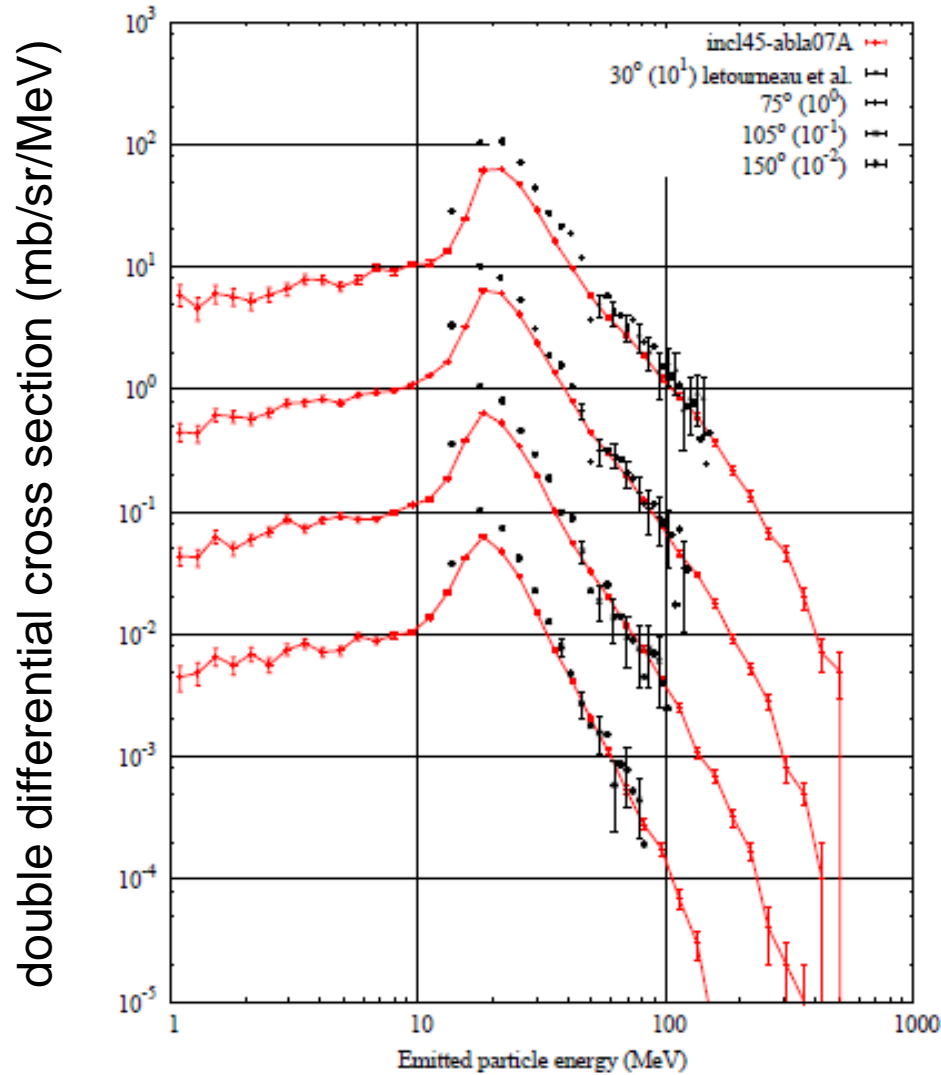
## ISABEL-ABLA07



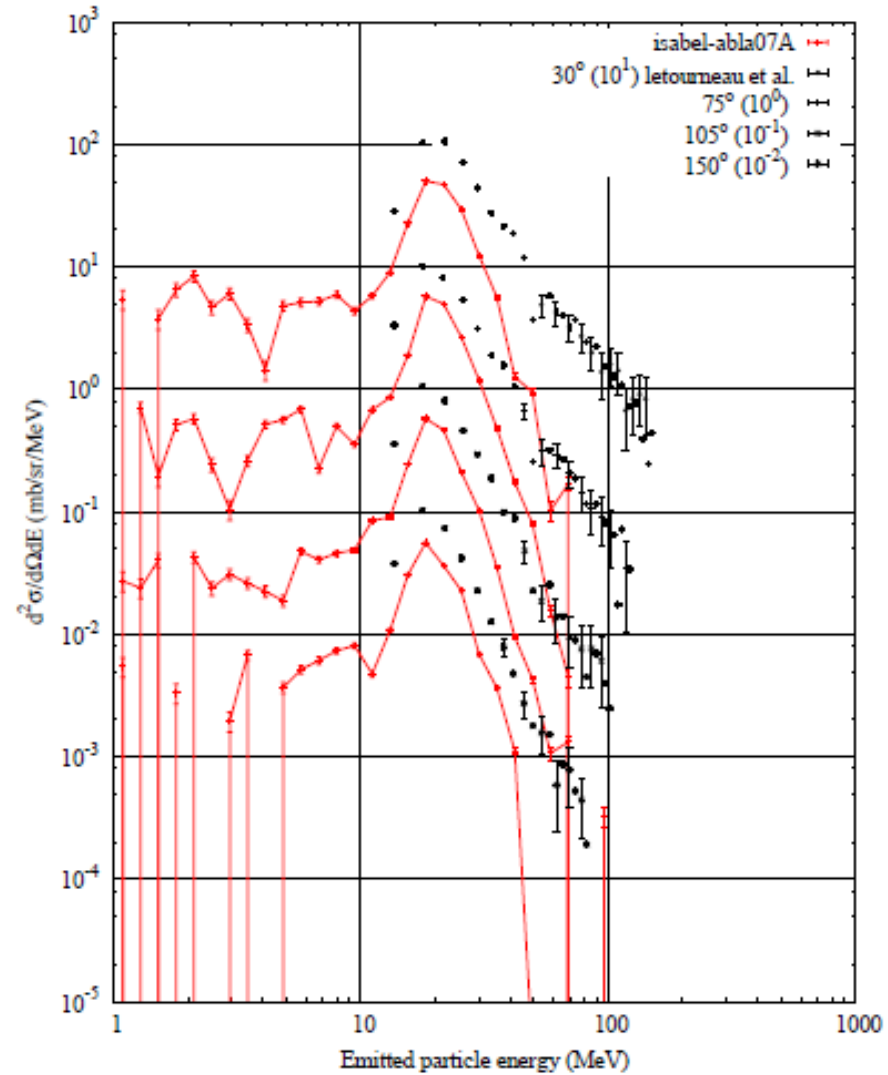
emitted-particle energy (MeV)

# p(2500 MeV) + Au – $^4\text{He}$ spectrum

## INCL45-ABLA07



## ISABEL-ABLA07



emitted-particle energy (MeV)

# Light Charged Particles

## LCP double differential cross sections (starting from $E_{proton} = 175$ MeV)

### Status

- Not so good with ISABEL + ABLA07
- Quite good with INCL45+ABLA07 for light target; not that good for heavy target
- Visible differences between INCL45 plus ABLA07 / SMM / GEMINI++
- But the spectra depend also on the INC (even in the low energy part of the spectrum)

**Improvement:** Complicate!

Concerning ABLA07:

- Tunneling through barrier (now taken into account only for calculating decay widths)
- Coulomb barriers (from Bass prescription) could be adjusted
- Apparently, in evaporation from deformed nuclei the emission could not be isotropic (to be demonstrated)

*The empirical nuclear potential of R. Bass*

$$-V_N(s) = \frac{C_1 \cdot C_2}{C_1 + C_2} \cdot \frac{1}{A \cdot \exp\left(\frac{s}{d_1}\right) + B \cdot \exp\left(\frac{s}{d_2}\right)}$$

$$A = 0.333 \text{ MeV}^{-1} \text{ fm},$$

$$d_1 = 3.5 \text{ fm},$$

$$B = 0.007 \text{ MeV}^{-1} \text{ fm},$$

$$d_2 = 0.65 \text{ fm}.$$

$$C_i = R_i \cdot \left(1 - \frac{(0.9984 \text{ fm})^2}{R_i^2}\right), \quad R_1 = \left(1.28 \cdot A_f^{1/3} - 0.76 + \frac{0.8}{A_f^{1/3}}\right) \text{ fm},$$

$$R_2 = \left(1.28 \cdot A_2^{1/3} - 0.76 + \frac{0.8}{A_2^{1/3}} + d\right) \text{ fm}, \quad d = \begin{cases} 3 \text{ fm}, & 1\text{H} \\ 0 \text{ fm}, & 2\text{H} \\ 0 \text{ fm}, & 3\text{H} \\ 0 \text{ fm}, & 3\text{He} \\ 1 \text{ fm}, & 4\text{He} \end{cases}$$

*Coulomb potential*

$$V_C = \begin{cases} 1.44 \cdot \frac{Z_1 \cdot Z_2}{r}, & r > R_C \\ 1.44 \cdot \frac{Z_1 \cdot Z_2}{2 \cdot R_C} \cdot \left(3 - \frac{r^2}{R_C^2}\right) & r \leq R_C \end{cases} \quad R_C = 1.3 \cdot (A_1^{1/3} + A_2^{1/3}) \text{ fm}$$

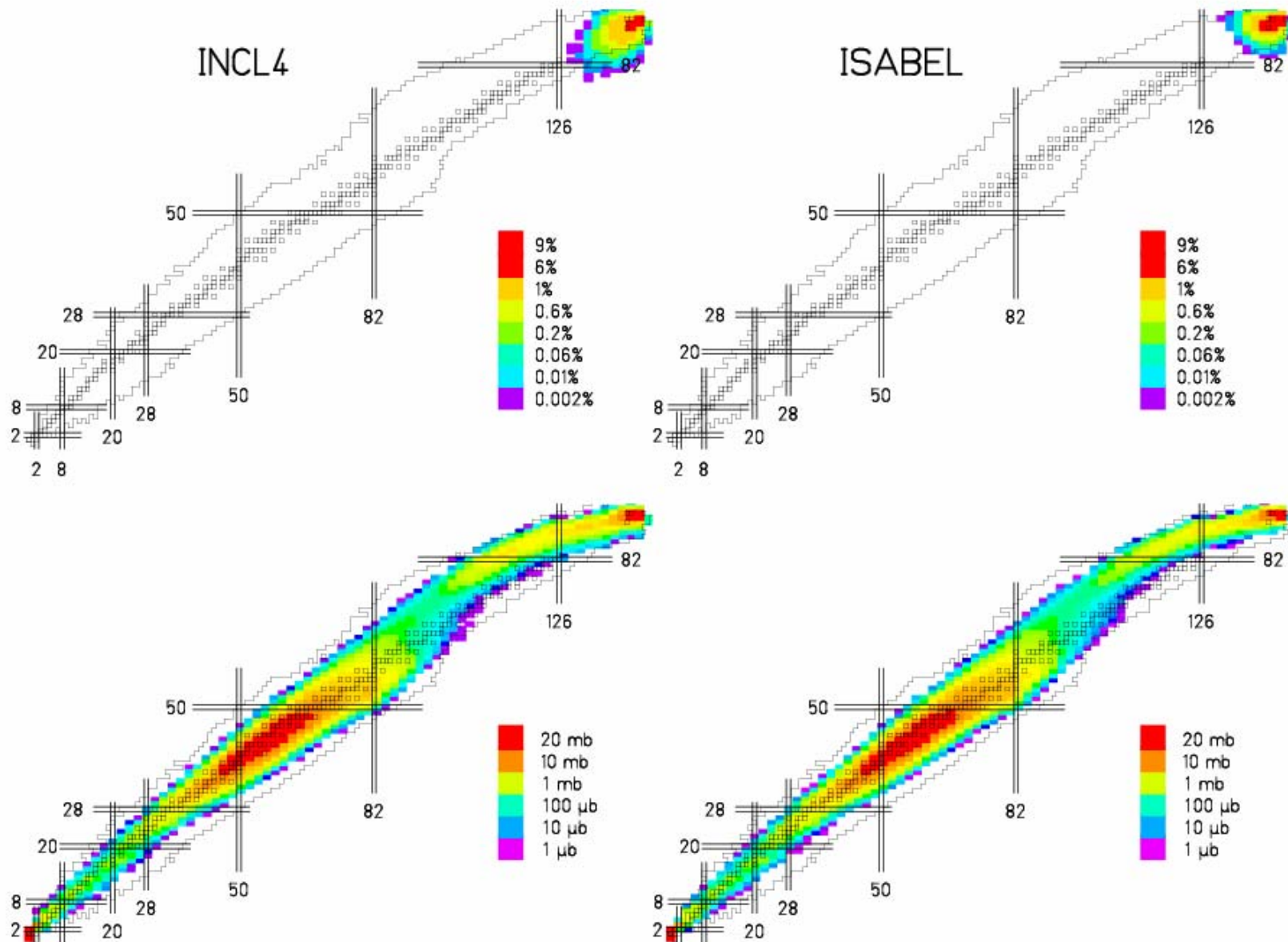
# Residues



# Fingerprints of the de-excitation process

1 GeV p +  $^{238}\text{U}$

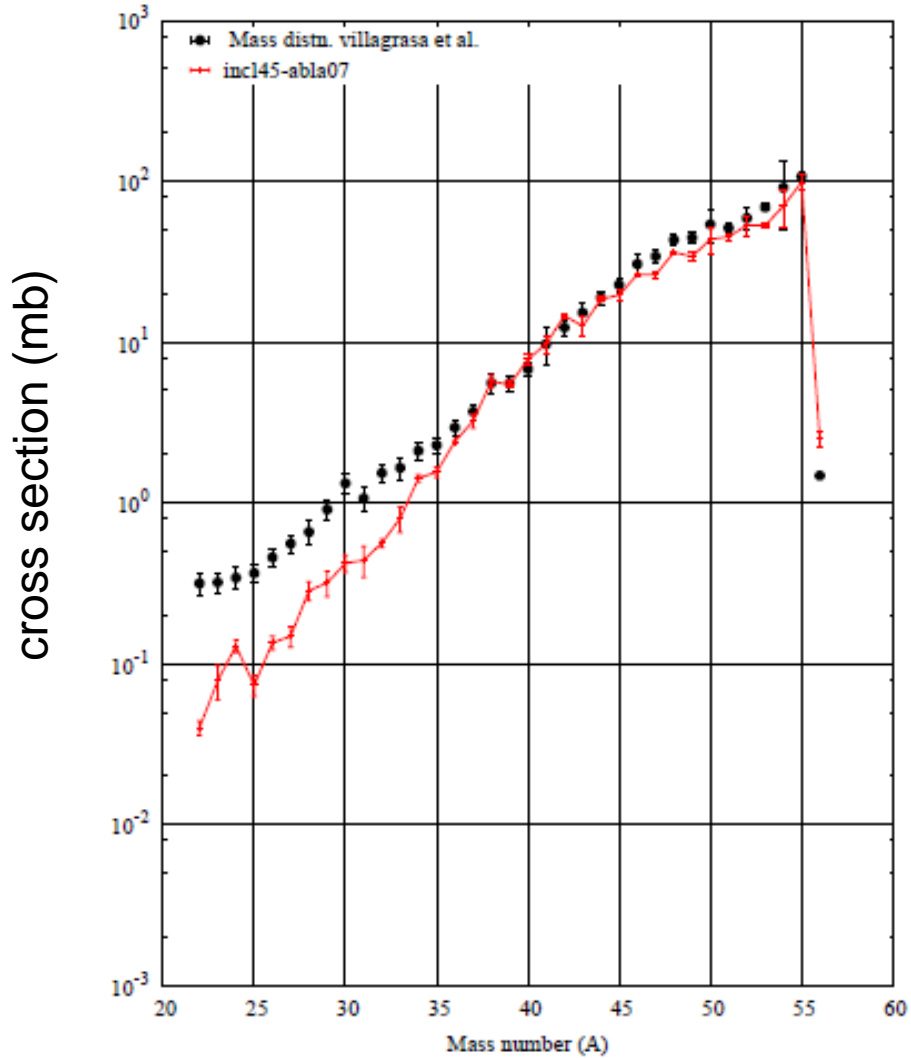
INCL4.5, ISABEL + ABLA07



# p(300 MeV) + $^{56}\text{Fe}$ – final residues

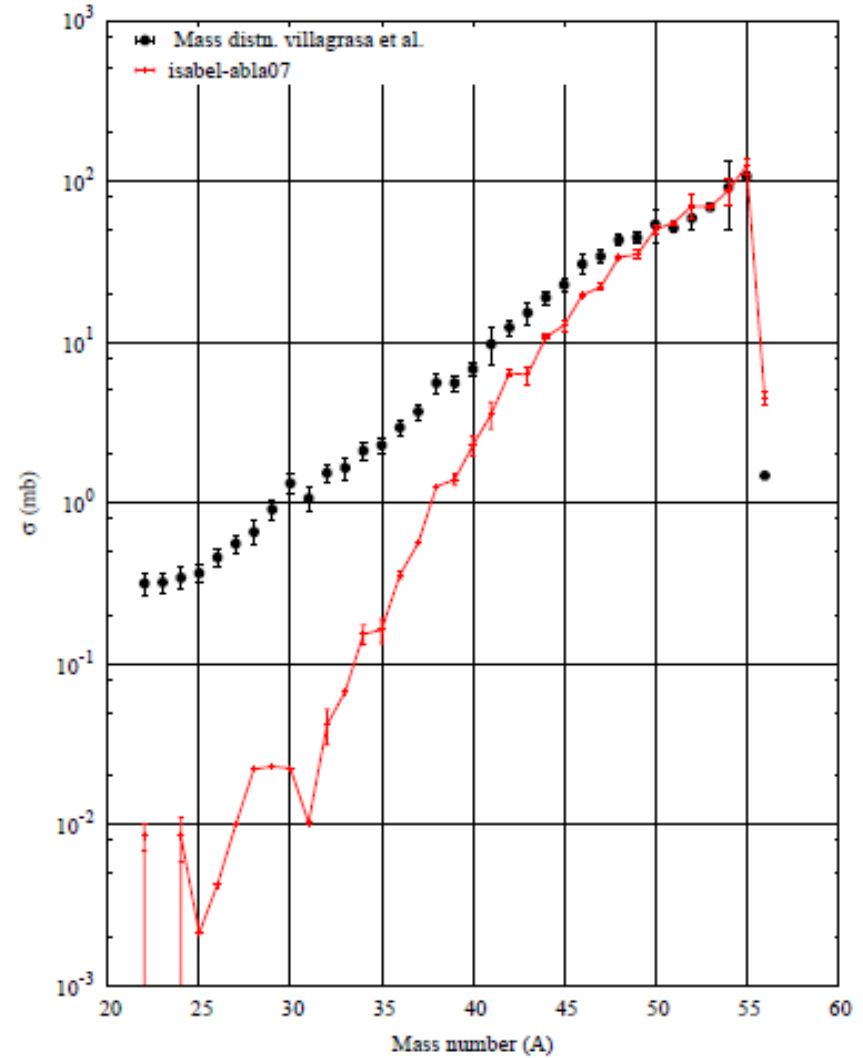
## INCL45-ABLA07

p (300 MeV) +  $^{56}\text{Fe}$  – Residue mass production



## ISABEL-ABLA07

p (300 MeV) +  $^{56}\text{Fe}$  – Residue mass production

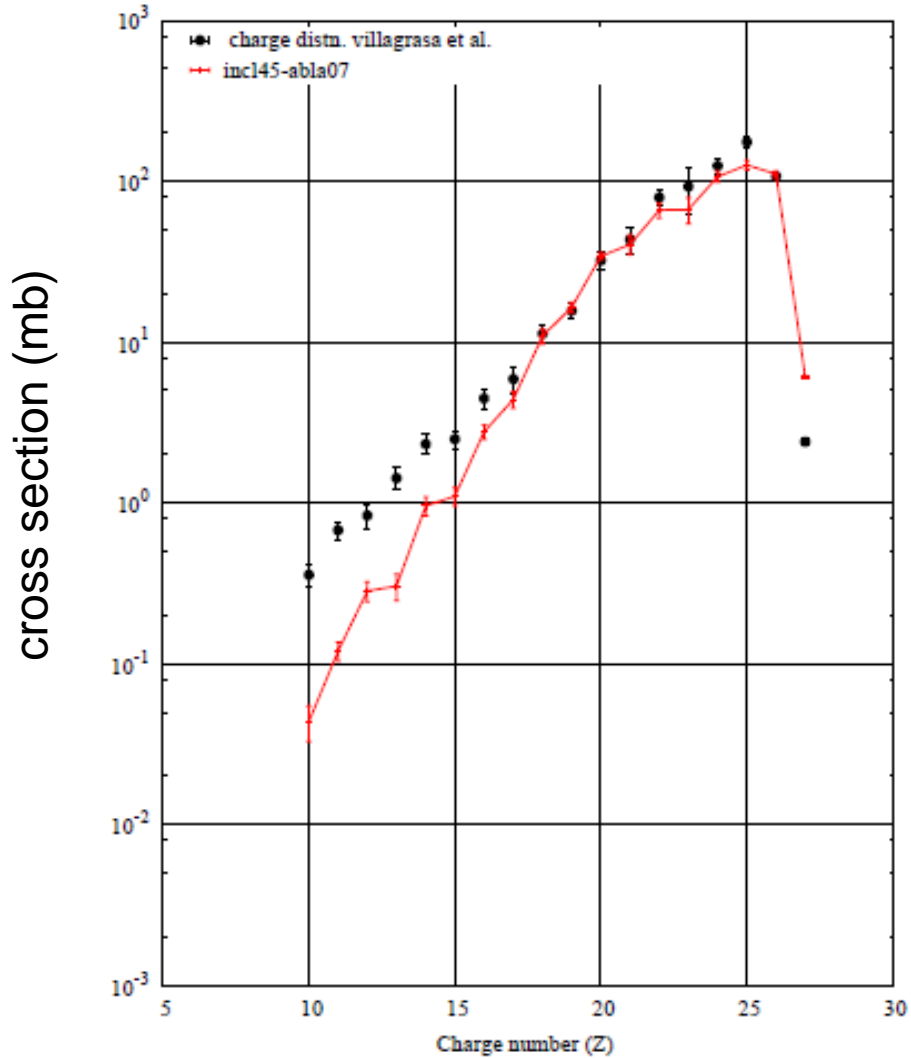


mass number A

# p(300 MeV) + $^{56}\text{Fe}$ – final residues

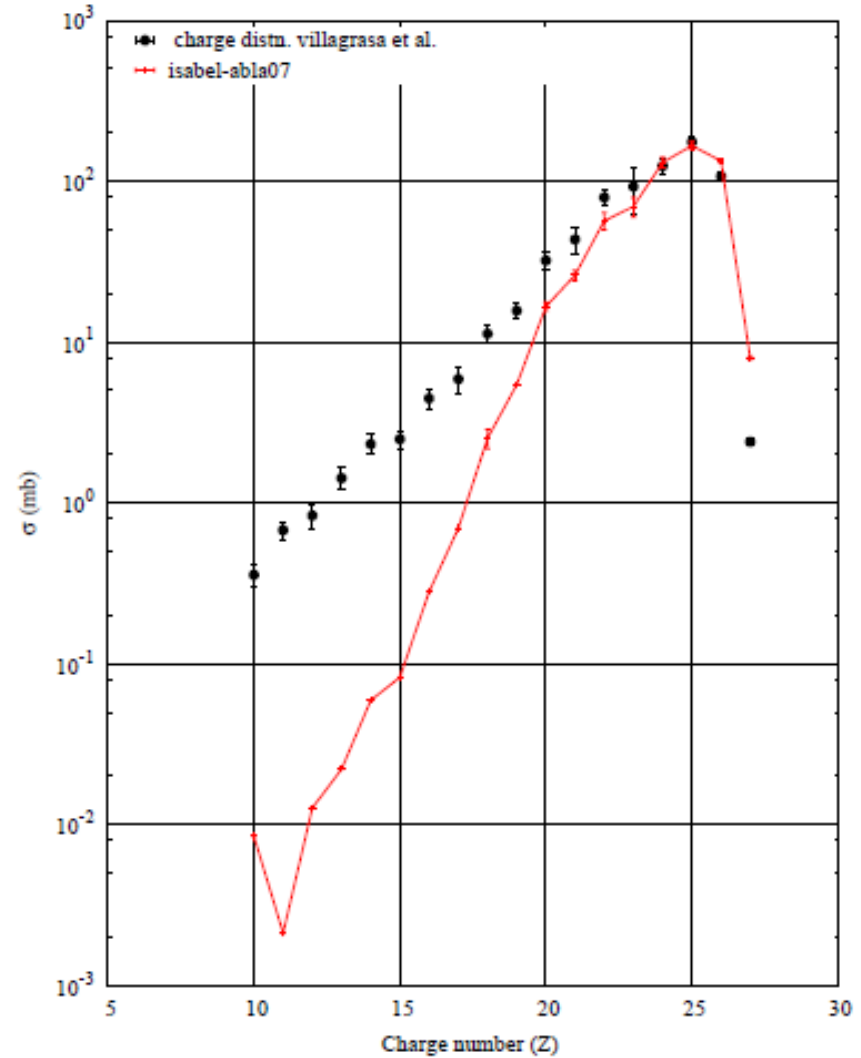
## INCL45-ABLA07

p (300 MeV) + Fe>0 -- Residue charge production



## ISABEL-ABLA07

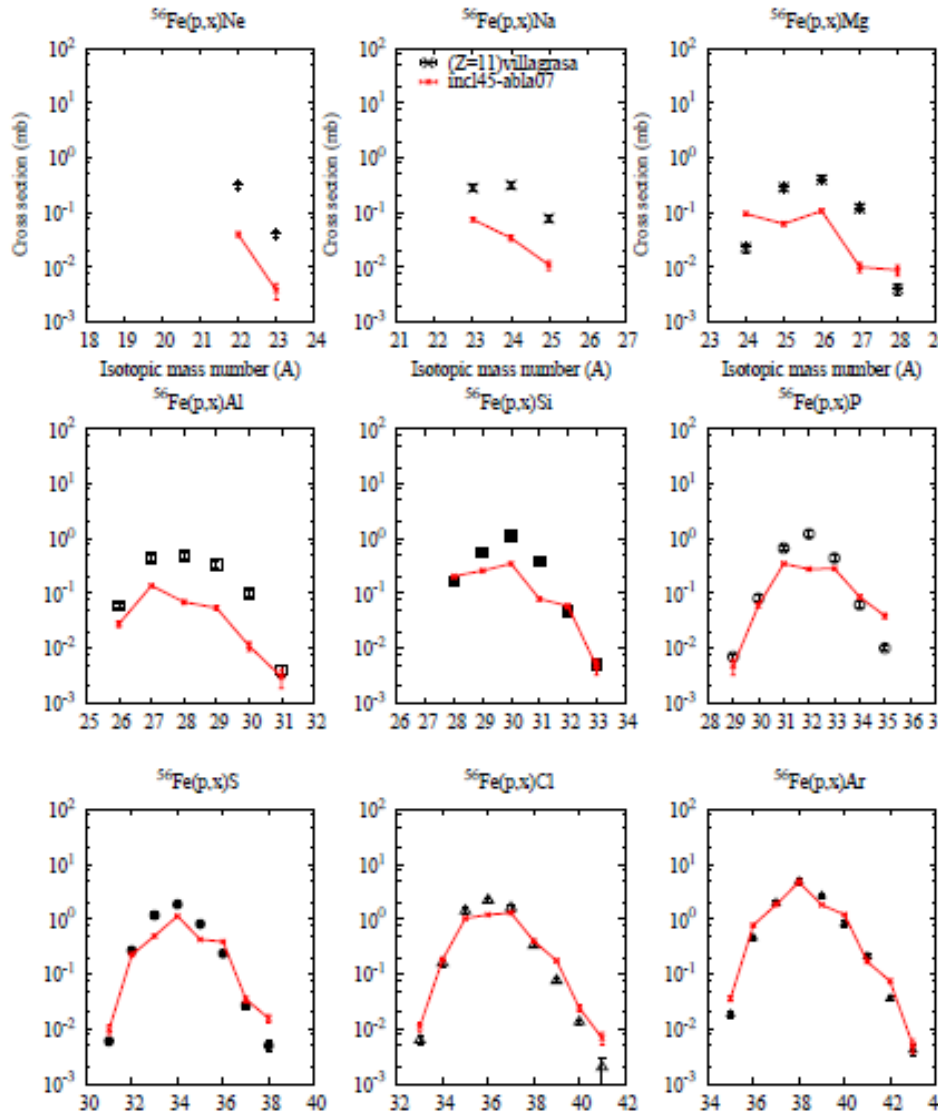
p (300 MeV) + Fe>0 -- Residue charge production



charge number Z

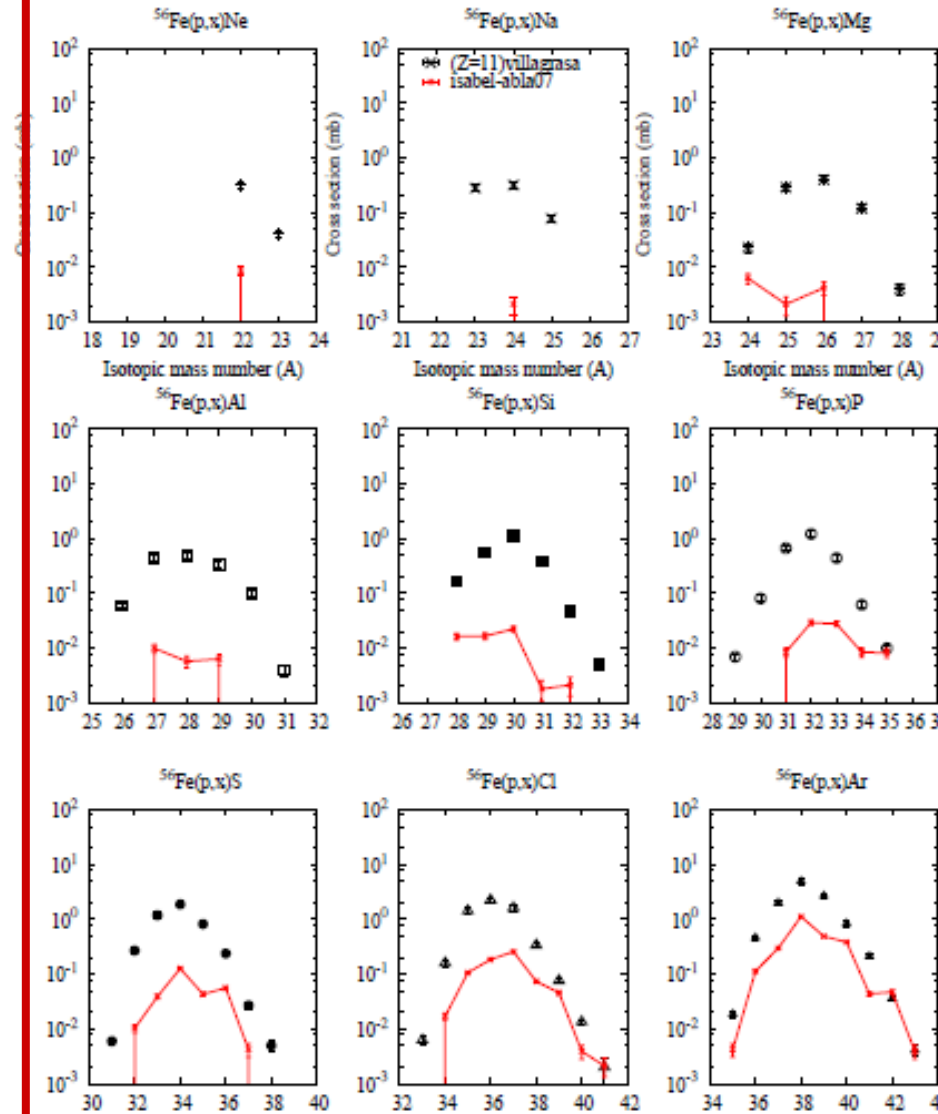
# p(300 MeV) + $^{56}\text{Fe}$ – final residues

## INCL45-ABLA07



mass number A

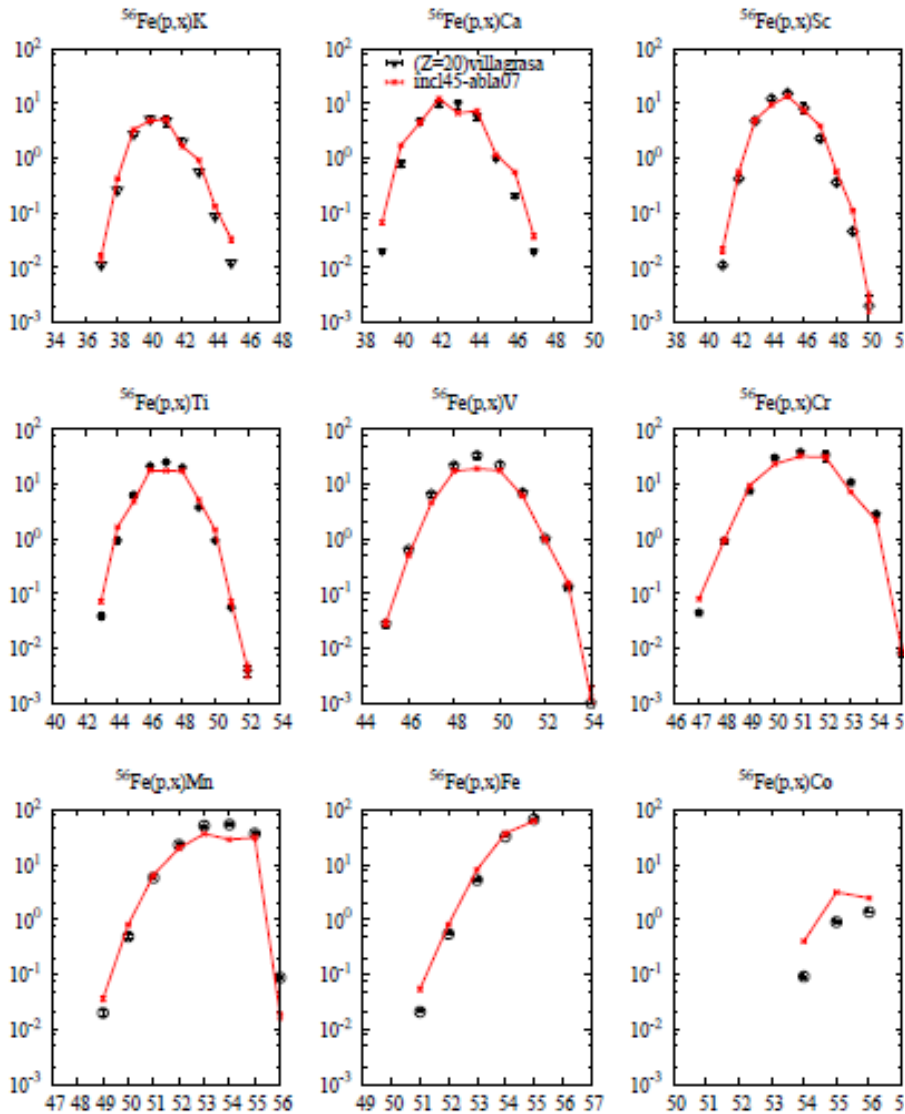
## ISABEL-ABLA07



mass number A

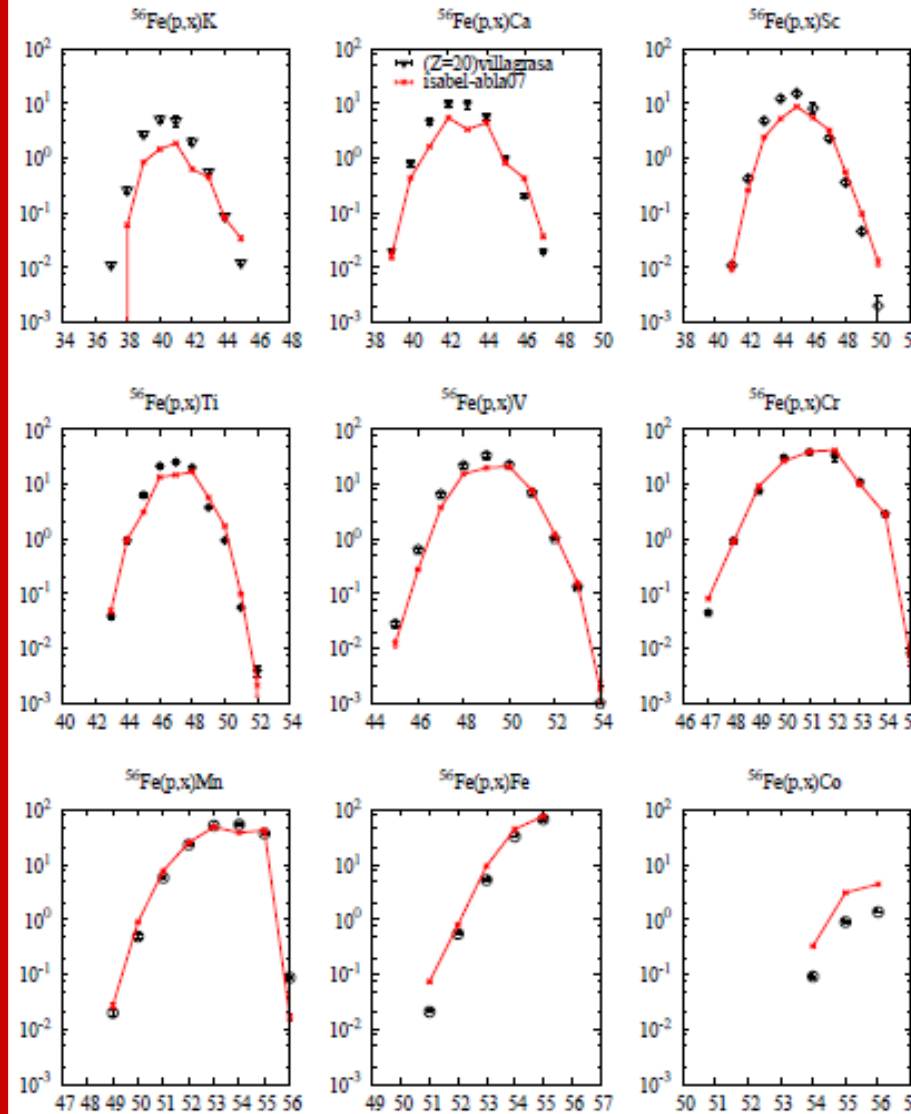
# $p(300 \text{ MeV}) + {}^{56}\text{Fe}$ – final residues

## INCL45-ABLA07



mass number A

## ISABEL-ABLA07

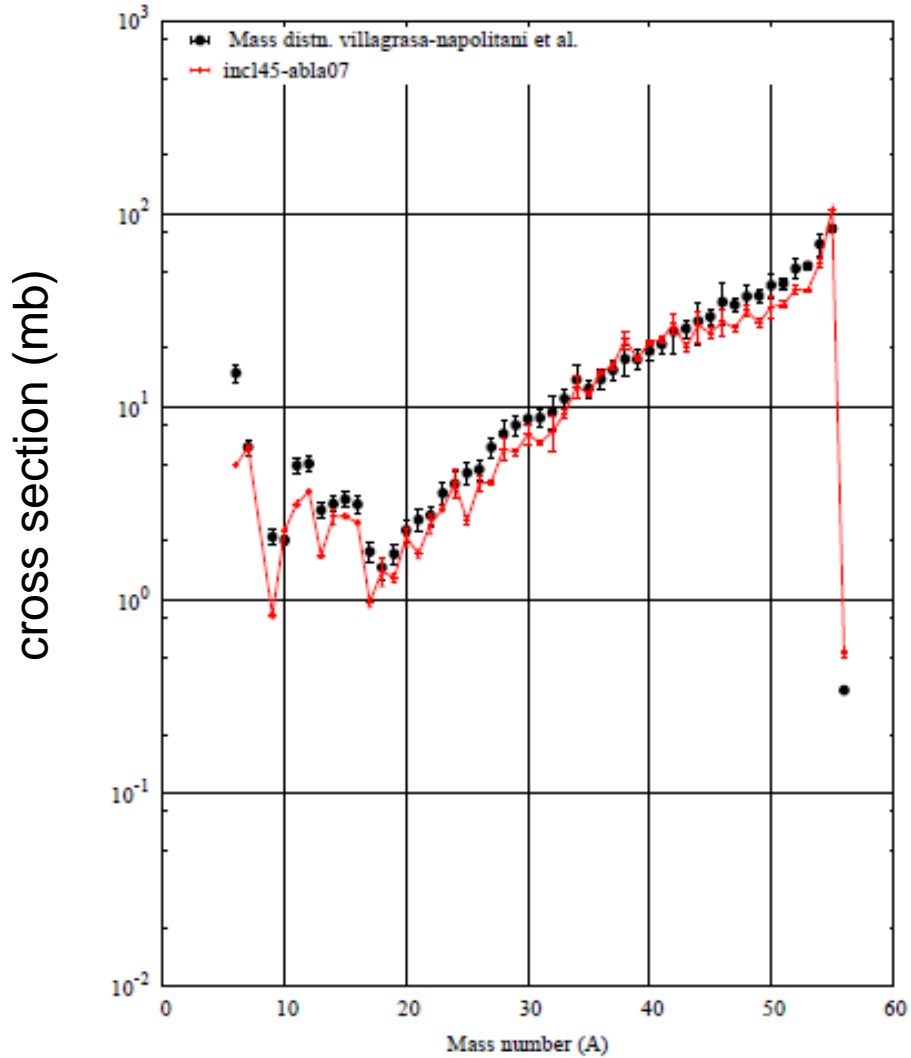


mass number A

# p(1000 MeV) + $^{56}\text{Fe}$ – final residues

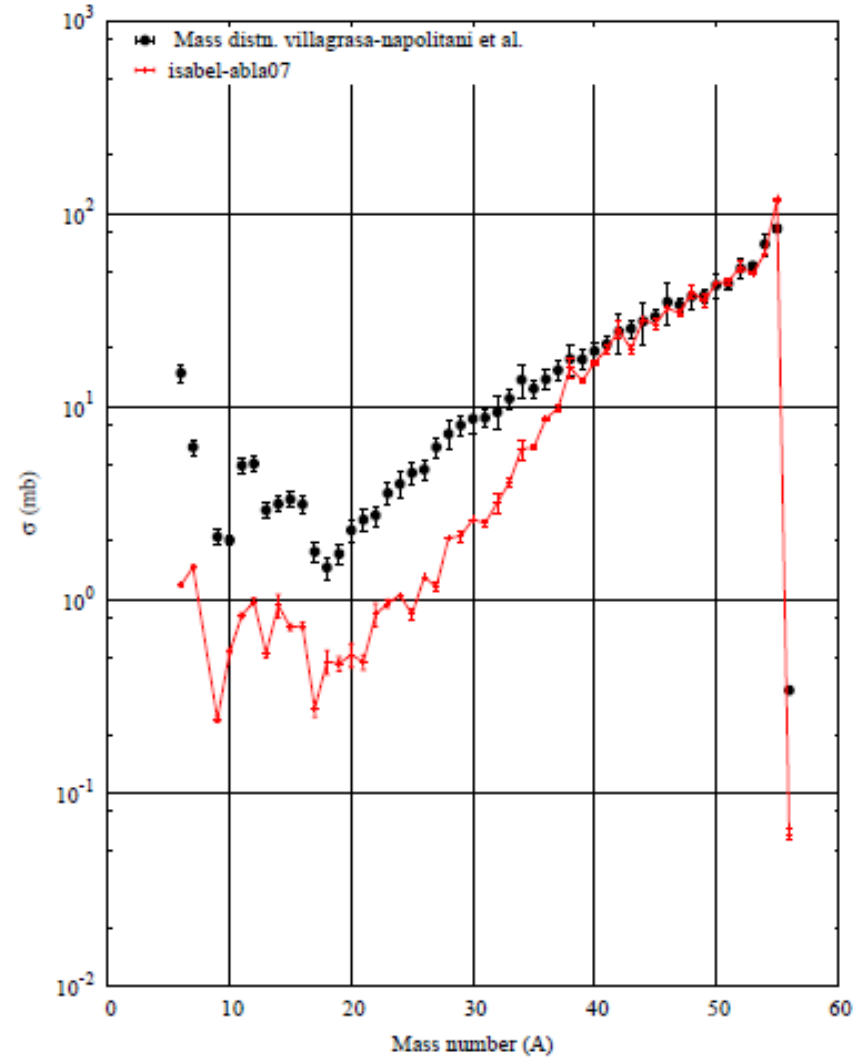
## INCL45-ABLA07

p (1000 MeV) +  $^{56}\text{Fe}$  -- Residue mass production



## ISABEL-ABLA07

p (1000 MeV) +  $^{56}\text{Fe}$  -- Residue mass production

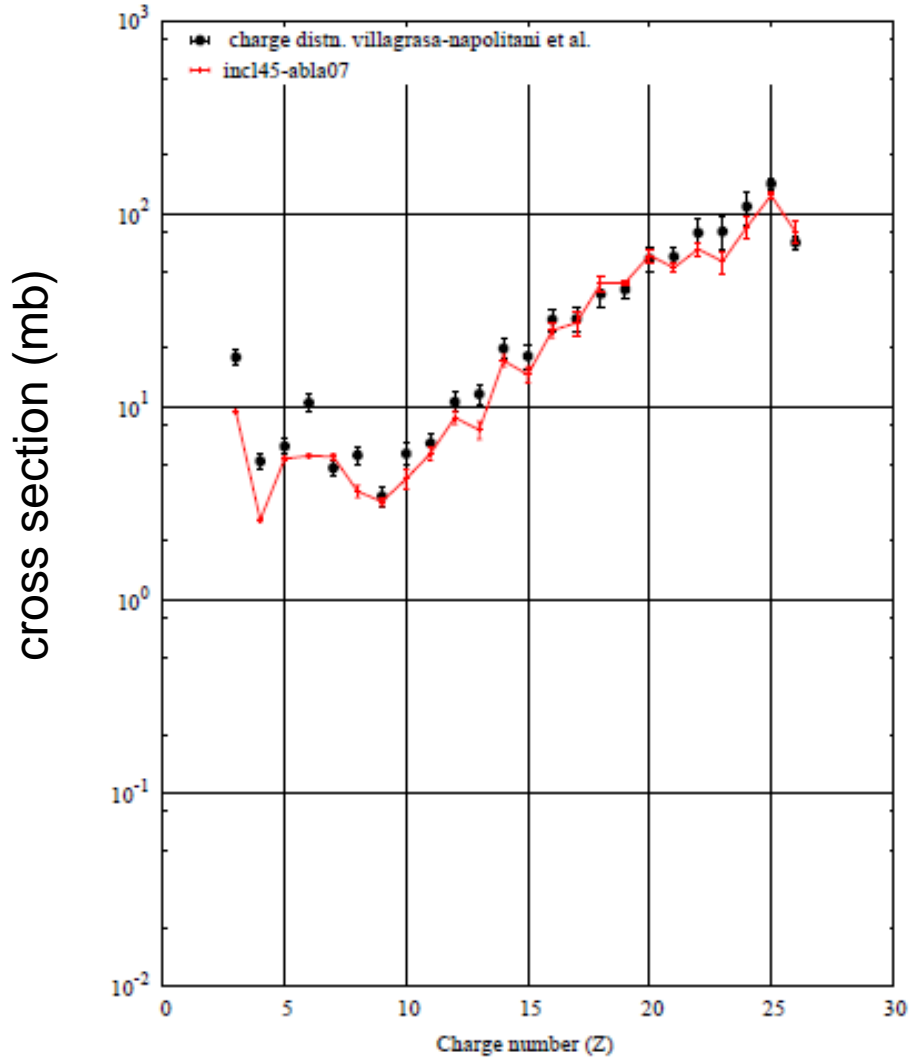


mass number A

# $p(1000 \text{ MeV}) + {}^{56}\text{Fe}$ – final residues

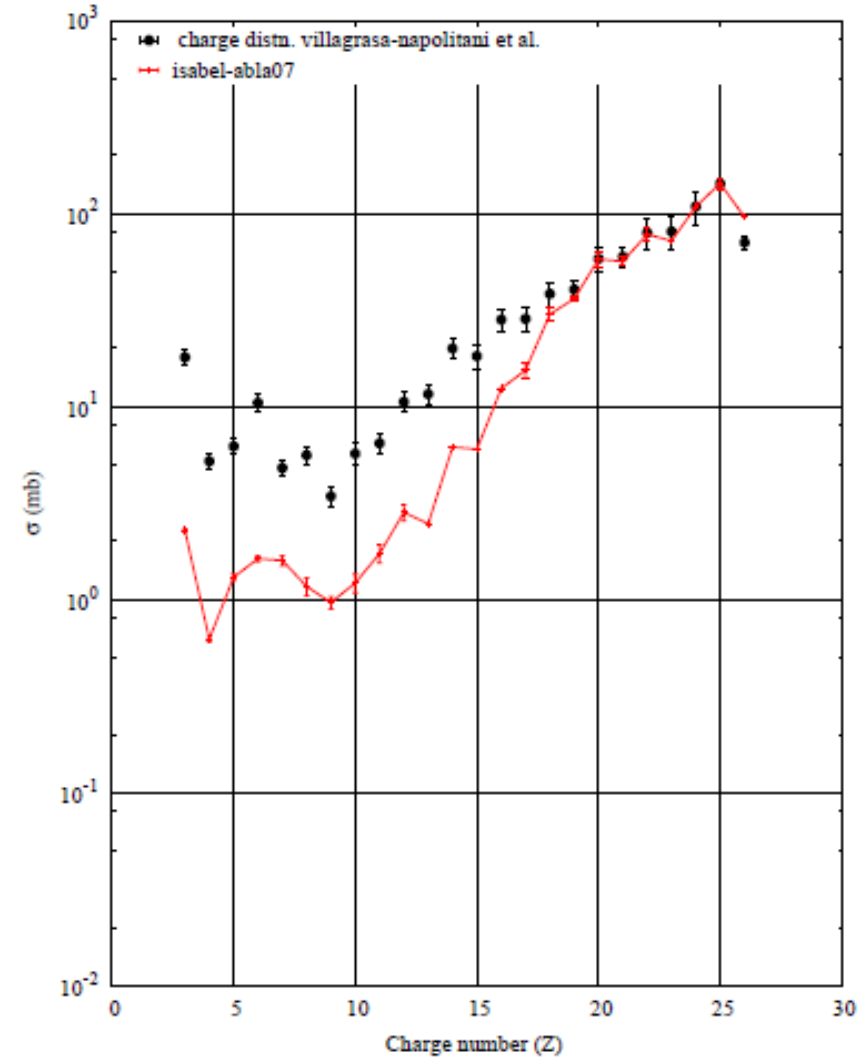
## INCL45-ABLA07

$p(1000 \text{ MeV}) + {}^{56}\text{Fe}$  – Residue charge production



## ISABEL-ABLA07

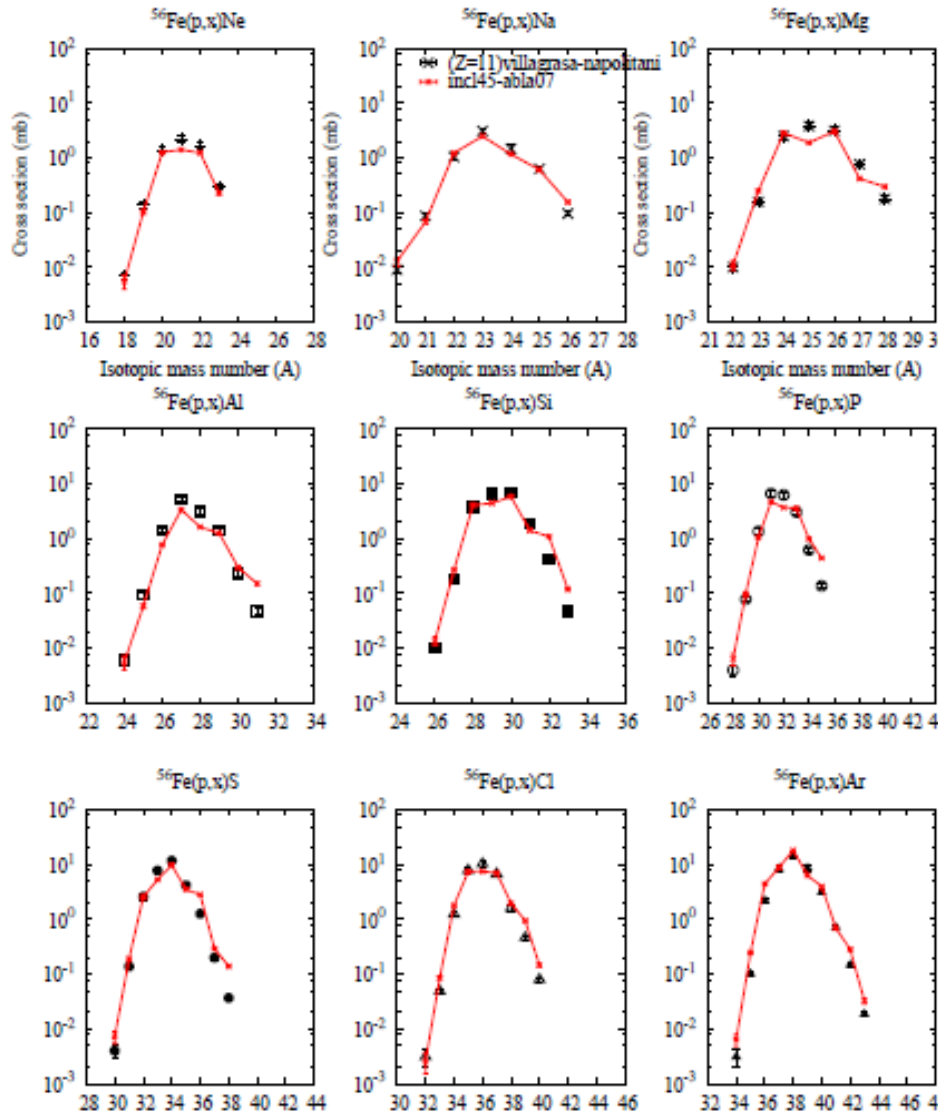
$p(1000 \text{ MeV}) + {}^{56}\text{Fe}$  – Residue charge production



charge number Z

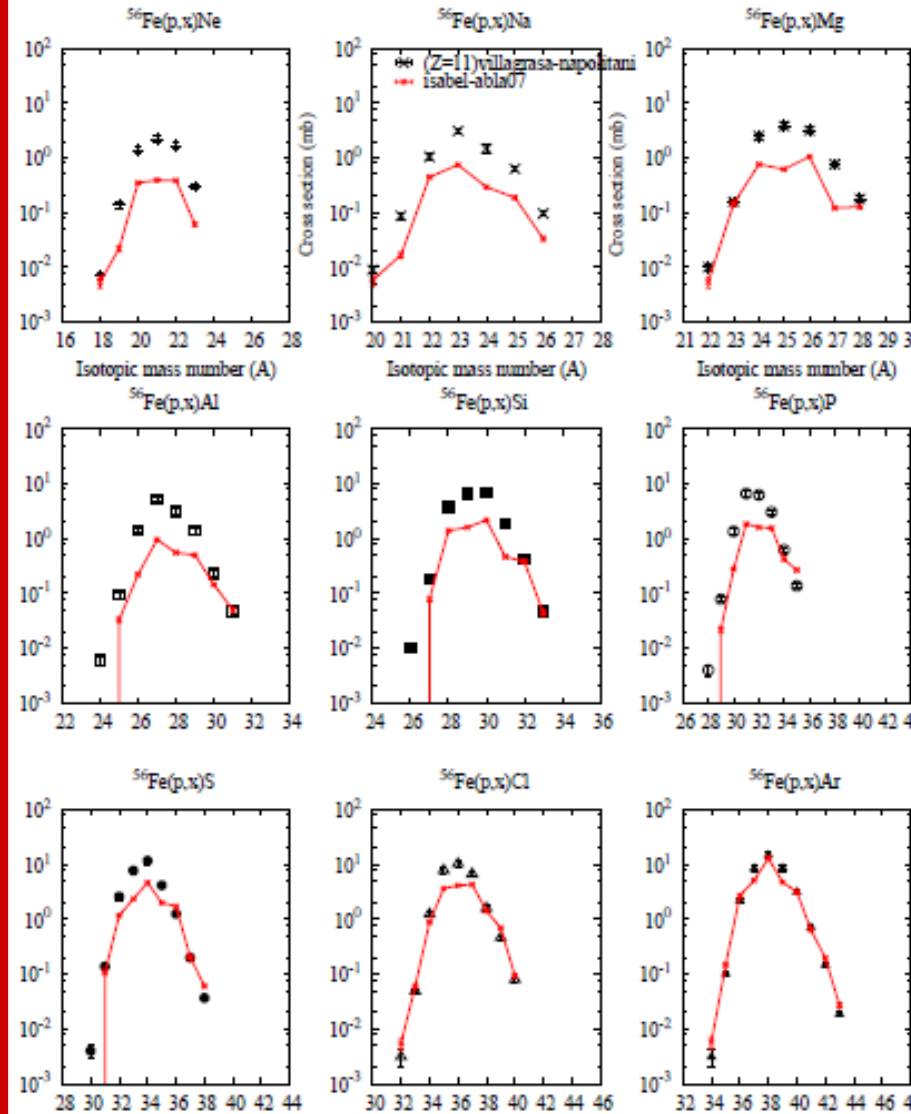
# $p(1000 \text{ MeV}) + {}^{56}\text{Fe}$ – final residues

## INCL45-ABLA07



mass number A

## ISABEL-ABLA07

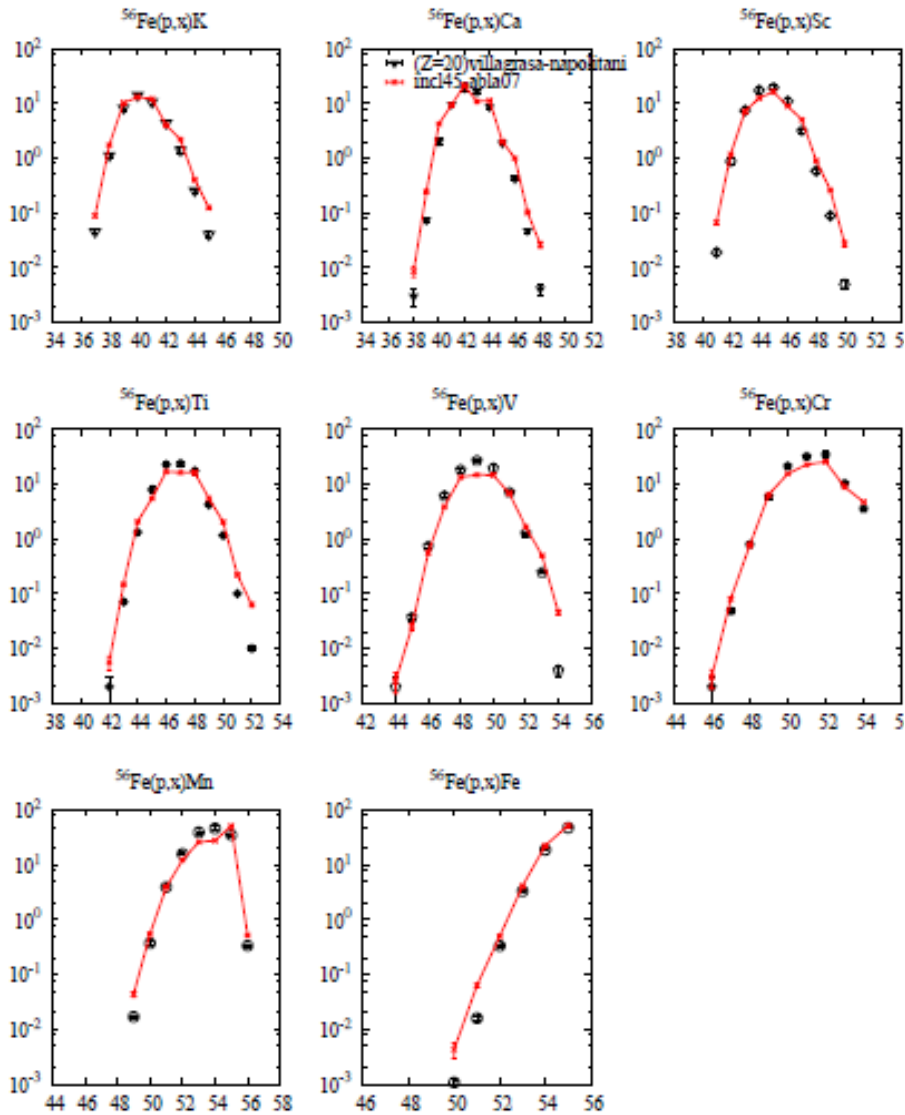


mass number A



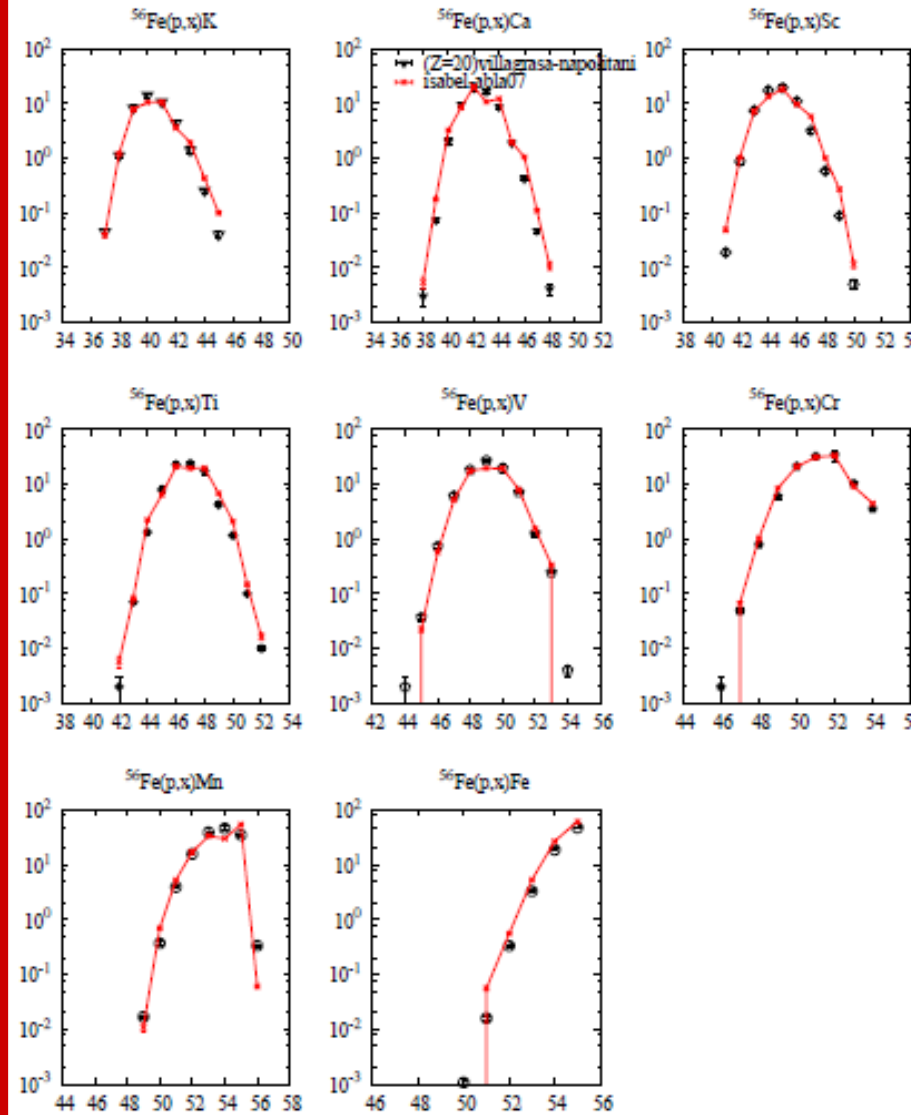
# $p(1000 \text{ MeV}) + {}^{56}\text{Fe}$ – final residues

## INCL45-ABLA07



mass number A

## ISABEL-ABLA07

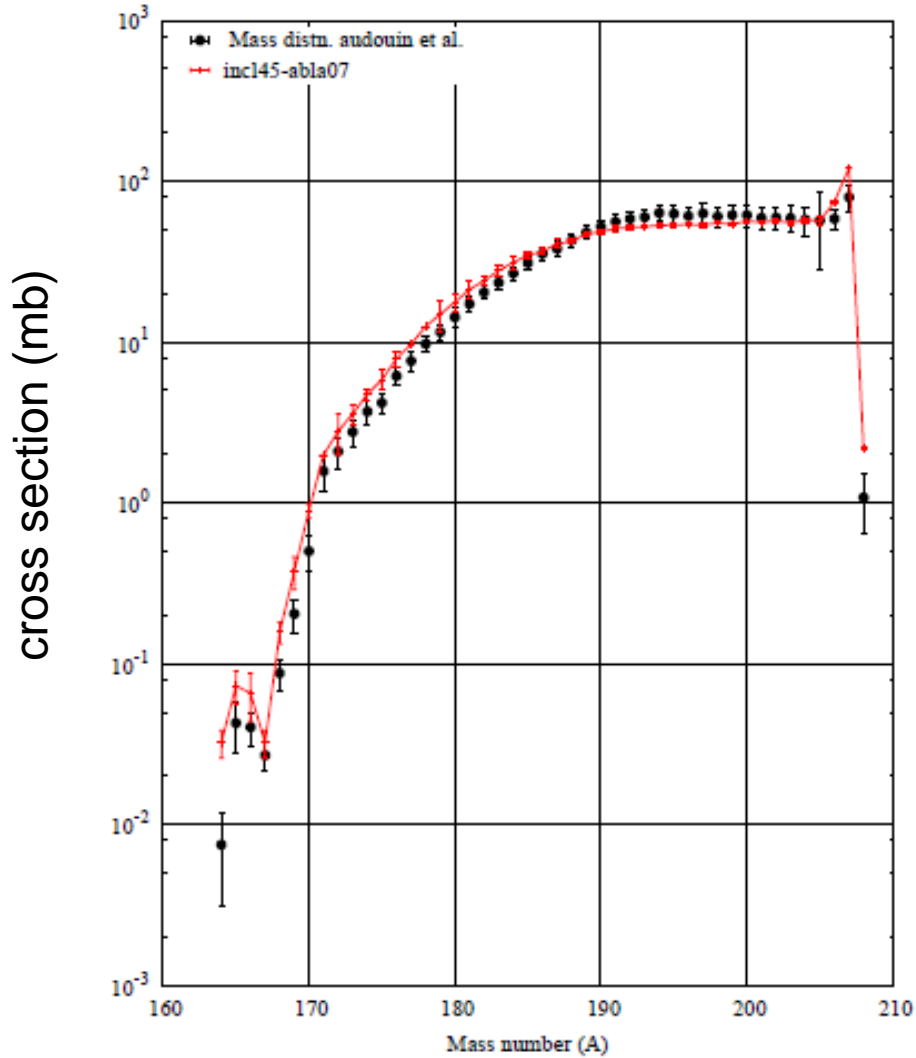


mass number A

# p(500 MeV) + $^{208}\text{Pb}$ – final residues

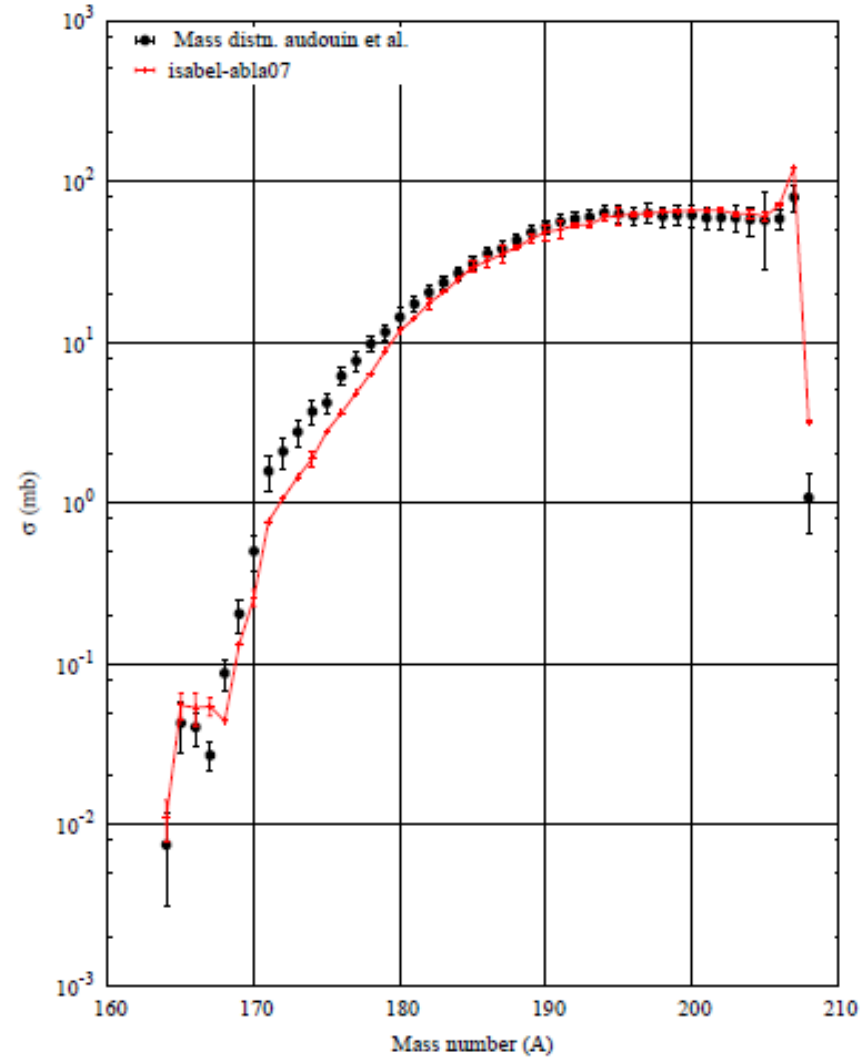
## INCL45-ABLA07

p (500 MeV) + Pb208 -- Residue mass production



## ISABEL-ABLA07

p (500 MeV) + Pb208 -- Residue mass production

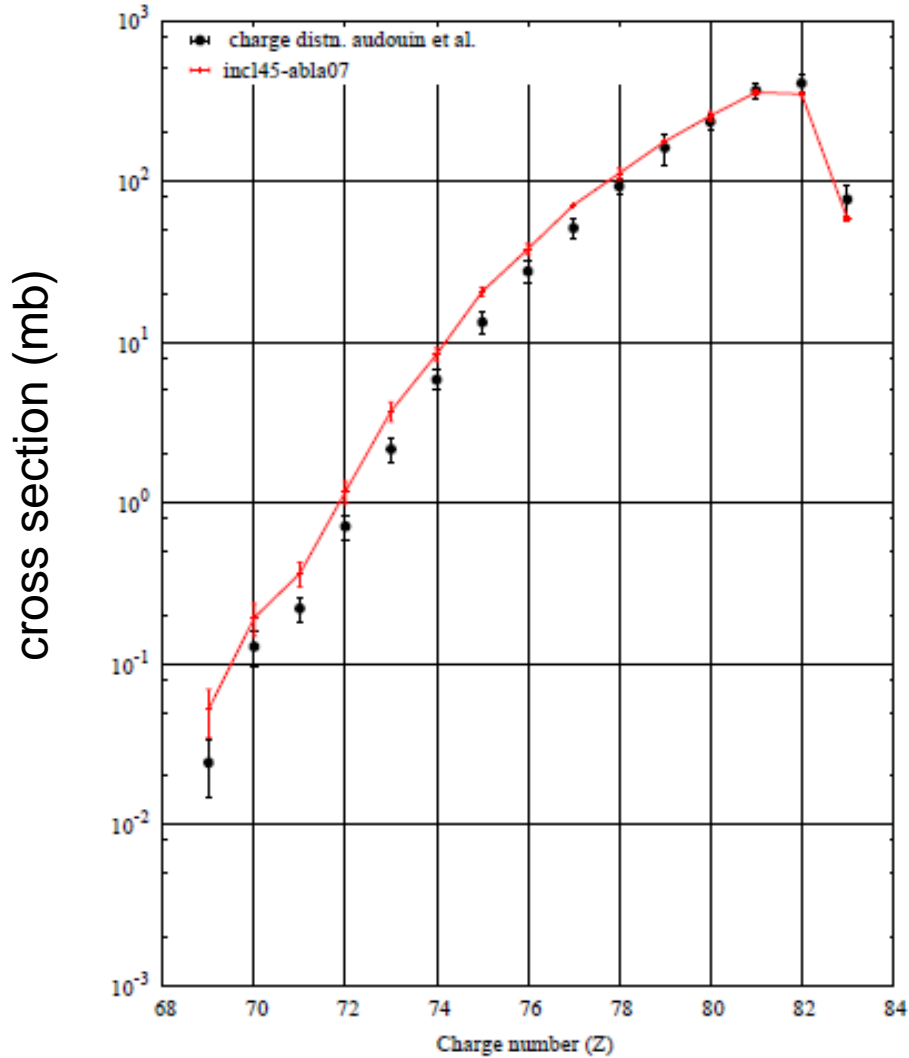


mass number A

# p(500 MeV) + $^{208}\text{Pb}$ – final residues

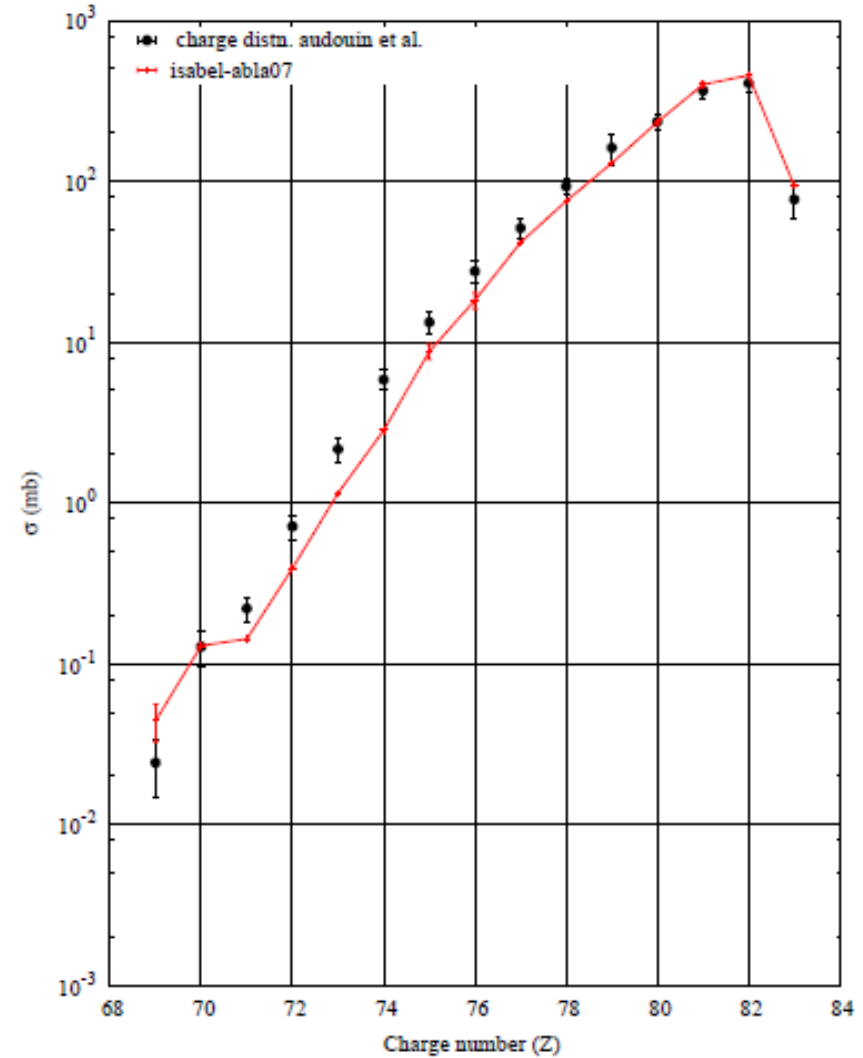
## INCL45-ABLA07

p (>500 MeV) + Pb208 -- Residue charge production



## ISABEL-ABLA07

p (>500 MeV) + Pb208 -- Residue charge production

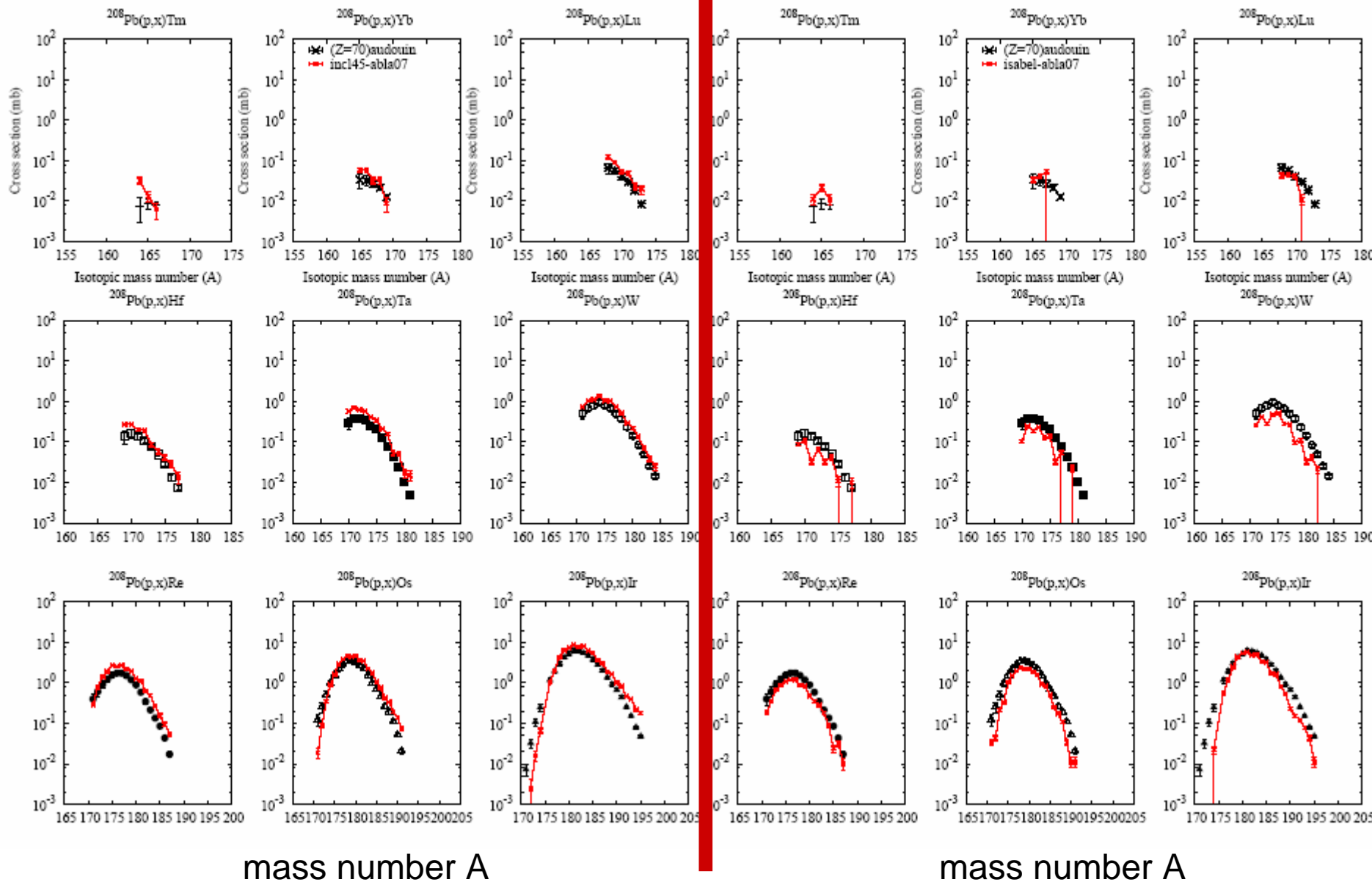


charge number Z

# $p(500 \text{ MeV}) + {}^{208}\text{Pb}$ – final residues

## INCL45-ABLA07

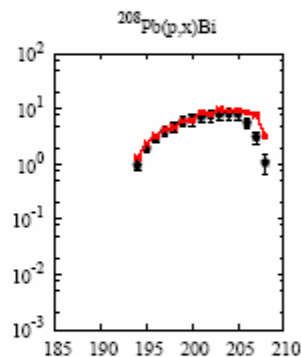
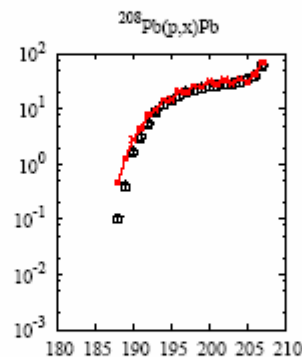
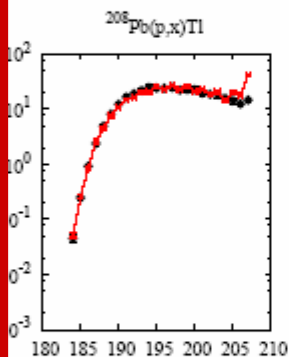
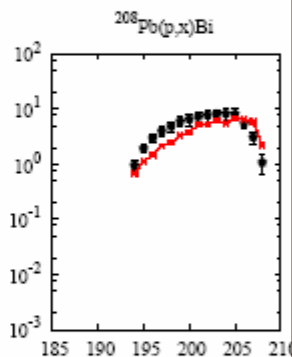
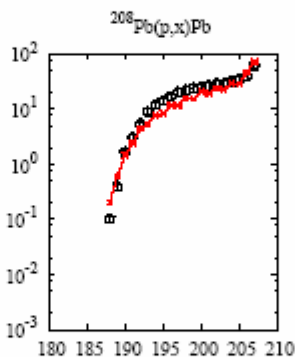
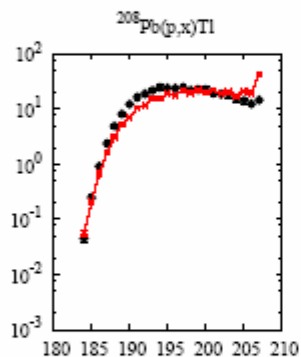
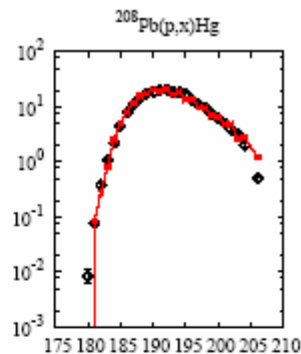
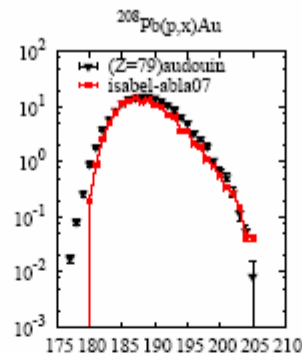
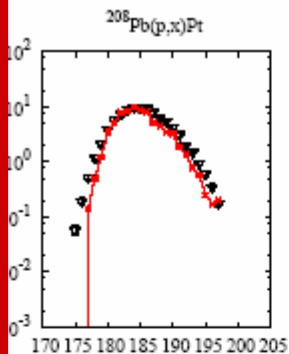
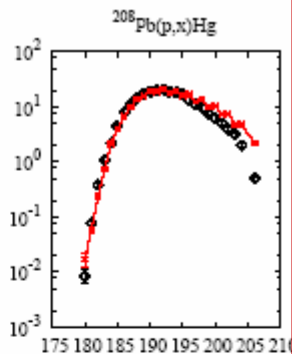
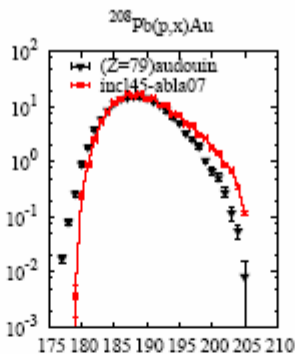
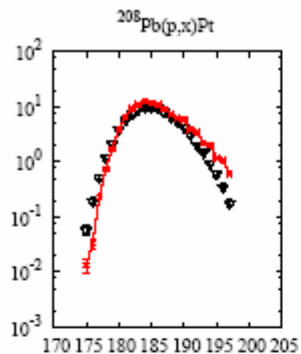
## ISABEL-ABLA07



# $p(500 \text{ MeV}) + {}^{208}\text{Pb}$ – final residues

## INCL45-ABLA07

## ISABEL-ABLA07



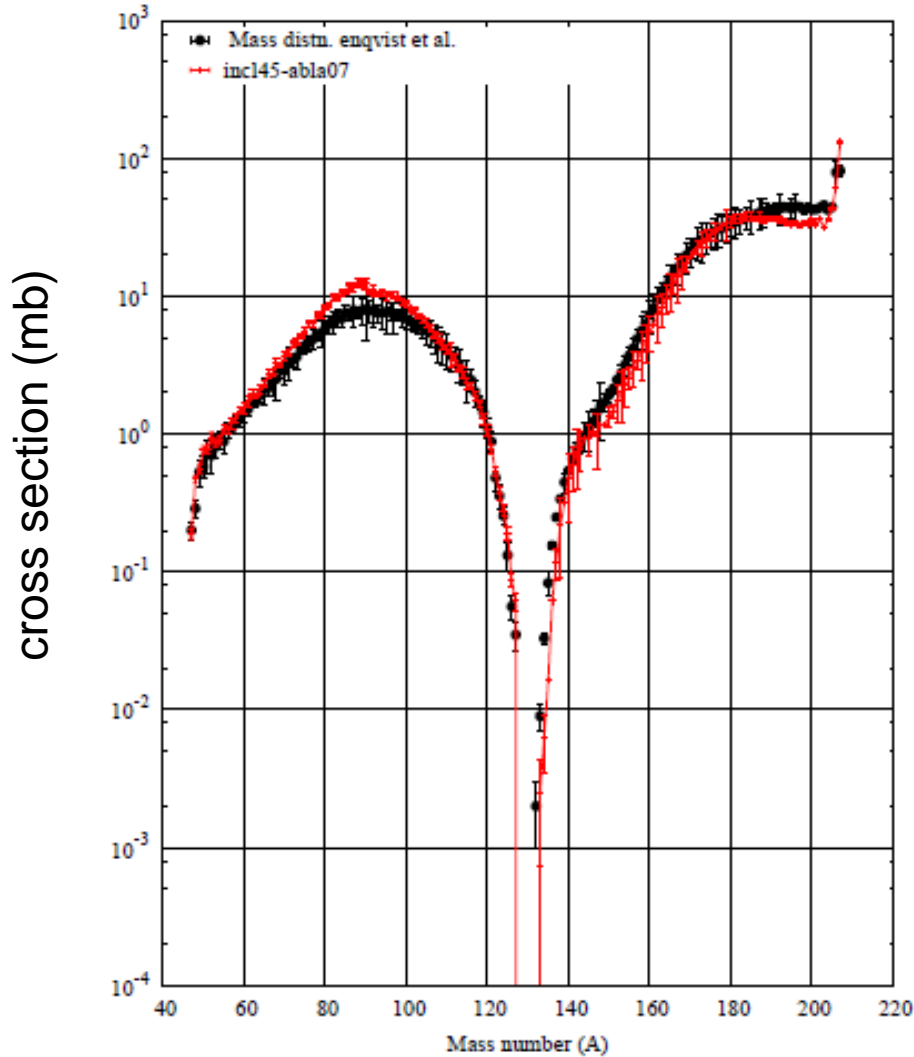
mass number A

mass number A

# p(1000 MeV) + $^{208}\text{Pb}$ – final residues

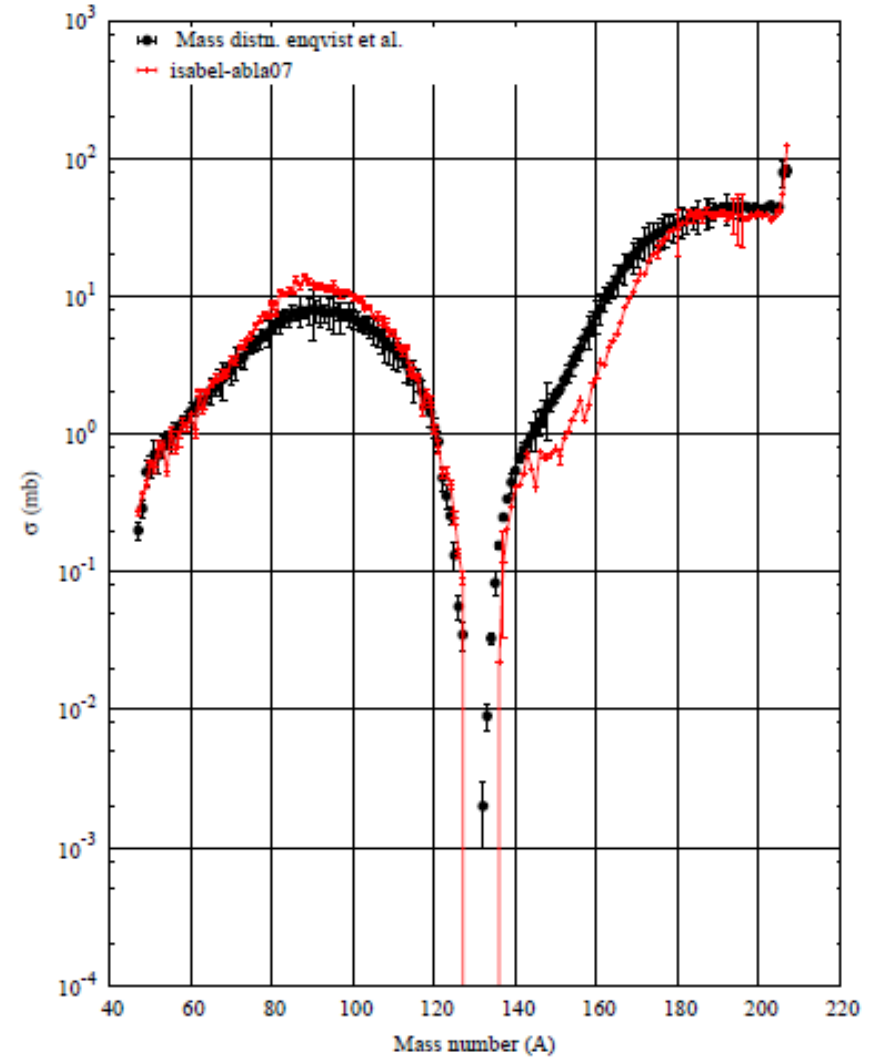
## INCL45-ABLA07

p (1000 MeV) +  $\text{Pb}208$  – Residue mass production



## ISABEL-ABLA07

p (1000 MeV) +  $\text{Pb}208$  – Residue mass production

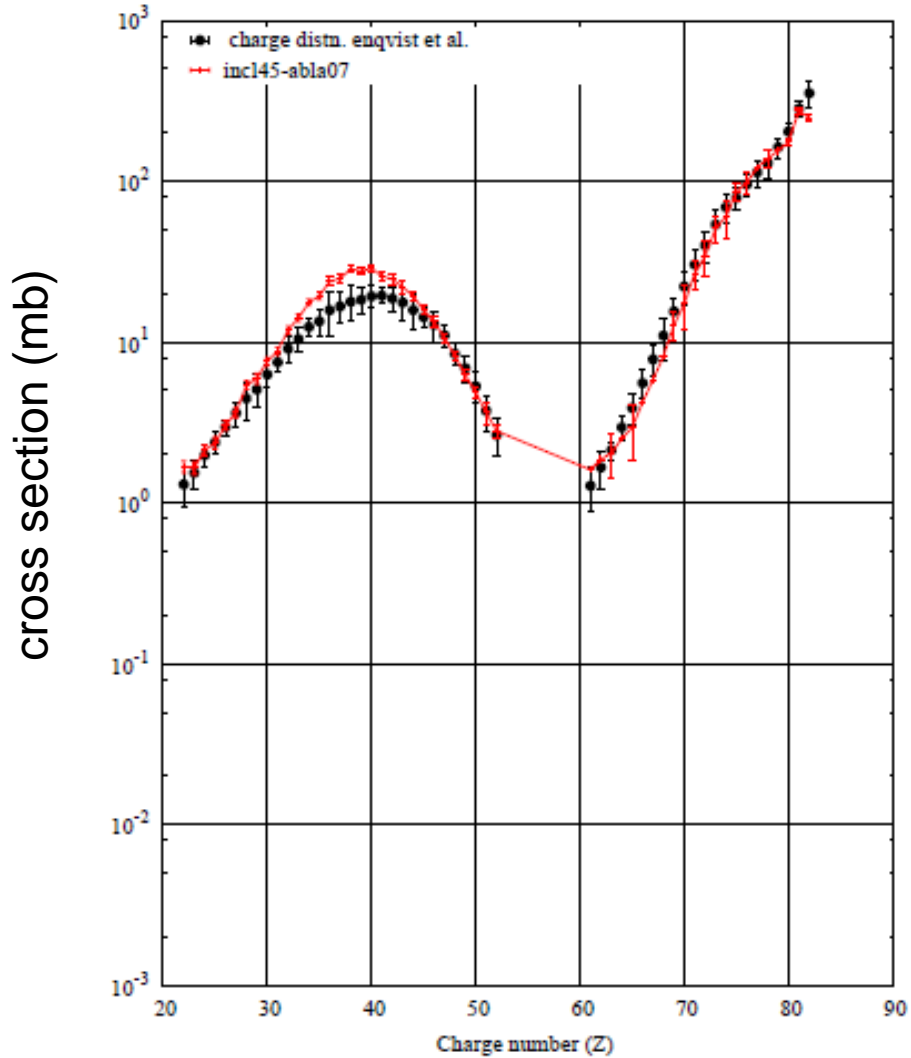


mass number A

# p(1000 MeV) + $^{208}\text{Pb}$ – final residues

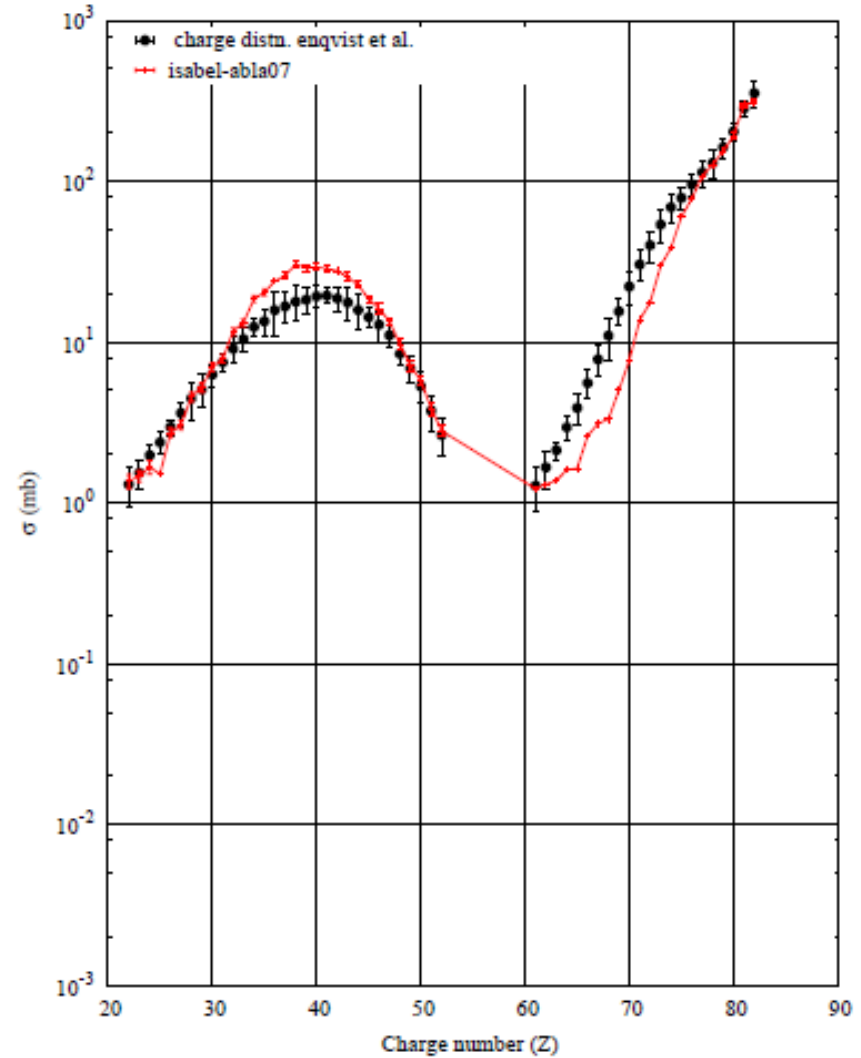
## INCL45-ABLA07

p (1000 MeV) + Pb208 – Residue charge production



## ISABEL-ABLA07

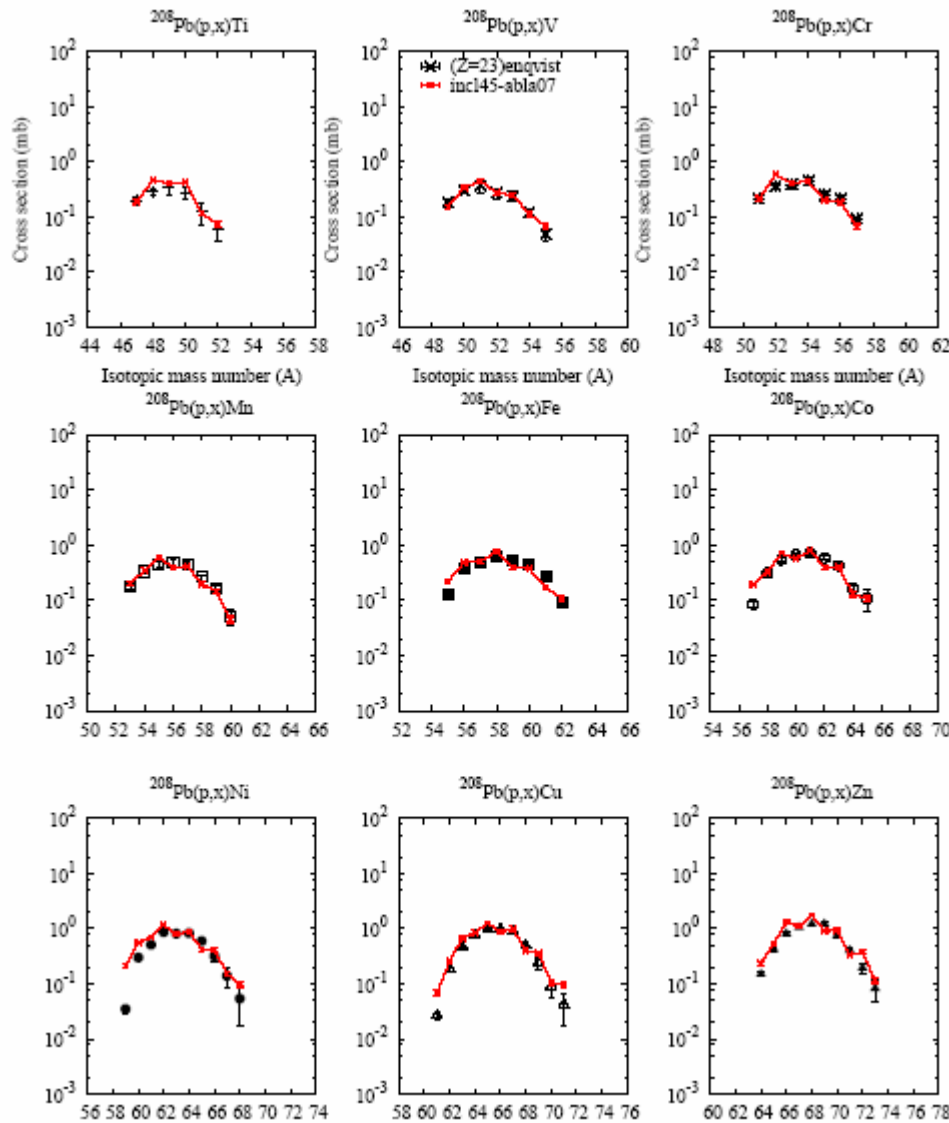
p (1000 MeV) + Pb208 – Residue charge production



charge number Z

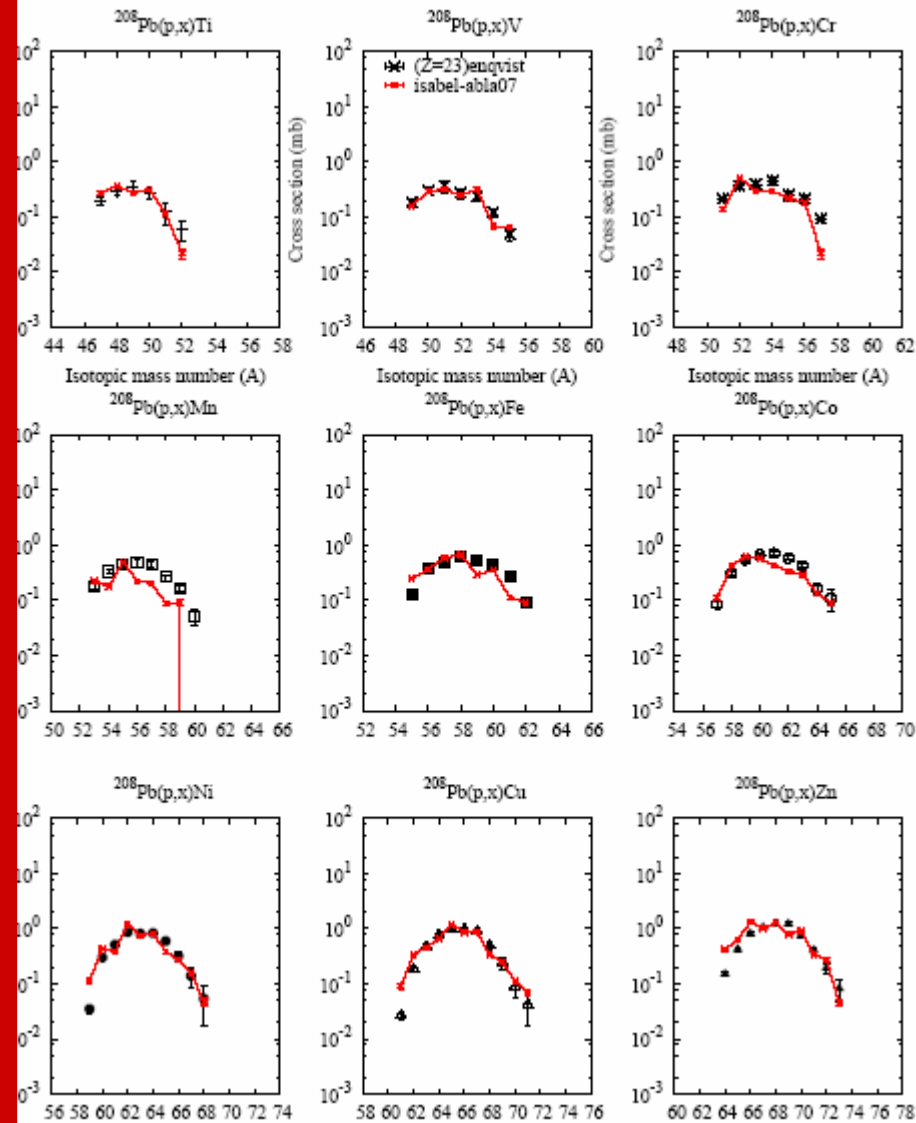
# $p(1000 \text{ MeV}) + {}^{208}\text{Pb}$ – final residues

## INCL45-ABLA07



mass number A

## ISABEL-ABLA07



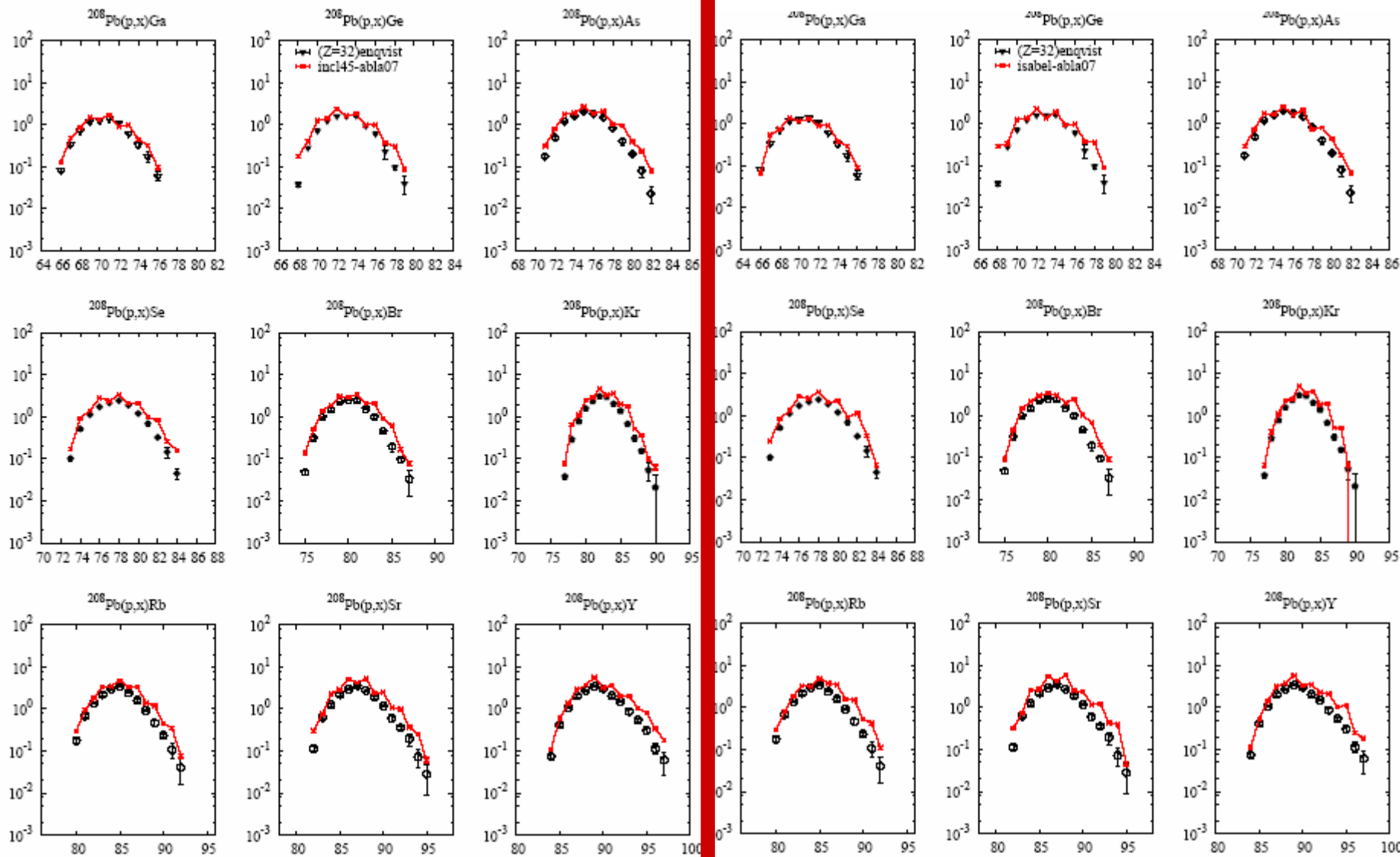
mass number A



# $p(1000 \text{ MeV}) + {}^{208}\text{Pb}$ – final residues

## INCL45-ABLA07

## ISABEL-ABLA07



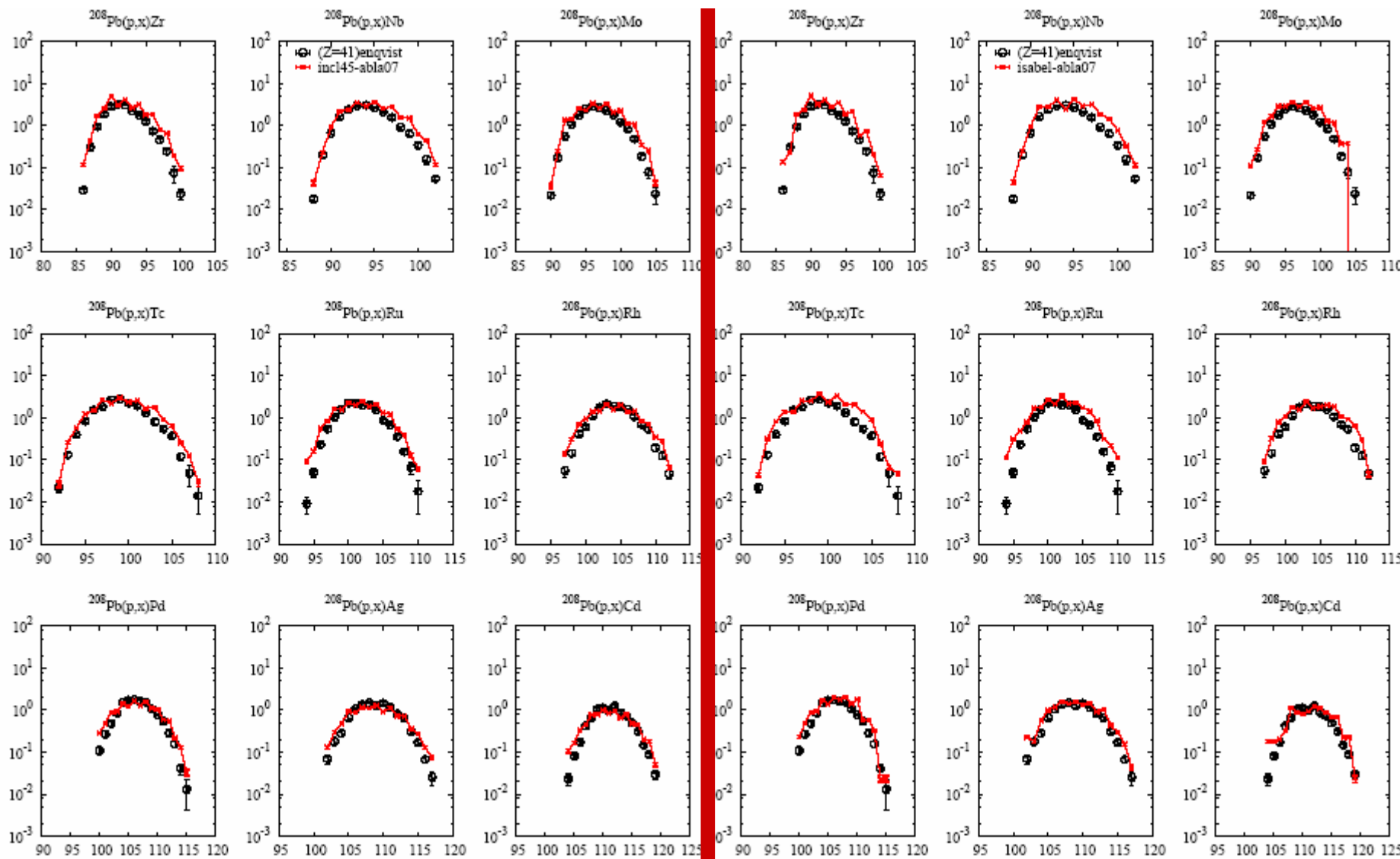
mass number A

mass number A

# $p(1000 \text{ MeV}) + {}^{208}\text{Pb}$ – final residues

## INCL45-ABLA07

## ISABEL-ABLA07



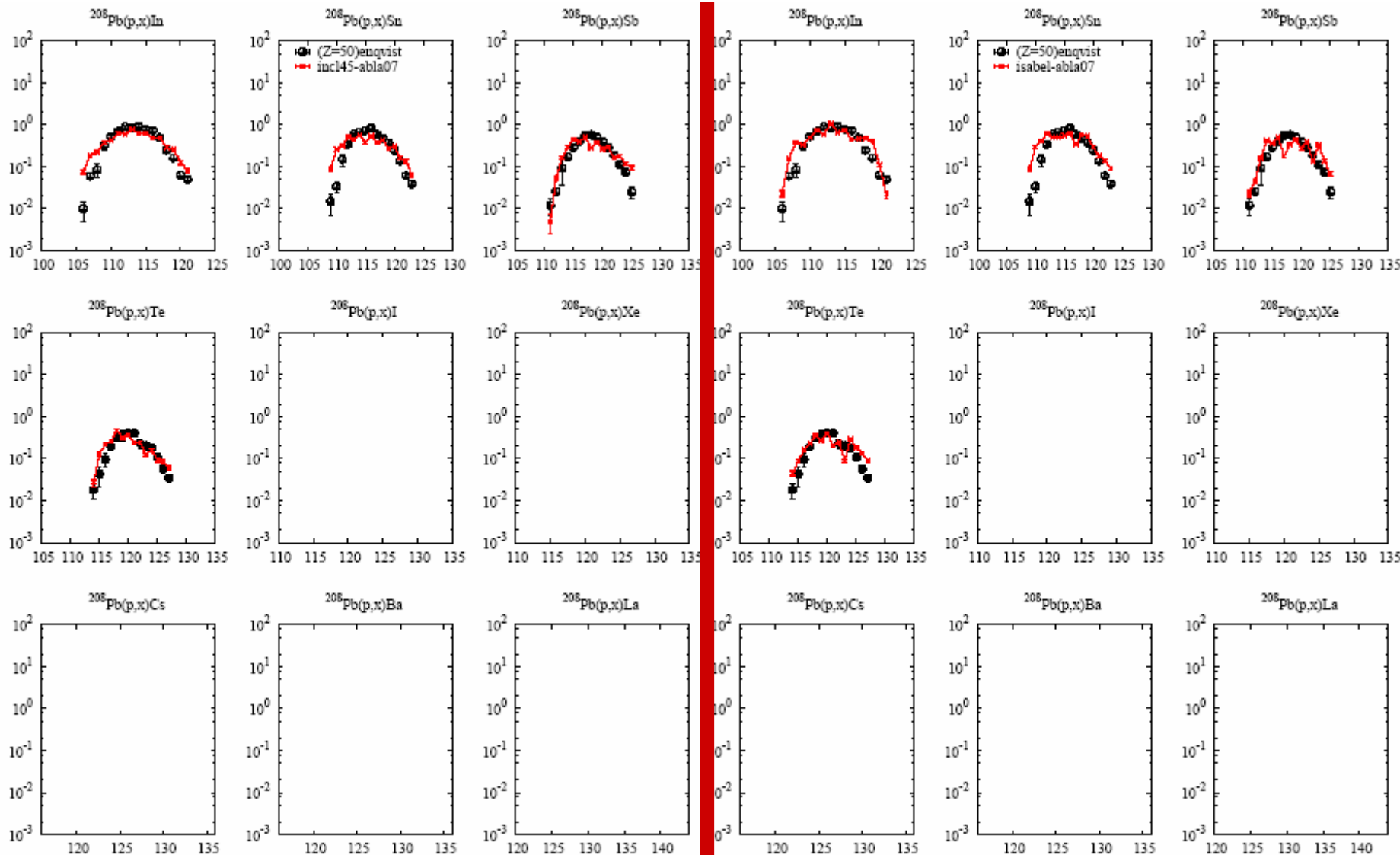
mass number A

mass number A

# $p(1000 \text{ MeV}) + {}^{208}\text{Pb}$ – final residues

## INCL45-ABLA07

## ISABEL-ABLA07



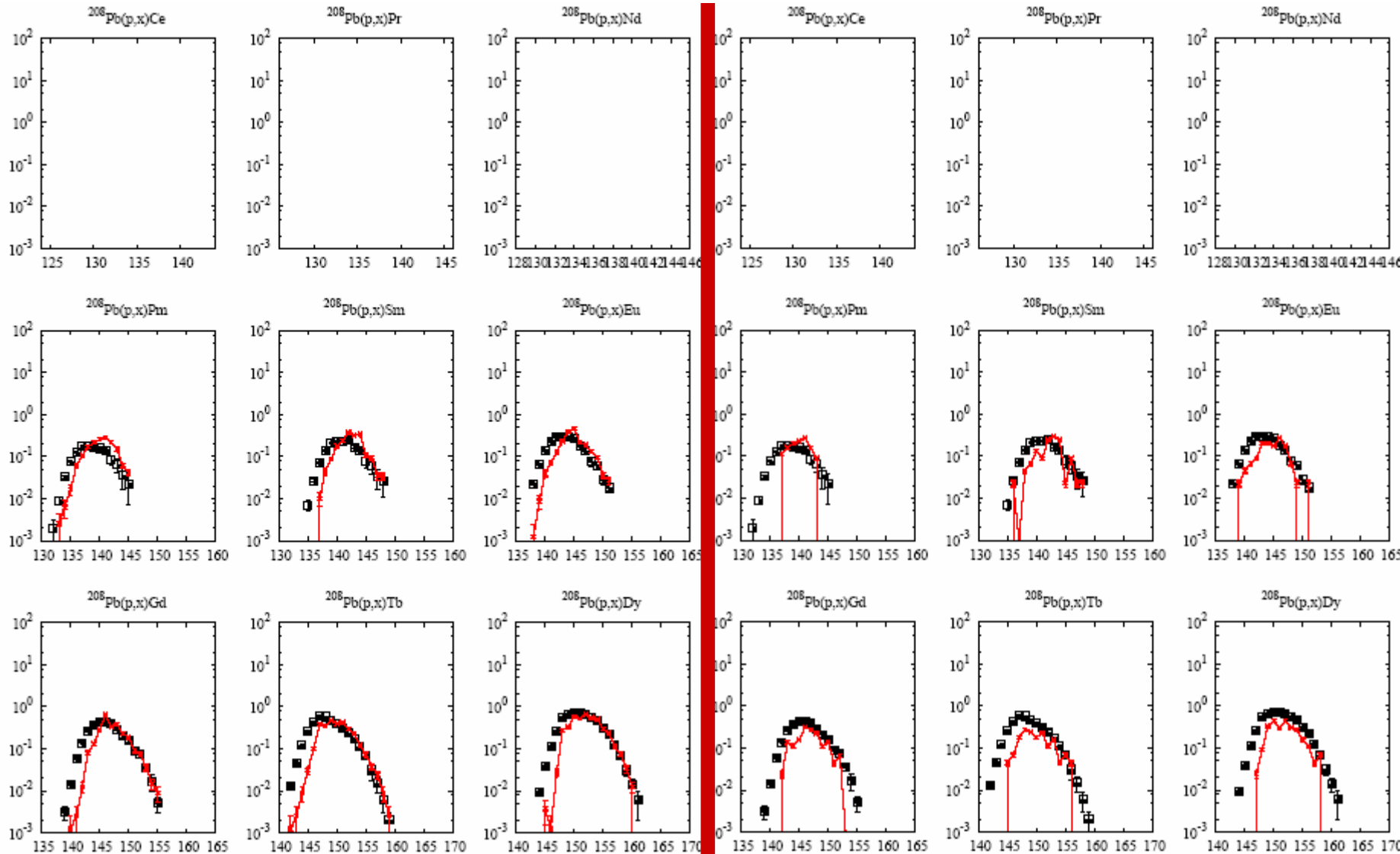
mass number A

mass number A

# $p(1000 \text{ MeV}) + {}^{208}\text{Pb}$ – final residues

## INCL45-ABLA07

## ISABEL-ABLA07

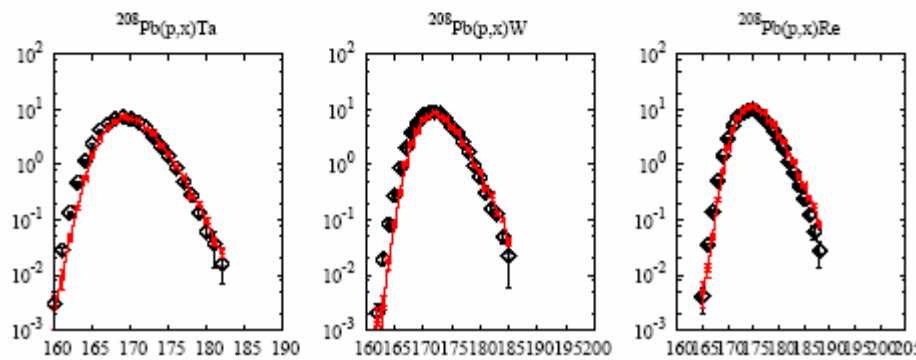
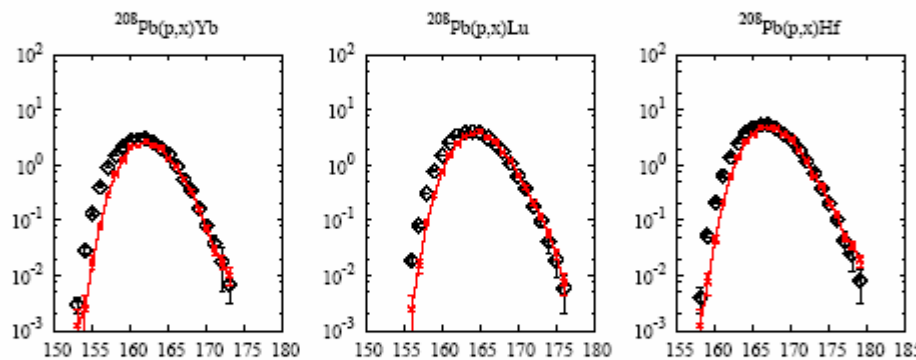
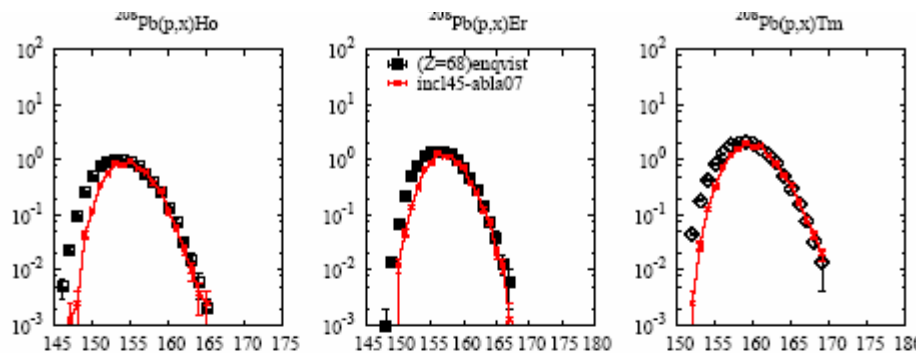


mass number A

mass number A

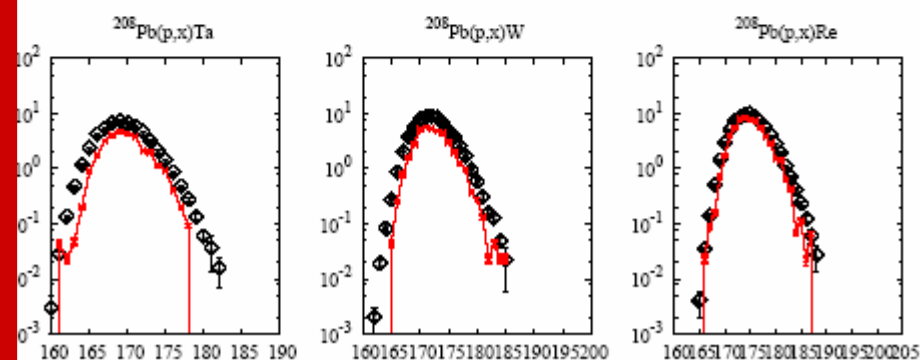
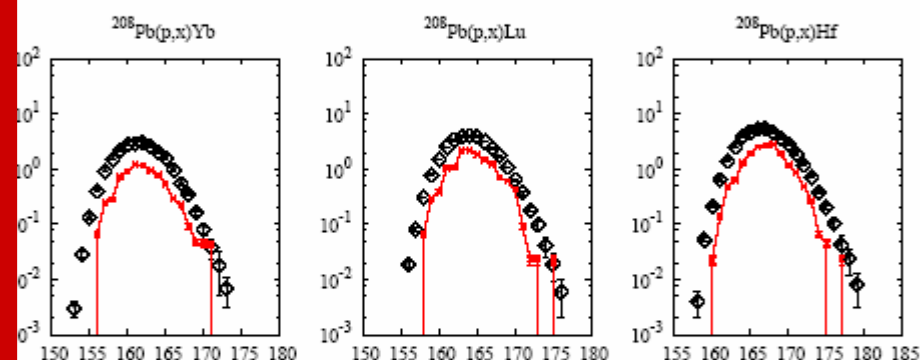
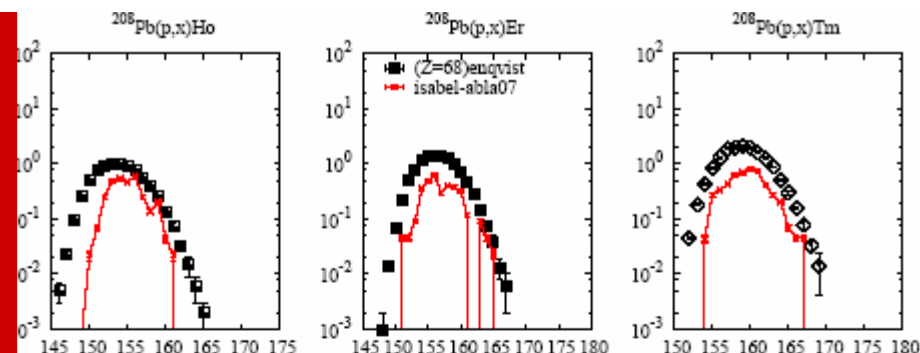
# $p(1000 \text{ MeV}) + {}^{208}\text{Pb}$ – final residues

## INCL45-ABLA07



mass number A

## ISABEL-ABLA07

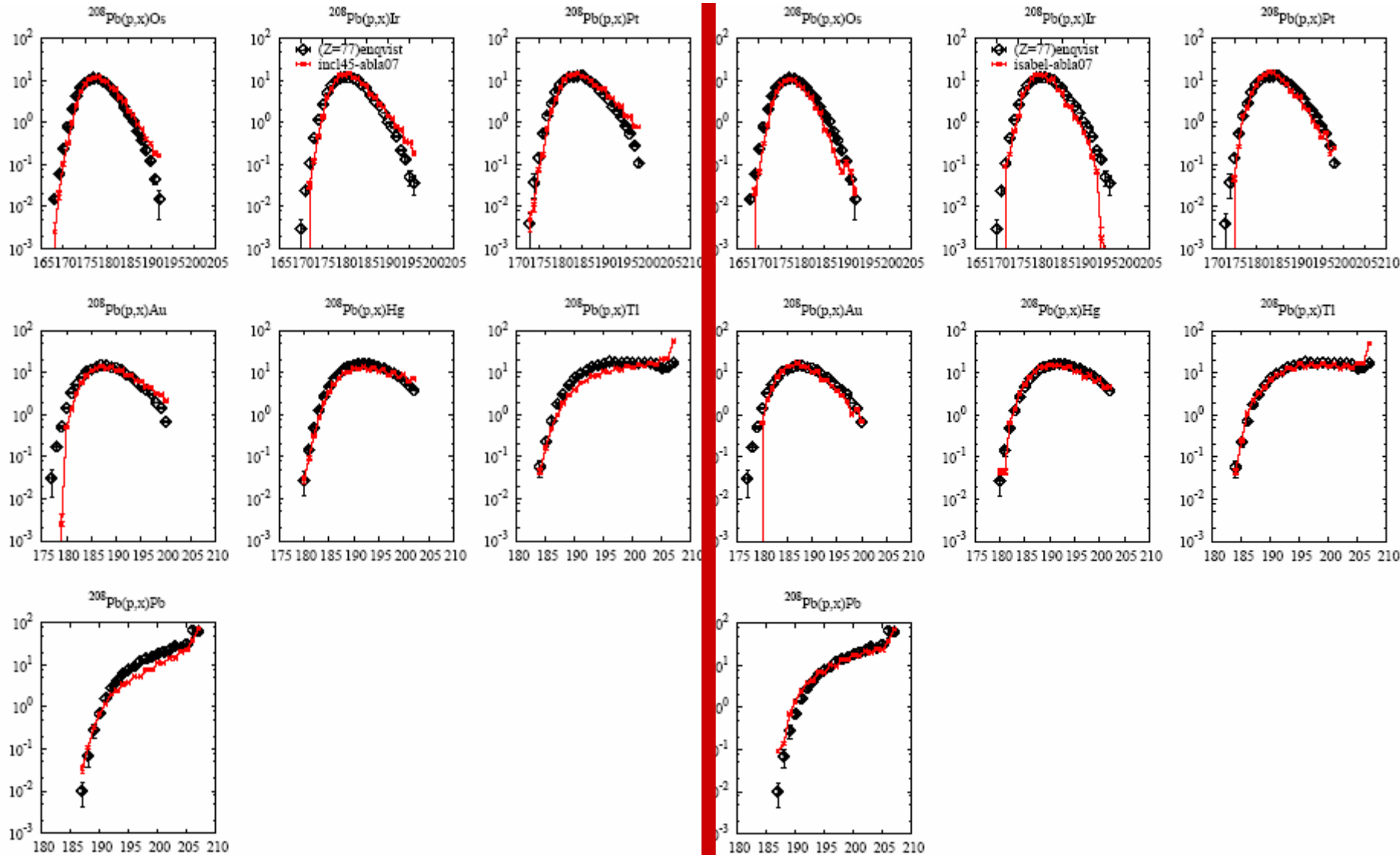


mass number A

# $p(1000 \text{ MeV}) + {}^{208}\text{Pb}$ – final residues

## INCL45-ABLA07

## ISABEL-ABLA07



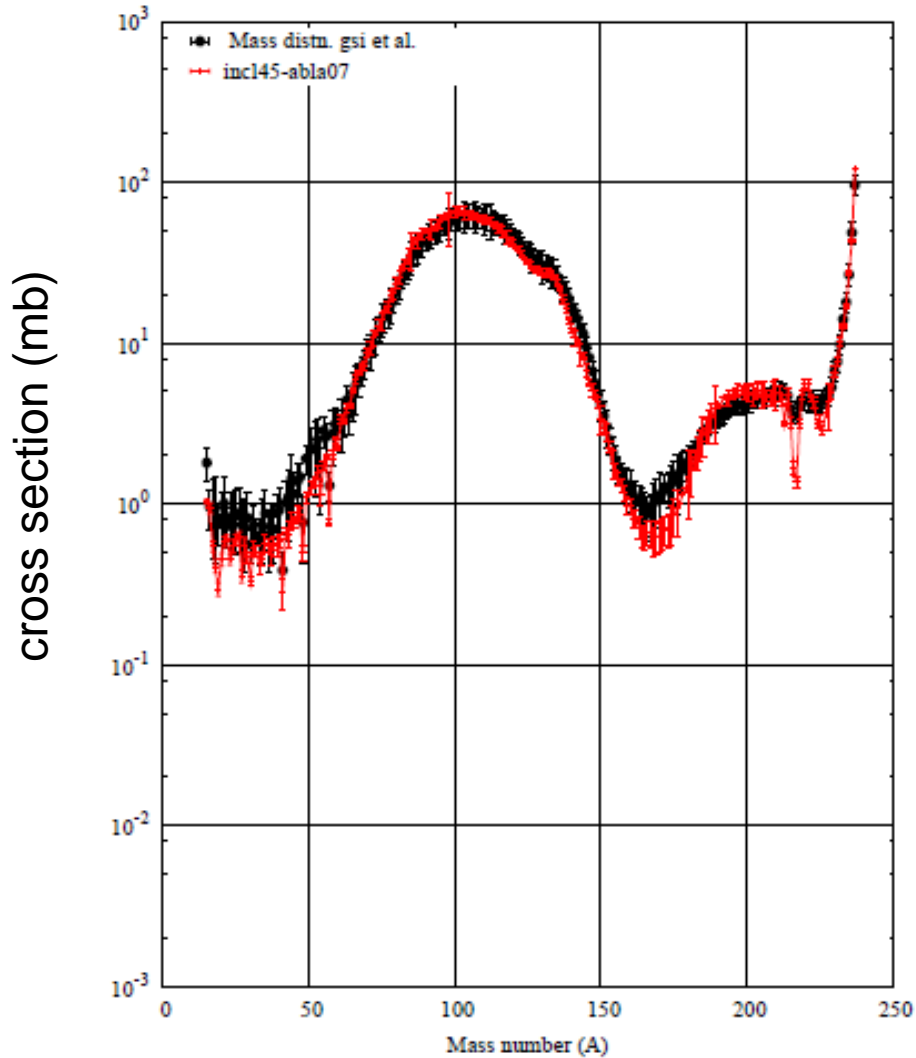
mass number A

mass number A

# p(1000 MeV) + $^{238}\text{U}$ – final residues

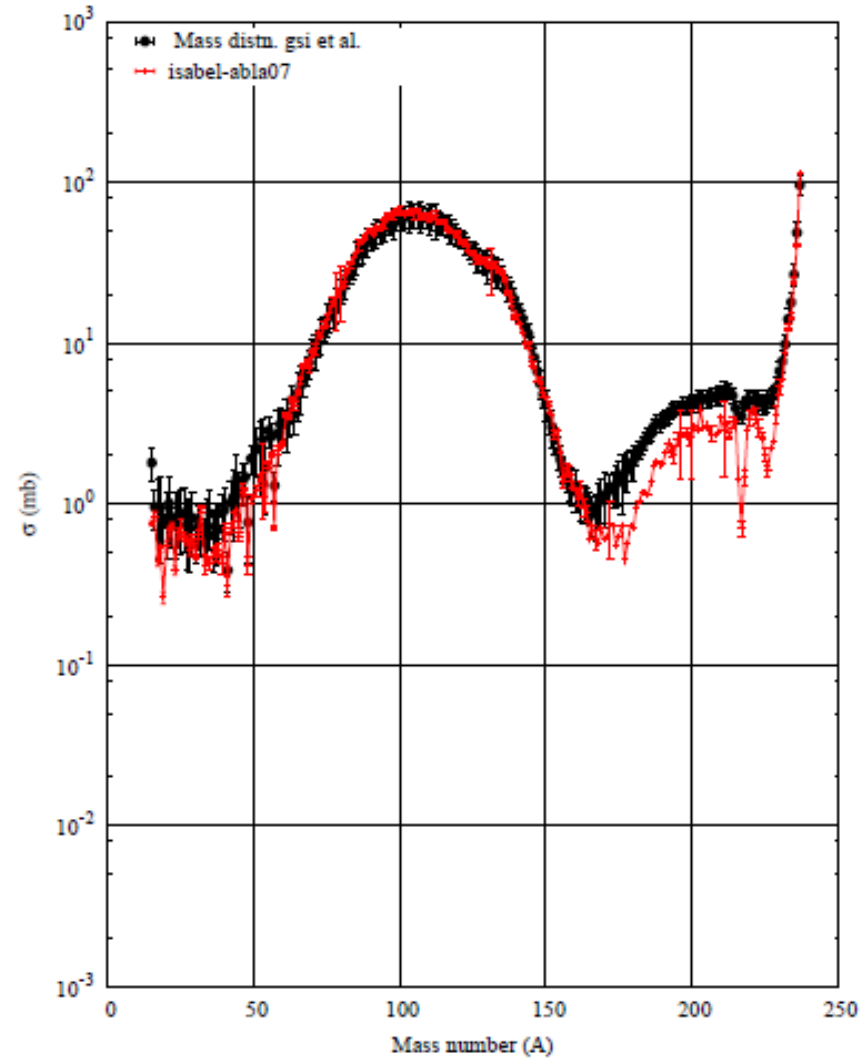
## INCL45-ABLA07

p (1000 MeV) +  $^{238}\text{U}$  – Residue mass production



## ISABEL-ABLA07

p (1000 MeV) +  $^{238}\text{U}$  – Residue mass production

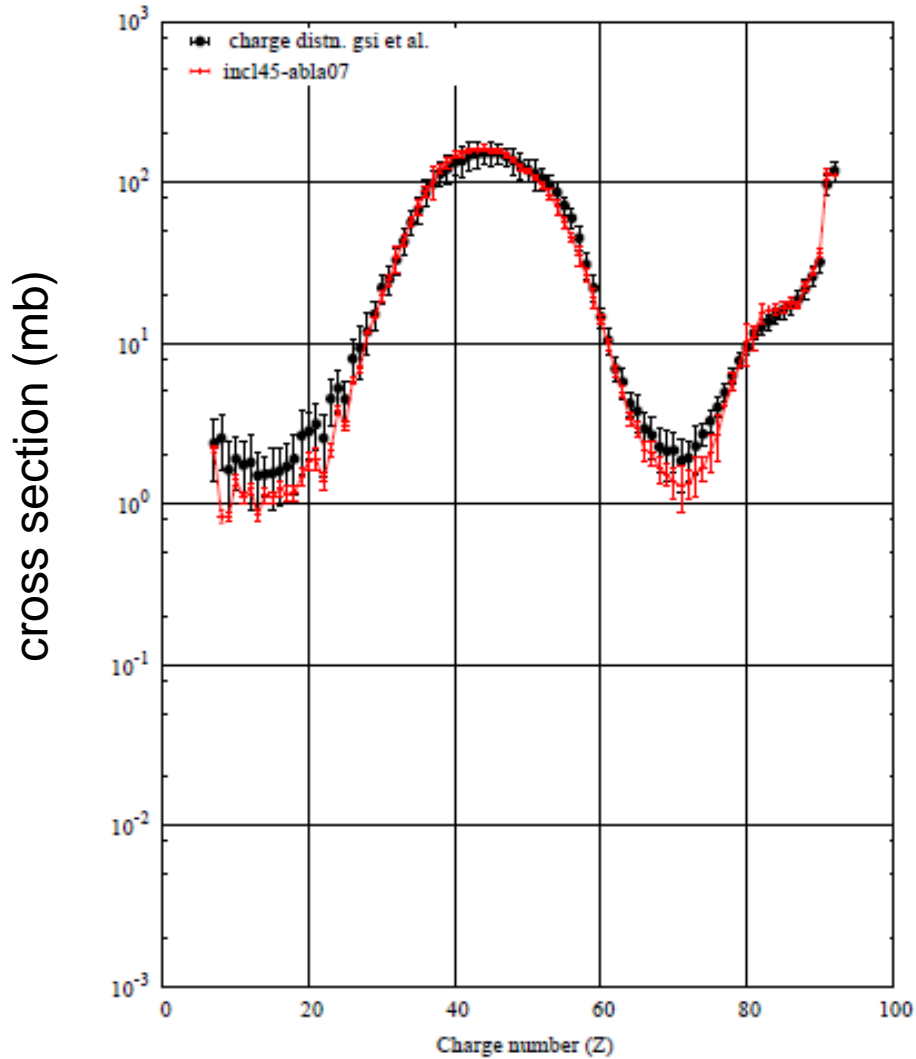


mass number A

# $p(1000 \text{ MeV}) + {}^{238}\text{U}$ – final residues

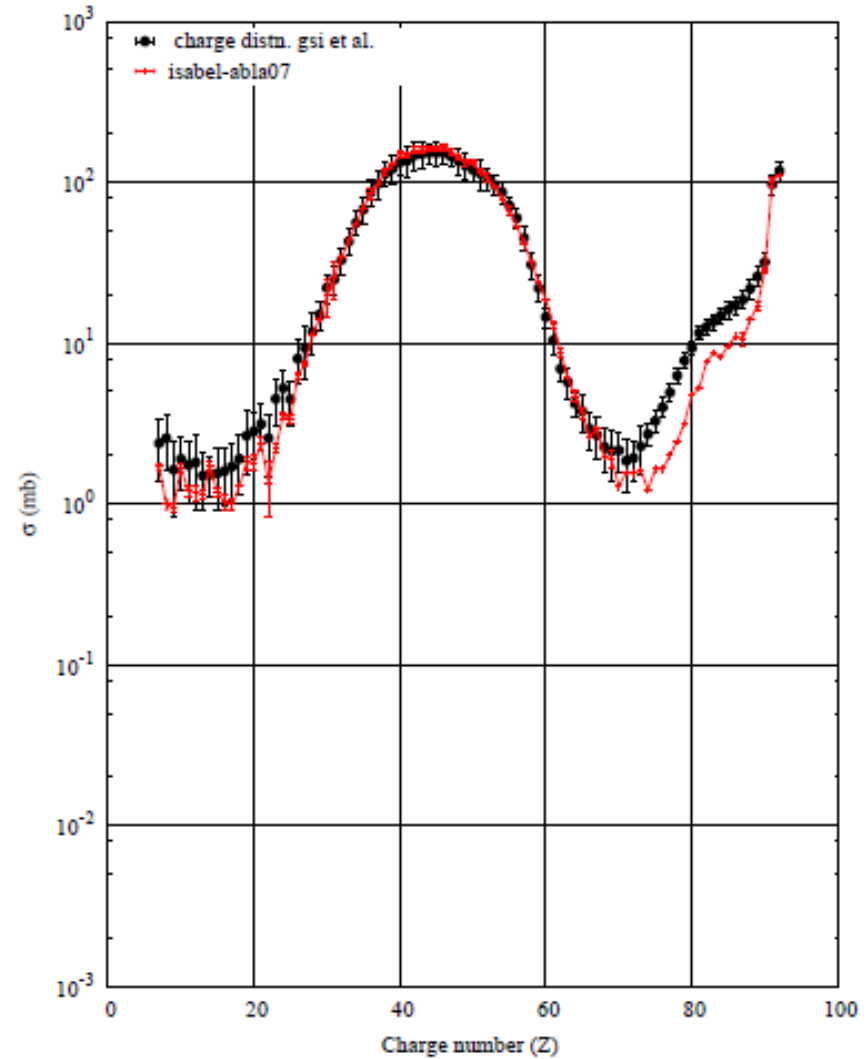
## INCL45-ABLA07

$p(1000 \text{ MeV}) + {}^{238}\text{U}$  -- Residue charge production



## ISABEL-ABLA07

$p(1000 \text{ MeV}) + {}^{238}\text{U}$  -- Residue charge production



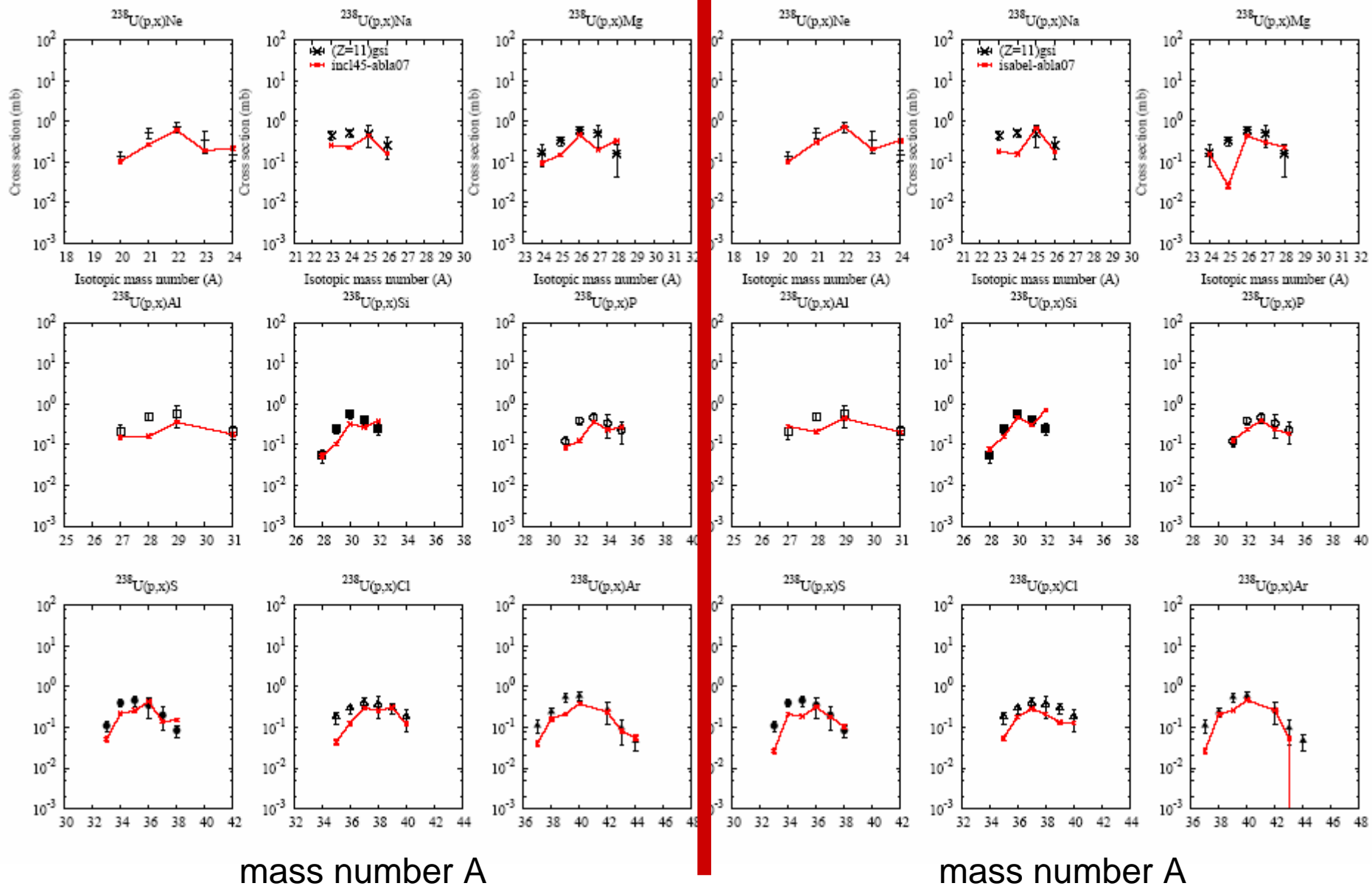
charge number Z



# $p(1000 \text{ MeV}) + {}^{238}\text{U}$ – final residues

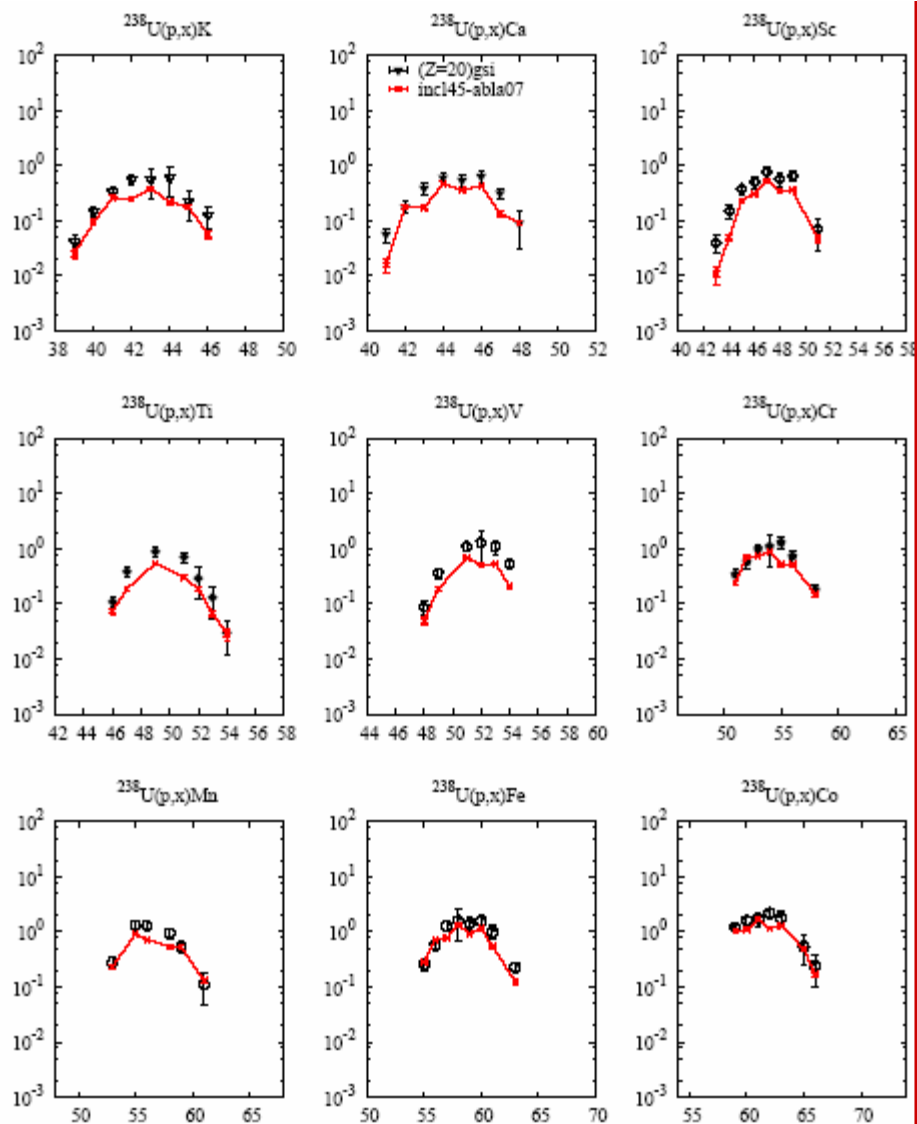
## INCL45-ABLA07

## ISABEL-ABLA07



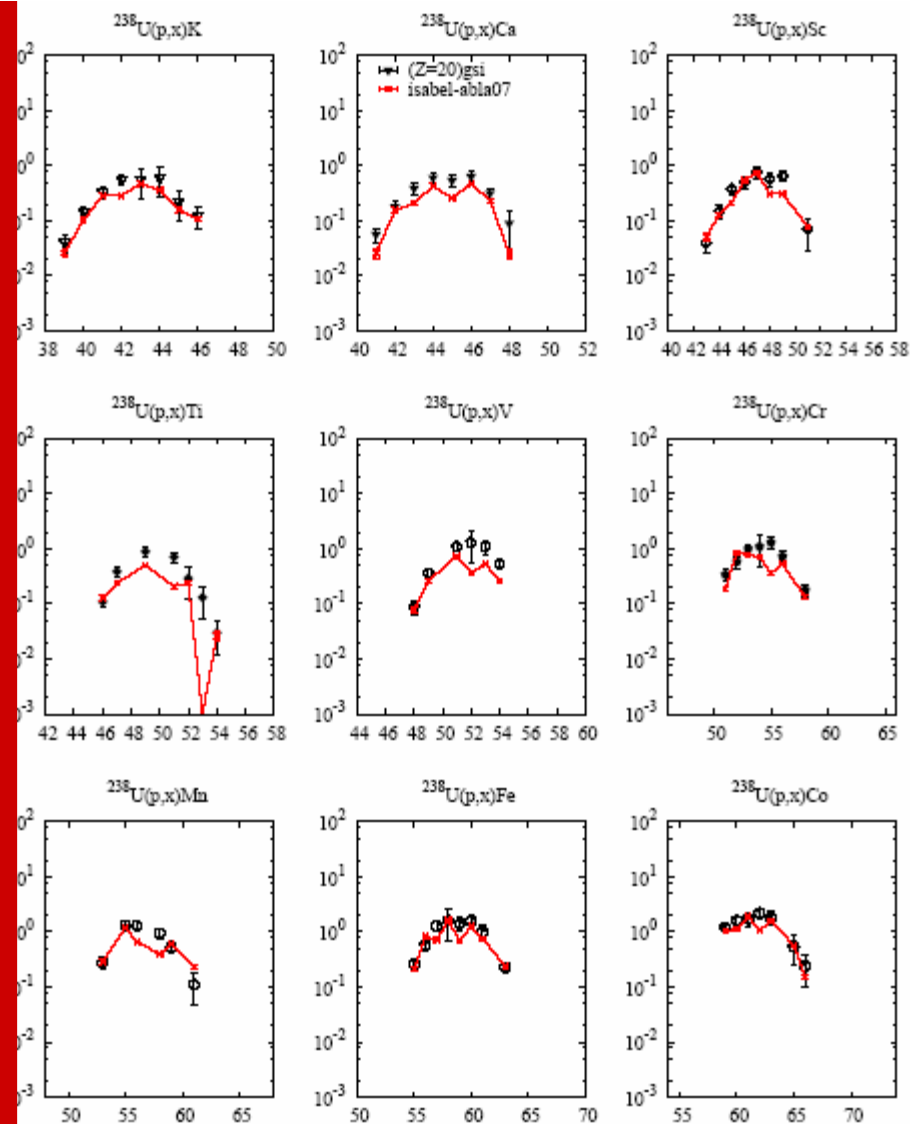
# $p(1000 \text{ MeV}) + {}^{238}\text{U}$ – final residues

## INCL45-ABLA07



mass number A

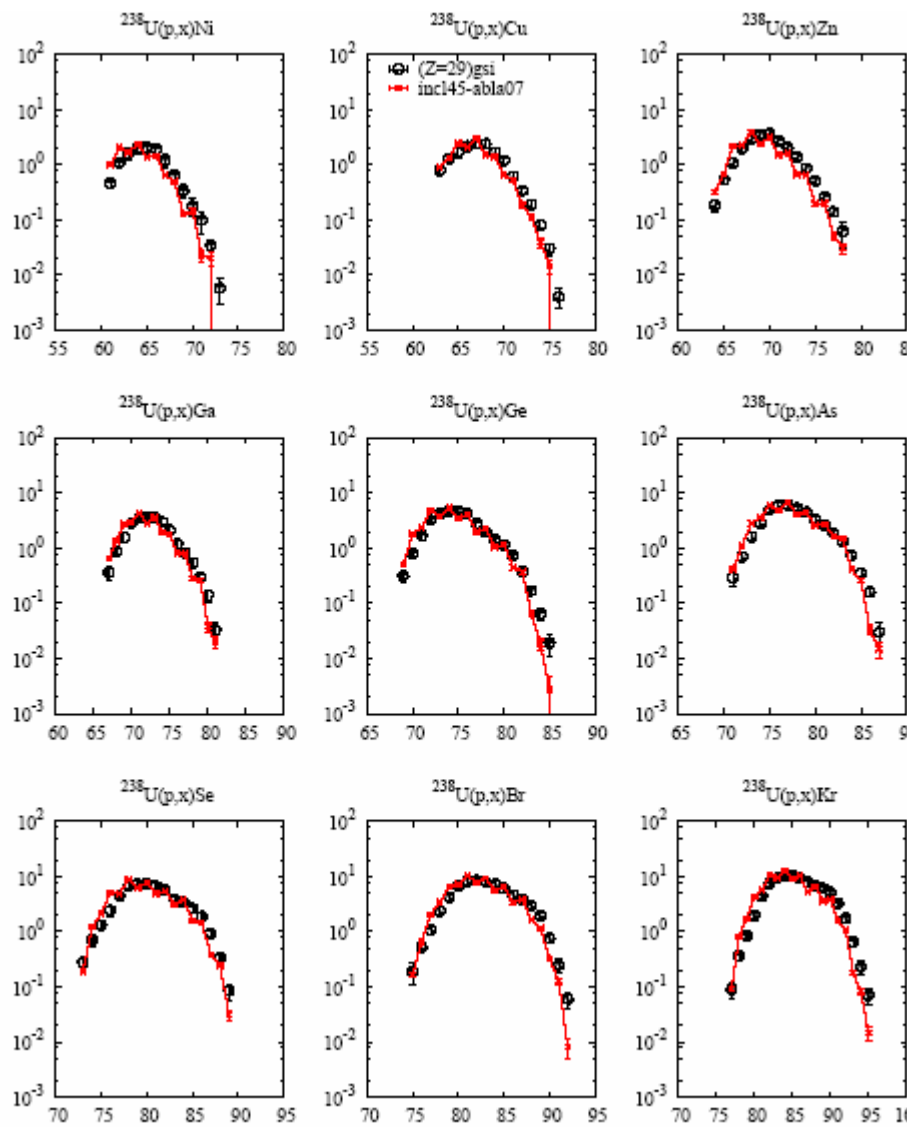
## ISABEL-ABLA07



mass number A

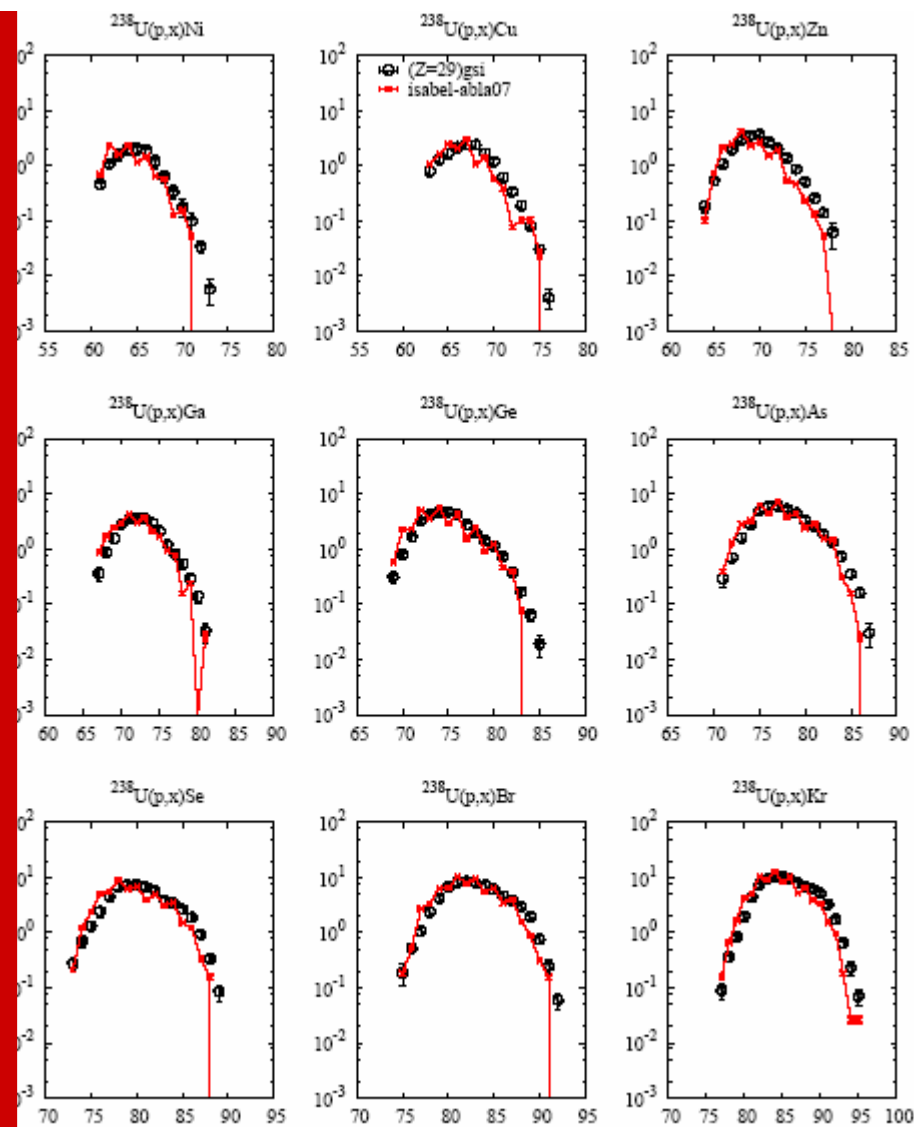
# $p(1000 \text{ MeV}) + {}^{238}\text{U} - \text{final residues}$

## INCL45-ABLA07



mass number A

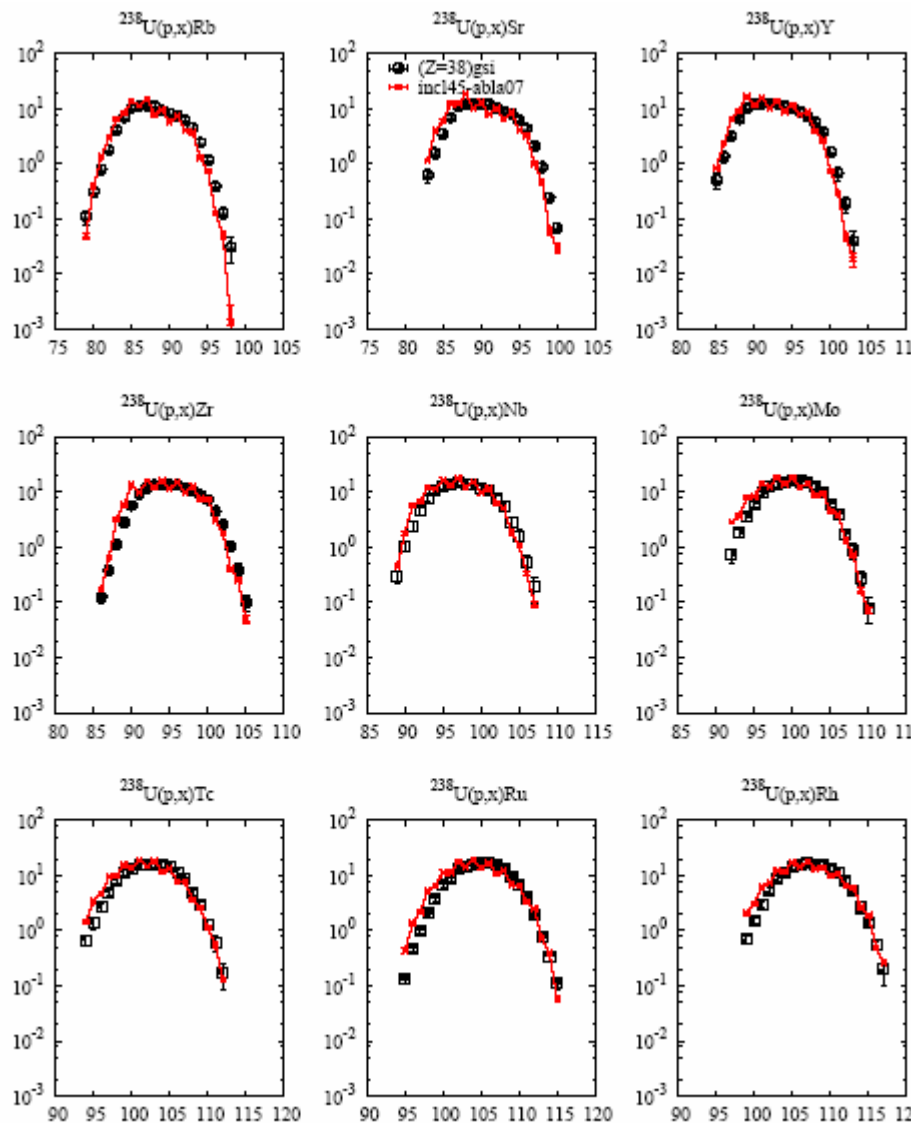
## ISABEL-ABLA07



mass number A

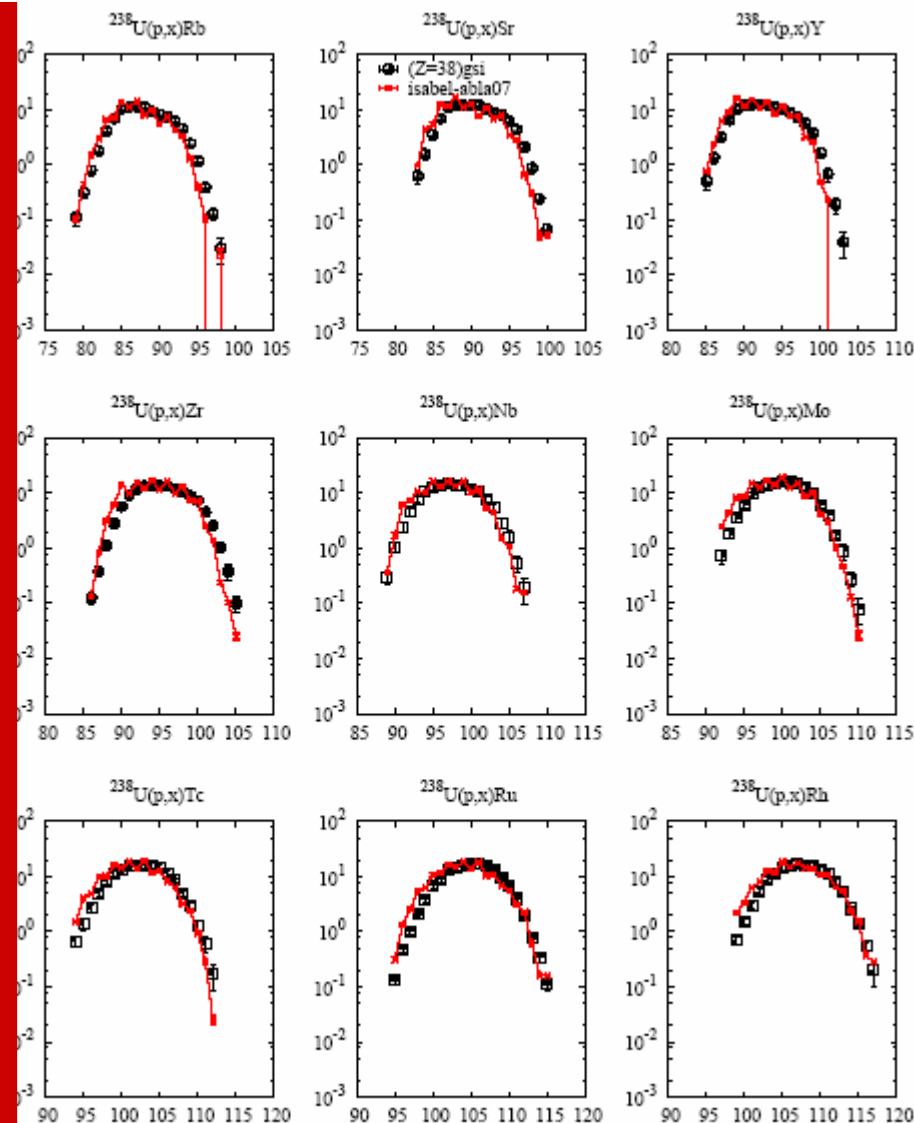
# $p(1000 \text{ MeV}) + {}^{238}\text{U}$ – final residues

## INCL45-ABLA07



mass number A

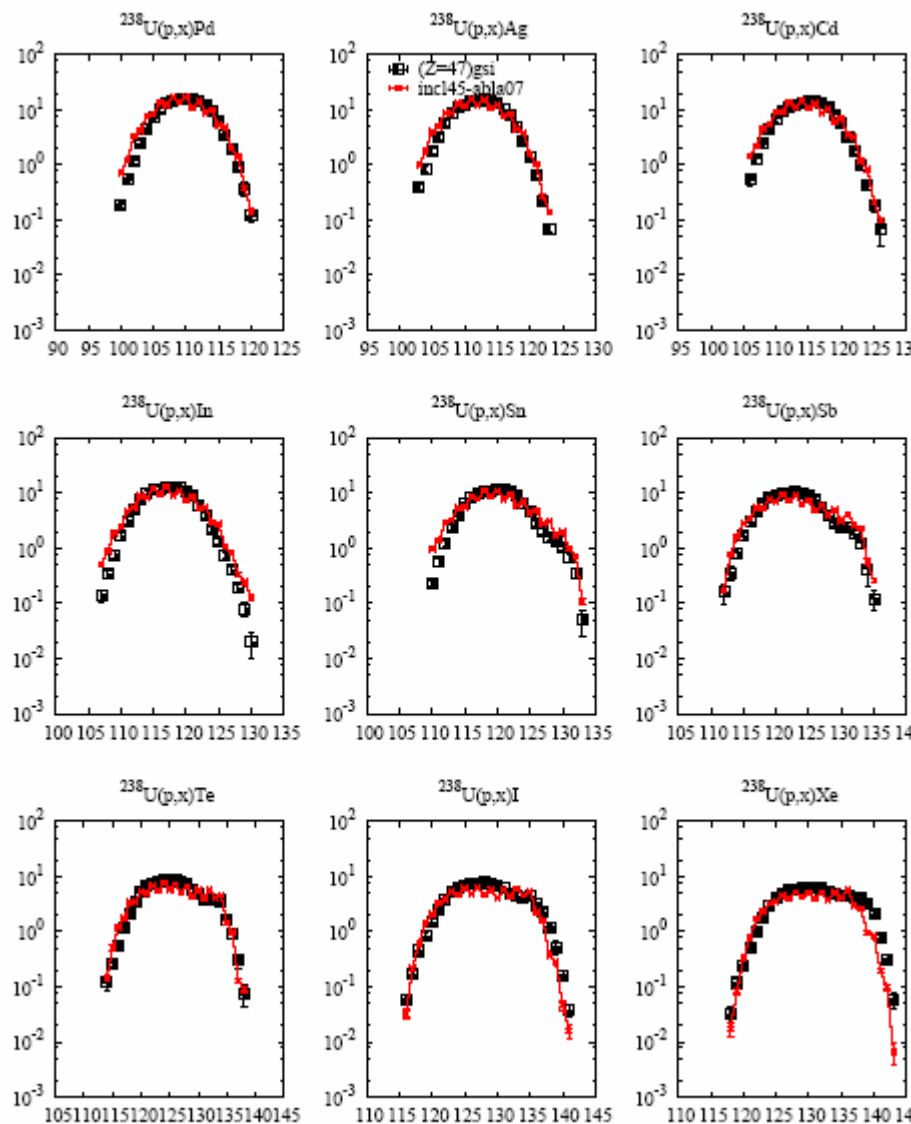
## ISABEL-ABLA07



mass number A

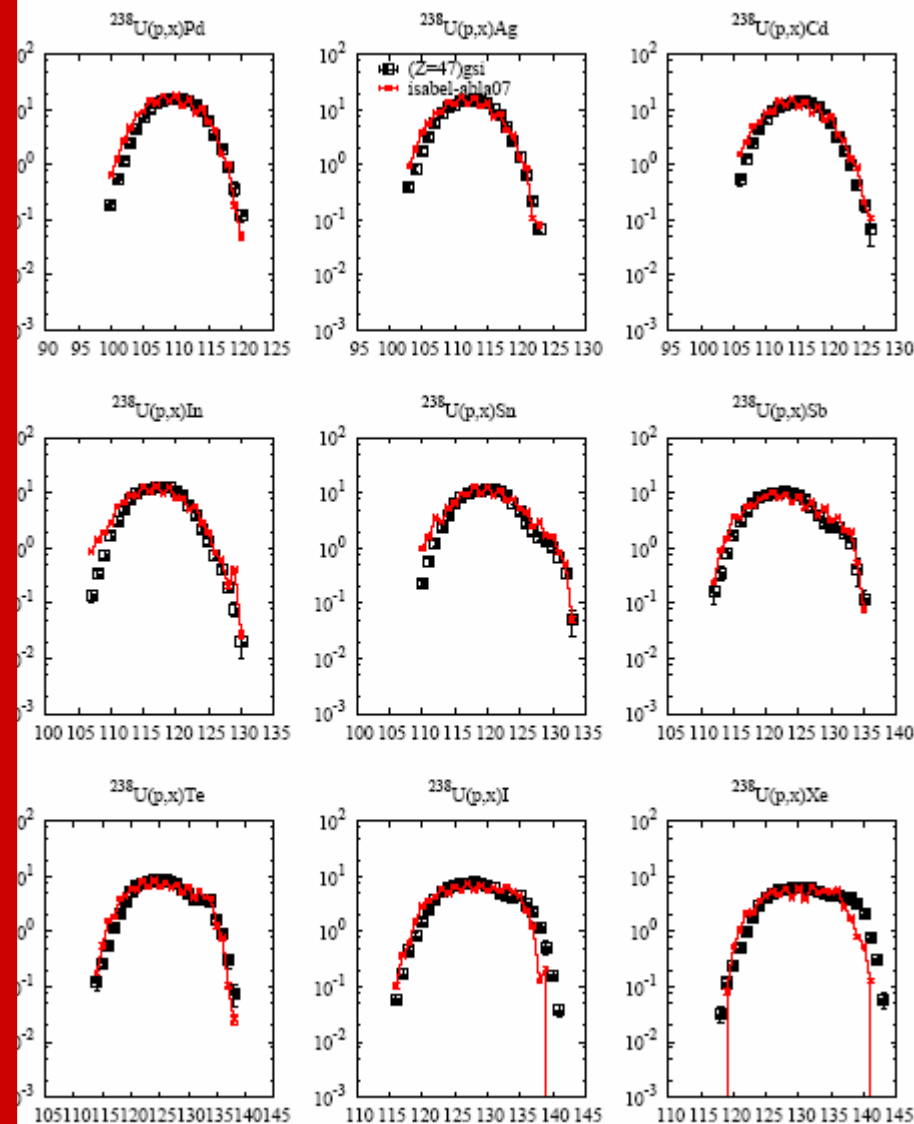
# $p(1000 \text{ MeV}) + {}^{238}\text{U}$ – final residues

## INCL45-ABLA07



mass number A

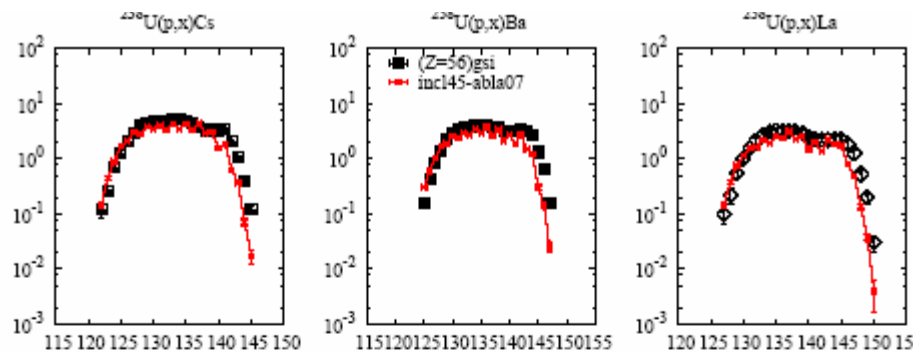
## ISABEL-ABLA07



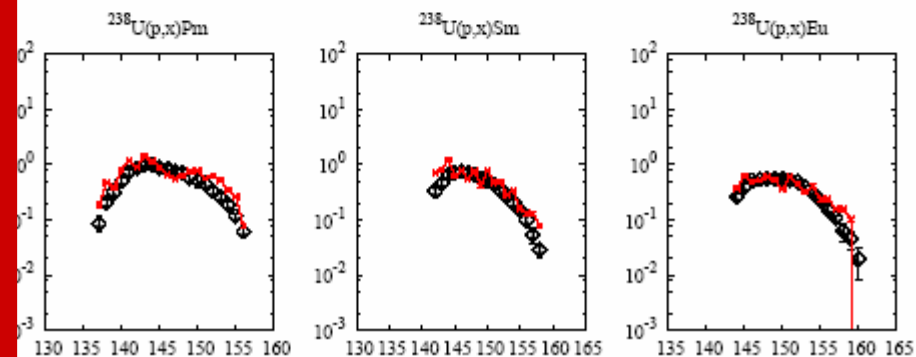
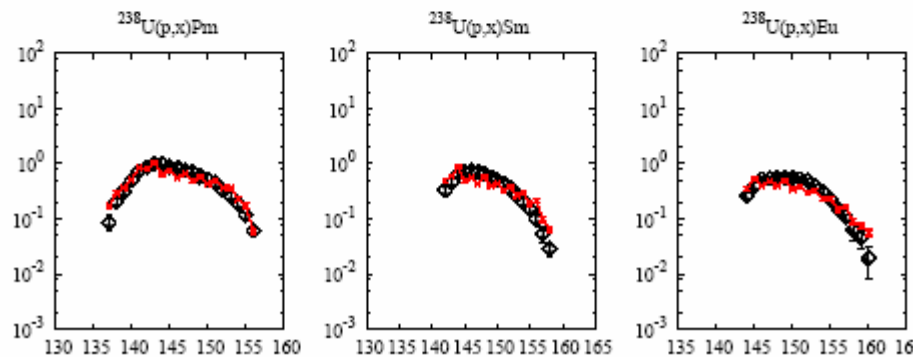
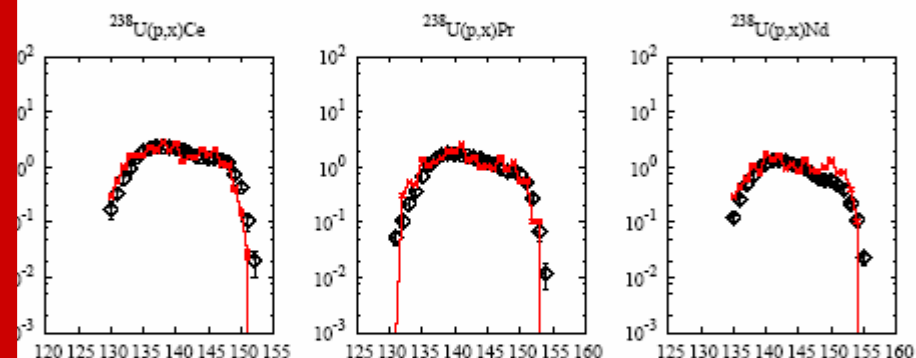
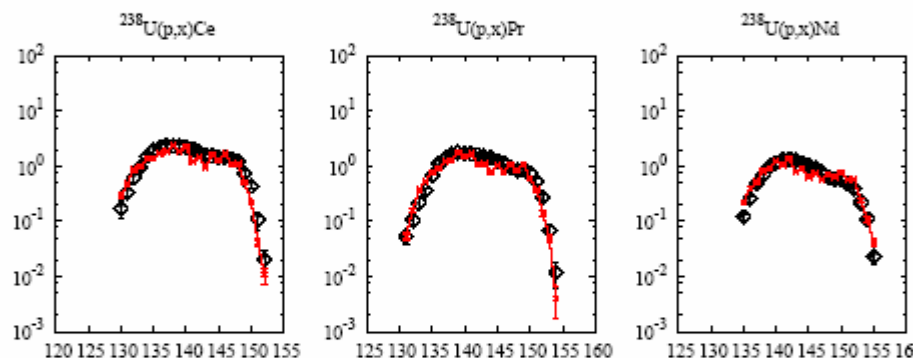
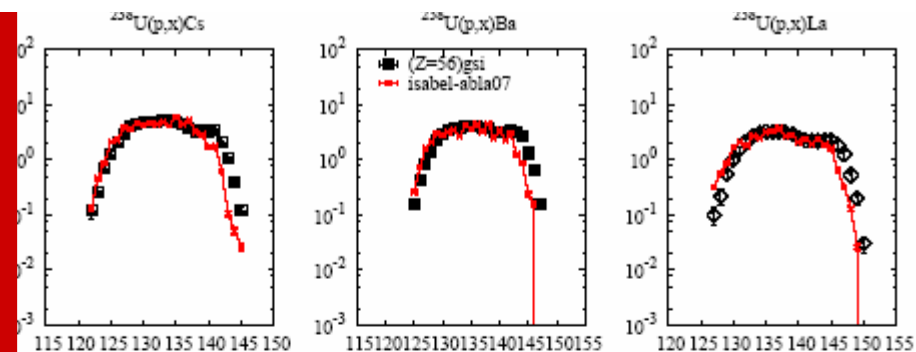
mass number A

# $p(1000 \text{ MeV}) + {}^{238}\text{U}$ – final residues

## INCL45-ABLA07



## ISABEL-ABLA07

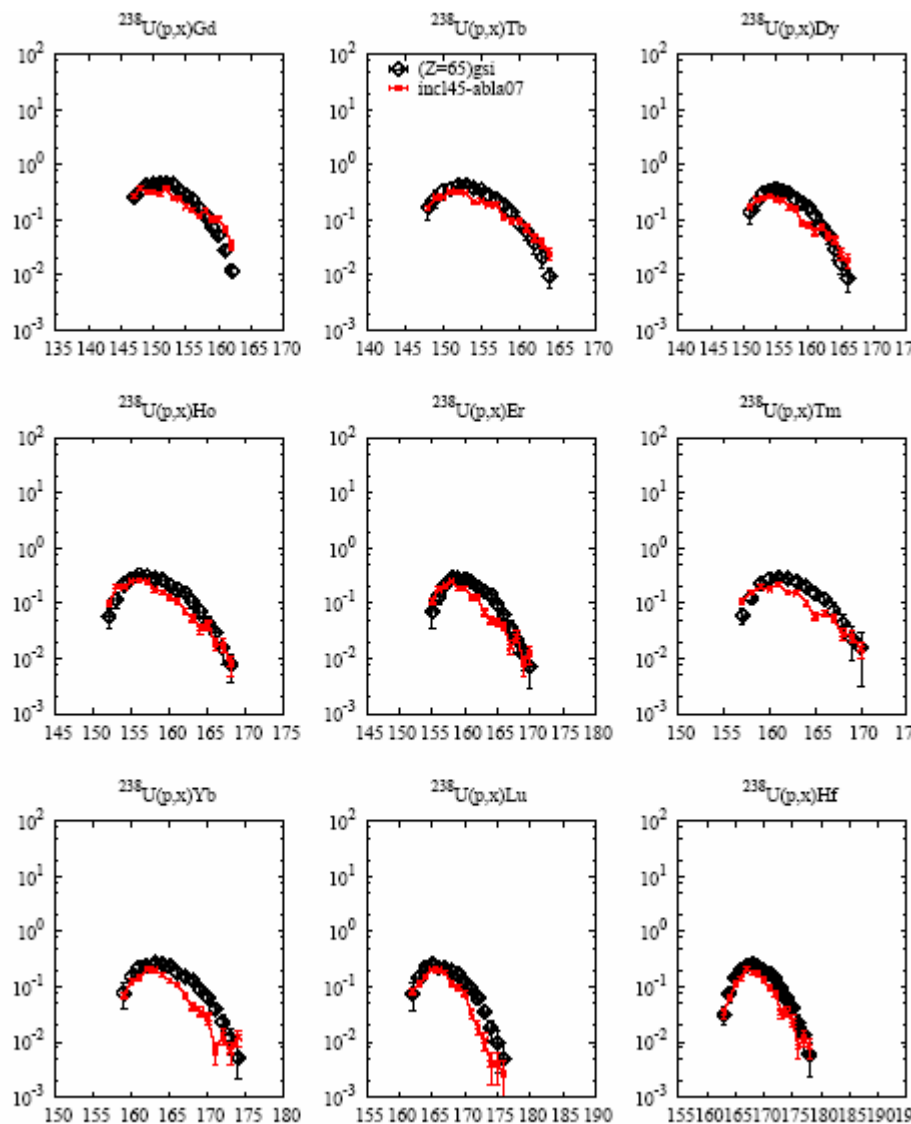


mass number A

mass number A

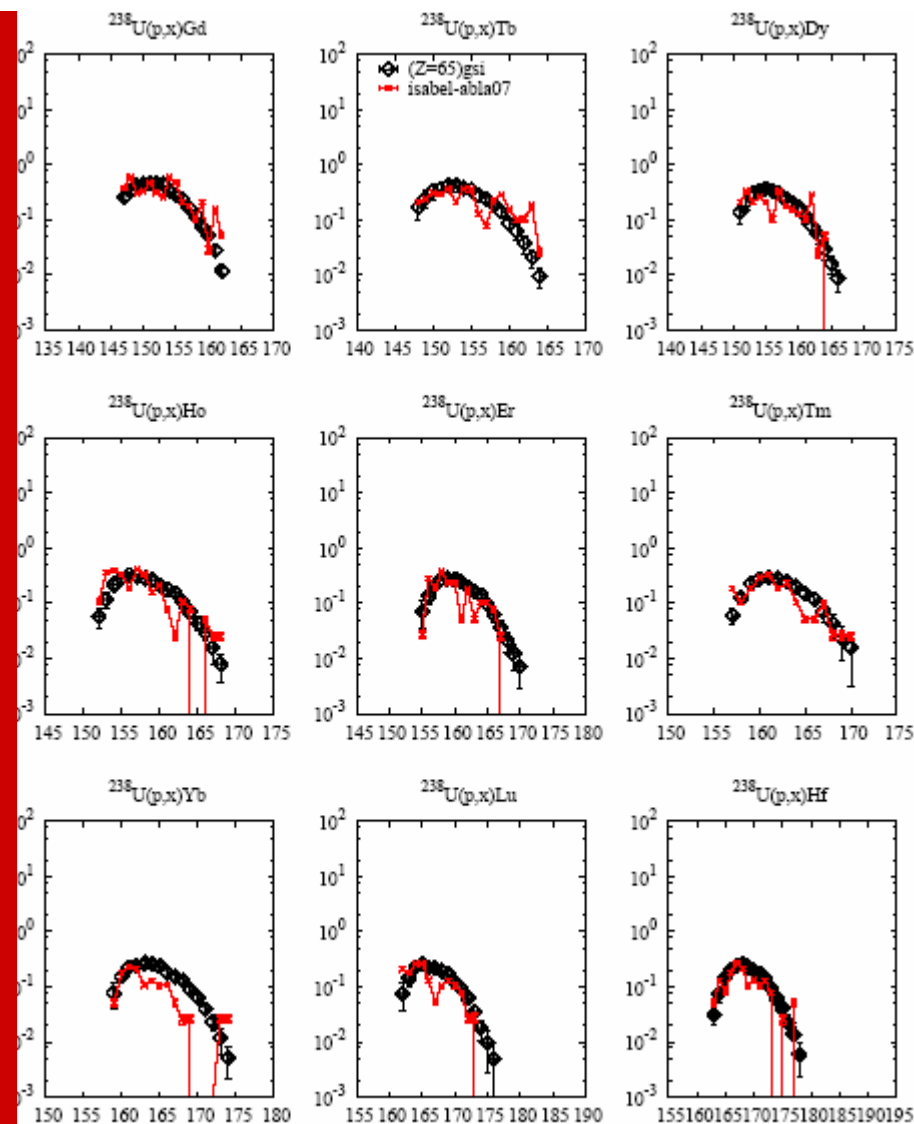
# $p(1000 \text{ MeV}) + {}^{238}\text{U}$ – final residues

## INCL45-ABLA07



mass number A

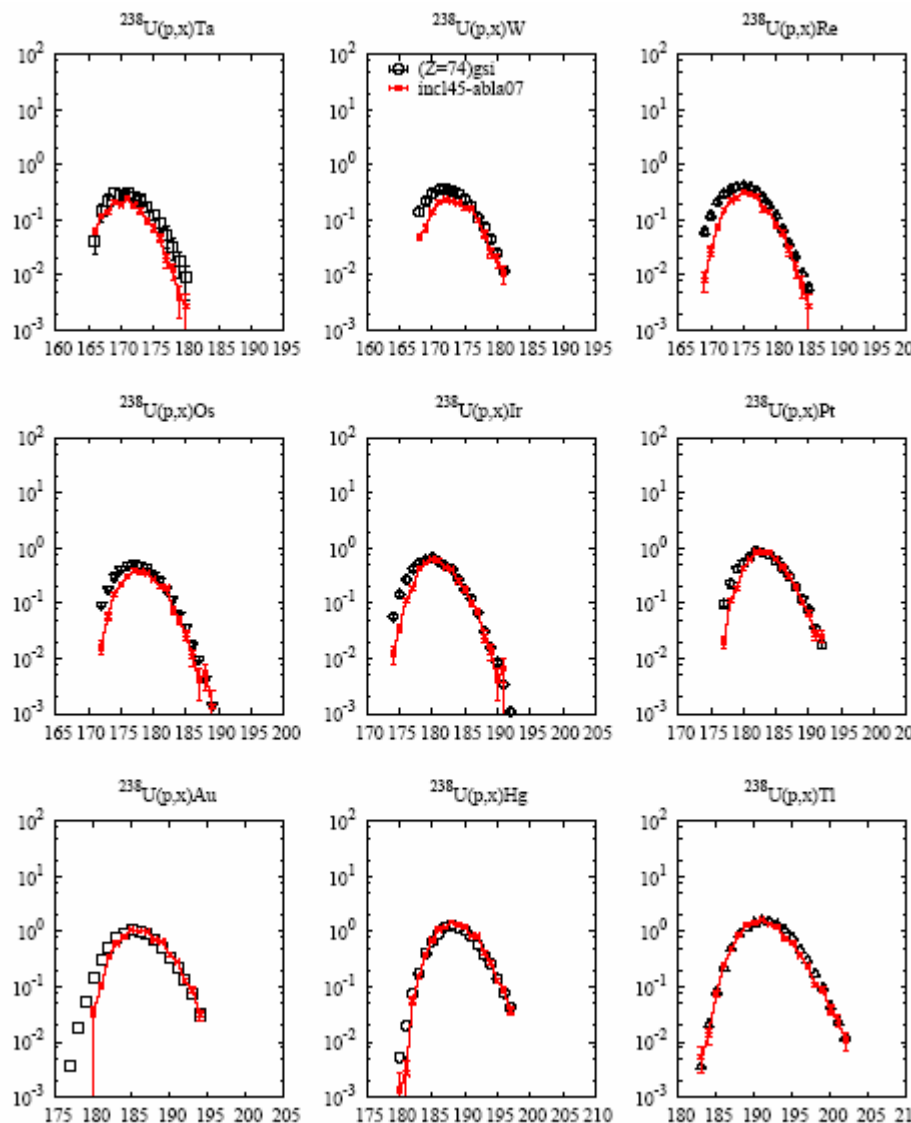
## ISABEL-ABLA07



mass number A

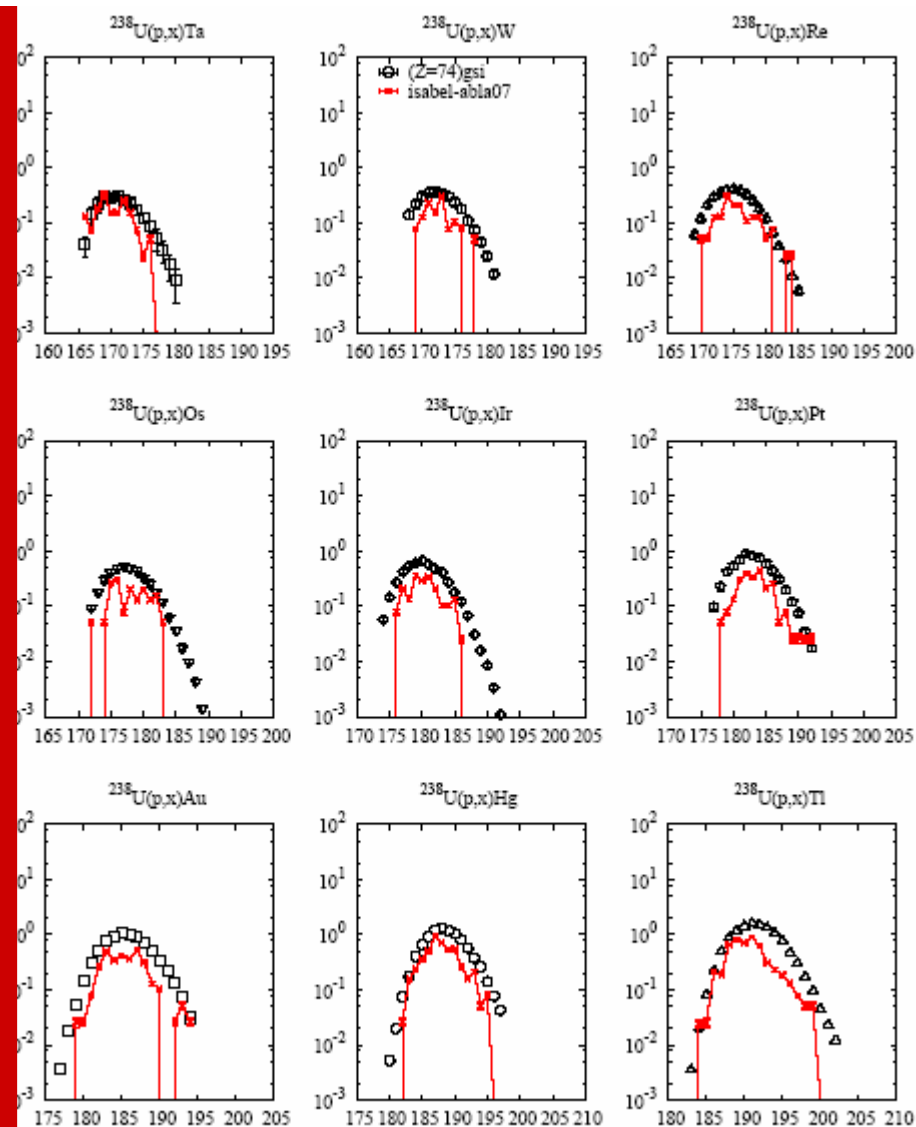
# $p(1000 \text{ MeV}) + {}^{238}\text{U}$ – final residues

## INCL45-ABLA07



mass number A

## ISABEL-ABLA07

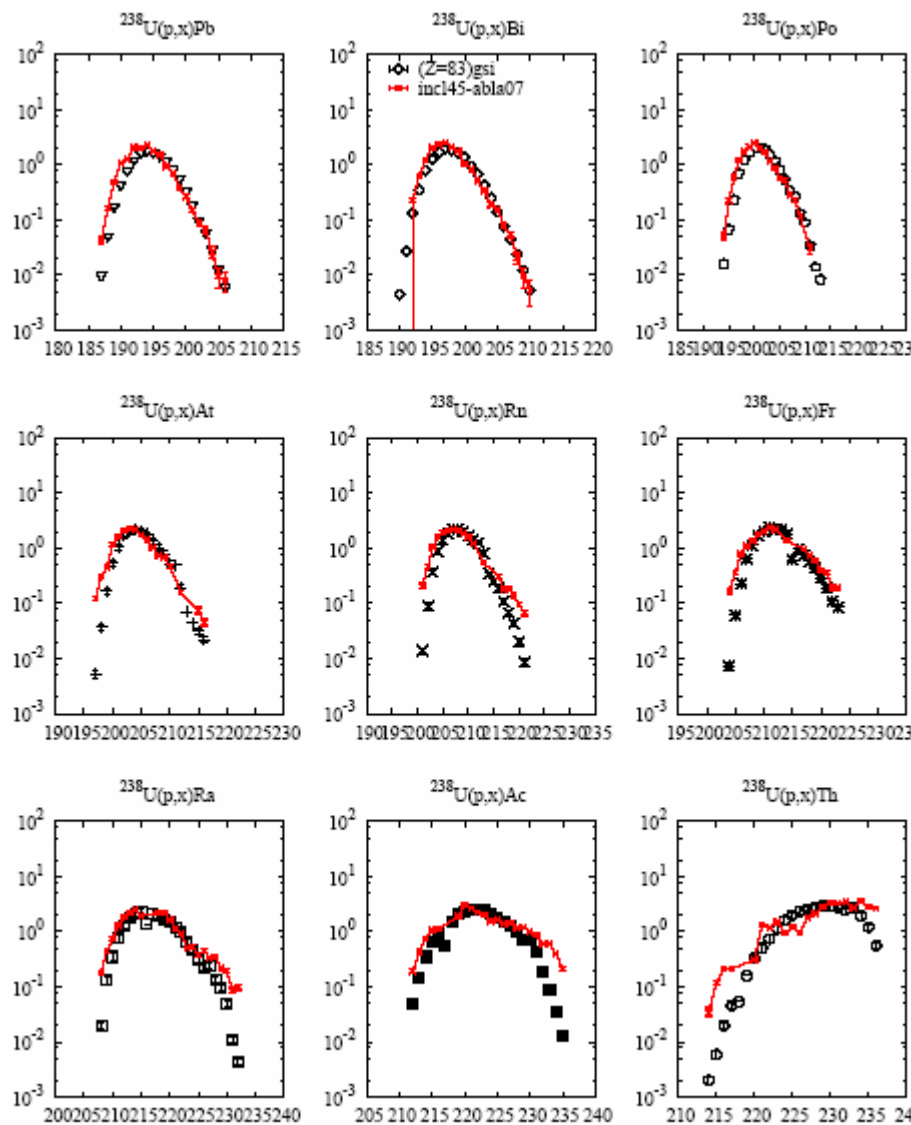


mass number A



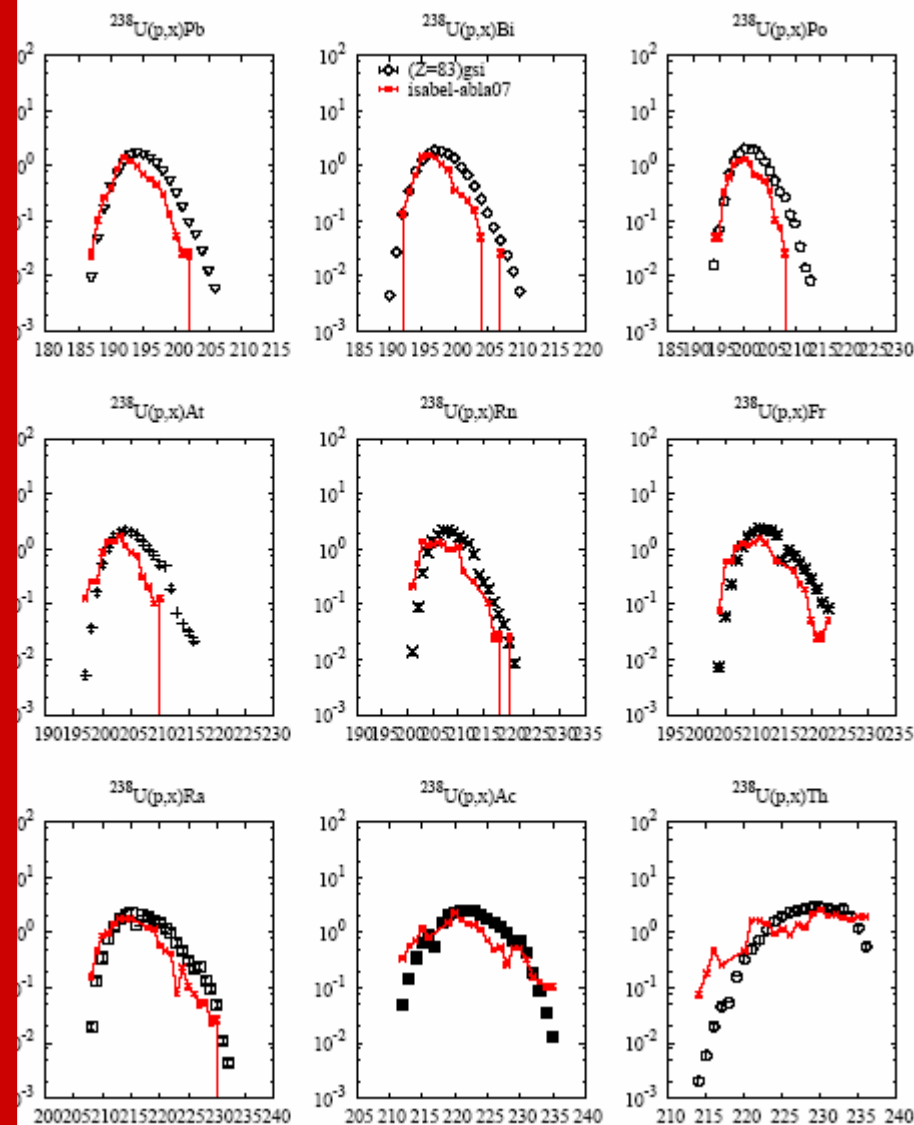
# $p(1000 \text{ MeV}) + {}^{238}\text{U}$ – final residues

## INCL45-ABLA07



mass number A

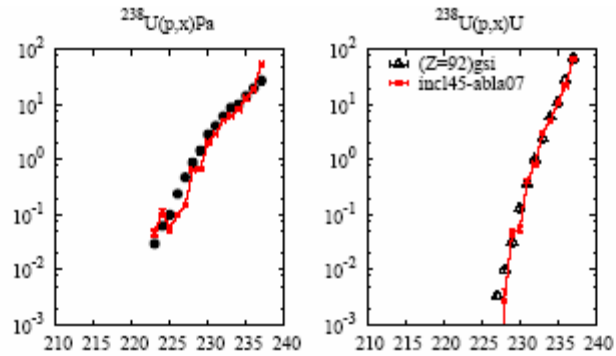
## ISABEL-ABLA07



mass number A

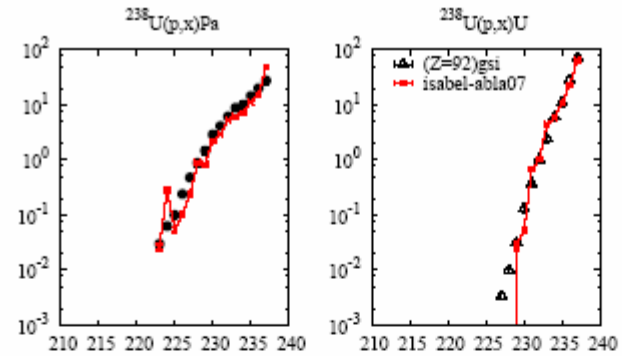
# $p(1000 \text{ MeV}) + {}^{238}\text{U}$ – final residues

## INCL45-ABLA07



mass number A

## ISABEL-ABLA07



mass number A

# Residues

**Status:** Good

**Improvement:** Difficult to establish how to disentangle INC and de-excitation...

Concerning ABLA:

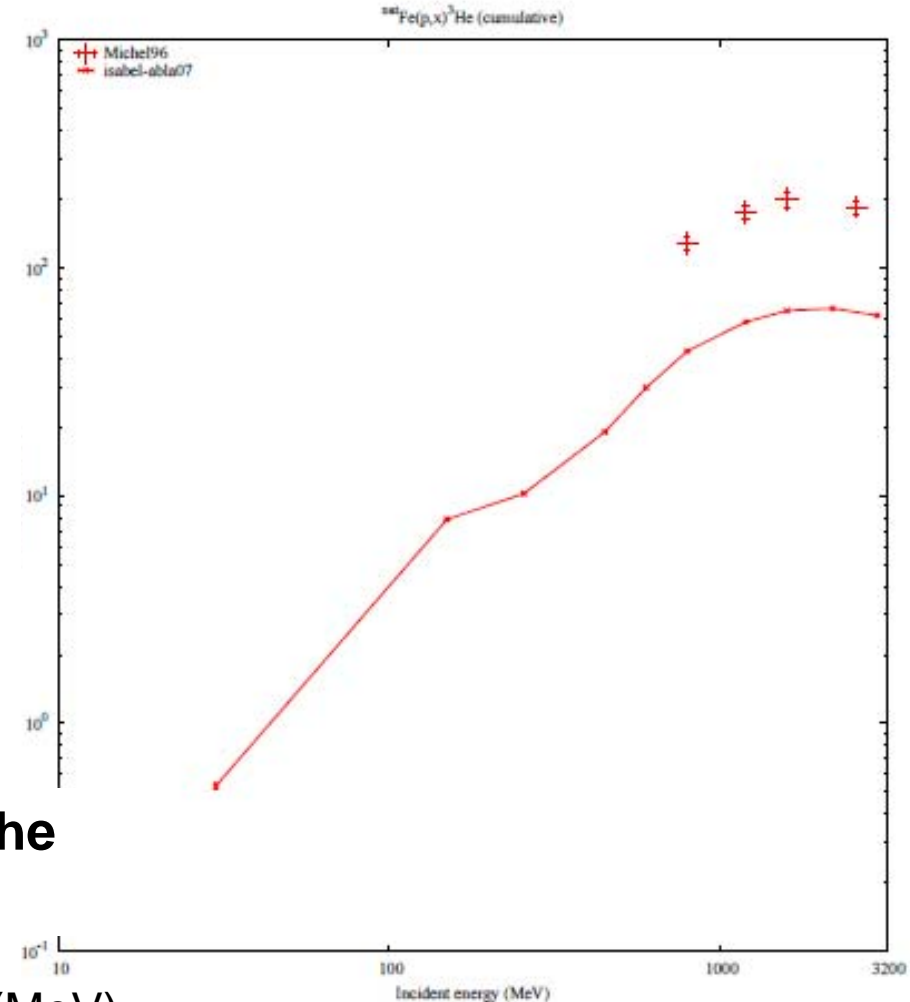
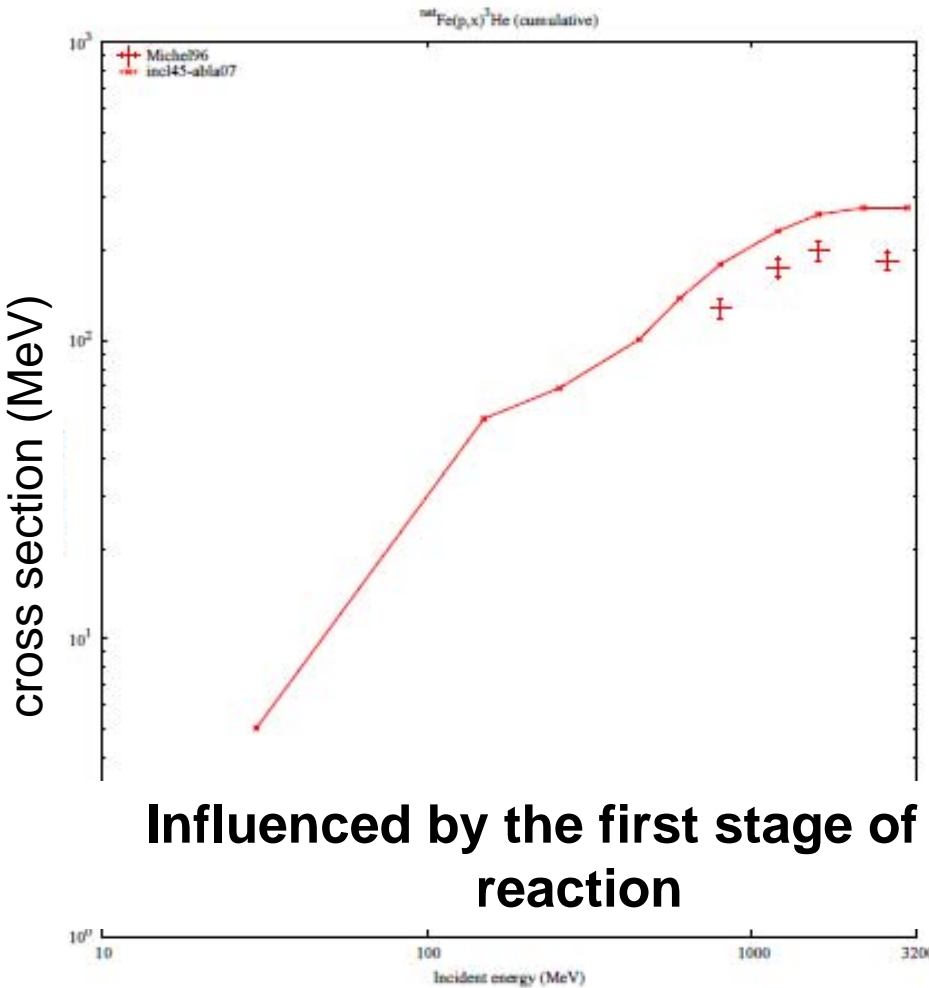
- 1) improve the description of even-odd ( $\rightarrow$  gamma decay strength)
- 2) improve structural effects (could be relevant for very light residues)
- 3) fission

# Excitation functions

# $^{nat}\text{Fe}(p,x)^3\text{He}$ (cumulative)

## INCL45-ABLA07

## ISABEL-ABLA07



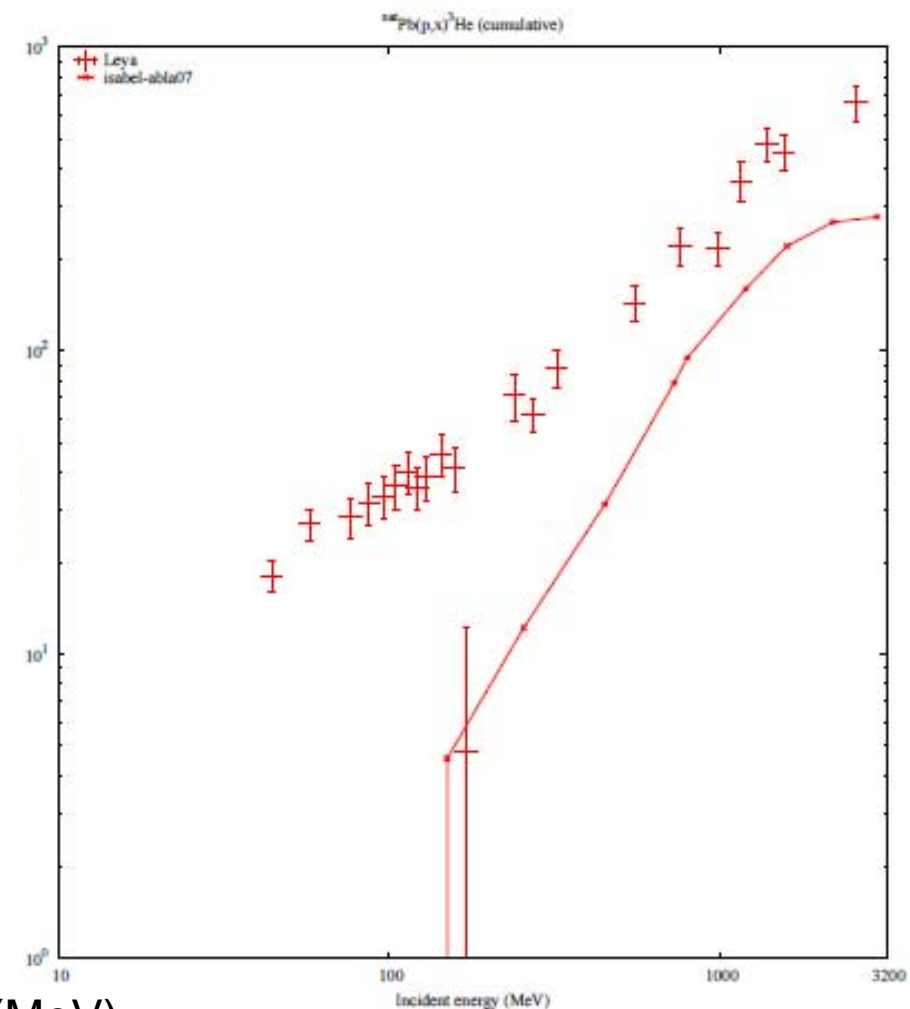
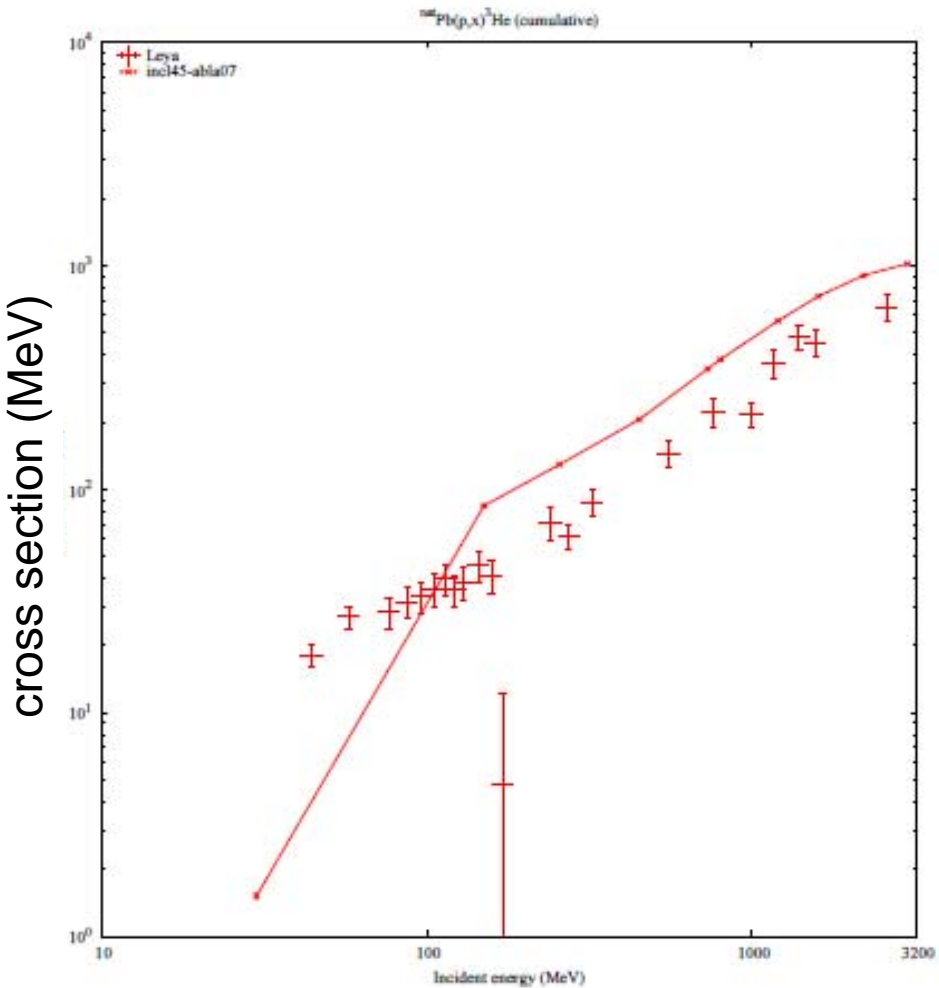
**Influenced by the first stage of the reaction**

energy (MeV)

# $^{nat}\text{Pb}(p,x)^3\text{He}$ (cumulative)

## INCL45-ABLA07

## ISABEL-ABLA07



energy (MeV)

# General conclusions

We (all here in this workshop) did a good job!

General **tendencies** and **behaviors** are well reproduced  
→ we have understood the main physics behind!

Left to do: refinements...

## Concerning ABLA07

Overall behavior satisfactorily, but there is still work to do

- neutron multiplicity distributions (INC or de-excitation?)
  - LCP spectra: barriers, tunneling, break-up?
- Residues: even-odd effect, structural effects, fission

Strength of ABLA07: high physics content  
relatively low computing time (we want to keep this feature)

**Thanks EU contribution (EUROTRANS)**